

TODAY'S COURSE

CHAPTER 1.5-1.7, 2.2.2 OAF-2

TOPICS:

 RECAP: ALIASING & NYQUIST THEOREM

 (OPTIMAL) FILTERING

 MEASURING MOMENTS OF A S.P.

 STOCHASTIC NATURE OF RADIATION PROCESSES

SIGNAL DETECTION INVOLVES:

 LIMITED TIME INTERVAL → WINDOWING

 NOT CONTINUOUS → SAMPLING

 SAMPLES NOT INSTANTANEOUS → AVERAGING

 DEALING WITH NOISE → FILTERING

 RESPONSE OF THE DETECTION SYSTEM



THUS THE DETECTED SIGNAL WILL ONLY
APPROXIMATE THE SOURCE SIGNAL

POWER SPECTRAL DENSITY

(\propto AMPLITUDE OF INDIVIDUAL SINUSOIDS)

(WILL RETURN IN MORE DEPTH IN CHAPTER 6)

CONTINUOUS FT: $F(f) = \int_{-\infty}^{\infty} f(t) e^{-2\pi i f t} dt$

CONTINUOUS PSD: $P(f) = F(\tilde{f})F(\tilde{f})^*$

FOR WSS SIGNALS: $P(f) = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i f \tau} d\tau$

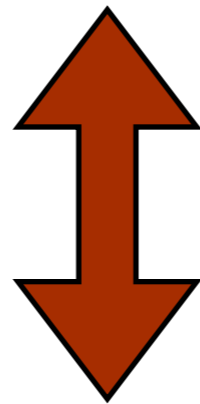
HENCE:

$$F(\tilde{f})F(\tilde{f}) = |F(f)|^2 = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i f \tau} d\tau$$

WIENER-KHINCHINE THEOREM

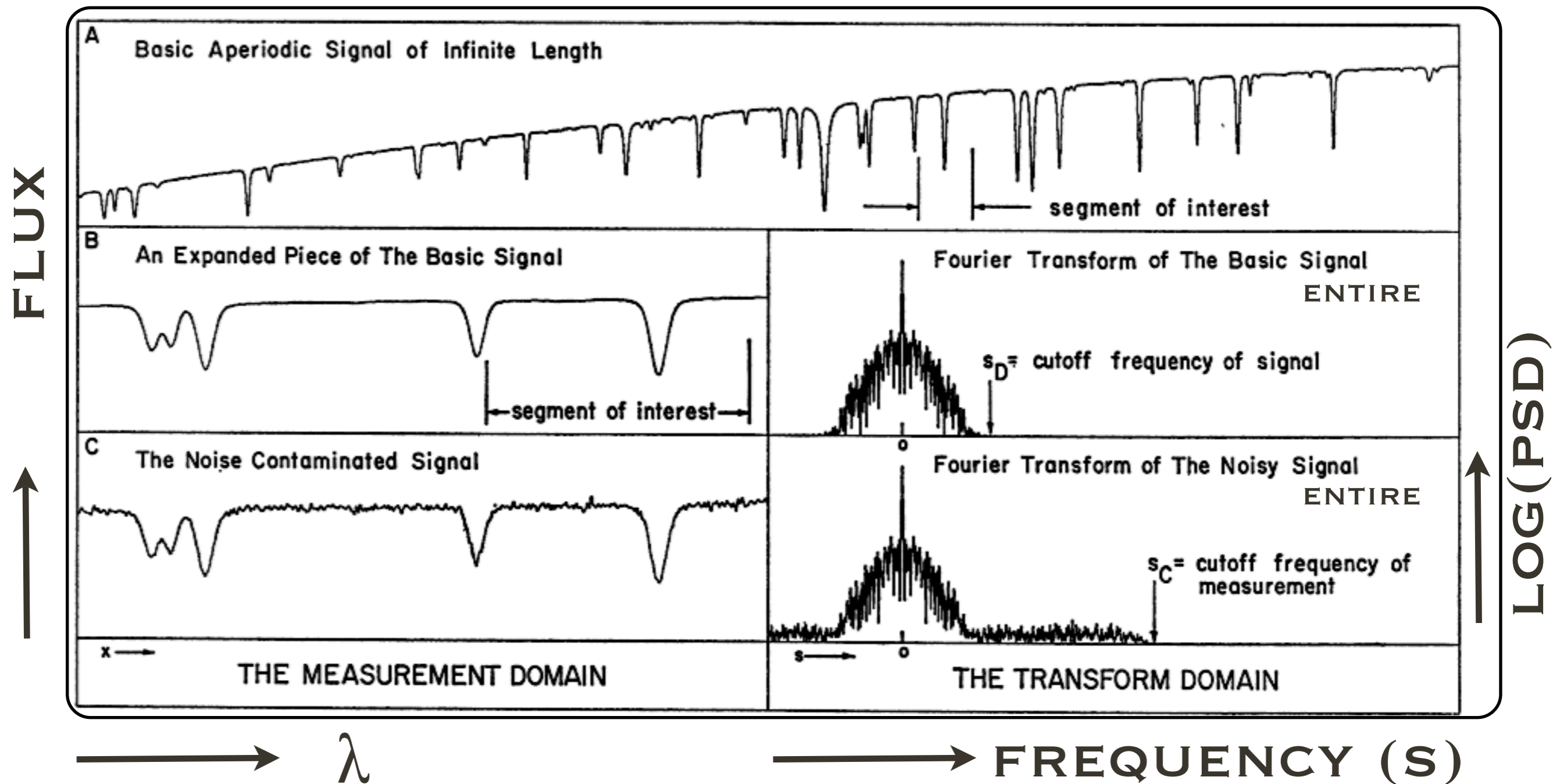
FOURIER TRANSFORM OF THE
AUTOCORRELATION OF $f(x)$ IS EQUAL TO THE
POWER DENSITY SPECTRUM $|F(s)|^2$

