# C<sub>6</sub>H and C<sub>6</sub>D: Electronic spectra and Renner-Teller analysis

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### I. INTRODUCTION

The hextriynyl radical C<sub>6</sub>H, a member of the linear carbon chain  $C_{2n}H$  series, is an important intermediate in interstellar chemistry. It was initially detected by radio astronomy in the cold dense cloud TMC-1,1,2 and later unambiguously identified in a laboratory microwave study.<sup>3</sup> In the latter study accurate spectroscopic constants for the  $X^2\Pi$  electronic ground state were derived that allowed further identifications in TMC-1<sup>4</sup> and around the carbon-rich star IRC+10216.<sup>2,5-8</sup> Following matrix isolation studies<sup>9</sup> and the improvement in detection sensitivity of optical spectroscopic techniques (see e.g. Refs. 10 and 11), it became possible to record the electronic spectra of C<sub>6</sub>H, <sup>12-15</sup> among many other carbon chain radicals. In a combined microwave and optical study 16 the rotationally fully resolved origin band spectra of the  $B^2\Pi$  -  $X^2\Pi$ electronic system of C<sub>6</sub>H and C<sub>6</sub>D were analyzed, resulting in accurate spectroscopic parameters for both  $^2\Pi$  states.

Recently, the detection of  $C_6H^-$ , the first interstellar anion,  $^{17}$  and vibrationally excited  $C_6H^{18,19}$  has renewed the spectroscopic and astronomical interest in the hextriynyl radical. In Ref. 19, a detailed laboratory microwave study of an excited  $C_6H$  and  $C_6D$  bending vibrational level ( $\nu_{11}=1$ ) has been reported for two vibronic states,  $^2\Sigma$  and  $^2\Delta$ . The resulting effective spectroscopic parameters in combination with the very low excitation energy ( $\sim 20~K$ ) of the  $\mu^2\Sigma$  vibronic component are indicative for a relatively strong Renner-Teller interaction, i.e., the coupling between the bending vibrational and degenerate electronic orbital angular momenta, for the  $\nu_{11}$  mode in the  $X^2\Pi$  state.

In the present work, an extended analysis is presented for the  $B^2\Pi$  -  $X^2\Pi$   $0_0^0$  origin band transitions of  $C_6H$  and

 $C_6D$ . The focus, however, is on the rotationally resolved  $11_1^1$   $\mu^2\Sigma$  -  $\mu^2\Sigma$  vibronic hot band transition that is presented here for the first time. The Renner-Teller (R-T) analysis for the lowest bending mode  $\nu_{11}$  is carried out. The precise excitation energy of the  $11_1$   $\mu^2\Sigma$  state follows from the R-T analysis and from experimental data on band intensities.

## **II. EXPERIMENTAL DETAILS**

The C<sub>6</sub>H and C<sub>6</sub>D radicals are produced (among other species) in a supersonically expanding plasma by discharging a mixture of acetylene in helium. The spectra are recorded in direct absorption using cavity ring-down spectroscopy and details of the setup and experimental procedures are available from Refs. 20–22. A 3 cm  $\times$  100  $\mu$ m slit discharge nozzle is employed to generate a planar plasma expansion which provides an essentially 'Doppler-free' environment and a relatively long effective absorption path length. Three gas mixtures,  $\sim 0.5\%$  C<sub>2</sub>H<sub>2</sub>/He,  $\sim 0.3\%$  C<sub>2</sub>D<sub>2</sub>/He, and  $\sim (0.2\%$  C<sub>2</sub>H<sub>2</sub> + 0.2% C<sub>2</sub>D<sub>2</sub>)/He, are used in the present experiment to create a C/H, C/D, and C/H/D containing plasma, respectively. Typical running conditions are: V/I  $\approx -750$  V/100 mA and  $\sim$ 10 bar backing pressure. The nozzle is mounted with its slit parallel to and a few mm away from the optical axis between two highly reflective mirrors that form an optical cavity with a total length of 58 cm. Cavity ring-down spectroscopy is used to measure spectra in direct absorption. The tunable output of a pulsed Nd: YAG pumped dye laser system (10 Hz) is injected into the cavity and the light leaking out is detected by a photo-multiplier visualizing separate ring-down events. Typical ring-down times are and amount to 60–80  $\mu$ s. A spectrum is subsequently obtained by measuring the ring-down time as a function of laser frequency. Laser bandwidths of better than 0.035 cm<sup>-1</sup> are obtained by running the dye laser in a 2nd

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FIG. 1. Experimentally observed spectra through: (a) C/H plasma, (b) mixed C/H/D plasma, and (c) C/D plasma. The vertical dotted lines indicate the band origins of the observed bands.  $\Delta$  denotes the isotopic shift.

order grating setting, as recently demonstrated in the study of HC<sub>7</sub>H.<sup>21</sup> An absolute frequency calibration with accuracy better than 0.02 cm<sup>-1</sup> is realized by recording simultaneously an iodine reference spectrum and an etalon to correct for the non-linear scanning of the dye laser.

In addition to the slit nozzle experiments spectra are recorded using a pinhole discharge nozzle, which has been described in detail previously. The best achievable resolution with this nozzle is less, but spectra with different rotvibrational temperatures can be more easily obtained by varying the distance of pinhole nozzle orifice to the cavity axis. Specifically, for the  $C_6H$  radical studied in the present experiment, spectra are obtained at six different distances between 15 and 2 mm resulting in rotational temperatures between  $\sim$ 7 and 21 K.

#### III. RESULTS AND ANALYSIS

#### A. Observed spectra

In Ref. 14 an optical survey of hot electronic spectra of C<sub>6</sub>H was presented using cavity ring-down spectroscopy in

a hollow cathode discharge of acetylene and helium. These spectra were not rotationally resolved as a consequence of the high rotational (and vibrational) temperature > 200 K (yielding band heads) and Doppler broadening in the cell. Only the  $C_6H\ 0_0^{\ 0}$  bands were unambiguously identified, following matrix work. <sup>12,13</sup> The focus of the present experiment is a high resolution study of vibronic hot transitions starting from the low-lying excited vibrations of  $C_6H\ (C_6D)$  and to link the optical spectra to recent microwave work. <sup>19</sup>

Figure 1 shows overview spectra recorded through an expanding hydrocarbon plasma. Three absorption bands, labeled  $[h_0]$ ,  $[h_0']$ , and  $[h_I]$ , are observed in C/H plasma, as well as the corresponding D-isotopic bands (labeled  $[d_0]$ ,  $[d_0']$  and  $[d_I]$ ) in C/D plasma. These six bands are also visible in the mixed C/H/D plasma.

