

Outline

- 1 Scattering Polarization
- 2 Zeeman Effect
- 3 Hanle Effect

Scattering Polarization

Single Particle Scattering

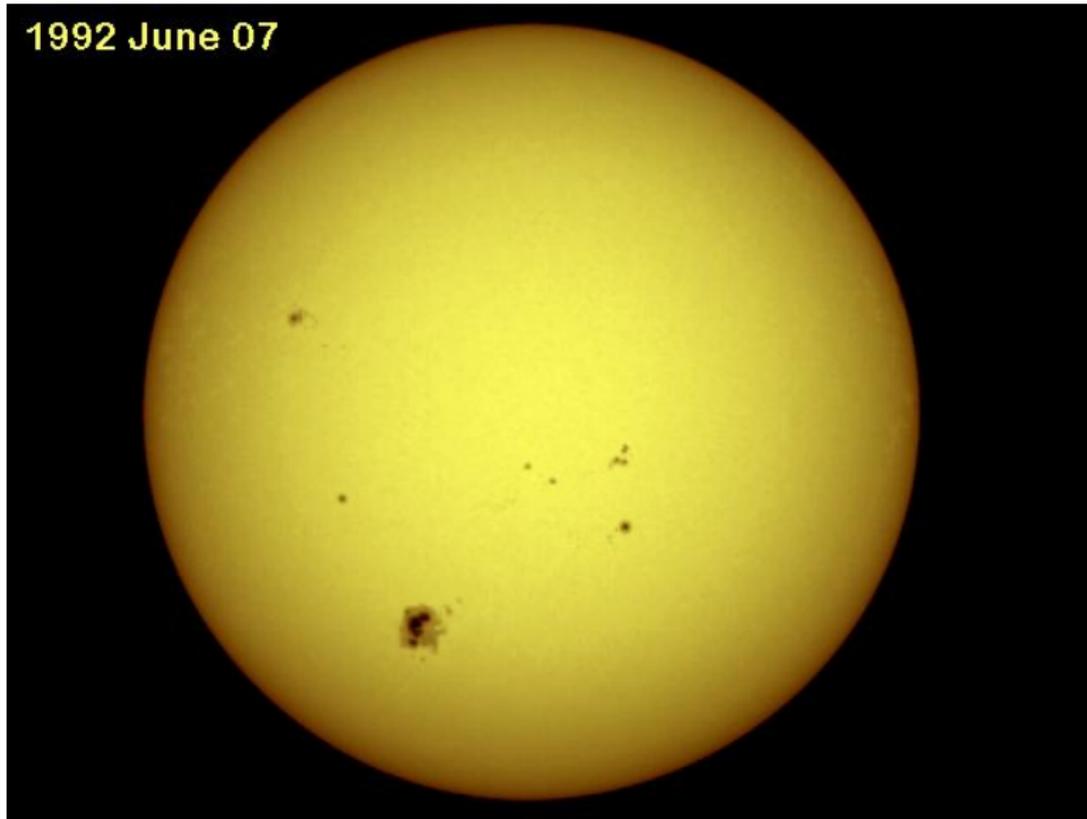
- light is absorbed and re-emitted
- if light has low enough energy, no energy transferred to electron, but photon changes direction \Rightarrow elastic scattering
- for high enough energy, photon transfers energy onto electron \Rightarrow inelastic (Compton) scattering
- Thomson scattering on free electrons
- Rayleigh scattering on bound electrons
- based on very basic physics, scattered light is linearly polarized