$$S(s) = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i s \tau} d\tau$$



$$R(t) = \int_{-\infty}^{\infty} S(s) e^{2\pi i t s} ds$$

THE NYQUIST THEOREM



NYQUIST THEOREM: CONT'D

DATA SAMPLING

DATA IS DISCRETE NOT CONTINUOUS

TIME DOMAIN MULTIPLY S.P. WITH SHAH FUNCTION

$$m_{samp,n} = m_s(x) = m(x) \frac{1}{\tau} \text{III}\left(\frac{x}{\tau}\right) = \sum_{n=0}^{N-1} m(n\tau) \delta(x - n\tau)$$

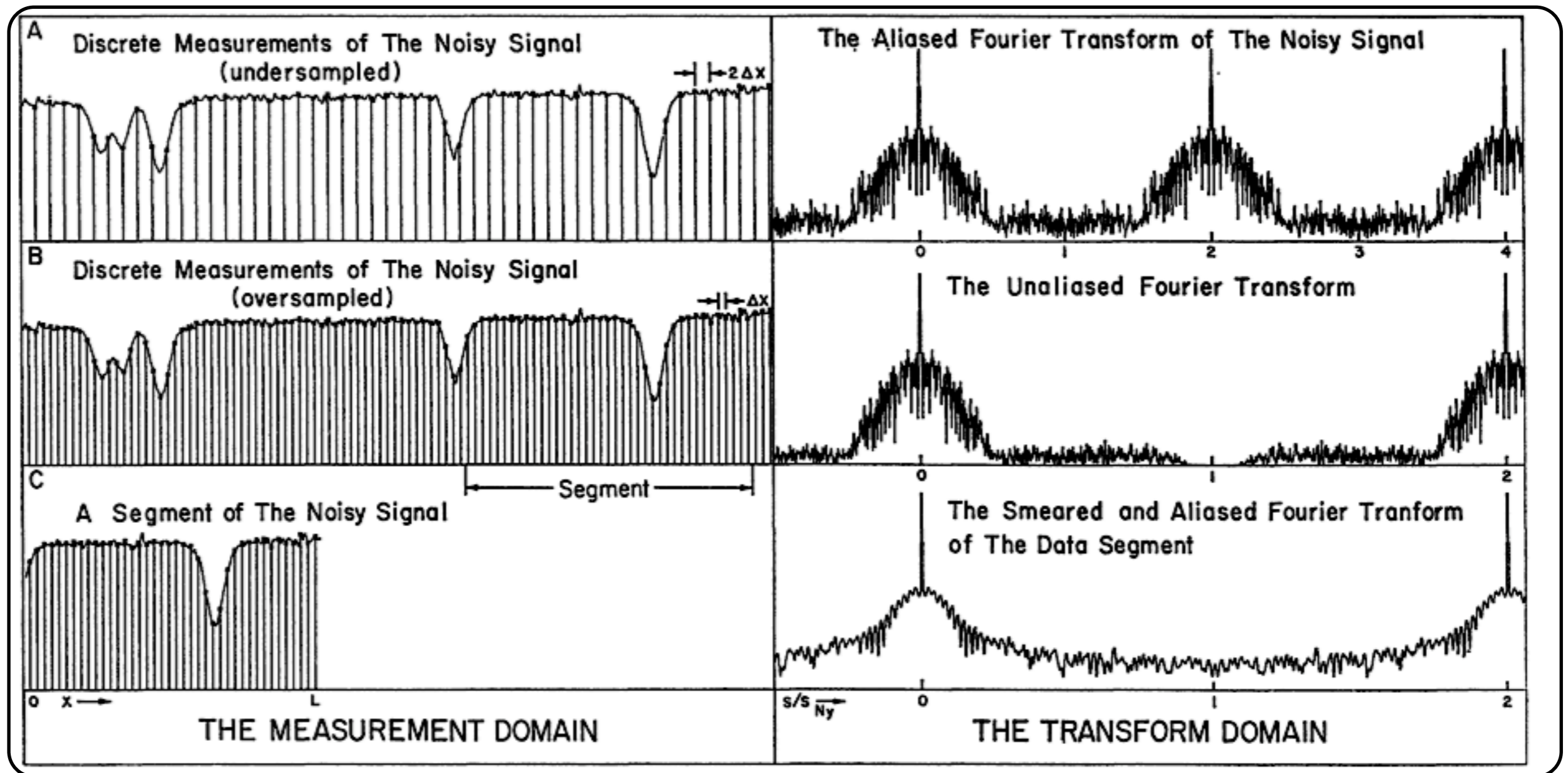
DISCRETE FT: $M_{samp,k} = \sum_{n=0}^{N-1} m_{samp,n} e^{2\pi i n k / N}$