Following previous work, <sup>14,16</sup> the band pairs  $[h_0, h_0']$  and  $[d_0, d_0']$  can be unambiguously assigned to  $B^2\Pi_{\Omega}$  -  $X^2\Pi_{\Omega}$  ( $\Omega = 3/2$  - 3/2 and 1/2 - 1/2)  $0_0^0$  electronic origin band transitions of  $C_6H$  and  $C_6D$ , respectively. The fully resolved spectra of these bands are shown in Figs. 2 and 3. The spectra are very similar to the ones previously presented in Ref. 14. The lower detection limit and the better laser

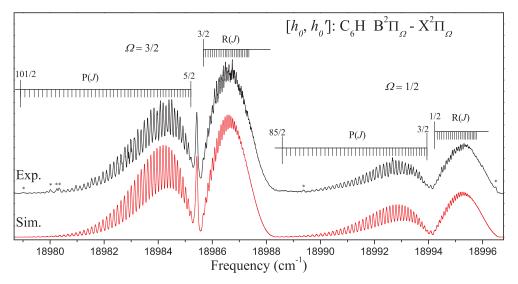


FIG. 2. The rotationally resolved spectrum of the  $B^2\Pi - X^2\Pi \ 0_0^0$  electronic origin transition of  $C_6H$ , corresponding to the band pair  $[h_0, h_0']$  in Fig. 1. The simulated spectrum for a rotational temperature of  $\sim 21$  K is shown in the lower trace. The irregular features marked with an asterisk are due to blending transitions by other small species like  $C_2$ , CH, etc.

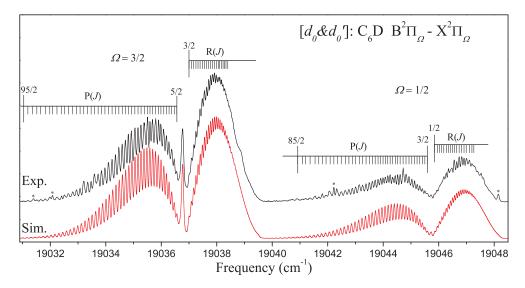


FIG. 3. The rotationally resolved spectrum of the  $B^2\Pi - X^2\Pi \ 0_0^0$  electronic origin transition of  $C_6D$ , corresponding to the band pair  $[d_0, d_0']$  in Fig. 1. The simulated spectrum for a rotational temperature of  $\sim$ 21 K is shown in the lower trace. The irregular features marked with an asterisk are due to blending transitions of other small species like  $C_2$ , CD, etc.

bandwidth in the present experiment allow the recording of higher *J*-value transitions (see the supplementary material<sup>23</sup>), and to derive more accurate values for the spectroscopic constants  $T_0$ , B', and A'. The parameter set can be extended with a centrifugal distortion constant D' and spin-rotational constant  $\gamma'$  for the excited state. The fit<sup>24</sup> is performed using the PGOPHER software<sup>25</sup> and the resulting values are listed in Table I.

The band  $[h_I]$  observed in the C/H plasma and the band  $[d_I]$  observed in the C/D plasma have not been reported before and are discussed here in detail. The newly observed band  $[h_I]$  coincides with more than three absorption peaks recorded in the hollow cathode discharge. The assignment of the bands  $[h_I]$  and  $[d_I]$  to vibrationally excited  $C_6H$  and

C<sub>6</sub>D, therefore, is logical and consistent with a number of observations:

- (i) As shown in Figs. 4 and 5, the rotational contours of the bands  $[h_I]$  and  $[d_I]$  are almost the same. The corresponding isotopic shift is  $\sim$ 49 cm<sup>-1</sup>, and this value is very close to the isotopic shift ( $\sim$ 52 cm<sup>-1</sup>) of the  $0_0^{0}$  electronic origin bands ( $[h_0, h_0']$  versus  $[d_0, d_0']$ ) of C<sub>6</sub>H and C<sub>6</sub>D, respectively. This indicates that  $[d_I]$  represents the D-isotopically shifted band of  $[h_I]$
- (ii) No additional bands between or around  $[h_I]$  and  $[d_I]$  are observed in the C/H/D plasma (see Fig. 1), indicating that the carrier must be of the form  $C_nH$  with only one H/D atom.

TABLE I.	Spectroscopic constants	(in cm-	$^{1}$ ) of $C_{6}H$ and $C_{6}D$ .
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	$C_6$ Н В $^2$ П - $X^2$ П		$C_6 D B^2 \Pi$ - $X^2 \Pi$	
	000	$11_1{}^1 \mu^2 \Sigma - \mu^2 \Sigma$	00 <sup>0a</sup>	$11_1^1 \mu^2 \Sigma - \mu^2 \Sigma$
<i>A</i> ″ <sup>b</sup>	-15.0424		-15.1276	
<i>B</i> ′′ <sup>b</sup>	0.04640497	0.04652018	0.04429243	0.0443852
$D''(10^{-9})^{\mathbf{b}}$	1.35	1.55	1.20	1.38
$\gamma''(10^{-3})^{\mathbf{b}}$	-7.12	-0.628	-3.85	-0.631
$T^{c}$	18989.7672(4)	19066.8372(6)	19041.2564(5)	19115.5251(10)
A'	-23.6924(7)		-24.0727(10)	
B'	0.0455952(5)	0.0456873(11)	0.0435921(39)	0.0436478(19)
$D'(10^{-9})$	1.58(28)	39.1(1.2)	1.20 <sup>d</sup>	13.4(1.7)
$\gamma'(10^{-3})$	-3.63(18)	38.29(6)	-6.3(7)	39.68(10)
$B^{\prime\prime}/B^{\prime}{}^{\mathbf{e}}$	1.0177	1.0182	1.0171	1.0171
$B_0^{\prime\prime}/B_{\mathrm{eff}}^{\prime\prime}{}^{\mathrm{e}}$	0.9975		0.9979	
$B_0'/B_{\mathrm{eff}}'^{\mathbf{e}}$	0.9980		0.9979	

<sup>&</sup>lt;sup>a</sup>Spectroscopic constants are obtained by fitting rotational transitions with  $J' \le 15.5$  for the  $\Omega = 3/2$  component and  $J' \le 10.5$  for the  $\Omega = 1/2$  component.<sup>24</sup>

<sup>&</sup>lt;sup>b</sup>Spectroscopic constants for the lower states are taken from Refs. 16 and 19 and fixed.

<sup>&</sup>lt;sup>c</sup>The errors of the band origins (T) represent the statistical uncertainties as obtained from the least squares fit. The maximum uncertainty of the absolute laser frequency in the present work is  $\sim 0.02 \text{ cm}^{-1}$ .

dFixed parameter in the least-squares fit.

eRatio of rotational constants is dimensionless.