Polarization as a Function of Scattering Angle

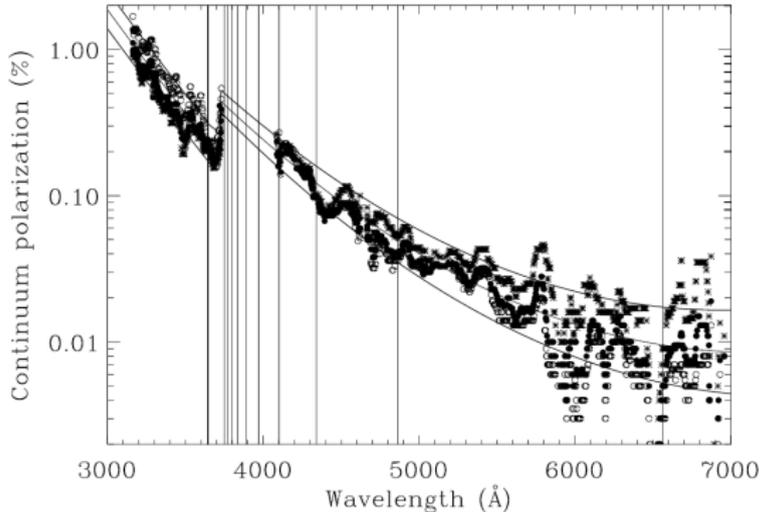
- same variation of polarization with scattering angle applies to Thomson and Rayleigh scattering
- scattering angle θ
- projection of amplitudes:
 - 1 for polarization direction perpendicular to scattering plane
 - $\cos \theta$ for linear polarization in scattering plane
- intensities = amplitudes squared
- ratio of $+Q$ to $-Q$ is $\cos^2 \theta$ (to 1)
- total scattered intensity (unpolarized = averaged over all polarization states) proportional to $\frac{1}{2} (1 + \cos^2 \theta)$