DISCRETE PSD: $P_j = \frac{2}{a_0} |a_j|^2$ POWER \propto AMPLITUDE SQUARED:

J=K LABEL $a_0 = M_{samp,k=0} = \sum_{n=0}^{N-1} m_{samp,n} \equiv N_0$

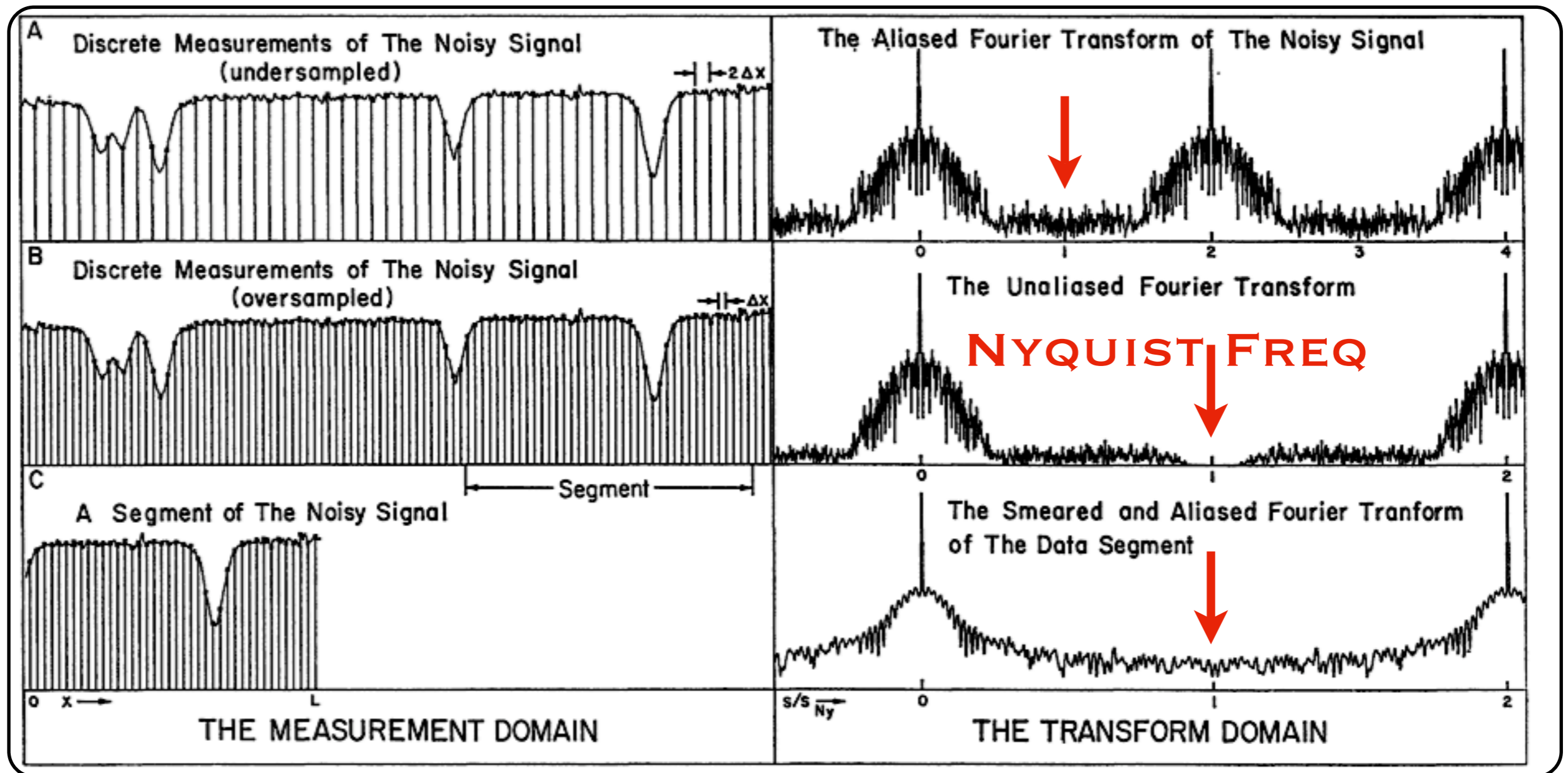
$$a_k = M_{samp,k} = \sum_{n=0}^{N-1} m_{samp,n} e^{2\pi i n k / N}$$

