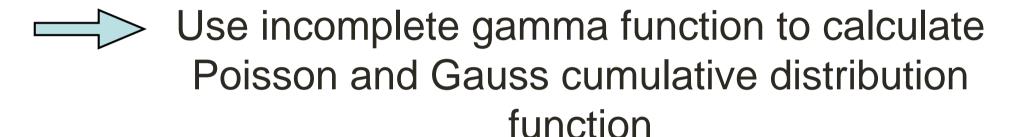
## Recap lecture 4

thermal limit of stochastic radiation processes  $h\nu << kT$  Thermal limit  $\Delta P^2(\nu) = \bar{P}^2(\nu)$ 



Propagation of errors under the assumption of independent variables

$$\bar{f} = f(\bar{u}, \bar{v}, ..)$$

$$\sigma_f^2 = \sigma_u^2 (\frac{\partial f}{\partial u})^2 + \sigma_v^2 (\frac{\partial f}{\partial v})^2 + ...$$

## Today:

Comparing data with a model: Leastsquares fitting, maximum likelihood method: Gaussian data

Monte Carlo simulations

"real" maximum likelihood method:

Poissonian data

OAF2 chapter 5.3+5.4 see also Num Res Chapter 15

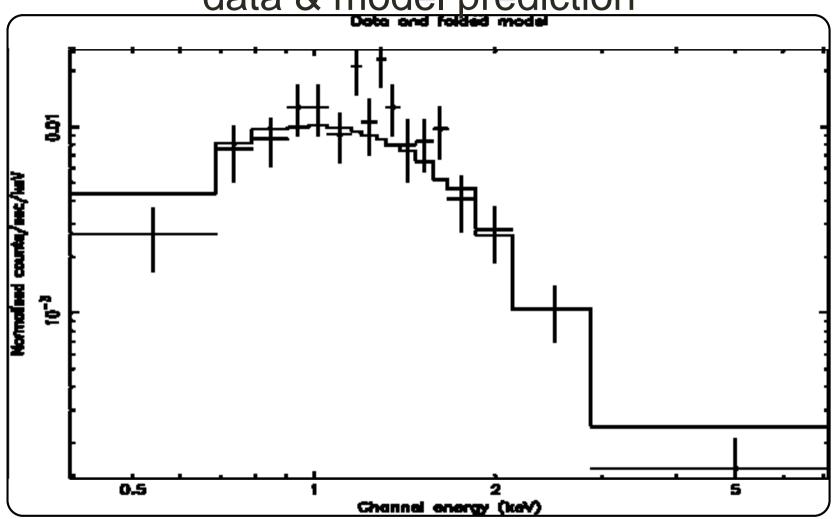
## Compare data with a model

Describe data in terms of a continuous function

Compare observations (data) with theoretical model prediction

Describe the data in a few parameters

Example: continuous model through discrete data & model prediction



X-ray binary in quiescence neutron star atmosphere model

# Maximum likelihood: most likely outcome is assumed to be the 'correct' one Method of least squares

$$P(y_i)\Delta y = \frac{1}{\sqrt{2\pi}\sigma_i} exp(-\frac{1}{2} \frac{(y_i - y_m)^2}{\sigma_i^2}) \Delta y$$

note:  $y_m$  = model value not mean here!

## Method of least squares

$$P \propto \prod_{i=1}^{N} \{exp[-rac{1}{2}(rac{y_i-y_m}{\sigma_i})^2]\}$$

$$\propto exp[-rac{1}{2}\sum_{i=1}^{N}(rac{y_i-y_m}{\sigma_i})^2]\}$$

minimise: 
$$\chi^2 \equiv \sum_{i=1}^{N} (\frac{y_i - y_m}{\sigma_i})^2$$

minimisation: root finding problem

1D: 
$$\frac{\partial}{\partial y_i} \sum_{i=1}^{N} (\frac{y_i - y_m}{\sigma_i})^2 = 0$$

## more about $\chi^2$

drawn from normal distribution distribution of  $x_i^2$  is a  $x^2$  distribution for N measurements described by M variables, there are N-M Degrees of Freedom (d.o.f.)

Probability of obtaining a certain  $\chi^2$  or higher by chance  $P(\chi^2_{obs}) = \operatorname{gammq}(\frac{N-M}{2}, \frac{\chi^2_{obs}}{2})$ 

x² fitting provides:

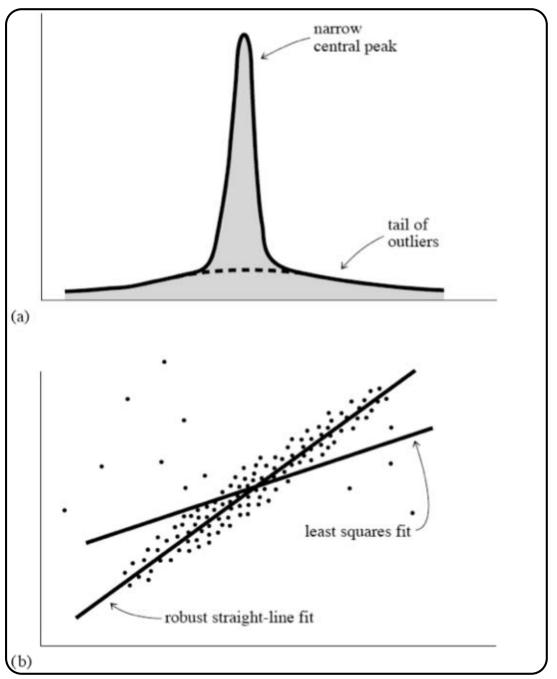
Best-fitting parameters

An error estimate of the uncertainty of the fitted parameters

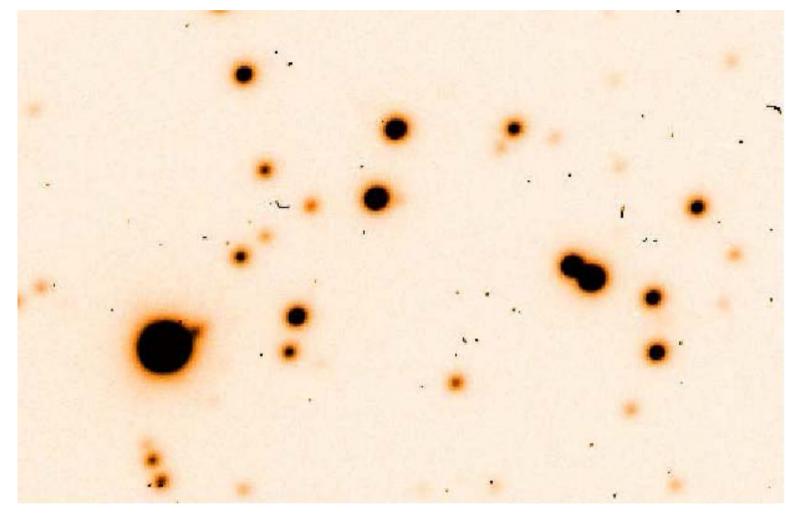
A probability that the data is drawn from a parent population described by the model parameters