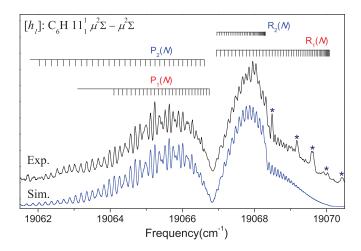


FIG. 4. The rotationally resolved spectrum of the  $11_1^1 \mu^2 \Sigma - \mu^2 \Sigma$  vibronic transition of C<sub>6</sub>H, corresponding to band [ $h_I$ ] in Fig. 1. The simulated spectrum for a rotational temperature of  $\sim$ 21 K is shown in the lower trace. The irregular features marked with an asterisk are due to blending transitions by other small species.

- (iii) As shown in Figs. 4 and 5, doublet P- and R-branch contours with large spin-splitting, i.e.,  $P_1(N)$ - $P_2(N)$ , and  $R_1(N)$ - $R_2(N)$ , can be seen in the rotationally resolved spectra of  $[h_I]$  and  $[d_I]$ . In contradiction to the origin band transitions (Figs. 2 and 3) no Q-branch is present. This is typical for a  ${}^2\Sigma$   ${}^2\Sigma$  transition of a linear molecule.
- (iv) From the combination differences, R(N-1) P(N+1) = B''(4N+2), the approximate rotational constant B'' for the lower  $^2\Sigma$  state is estimated as  $\sim 0.05$  cm $^{-1}$ . This deviates substantially from B'' values found for  $C_5H^{(+/-)}$  ( $\sim 0.08$  cm $^{-1}$ ) $^{26,27}$  and  $C_7H^{(+/-)}$  ( $\sim 0.03$  cm $^{-1}$ ), $^{28}$  but is of the same magnitude as that of  $C_6H^{(+/-)}$  ( $\sim 0.05$  cm $^{-1}$ ). $^{19}$  Since electronic spin multiplicities for the  $C_6H^+$  and  $C_6H^-$  are triplet and singlet, respectively,  $C_6H$  remains as only candidate.

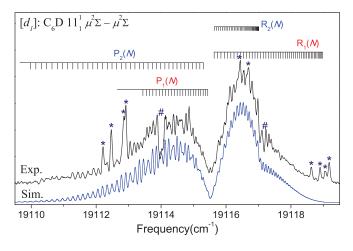


FIG. 5. The rotationally resolved spectrum of the  $11_1^1 \mu^2 \Sigma - \mu^2 \Sigma$  vibronic transition of C<sub>6</sub>D, corresponding to the band [ $d_I$ ] in Fig. 1. The simulated spectrum for a rotational temperature of  $\sim$ 21 K is shown in the lower trace. The irregular features marked with an asterisk are due to blending transitions by other small species. The irregular features marked with a hash are reproducible, and may be due to a local perturbation in the upper state.

(v) The lowest  $\mu^2 \Sigma$  vibronic state of  $C_6H$ , which is split from the  $\nu_{11}$  vibrational levels due to the R-T effect, was estimated in a recent microwave experiment<sup>19</sup> to be only  $\sim 14 \pm 7$  cm<sup>-1</sup> above the  $X^2\Pi_{3/2}$  level. In a plasma expansion such a low-lying state is sufficiently populated to be observable.

These observations hint at an assignment of the observed  $[h_I]$  and  $[d_I]$  bands to transitions starting from a vibrationally excited level of C<sub>6</sub>H and C<sub>6</sub>D. Since no other bands are observed besides the electronic origin transition of C<sub>6</sub>H/C<sub>6</sub>D, the vibrational mode exited in the upper state is likely the same  $\nu_{11}$  bending mode. Thus, summarizing, the most likely assignment of the bands  $[h_I]$  and  $[d_I]$  is the  $11_1^1 \mu^2 \Sigma - \mu^2 \Sigma$  vibronic hot band transitions of C<sub>6</sub>H and C<sub>6</sub>D, respectively. An unambiguous proof is given in Sec. III B, where it is shown, that the new bands  $[h_I]$  and  $[d_I]$  can be fitted using the lower  $\nu_{11} \mu^2 \Sigma$  vibronic state constants of C<sub>6</sub>H and C<sub>6</sub>D as derived in Ref. 19.

## B. Rotational analysis

As shown in Figs. 4 and 5, both bands  $[h_I]$  and  $[d_I]$  exhibit four main branches,  $R_1(N)$  and  $R_2(N)$ ,  $P_1(N)$  and  $P_2(N)$ , consisting of transitions with  $\Delta N = \pm 1$ , respectively. Most rotational transitions between the doublet P- and R-branches, particularly in the band origin regions are blended. This blending increases the uncertainty of the transition frequency determinations, particularly for the  $R_2$ -branch, compared to the  $Q_0$ 0 electronic origin band. The measured line positions for the (partially) resolved rotational transitions of  $[h_I]$  and  $[d_I]$  are summarized in the supplementary material. All these transition frequencies are included in a least-squares fitting routine for a  $^2\Sigma$  -  $^2\Sigma$  transition. The rotational levels in both lower and upper  $^2\Sigma$  states are described by Hund's case (b) expressions:

$$F_1(^2\Sigma, N) = T_{\Sigma} + [B_{eff} - DN(N+1)]N(N+1)$$

$$+ \frac{\gamma_{eff}}{2}N$$
(1)

$$F_2(^2\Sigma, N) = T_{\Sigma} + [B_{eff} - DN(N+1)]N(N+1) - \frac{\gamma_{eff}}{2}(N+1)$$
 (2)

where  $B_{eff}$  is the effective rotational constant, D is the centrifugal distortion constant, and  $\gamma_{eff}$  is an effective spin splitting constant, which does not necessarily correspond to the spin-rotational constant in an isolated  $^2\Sigma$  state. In the case of a sufficiently large R-T effect, as derived by Hougen,  $^{29}$  the parameter  $\gamma_{eff}$  depends on the magnitude of the R-T interaction relative to the spin-orbit (SO) interaction in the  $^2\Pi$  electronic state. Such an effective spin splitting constant generally has a value of  $0 \le \gamma_{eff} \le 2B_{eff}$ . Its value is close to zero when the R-T interaction is large, and approaches a value of  $2B_{eff}$  when the R-T interaction is very small. For the lower  $\nu_{11}$   $\mu^2\Sigma$  vibronic state of  $C_6H/C_6D$ , the R-T effect is relatively large, as has been discussed in Ref. 19. This state can be analyzed separately using the standard rotational Hamiltonian

for an isolated  $^2\Sigma$  state. Consequently, the value of the effective spin splitting constant  $\gamma_{eff}''$  is predominantly determined by the spin-rotational interaction. Large spin splittings in the doublet branch structure of the bands  $[h_I]$  and  $[d_I]$  predict a smaller R-T effect in the upper state, and the constant  $\gamma_{eff}'$  is most likely determined by a joint SO and R-T interaction.