NYQUIST THEOREM: CONT'D



SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

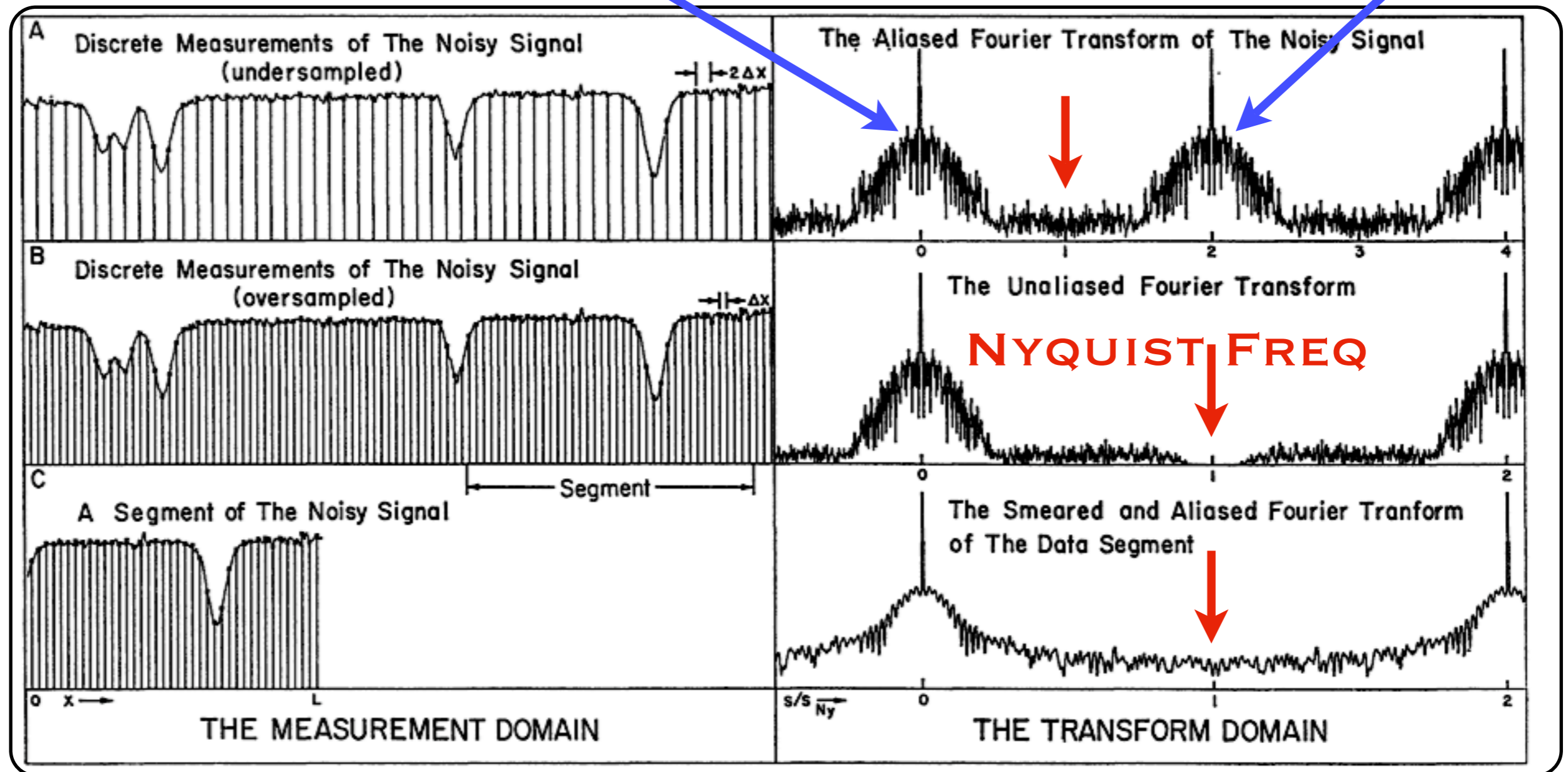
NYQUIST THEOREM: CONT'D



SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

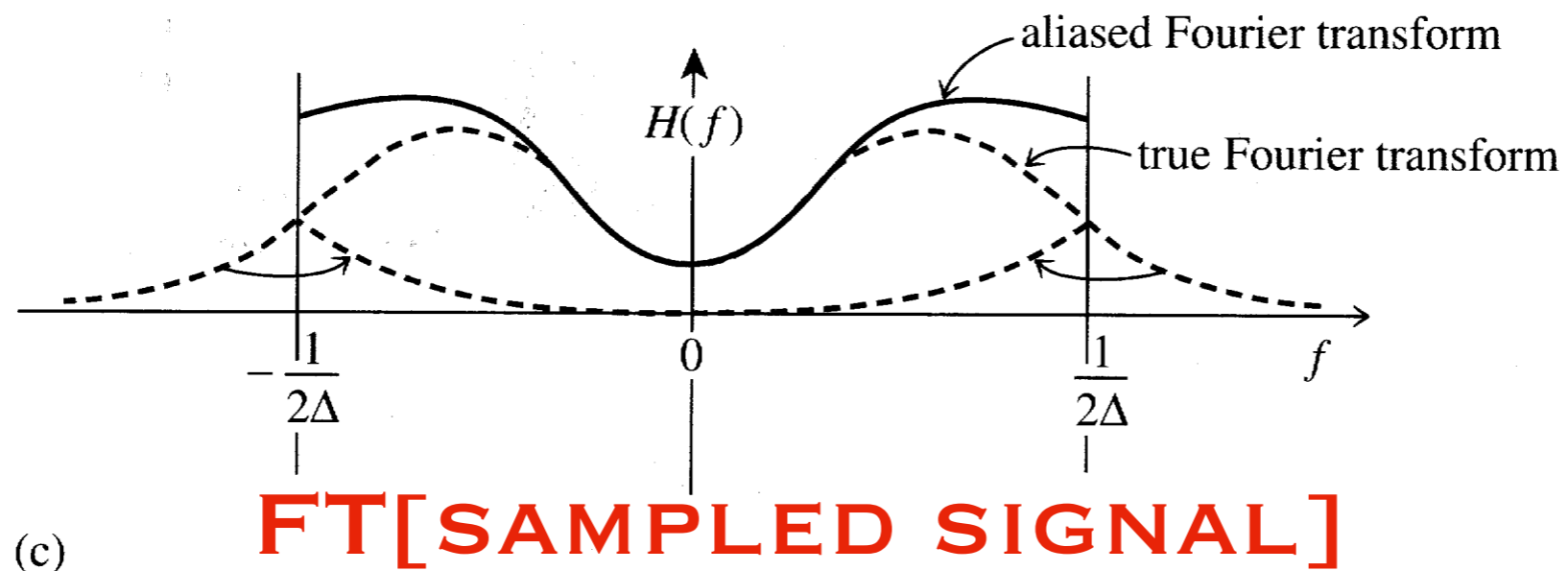
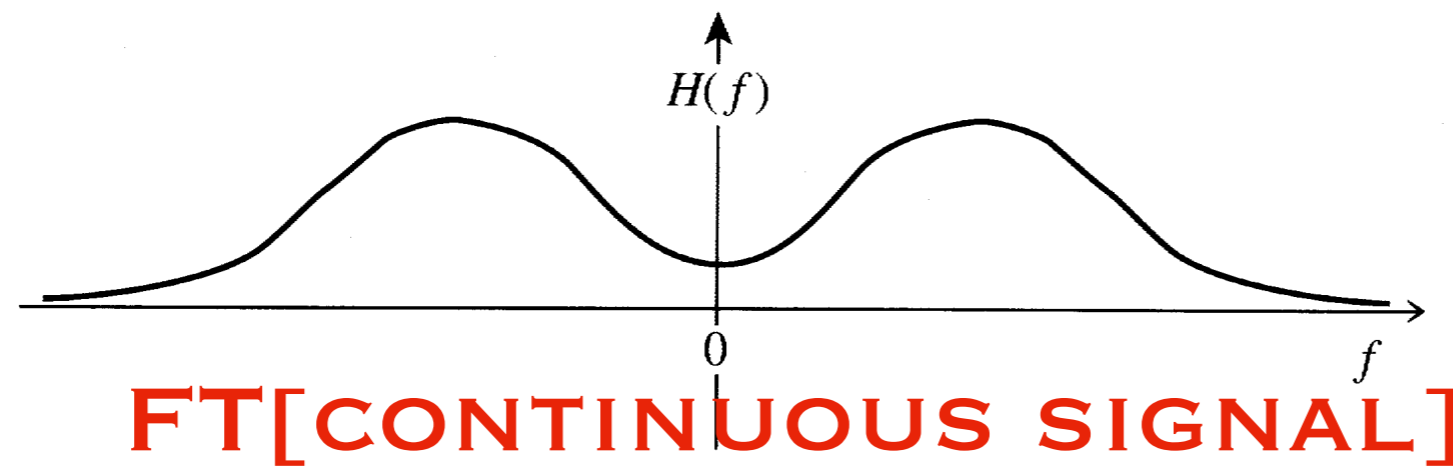
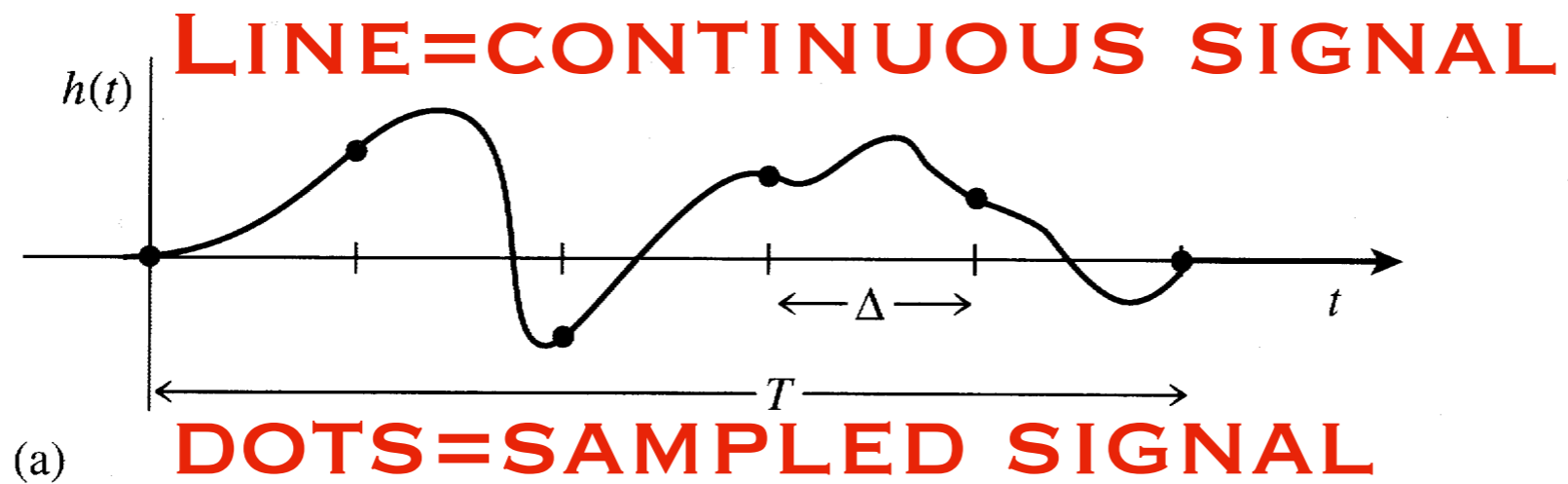
NYQUIST THEOREM: CONT'D