Note that outliers make this probability generally low

## Be aware of non-gaussian distributions

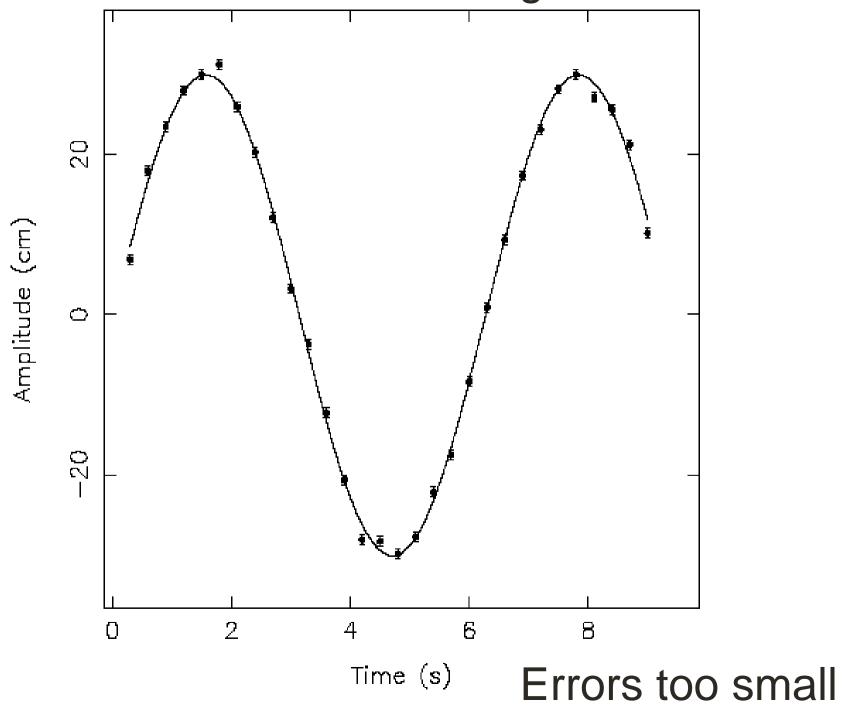


Part of U-band image VLT

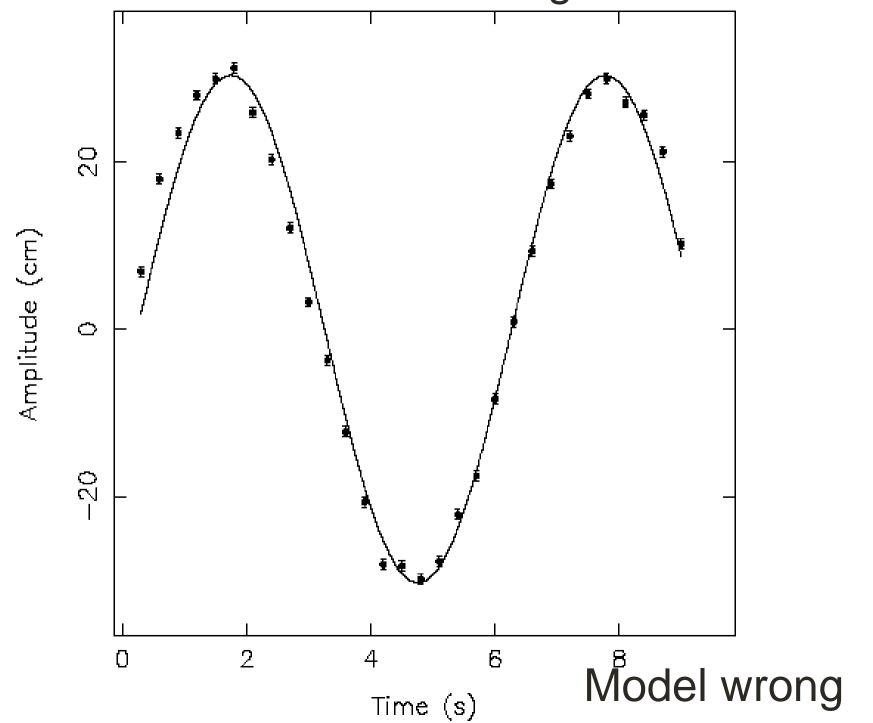


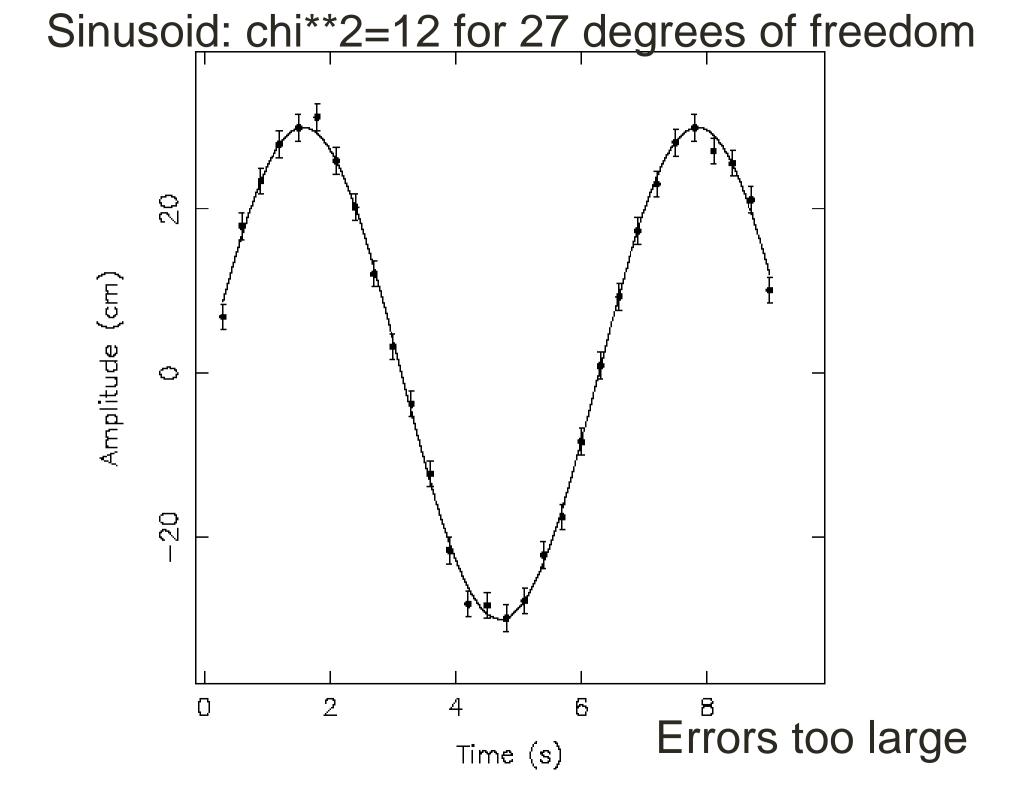
Remove outliers via (sigma) clipping

## Sinusoid: chi\*\*2=81.2 for 26 degrees of freedom

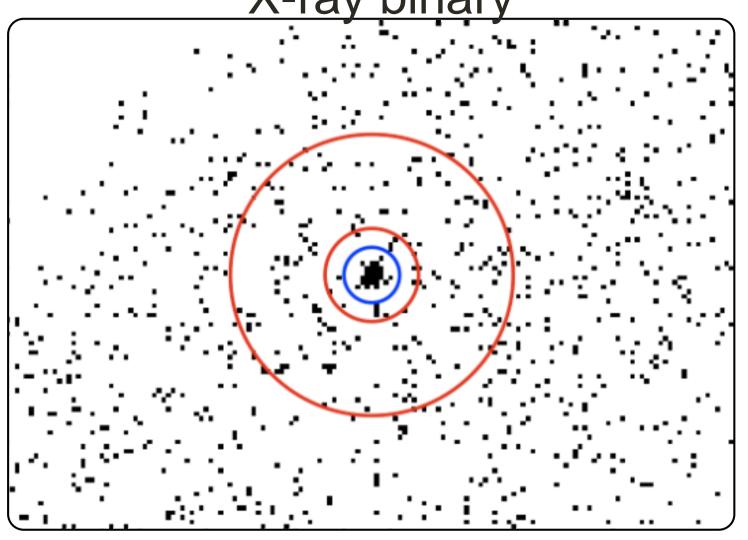


## Sinusoid: chi\*\*2=580 for 28 degrees of freedom

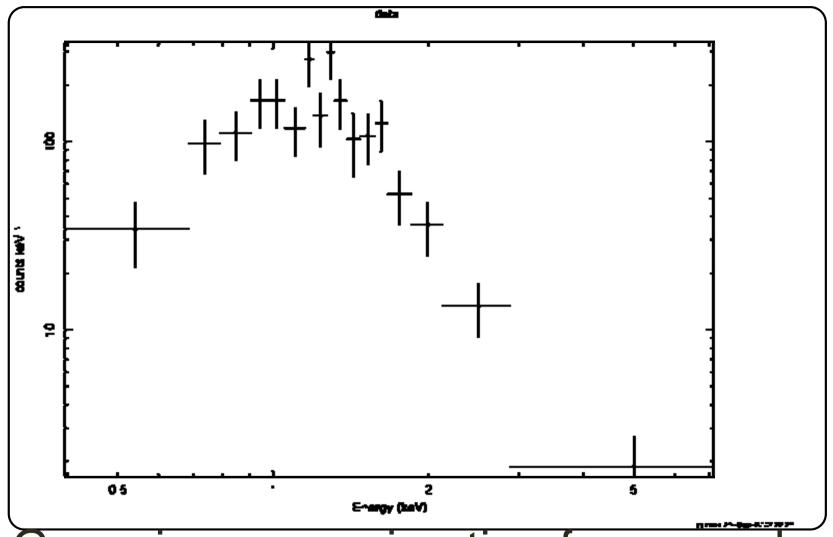




Chandra CCD (ACIS) observation of an X-ray binary

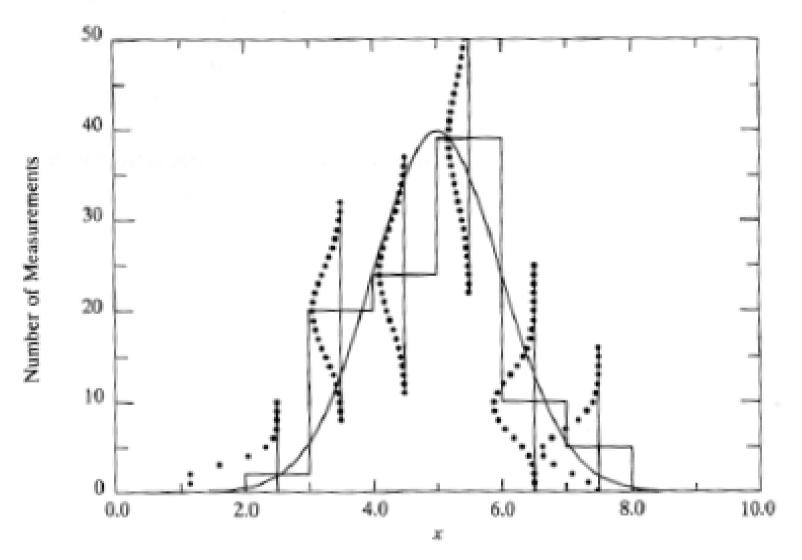


## Same data as before



Gaussian approximation for errors but at low counts Gauss and Poisson errors differ

### VALUE AND POISSON ERRORS



EXAMPLE FROM BEVINGTON & ROBERTSON 1992

## Fitting a straight line to the data

$$y_m(x_i, a, b) = a + bx_i$$

minimise  $\chi_i^2$  to find best-fitting parameters

$$\frac{\partial \sum_{i=1}^{N} (\frac{y_i - a - bx_i}{\sigma_i})^2}{\partial a} = 0$$

$$\sum \frac{y_i}{\sigma_i^2} - a \sum \frac{1}{\sigma_i^2} - b \sum \frac{x_i}{\sigma_i^2} = 0$$