The least-squares fits are performed using the PGOPHER software. The spectroscopic constants,  $B_{eff}''$ , D'', and  $\gamma_{eff}''$ , in the lower  $\nu_{11}$   $\mu^2\Sigma$  states of C<sub>6</sub>H and C<sub>6</sub>D for the bands  $[h_I]$  and  $[d_I]$ , are taken from Ref. 19 and fixed, while the band origin T, and spectroscopic constants  $B_{eff}'$ , D', and  $\gamma_{eff}'$  in the upper  $^2\Sigma$  state are set as variable parameters.

The fit results are listed in Table I. The resulting simulated spectra for both bands  $[h_I]$  and  $[d_I]$  are shown in Figs. 4 and 5. It is found that very characteristic features, such as spectral interferences between  $P_1$ -and  $P_2$ -branches, and a sharp decrease in intensity around 19068 (19117) cm<sup>-1</sup> due to the  $R_2$ -branch band head are well reproduced in the  $C_6H$  ( $C_6D$ ) simulated spectra.

The derived rotational constants in the upper  $\mu^2 \Sigma$  states are compared to the constants in the  $B^2\Pi - X^2\Pi \ 0_0^0$  electronic origin bands of C<sub>6</sub>H and C<sub>6</sub>D. The change of the rotational constants upon electronic or vibrational excitation reflects the difference of the overall length of the chain. Two sets of rotational constant ratios, B''/B' for the  $B^2\Pi - X^2\Pi$ electronic excitation, and  $B_0/B_{eff}(^2\Sigma)$  for the vibrational excitation in the same electronic state, are compared in Table I. The similar value of the four B''/B' ratios,  $\sim 1.017$ , agrees with the assumption that the electronic excitation for bands  $[h_1]$ and  $[d_I]$  is the same as that for the  $0_0^0$  band of  $C_6H$  and  $C_6D$ , i.e.,  $B^2\Pi - X^2\Pi$ . The value indicates that the overall length of C<sub>6</sub>H and C<sub>6</sub>D somewhat increases upon electronic excitation. The similar value of the  $B_0''/B_{eff}''(^2\Sigma)$  and  $B_0'/B_{eff}'(^2\Sigma)$  ratios of C<sub>6</sub>H and C<sub>6</sub>D indicates that the excited vibrational mode in both lower and upper electronic state is the same, i.e., the  $v_{11}$ bending vibration. The value of  $B_0/B_{eff}(^2\Sigma)$  is  $\sim 0.2\%$  smaller than unity and due to the fact that the effective chain length of a linear polyatomic molecule slightly decreases upon excitation of a bending vibration.

#### C. Renner-Teller analysis

Owing to the R-T interaction, each bending fundamental mode in a  $^2\Pi$  electronic state is split into one  $^2\Delta$  and two  $^2\Sigma$  vibronic states. An effective Hamiltonian that considers both the R-T effect and the SO coupling is required to reproduce the vibronic structure in degenerated  $^2\Pi$  electronic states of C<sub>6</sub>H. The effective Hamiltonian without the rotational part is expressed as

$$H = H_{vib} + H_{RT} + H_{SO} \tag{3}$$

where  $H_{vib}$  is the vibrational Hamiltonian for the bending mode,  $H_{RT}$  the R-T Hamiltonian, and  $H_{SO}$  the SO coupling term. Because only the lowest bending vibration  $v_{11}$  is considered, the analysis here is simplified by following the assumption that high-order and intermode interactions can be neglected. The representation of the Hamiltonian in Eq. (3) is

constructed by a basis set of<sup>30</sup>

$$|\Lambda, v, K, \Sigma\rangle$$
 (4)

where  $\Lambda = \pm 1$  for a  ${}^2\Pi$  state, electronic spin  $\Sigma = \pm 1/2$ , and  $K = |\Lambda + l| = 0$ , 1, and 2 corresponding to the  $\Sigma$ ,  $\Pi$  and  $\Delta$  vibronic states, respectively. The non-zero matrix elements of the Hamiltonian in Eq. (3) are  ${}^{30}$ 

$$\langle v | H_{vib} | v \rangle = \omega (v+1) \tag{5}$$

$$\langle -1, v, K | H_{RT} | +1, v-2, K \rangle$$

$$= \frac{\varepsilon \omega}{4} \sqrt{(v-1+K)(v+1+K)}$$
(6)

$$\langle +1, v, K | H_{RT} | -1, v - 2, K \rangle$$

$$= \frac{\varepsilon \omega}{4} \sqrt{(v - 1 - K)(v + 1 - K)}$$
(7)

$$\langle +1, v, K | H_{RT} | -1, v, K \rangle = \frac{\varepsilon \omega}{2} \sqrt{(v+1)^2 - K^2}$$
 (8)

$$\langle \Lambda, \Sigma | H_{SO} | \Lambda, \Sigma \rangle = A_{SO} \Lambda \Sigma \tag{9}$$

where  $\omega$  is the vibrational frequency of the bending mode,  $\varepsilon$  the Renner parameter and  $A_{SO}$  the true SO splitting constant.

The previous microwave studies  $^{16,19}$  have resulted in effective spin splittings in the  $X^2\Pi$  ground state of  $C_6H$ :  $A_0''(0_0\ ^2\Pi) \sim -15.04\ cm^{-1}$ , and  $A_\Delta''(11_1\ ^2\Delta) \sim -11.90\ cm^{-1}$ , as well as  $A_0'(0^0\ ^2\Pi) \sim -23.69\ cm^{-1}$  in the  $B^2\Pi$  state. However, to reproduce the vibronic structure in both  $^2\Pi$  electronic states knowledge of the vibrational frequencies ( $\omega_{11}''$  and  $\omega_{11}'$ ) of the  $\nu_{11}$  bending mode is still required. Considering that the  $B_0''/B_{eff}''(\nu_{11})$  ratio is almost the same as  $B_0'/B_{eff}''(\nu_{11})$ , we assume that

$$0.9\omega_{11}'' \le \omega_{11}' \le 1.1\omega_{11}'' \tag{10}$$

Under this assumption, the parameters  $\varepsilon_{11}$ ,  $\omega_{11}$  and  $A_{SO}$  for both  $^2\Pi$  electronic states are fitted using the transition energies of the bands  $[h_0]$ ,  $[h_0']$  and  $[h_I]$  listed in Table I. The resulting parameters from the R-T analysis are summarized in Table II. The derived vibrational frequency  $\omega_{11} \sim 95 \text{ cm}^{-1}$  is in good agreement with the previously predicted values from theoretical calculations. The schematic diagram in Fig. 6 shows the vibronic structures of C<sub>6</sub>H derived from the R-T analysis, where the corresponding transitions to the bands  $[h_0]$ ,  $[h_0']$  and  $[h_I]$  are indicated as well.