SAMPLING CAUSES REPLICATION OF SIGNAL



SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

ALIASING



RECAP NYQUIST & ALIASING

DATA SAMPLING: **IF** THE NYQUIST CRITERIUM IS FULFILLED (I.E. SAMPLE AT A RATE HIGHER THAN TWO TIMES THE HIGHEST FREQUENCY IN THE SIGNAL) **THEN** SAMPLING DOES NOT LEAD TO LOSS OF INFORMATION

CONDITIONS:

- ➔ BAND-LIMITED RESPONSE OF THE DETECTOR REMOVES HIGHEST NOISE POWERS AND THE SAMPLING IS FAST ENOUGH TO COVER THE BAND LIMIT OF THE DETECTOR
- ➔ SIGNAL IS BAND-LIMITED ALSO AND

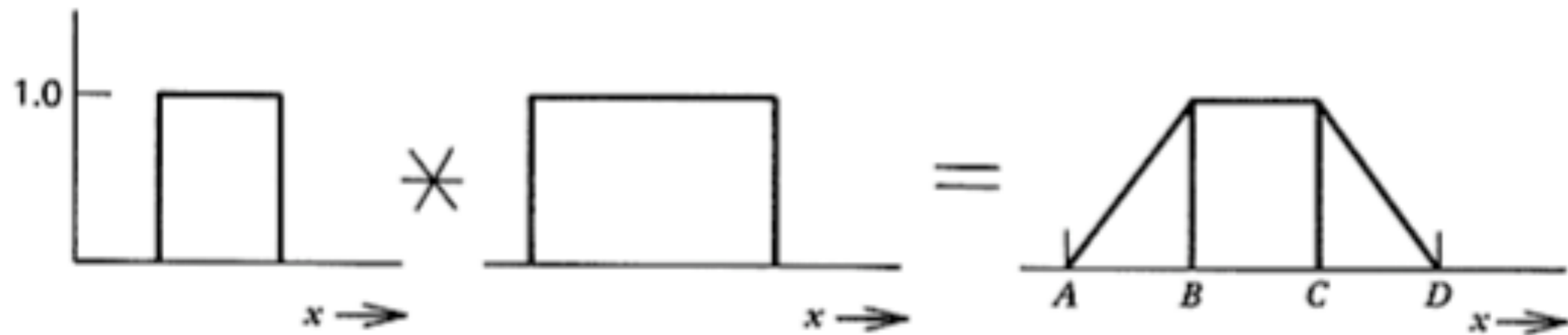
$$\nu_{\text{sampling}} > \nu_{\text{max,detector}} > \nu_{\text{max,signal}}$$