$$\sum \frac{x_i y_i}{\sigma_i^2} - a \sum \frac{x_i}{\sigma_i^2} - b \sum \frac{x_i^2}{\sigma_i^2} = 0$$

$$a = \frac{\sum_{i} \frac{x_i^2}{\sigma_i^2} \sum_{i} \frac{y_i}{\sigma_i^2} - \sum_{i} \frac{x_i}{\sigma_i^2} \sum_{i} \frac{x_i y_i}{\sigma_i^2}}{\sum_{i} \frac{1}{\sigma_i^2} \sum_{i} \frac{x_i^2}{\sigma_i^2} - (\sum_{i} \frac{x_i}{\sigma_i^2})^2}$$

determine errors on the best-fitting parameters

remember 
$$\sigma_f^2 = \sigma_u^2 (\frac{\partial f}{\partial u})^2 + \sigma_v^2 (\frac{\partial f}{\partial v})^2 + \dots$$

$$\sigma_a^2 = \sum_{i=1}^N [\sigma_i^2 \frac{\partial a}{\partial y_i}]^2$$

 $\partial u \& \partial v$  etc are the different measurement values  $y_i$ 

$$\sigma_a^2 = \frac{\sum_i \frac{x_i^2}{\sigma_i^2}}{\sum_i \frac{1}{\sigma_i^2} \sum_i \frac{x_i^2}{\sigma_i^2} - (\sum_i \frac{x_i}{\sigma_i^2})^2}$$
similarly
$$\sigma_b^2 = \frac{\sum_i \frac{1}{\sigma_i^2}}{\sum_i \frac{1}{\sigma_i^2} \sum_i \frac{x_i^2}{\sigma_i^2} - (\sum_i \frac{x_i}{\sigma_i^2})^2}$$