The large value for the Renner parameter  $\varepsilon_{11}''$  ( $\sim$ 0.995) indicates that the potential surface of the  $X^2\Pi$  ground state has an energy minimum in a non-linear geometry due to strong coupling between vibrational and electronic orbital angular momenta. Further, the R-T analysis shows that, for such a large  $\varepsilon$  value the  $0_0$   $^2\Pi_{1/2}$  level is significantly mixed with the  $11^2$   $^2\Pi_{1/2}$  levels. This behavior has also been found for the  $A^2\Pi$  state of  $C_4H$ ,  $^{34}$  which has a similar electronic configuration and a similarly large Renner parameter ( $\varepsilon$   $\sim$ 0.99) as the  $X^2\Pi$  state of  $C_6H$ .

TABLE II. Vibronic parameters of  $C_6H$  and  $C_6D$  in both  $^2\Pi$  states.

	C <sub>6</sub> H		Ce	C <sub>6</sub> D	
	$B^2\Pi$	Х2П	$B^2\Pi$	X <sup>2</sup> Π	
$\varepsilon_{II}^{a}$ $\omega/\text{cm}^{-1}$ $A_{so}/\text{cm}^{-1}$	0.08869(2) 95(1) -23.74(1)	0.995(5) 100(10) -47.5(3)	0.08852 92° -24.12 <sup>d</sup>	0.99(1) <sup>b</sup> 97 <sup>c</sup> -47	

<sup>&</sup>lt;sup>a</sup>The sign of the parameter  $\varepsilon_{II}$  could not be unambiguously determined in the present study.

In contrast, the R-T effect for the  $\nu_{11}$  bending mode in the  $B^2\Pi$  state of  $C_6H$  is much smaller ( $\varepsilon'\sim0.09$ ) than in the  $X^2\Pi$  ground state. According to Hougen, the values of  $B_{eff}$  and  $\gamma_{eff}$  for  $^2\Sigma$  vibronic states can be approximately expressed by  $^{29,35}$ 

$$B_{eff}^{\mu/k} = B \left( 1 \pm \frac{A_{SO}^2 B}{A_{\Sigma}^3} \right)$$
 (11)

$$\gamma_{eff}^{\mu/k} = 2B \left( 1 - \frac{2\varepsilon\omega}{A_{\Sigma}} \pm \frac{A_{SO}^2 B}{A_{\Sigma}^3} \right)$$
 (12)

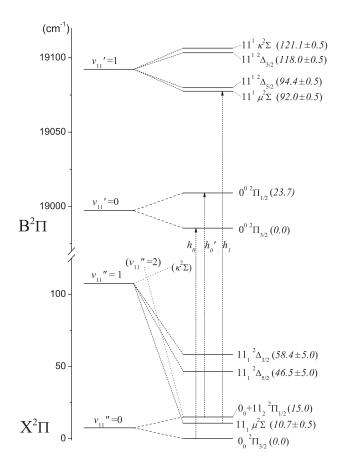


FIG. 6. Schematic diagram of the vibronic structure in both  $^2\Pi$  electronic states of C<sub>6</sub>H derived from the R-T analysis. The relative energies (in cm<sup>-1</sup>) with respect to the  $\nu=0\,^2\Pi_{3/2}$  level in each electronic state are indicated by italic numbers in brackets.

TABLE III. Spin-orbit coupling constants and Renner parameters of  $C_6H$  and  $C_4H$ .

	C <sub>6</sub> H		C <sub>4</sub> H	
	$B^2\Pi$	Х <sup>2</sup> П	$B^2\Pi^a$	A <sup>2</sup> Π <sup>b</sup>
$A_0$ /cm <sup>-1</sup>	-23.69	-15.04	-14.8	-28
$A_{SO}/\mathrm{cm}^{-1}$	-23.74	-47.5	-22.1	-45.5
$ \varepsilon_{(CCC\ bending)} $	$ \varepsilon_{11}  \sim 0.089$	$ \varepsilon_{11} $ $\sim$ 0.995	$ \varepsilon_7  \sim 0.016$	$ \varepsilon_7  \sim 0.061$
$ \varepsilon_{(\text{CCH bending})} $	_	_	$ \varepsilon_6  \sim 0.657$	$ \varepsilon_6 $ $\sim$ 0.99

a Values taken from Ref. 38.

where  $A_{\Sigma}$  is the effective spin splitting between the two  ${}^{2}\Sigma$  states:

$$A_{\Sigma} = \sqrt{A_{SO}^2 + 4\varepsilon^2 \omega^2} \tag{13}$$

In the case of the 11<sup>1</sup>  $\mu^2 \Sigma$  vibronic state of C<sub>6</sub>H, using the values of  $B' \approx B_{eff}'$ ,  $\varepsilon'$ ,  $\omega_{11}'$ , and  $A_{SO}'$  listed in Tables I and II, we obtain  $\gamma_{eff}' \approx 0.0386 \text{ cm}^{-1}$ , which is in good agreement with the  $\gamma_{eff}' = 0.038285 \text{ cm}^{-1}$  as derived from the rotational analysis on band  $[h_I]$ .

Since the microwave data for the  $11_1^2\Delta$  vibronic state of  $C_6D$  have not been reported up to now, it remains a challenge to reproduce the vibronic structure of  $C_6D$ . Using  $\gamma_{eff}'\sim 0.0399~\text{cm}^{-1}$  in the  $11^1~\mu^2\Sigma$  vibronic state of  $C_6D$ , the parameter  $\varepsilon_{11}'$  can be estimated to be  $\sim 0.0885$  from Eqs. (11)–(13). Since the R-T effect in the ground state of  $C_6H$  is very large, it is expected that the vibronic structure of  $C_6D$  in the low energy range ( $<100~\text{cm}^{-1}$ ) may be very similar to that in the  $X^2\Pi$  ground state of  $C_6H$ . The estimated parameters in both  $^2\Pi$  states of  $C_6D$  derived from this approximate R-T analysis are also summarized in Table II.