(OPTIMAL) FILTERING

NUM RES CHAPTER 13.0-13.3

DECONVOLVE MEASURED SIGNAL AND
RESPONSE FUNCTION OF SAMPLED DATA

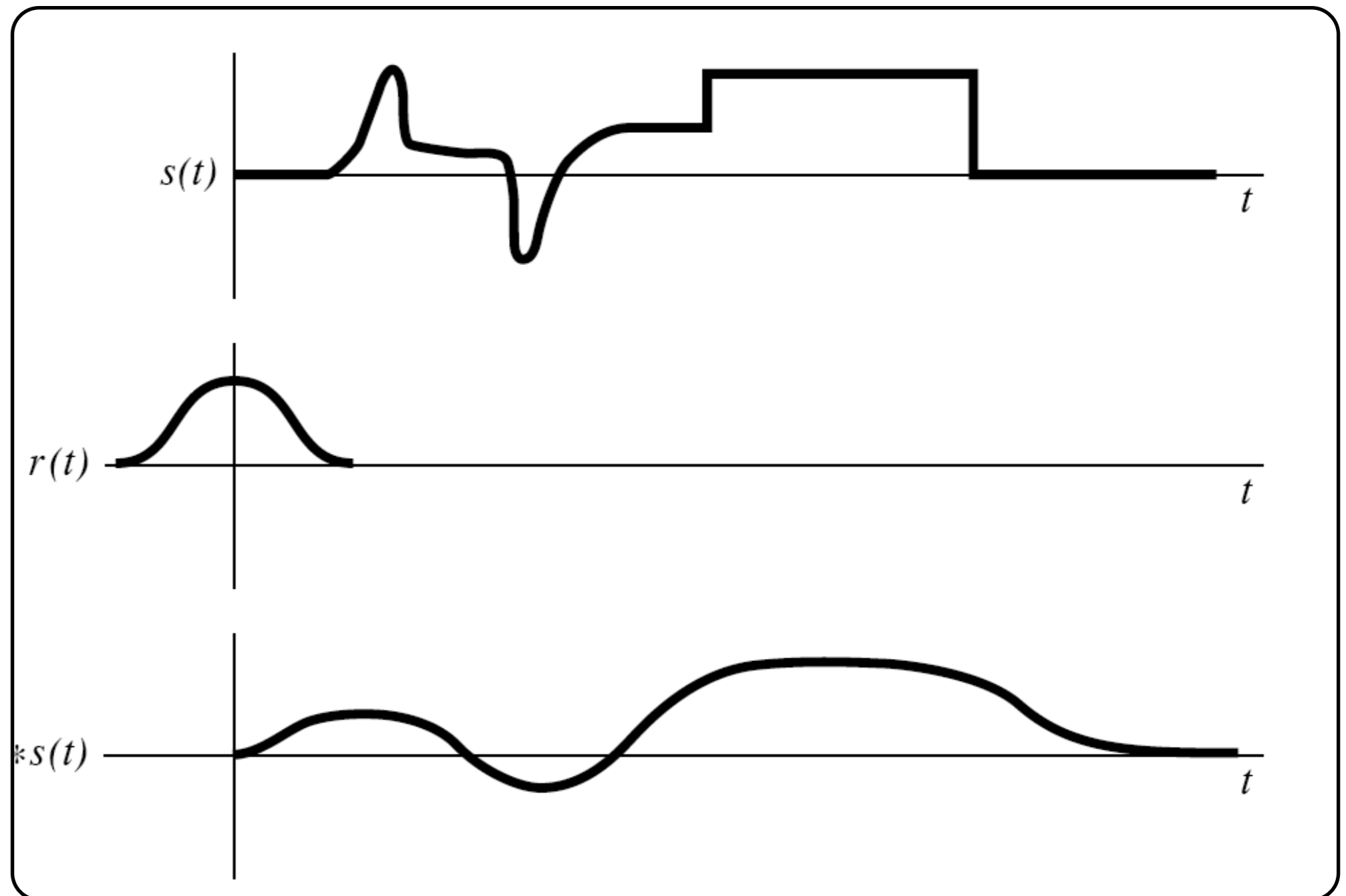
REMINDER CONVOLUTION



(OPTIMAL) FILTERING

NUM RES CHAPTER 13.0-13.3

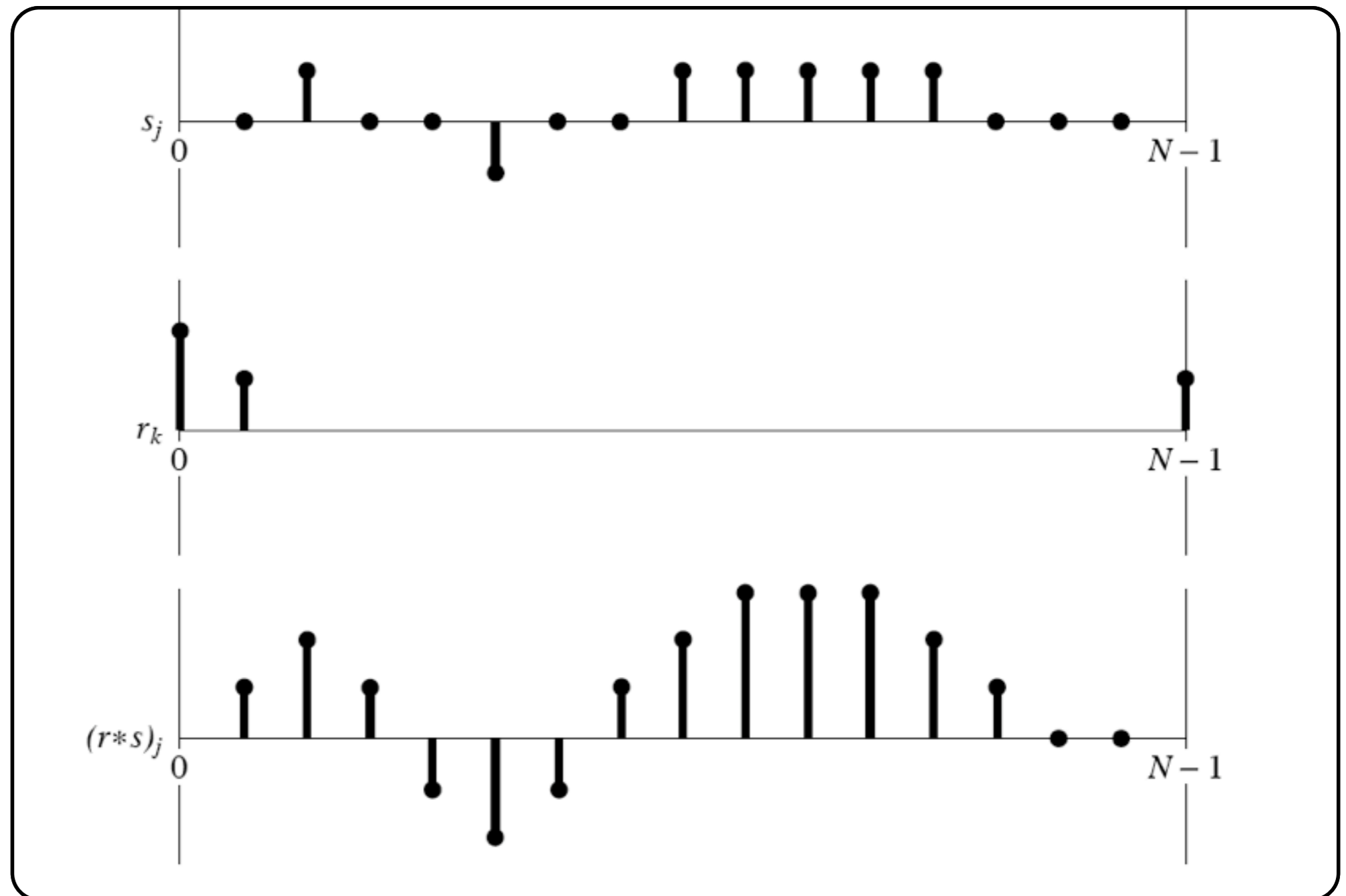