Finally calculate the probability of obtaining the by chance

$$P(\chi_{obs}^2) = \operatorname{gammq}(\frac{N-M}{2}, \frac{\chi_{obs}^2}{2})$$

for the straight line fit m=2

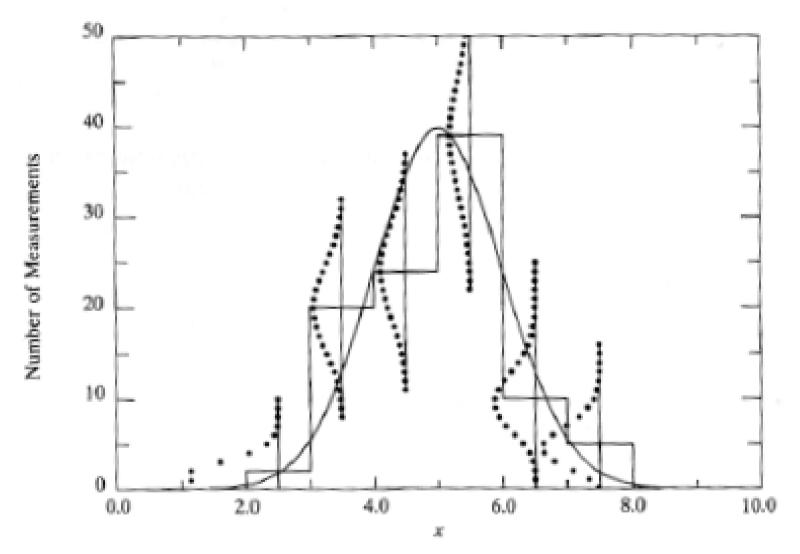
$$\nu = N - M$$
 degrees of freedom reduced  $\chi^2_{\nu}$ 

$$\chi^2_
u \equiv rac{\chi^2}{
u}$$

for data fitting:

$$\chi^2_{
u} \sim 1$$
  $\chi^2 \approx 
u$   $\sigma_{\chi^2} = \sqrt{2
u}$ 

#### VALUE AND POISSON ERRORS



EXAMPLE FROM BEVINGTON & ROBERTSON 1992

Estimating confidence regions via Monte Carlo simulations

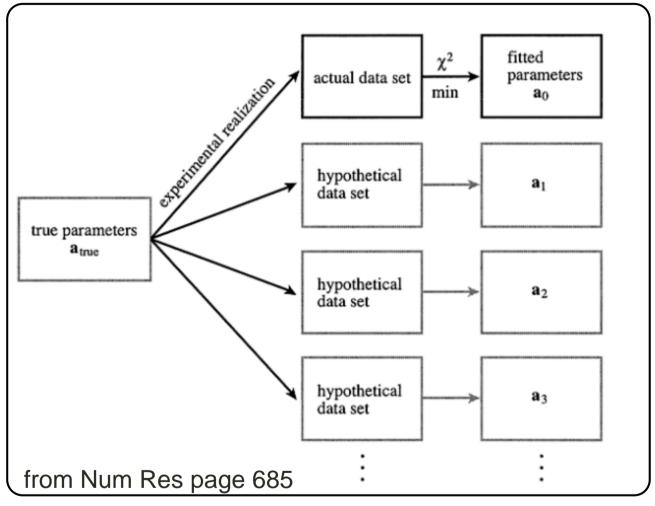
## Monte Carlo simulations

1 Replace the observed values by another random selected value from the range y+-sigma

2 Repeat fitting (chi\*\*2 minimisation etc)

3 Repeat 1 & 2 N times to build up a distribution in the determined parameters and from that determine the mean, variance etc

## Estimating confidence limits Monte Carlo simulations



Measurement one draw from the distribution of a's

assume that the distribution of  $a_i-a_0$  is close to the probability distribution  $a_i-a_{true}$ 

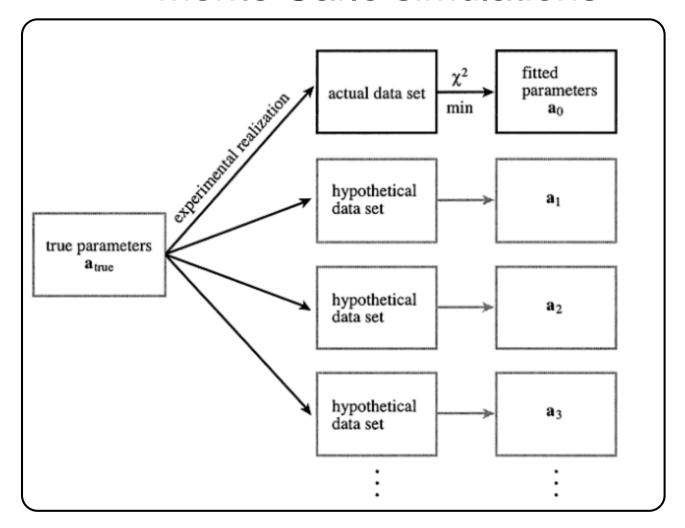
# $a_i - a_0$ distribution we can determine via Monte Carlo simulations