## D. Excitation energy of the 11<sub>1</sub> $\mu^2\Sigma$ vibronic state

From the R-T analysis, the excitation energy of the lowest-lying  $11_1 \ \mu^2 \Sigma$  vibronic state of  $C_6H$  is obtained as  $10.7 \pm 0.5 \ \mathrm{cm}^{-1}$ . Additional measurements have been performed to experimentally determine the excitation energy of the  $11_1 \ \mu^2 \Sigma$  vibronic state. This is accomplished by measuring a series of spectra with different rotational temperatures. For this a pinhole discharge nozzle is used, as described in the experimental section. The recorded spectra for the bands  $[h_0]$ ,  $[h_0']$ , and  $[h_1]$  are shown in Fig. 7. It is found that, with the ongoing adiabatic cooling the relative intensities of the bands  $[h_0']$  and  $[h_1]$  become smaller with respect to the band  $[h_0]$ . The relative intensities of the bands  $[h_0']$  and  $[h_1]$  are expressed by

$$\frac{I_{[h_0']}}{I_{[h_0]}} \sim \frac{N((0_0 + 11_2)^2 \Pi_{1/2})}{N(0_0 {}^2 \Pi_{3/2})} = \exp\left(-\frac{|A_0''|}{kT}\right)$$
(14)

$$\frac{I_{[h_1]}}{I_{[h_2]}} \sim \frac{N(11_1 \,\mu^2 \,\Sigma)}{N(0_0 \,^2 \Pi_{3/2})} = \exp\left(-\frac{E(11_1 \,\mu^2 \,\Sigma)}{kT}\right) \tag{15}$$

where N is the population density of  $C_6H$  in separate lowlying vibronic levels, and T is the temperature (in Kelvin).

<sup>&</sup>lt;sup>b</sup>Estimated value.

<sup>&</sup>lt;sup>c</sup>Vibrational frequencies estimated by  $\omega_{\rm D}/\omega_{\rm H} \approx (B_{\rm D}/B_{\rm H})^{1/2}$ .

<sup>&</sup>lt;sup>d</sup>Assume  $A_{SO}' \approx A_{O}'$  in C<sub>6</sub>D.

<sup>&</sup>lt;sup>b</sup>Values taken from Ref. 34.

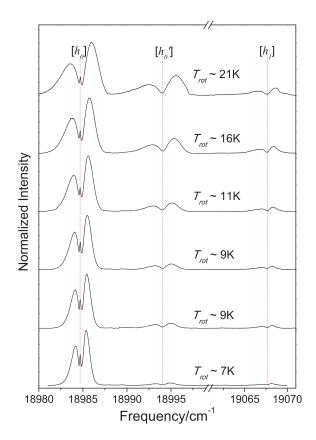


FIG. 7. Low resolution experimental spectra of the bands  $[h_0]$ ,  $[h_0']$ , and  $[h_1]$  with different rotational and vibrational temperatures through C/H plasma generated in a pinhole discharge nozzle.

As discussed in Sec. III C, the  $0_0$   $^2\Pi_{1/2}$  level is significantly mixed with the  $11_2$   $^2\Pi_{1/2}$  level. This mixing suggests that the C<sub>6</sub>H populations in the three levels,  $0_0$   $^2\Pi_{3/2}$ ,  $(0_0+11_2)$   $^2\Pi_{1/2}$ , and  $11_1$   $\mu^2\Sigma$ , should be determined by the same effective vibronic temperature, i.e., the temperatures T in Eqs. (14) and (15) are most likely the same. Under this assumption,

$$\ln\left(\frac{I_{[h_1]}}{I_{[h_0]}}\right) = \frac{\mathrm{E}(11_1 \,\mu^2 \Sigma)}{|A_0|} * \ln\left(\frac{I_{[h'_0]}}{I_{[h_0]}}\right) + \mathrm{C}$$
 (16)

applies, where C is a constant related to the Frank-Condon factors of the  $0_0^0$  and  $11_1^1$  transitions. In Fig. 8, a plot of relative intensities of the bands  $[h_0']$  vs  $[h_I]$  is shown, displaying a perfect linearity. The observation of this linearity provides proof for the validity of the assumption that spin and vibrational temperatures are equal. Using the value of  $|A_0''| = 15.04 \text{ cm}^{-1}$ , <sup>16</sup> we obtain the excitation energy  $E(11_1 \mu^2 \Sigma) \approx 11.4 \pm 0.5 \text{ cm}^{-1}$ , which is in reasonable agreement with the value resulting from the R-T analysis.

The final value for the excitation energy of the lowest-lying  $11_1~\mu^2\Sigma$  vibronic state is derived as average value from the experimental measurements and the R-T analysis, yielding  $11.0\pm0.8~\text{cm}^{-1}$ . This value is in agreement with the value of  $\sim\!14\pm7~\text{cm}^{-1}$  that was estimated previously from the microwave measurements,  $^{19}$  but much more accurate.

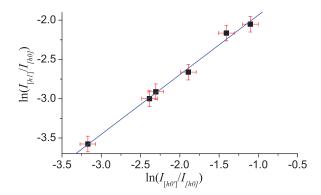


FIG. 8. The relative intensity of the bands  $[h_I]$  and  $[h_0]$  with respect to the band  $[h_0]$ . The uncertainty of the intensity determinations is estimated to be  $\sim \pm 10\%$ .

#### IV. DISCUSSION

The R-T analysis for the  $X^2\Pi$  ground state results in a large value of the true SO coupling constant,  $A_{SO}^{"}$  $= -47 \text{ cm}^{-1}$ , which is much larger than the one-electron spin-orbit coupling constant for atomic carbon,  $\zeta_{\rm C}=29$ cm<sup>-1</sup>. In the  $X^2\Pi$  ground state of  $HC_6H^+$ ,  $^{36,37}$  an ion isoelectronic with  $C_6H$ , the effective SO splitting constant  $A_{eff}$ "  $\sim -31 \text{ cm}^{-1}$  is also larger than  $\zeta_{\rm C}$ . Such large SO coupling constants in long conjugated carbon chain systems cannot be explained based on the orbital approximation of LCAO (linear combination of atomic orbitals) theory. This means, that the LCAO approximation may become invalid for long conjugated  $C_{2n}H$  systems. Ab initio calculations <sup>32,38,39</sup> show that dynamic electron correlation effects neglected in the LCAO approximation, are significant in the low-lying electronic states of the homologuous  $C_{2n}H$  radicals and the  $HC_6H^+$  ion. In addition, estimation of the SO coupling constants and hyperfine parameters by LCAO theory is usually applied for diatomic or tri-atomic molecules which can be simplified as a two-center system. 40 While the pattern of the conjugated  $\Pi$ bond in C<sub>6</sub>H distributes along the molecular axis, it cannot be simplified as a two-center system. This may be another possible reason explaining why the LCAO approximation may become invalid for long  $C_{2n}H$  radicals.