Limb Darkening

1992 June 07



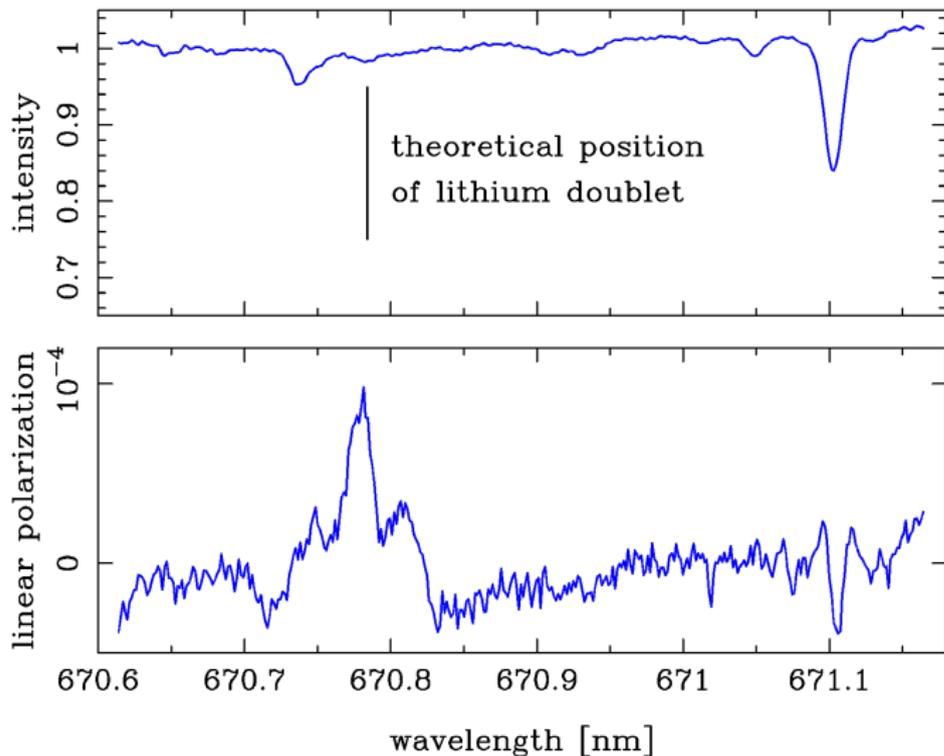
Solar Continuum Scattering Polarization



Stenflo 2005

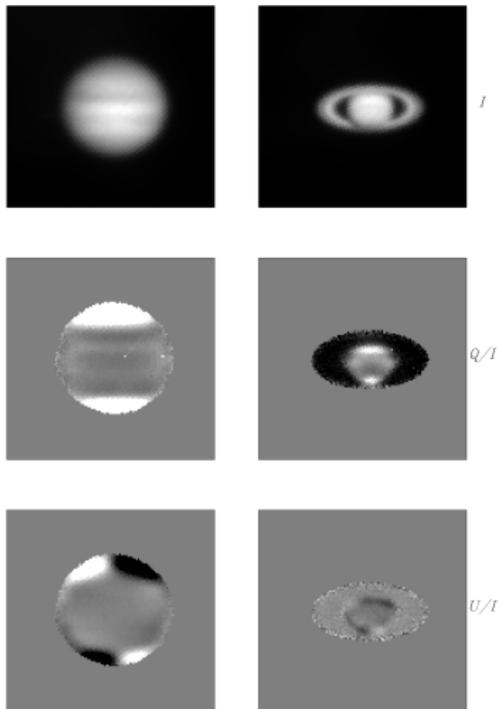
- due to anisotropy of the radiation field
- anisotropy due to limb darkening
- limb darkening due to decreasing temperature with height
- last scattering approximation without radiative transfer

Solar Spectral Line Scattering Polarization



resonance lines exhibit “large” scattering polarization signals

Jupiter and Saturn



(courtesy H.M.Schmid and D.Gisler)

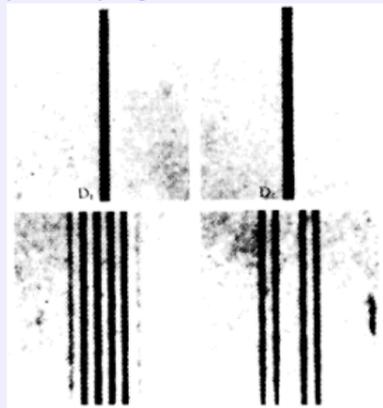
Planetary Scattered Light

- Jupiter, Saturn show scattering polarization
- multiple scattering changes polarization as compared to single scattering
- much depends on cloud height
- can be used to study extrasolar planetary systems
- ExPo instrument development at UU

Zeeman Effect



photos.aip.org/



Splitting/Polarization of Spectral Lines

- discovered in 1896 by Dutch physicist Pieter Zeeman
- different spectral lines show different splitting patterns
- splitting proportional to magnetic field
- split components are polarized
- *normal Zeeman effect* with 3 components explained by H.A.Lorentz using classical physics
- splitting of sodium D doublet could not be explained by classical physics (*anomalous Zeeman effect*)
- quantum theory and electron's intrinsic spin led to satisfactory explanation

Quantum-Mechanical Hamiltonian

- classical interaction of magnetic dipol moment $\vec{\mu}$ and magnetic field given by magnetic potential energy

$$U = -\vec{\mu} \cdot \vec{B}$$

$\vec{\mu}$ the magnetic moment and \vec{B} the magnetic field vector

- magnetic moment of electron due to orbit and spin
- Hamiltonian for quantum mechanics

$$H = H_0 + H_1 = H_0 + \frac{e}{2mc} (\vec{L} + 2\vec{S}) \cdot \vec{B}$$

H_0 Hamiltonian of atom without magnetic field

H_1 Hamiltonian component due to magnetic field

e charge of electron

m electron rest mass

\vec{L} the orbital angular momentum operator

\vec{S} the spin operator

Energy States in a Magnetic Field

- energy state $\langle E_{NLSJ} |$ characterized by
 - main quantum number N of energy state
 - $L(L + 1)$, the eigenvalue of \vec{L}^2
 - $S(S + 1)$, the eigenvalue of \vec{S}^2
 - $J(J + 1)$, the eigenvalue of \vec{J}^2 ,
 $\vec{J} = \vec{L} + \vec{S}$ being the total angular momentum
 - M , the eigenvalue of J_z in the state $\langle NLSJM |$
- for the magnetic field in the z-direction, the change in energy is given by