Many Monte Carlo methods

Basis: (pseudo) random draws

Also simulate an experiment! a.o. useful for proposal writing Computer exercise

### Monte Carlo simulations



calculate distribution of  $a_i - a_0$ by simulating many sets of data and using  $\chi^2$  fitting to determine  $a_i$ 

## Special MC: bootstrapping

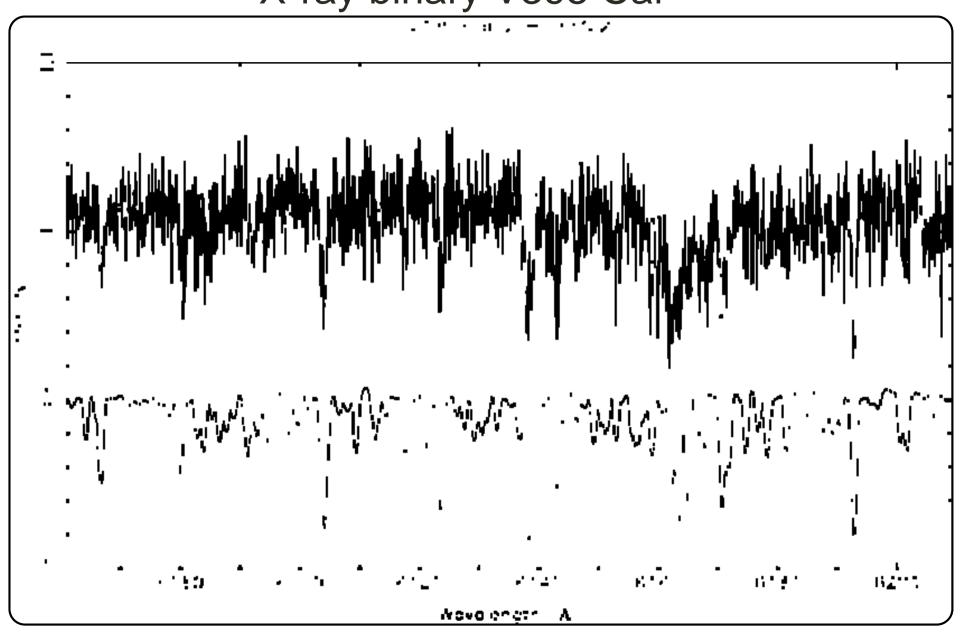
1 Replace a random number of observed values by another random selected observed value

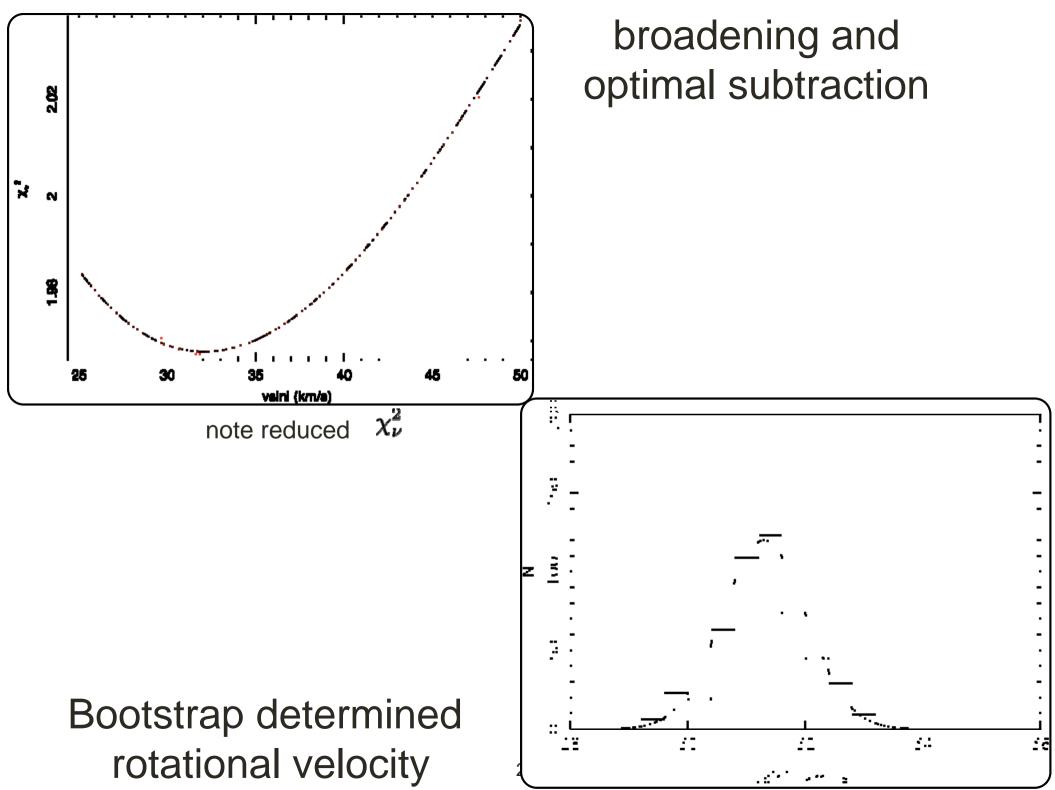
2 Repeat fitting (chi\*\*2 minimisation etc)