It is of interest to compare the values for the electronic SO coupling and R-T effect for the linear radicals C<sub>4</sub>H and C<sub>6</sub>H, which may reveal the behavior of angular momentum couplings in linear species with increasing chain length. Previous studies<sup>15,16,34,38,39,41</sup> show that the order of the first two electronic states,  $X^2\Pi$  and  $A^2\Sigma$  for  $C_6H$ , is reversed for the smaller species  $C_4H$ . The  $C_4H$   $A^2\Pi$  and  $C_6H$   $X^2\Pi$  states have a similar electronic configuration. As shown in Table III, the effective SO splitting constants  $(A_0)$  in the X/A<sup>2</sup> $\Pi$  states and  $B^2\Pi$  states of  $C_6H$  and  $C_4H$  are rather different. However, the values of the true SO coupling constants  $(A_{SO})$  determined from the R-T analyses are very close for the two chains. Specifically, the  $A_{SO}$  values in both  $^2\Pi$  states of C<sub>6</sub>H are  $\sim$ 5% larger than those of C<sub>4</sub>H. Ab initio calculations<sup>32,38</sup> have shown that conjugated  $\Pi$  bond systems become more stable with increasing the chain length of the homologues  $C_{2n}H$  radicals. The slightly stronger electronic SO coupling in C<sub>6</sub>H may arise from the higher stability of the conjugation in a longer chain.

From Table III, we also deduce that the coupling between the electronic orbital and CCC bending  $(v_{11})$  vibrational angular momenta in the  ${}^{2}\Pi$  states of  $C_{6}H$  and  $C_{4}H$  must be rather different. For C<sub>6</sub>H, the coupling between the CCC bending vibration  $(v_{11})$  and the  $X^2\Pi$  electronic orbital motion is very strong ( $\varepsilon_{11} \sim 0.995$ ), whereas it is much weaker in the B<sup>2</sup> $\Pi$  state ( $\varepsilon_{11} \sim 0.089$ ). In contrast, the couplings between the CCC bending vibration  $(v_7)$  and both  ${}^2\Pi$  states of  $C_4H$  are very weak ( $\varepsilon_7 \sim 0.016$  and 0.061), but very strong for the CCH bending vibration ( $\nu_6$ ) in both  $^2\Pi$  states ( $\varepsilon_6 \sim 0.99$ and 0.656). For two members of the same linear chain family  $C_{2n}H$ , the lowest CCC bending vibrational modes are very similar. This suggests that, despite the fact that the lowest  $(A/X)^2\Pi$  electronic states have the same open shell configuration, the electronic orbital motion may have changed from  $C_4H$  to  $C_6H$ , while in the second  $B^2\Pi$  electronic state, the electronic orbital motion may be similar.

The present work also may have astronomical implications. In the R-T analysis of this work, contributions from other vibrational modes and intermode interactions are neglected. Since the parameter  $\varepsilon_{11}$  (~0.995) in the X<sup>2</sup> $\Pi$  ground state of C<sub>6</sub>H is large, it is expected that some other low-lying vibronic states exist that originate from combined vibrational excitation of  $v_{11}$  and other bending modes, such as the CCH bending vibration. Such a complicated vibronic structure has been found in the  $A^2\Pi$  state of  $C_4H$ .<sup>34</sup> Following the spectroscopic identifications of the  $11_1 \mu^2 \Sigma$  and  $^2 \Delta$  vibronic states of C<sub>6</sub>H, it was pointed out in Ref. 19 that there remain a number of unidentified features close in frequency to the identified C<sub>6</sub>H rotational lines in both the laboratory plasma experiment and in the outer molecular envelope of IRC+10216. Some of these features, consequently, may be due to other low-lying vibronic states of C<sub>6</sub>H. To reproduce the rotational structure in such vibronic states, a precise and comprehensive theoretical model, including R-T effect is needed. 40 Extending the laboratory work on the R-T effect to other vibrational modes of C<sub>6</sub>H may help astronomers to assign the remaining unidentified C<sub>6</sub>H features.

Since the electronic configuration and state order of C<sub>6</sub>H and  $C_8H$  are the same, it is predicted that the true SO constant in the  $X^2\Pi$  ground state of  $C_8H$  might be  $A_{SO}'' \sim -50$  cm<sup>-1</sup>, which is much larger than the effective SO splitting of  $A_0'' \sim -19.3 \text{ cm}^{-1}$  determined from spectroscopic observation. 42,43 This large difference predicts a relatively large R-T effect (probably  $\varepsilon > 0.9$ ) in the X<sup>2</sup> $\Pi$  ground state of C<sub>8</sub>H as well. On the other hand, as a longer chain, C<sub>8</sub>H has more bending vibrational modes and a smaller vibrational energy for the CCC bending mode. As a consequence several vibronic states may lie very close to the  $0_0$   ${}^2\Pi_{3/2}$  ground level, as for C<sub>6</sub>H but may exhibit a more complex structure. Up to now, both vibrationally excited C<sub>4</sub>H and C<sub>6</sub>H have been identified in IRC+10216,  $^{18,44}$  as well as C<sub>8</sub>H in its  $X^2\Pi$  ground state. 45 It is very likely that vibrationally excited C<sub>8</sub>H also exists in IRC+10216.

#### V. CONCLUSIONS

Rotationally resolved spectra of the  $B^2\Pi$  -  $X^2\Pi$   $0_0^0$  electronic origin band and  $11_1^1$   $\mu^2\Sigma$  -  $\mu^2\Sigma$  vibronic hot band

transitions of both  $C_6H$  and  $C_6D$  are recorded by cavity ringdown spectroscopy, resulting in precise spectroscopic parameters. The Renner-Teller analysis indicates a large R-T interaction for the  $\nu_{11}$  bending mode in the  $X^2\Pi$  ground state of  $C_6H$ , and relatively small in the  $B^2\Pi$  state. The excitation energy of the lowest-lying  $11_1~\mu^2\Sigma$  vibronic states is determined to be  $11.0\pm0.8~cm^{-1}$ .