$$\Delta E_{NLSJ}(M) = \langle NLSJM | H_1 | NLSJM \rangle$$

The Landé g Factor

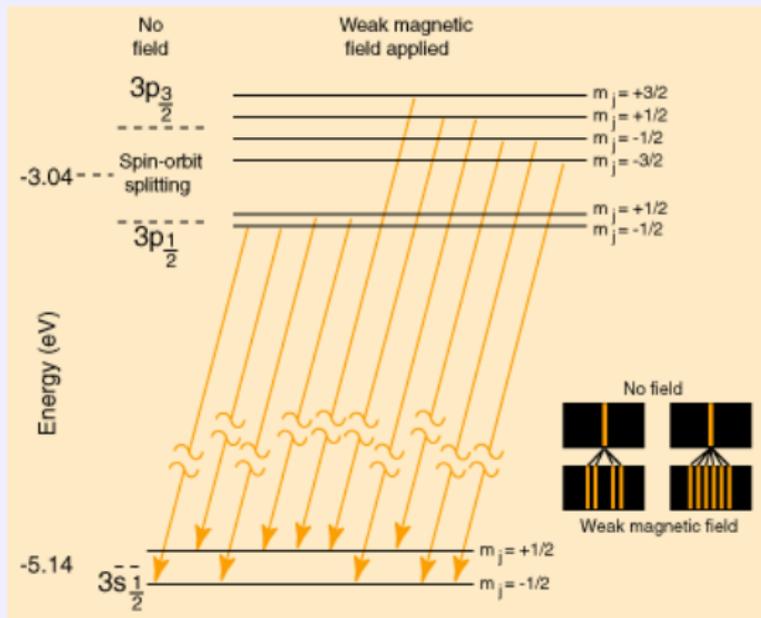
- based on pure mathematics (group theory, Wigner-Eckart theorem), one obtains

$$\Delta E_{NLSJ}(M) = \mu_0 g_L B M$$

with $\mu_0 = \frac{e\hbar}{2m}$ the Bohr magneton, and g_L the Landé g-factor

- in LS coupling where B sufficiently small compared to spin-orbit splitting field

$$g_L = 1 + \frac{J(J+1) + L(L+1) - S(S+1)}{2J(J+1)}$$



hyperphysics.phy-astr.gsu.edu/hbase/quantum/sodzee.html

Spectral Lines - Transitions between Energy States

- spectral lines are due to transitions between energy states:
 - lower level with $2J_l + 1$ sublevels M_l
 - upper level with $2J_u + 1$ sublevels M_u
- not all transitions occur

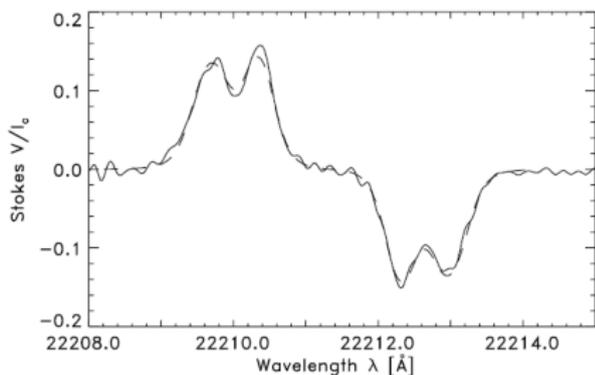
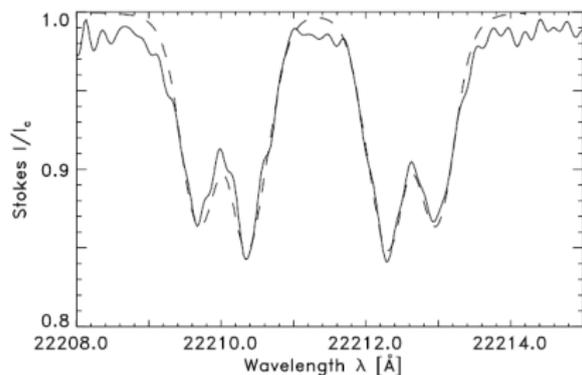
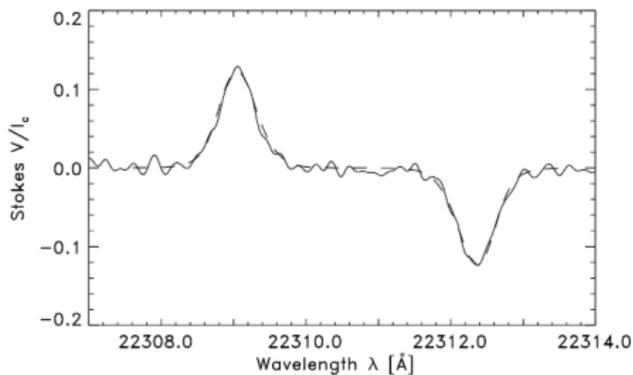
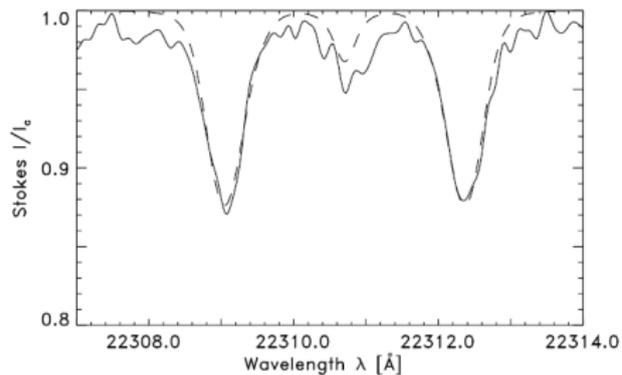
Selection rule