3 Repeat 1 & 2 N times to build up a distribution in the determined parameters and from that determine the mean, variance etc

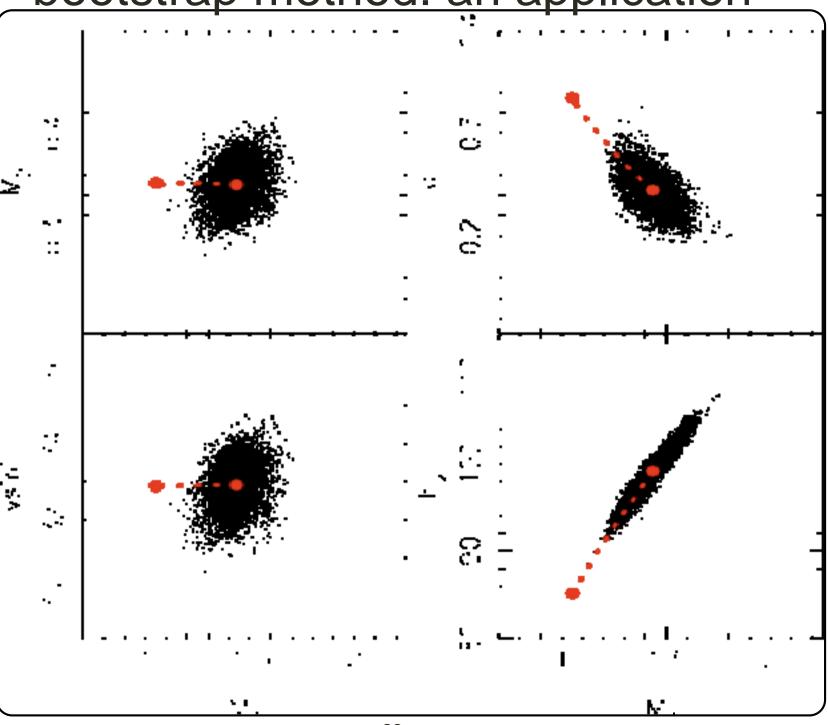
## bootstrap method and application

## X-ray binary V395 Car

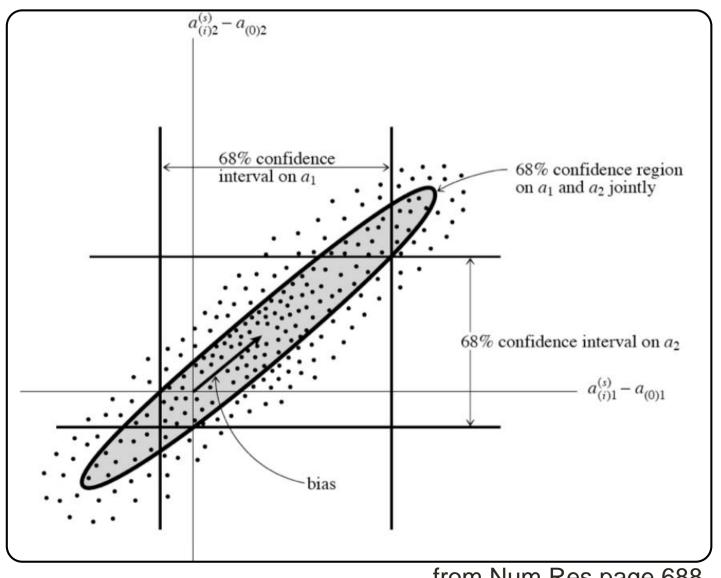




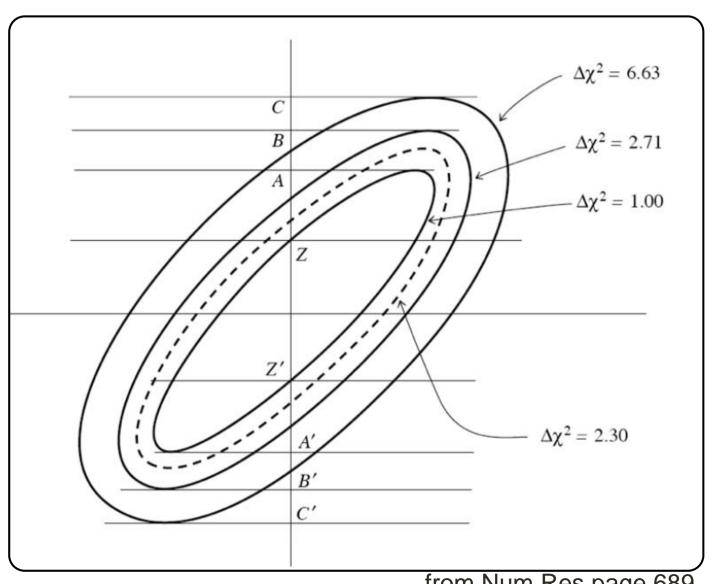
bootstrap method: an application



## Confidence limits single vs. multiple parameter confidence region



## **Projections**



## Maximum likelihood method (Poisson noise, unbinned data)

probability to find  $n_i$  photons when  $m_i$  expected

for each pixel in an image

$$P_i = \frac{m_i^{n_i} e^{-m_i}}{n_i!}$$

total probability  $L' \equiv \prod P_i$ 

$$L' \equiv \prod_i P_i$$

$$\ln L' \equiv \sum_{i} \ln P_{i} = \sum_{i} n_{i} \ln m_{i} - \sum_{i} m_{i} - \sum_{i} \ln n_{i}!$$

minimise 
$$\ln L \equiv -2(\sum_i n_i \ln m_i - \sum_i m_i)$$