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- <sup>1</sup>H. Suzuki, M. Ohishi, N. Kaifu, S. Ishikaw, and T. Kasuga, Publ. Astron. Soc. Jpn. 38, 911 (1986).
- <sup>2</sup>J. Cernicharo, M. Guelin, K. M. Menten, and C. M. Walmsley, Astron. Astrophys. **181**, L1 (1987).
- <sup>3</sup>J. C. Pearson, C. A. Gottlieb, D. R. Woodward, and P. Thaddeus, Astron. Astrophys. 189, L13 (1988).
- <sup>4</sup>N. Kaifu, M. Ohishi, K. Kawaguchi, S. Saito, S. Yamamoto, T. Miyaji, K. Miyazawa, S. Ishikawa, C. Noumaru, S. Harasawa, M. Okuda, and H. Suzuki, Publ. Astron. Soc. Jpn. **56**, 69 (2004).
- <sup>5</sup>M. Guelin, J. Cernicharo, C. Kahane, J. Gomez-Gonzalez, and C. M. Walmsley, Astron. Astrophys. 175, L5 (1987).
- <sup>6</sup>S. Saito, K. Kawaguchi, H. Suzuki, M. Ohishi, N. Kaifu, and S. Ishikawa, Publ. Astron. Soc. Jpn. 39, 193 (1987).
- <sup>7</sup>K. Kawaguchi, Y. Kasai, S. Ishikawa, and N. Kaifu, Publ. Astron. Soc. Jpn. 47, 853 (1995).
- <sup>8</sup>J. Cernicharo, M. Guelin, and C. Kahane, Astron. Astrophys. Suppl. Ser. 142, 181 (2000).
- <sup>9</sup>J. P. Maier, Chem. Soc. Rev. **26**, 21 (1997).
- <sup>10</sup>E. B. Jochnowitz, and J. P. Maier, Ann. Rev. Phys. Chem. **59**, 519 (2008).
- <sup>11</sup>R. Nagarajan, and J. P. Maier, Int. Rev. Phys. Chem. 29, 521 (2010).
- <sup>12</sup>D. Forney, J. Fulara, P. Freicogel, M. Jakobi, D. Lessen, and J. P. Maier, J. Chem. Phys. **103**, 48 (1995).
- <sup>13</sup>P. Freivogel, J. Fulara, M. Jakobi, D. Forney, and J. P. Maier, J. Chem. Phys. **103**, 54 (1998).
- <sup>14</sup>M. Kotterer, and J. P. Maier, Chem. Phys. Lett. **266**, 342 (1997).
- <sup>15</sup>T. R. Taylor, C. Xu, and D. M. Neumark, J. Chem. Phys. **108**, 10018 (1998).
- <sup>16</sup>H. Linnartz, T. Motylewski, O. Vaizert, J. P. Maier, A. J. Apponi, M. C. Mc-Carthy, C. A. Gottlieb, and P. Thaddeus, J. Mol. Spectrosc. 197, 1 (1999).
- <sup>17</sup>M. C. McCarthy, C. A. Gottlieb, H. Gupta, and P. Thaddeus, Astrophys. J. 652, L141 (2006).
- <sup>18</sup>J. Cernicharo, M. Guelin, M. Agundez, M. C. McCarthy, and P. Thaddeus, Astrophys. J. 688, L83 (2008).
- <sup>19</sup>C. A. Gottlieb, M. C. McCarthy, and P. Thaddeus, Astrophys. J. Suppl. **189**, 261 (2010)
- <sup>20</sup>T. Motylewski, and H. Linnartz, Rev. Sci. Instrum. **70**, 1305 (1999).
- <sup>21</sup>N. Wehres, D. Zhao, H. Linnartz, and W. Ubachs, Chem. Phys. Lett. 497, 30 (2010).
- <sup>22</sup>D. Zhao, N. Wehres, H. Linnartz, and W. Ubachs, Chem. Phys. Lett. **501**, 232–237 (2011).
- <sup>23</sup>See supplementary material at http://dx.doi.org/10.1063/1.3609112 for the measured rotational transition frequencies of the observed bands in the present work.
- <sup>24</sup>Rotational levels in the upper state of  $C_6D$   $0_0^0$  electronic origin band are found to be heavily perturbed (see Table S2 in the supplementary material (Ref. 23)). In the least squares fit, the constant D' is set as D'' and fixed, and only levels with  $J' \le 15.5$  for the  $\Omega = 3/2$  component and  $J' \le 10.5$  for the  $\Omega = 1/2$  component are included to derive the spectroscopic constants  $T_0$ , B', A' and  $\gamma'$ . As a consequence, the rms of the fit and the resulted values of spectroscopic constants of  $C_6D$  are comparable to those of  $C_6H$ .
- <sup>25</sup>PGOPHER, a Program for Simulating Rotational Structure, C. M. Western, University of Bristol, http://pgopher.chm.bris.ac.uk.

<sup>27</sup>M. Tulej, F. J. Mazzotti, and J. P. Maier, J. Phys. Chem. A **115**, 6878 (2011).

<sup>28</sup>M. J. Travers, M. C. McCarthy, C. A. Gottlieb, and P. Thaddeus, Astrophysic. J. 465, L77 (1996).

<sup>29</sup>J. T. Hougen, J. Chem. Phys. **36**, 519 (1962).

<sup>30</sup>J. M. Frye, and T. J. Sears, Mol. Phys. **62**, 919 (1987).

<sup>31</sup>S. T. Brown, J. C. Rienstra-Kiracofe, and H. F. Schaefer, III, J. Phys. Chem. A **103**, 4065 (1999).

<sup>32</sup>Z. Cao, and S. P. Peyerimho, Phys. Chem. Chem. Phys. **3**, 1403 (2001).

<sup>33</sup>Y. Wu, Comput. Theor. Chem. **963**, 104 (2011).

<sup>34</sup>F. J. Mazzotti, R. Raghunandan, A. M. Esmail, M. Tulej, and J. P. Maier, J. Chem. Phys. **134**, 164303 (2011).

<sup>35</sup>G. Herzberg, Molecular Spectra and Molecular Structure III. Electronic Spectra and Electronic Structure of Polyatomic Molecules (D. Van Nostrand, New York, 1966). <sup>36</sup>W. E. Sinclair, D. Pfluger, H. Linnartz, J. P. Maier, J. Chem. Phys. **110**, 296 (1999).

J. Chem. Phys. 135, 044307 (2011)

- <sup>37</sup>D. Pfluger, W. E. Sinclair, H. Linnartz, J. P. Maier, Chem. Phys. Lett. **313**, 171 (1999).
- <sup>38</sup> A. L. Sobolewski, and L. Adamowicz, J. Chem. Phys. **102**, 394 (1995).
- <sup>39</sup>D. E. Woon, Chem. Phys. Lett. **244**, 45 (1995).
- <sup>40</sup>M. D. Allen, K. M. Evenson, D. A. Gillett, and J. M. Brown, J. Mol. Spectrosc. 201, 18 (2000).
- <sup>41</sup>K. Hoshina, H. Kohguchi, Y. Ohshim, and Y. Endo, J. Chem. Phys. **108**, 3465 (1998).
- <sup>42</sup>M. C. McCarthy, M. J. Travers, A. Kovacs, C. A. Gottlieb, and P. Thaddeus, Astron. Astrophys. 309, L31 (1996).
- <sup>43</sup>H. Linnartz, T. Motylewski, and J. P. Maier, J. Chem. Phys. **109**, 3819 (1998).
- <sup>44</sup>M. Guelin, J. Cernicharo, S. Navarro, D. R. Woodward, C. A. Gottlieb, and P. Thaddeus, Astron. Astrophys. **182**, L37 (1987).
- <sup>45</sup>J. Cernicharo, and M. Guelin, Astron. Astrophys. **309**, L27 (1996).