- not all transitions between two levels are allowed
- assuming dipole radiation, quantum mechanics gives us the *selection rules*:
 - $L_u - L_l = \Delta L = \pm 1$
 - $M_u - M_l = \Delta M = 0, \pm 1$
 - $M_u = 0$ to $M_l = 0$ is forbidden for $J_u - J_l = 0$
- total angular momentum conservation: photon always carries $J_{\text{photon}} = 1$
- *normal Zeeman effect*: line splits into three components because
 - Landé g-factors of upper and lower levels are identical
 - $J_u = 1$ to $J_l = 0$ transition
- *anomalous Zeeman effect* in all other cases

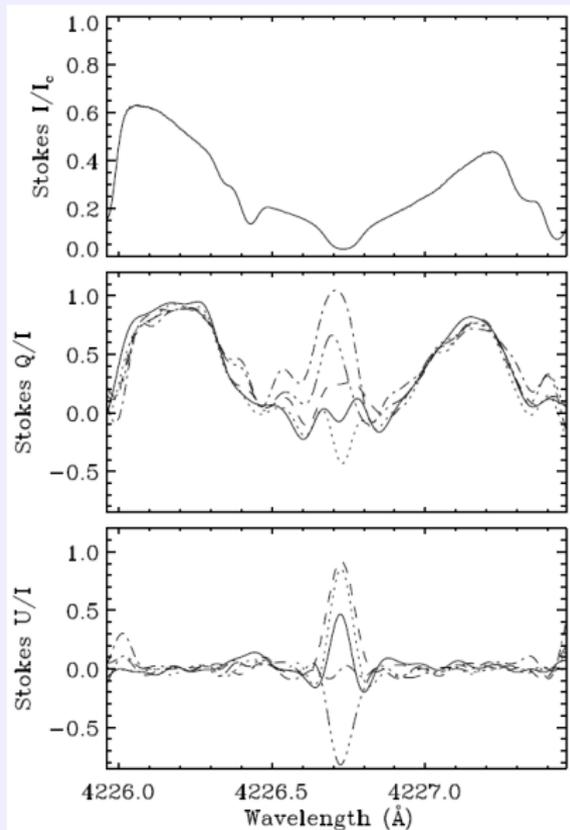
Effective Landé Factor and Polarized Components

- each component can be assigned an effective Landé g-factor, corresponding to how much the component shifts in wavelength for a given field strength
- components are also grouped according to the linear polarization direction for a magnetic field perpendicular to the line of sight
 - π components are polarized parallel to the magnetic field (**pi** for *parallel*)
 - σ components are polarized perpendicular to the magnetic field (**sigma** for German *senkrecht*)
- for a field parallel to the line of sight, the π-components are not visible, and the σ components are circularly polarized

Fully Split Titanium Lines at 2.2 μm



Hanle Effect



Depolarization and Rotation

- scattering polarization modified by magnetic field
- precession around magnetic field depolarizes and rotates polarization
- sensitive $\sim 10^3$ times smaller field strengths than Zeeman effect
- measurable effects even for isotropic field vector orientations