Detection of a constant background, A, plus a source of strength B of which a fraction  $f_i$  falls on pixel i

$$-0.5 \ln L = \sum_{i} n_{i} \ln(A + Bf_{i}) - \sum_{i} (A + Bf_{i})$$

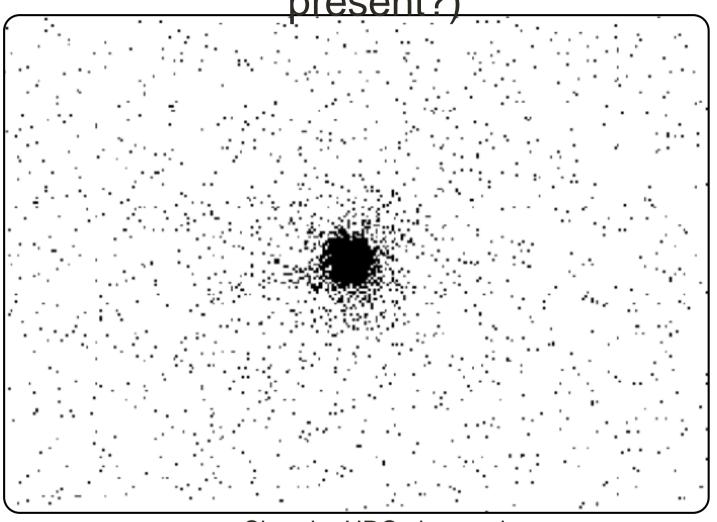
again search for the minimum of L for variations in A and B

determined independently in some cases total pixels Z

$$\frac{\partial \ln L}{\partial A} = 0 \Rightarrow \sum_{i} \frac{n_i}{A + Bf_i} - \sum_{i} (1) = \sum_{i} \frac{n_i}{A + Bf_i} - Z = 0$$

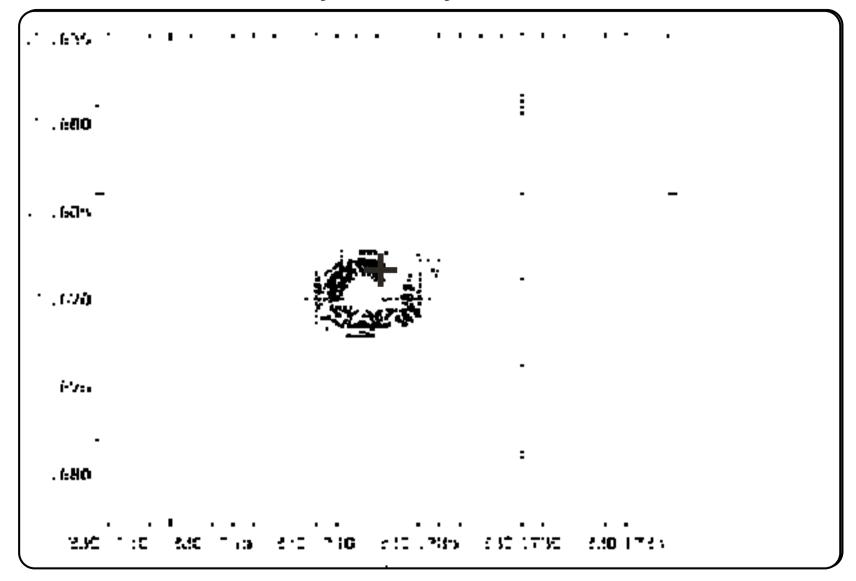
$$\frac{\partial \ln L}{\partial B} = 0 \Rightarrow \sum_{i} \frac{n_i f_i}{A + B f_i} - \sum_{i} (f_i) = \sum_{i} \frac{n_i f_i}{A + B f_i} - 1 = 0$$

Maximum likelihood method (application X-ray binary Cir X-1, a jet present?)



Chandra HRC observation model and subsequently subtract PSF only close to the source the assumption of a constant background is valid

# application maximum likelihood method X-ray binary Cir X-1



one source subtracted