DECONVOLVE MEASURED SIGNAL AND
RESPONSE FUNCTION OF SAMPLED DATA



(OPTIMAL) FILTERING

NUM RES CHAPTER 13.0-13.3

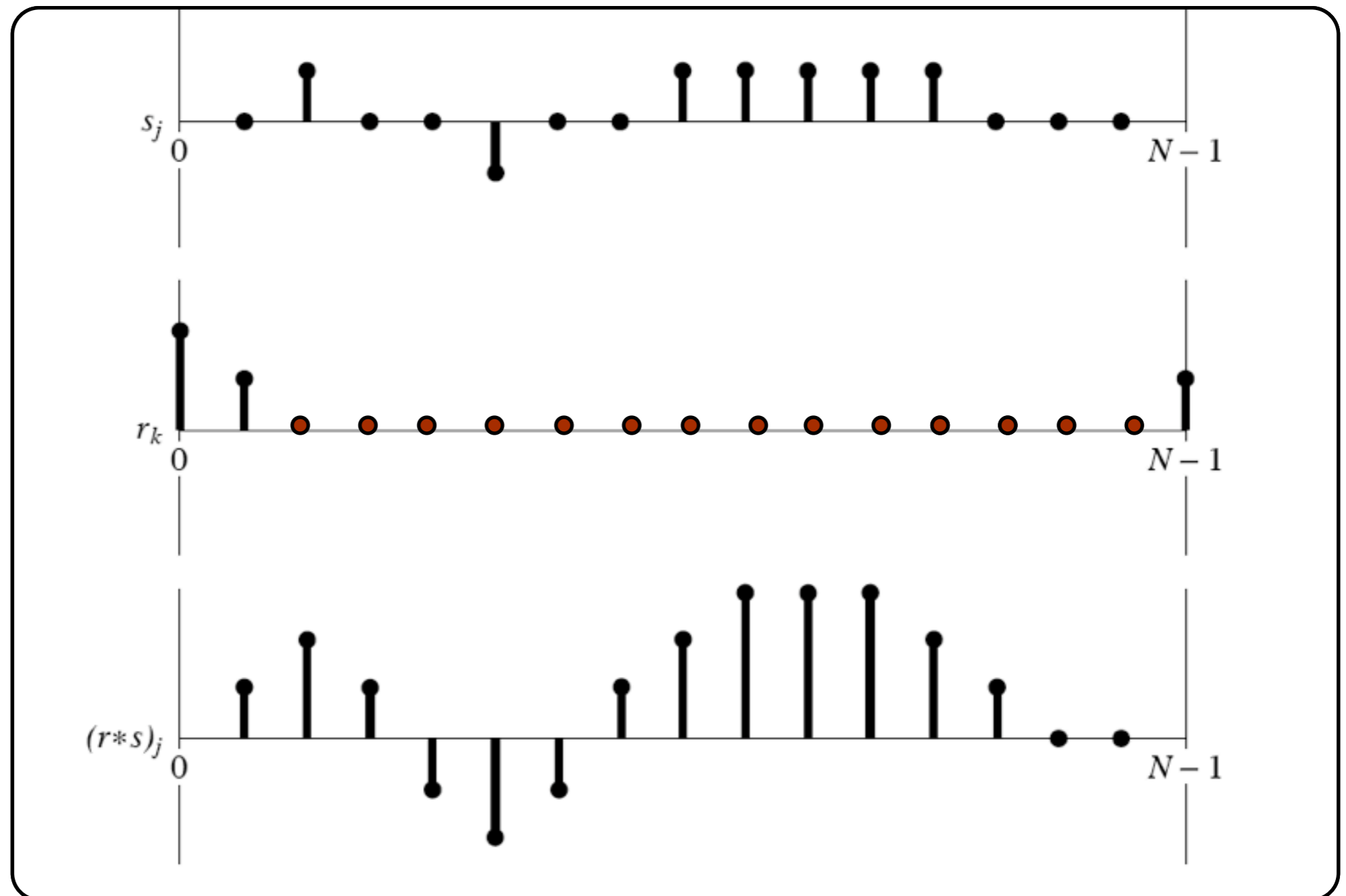
DECONVOLVE MEASURED SIGNAL AND
RESPONSE FUNCTION OF SAMPLED DATA



(OPTIMAL) FILTERING

NUM RES CHAPTER 13.0-13.3

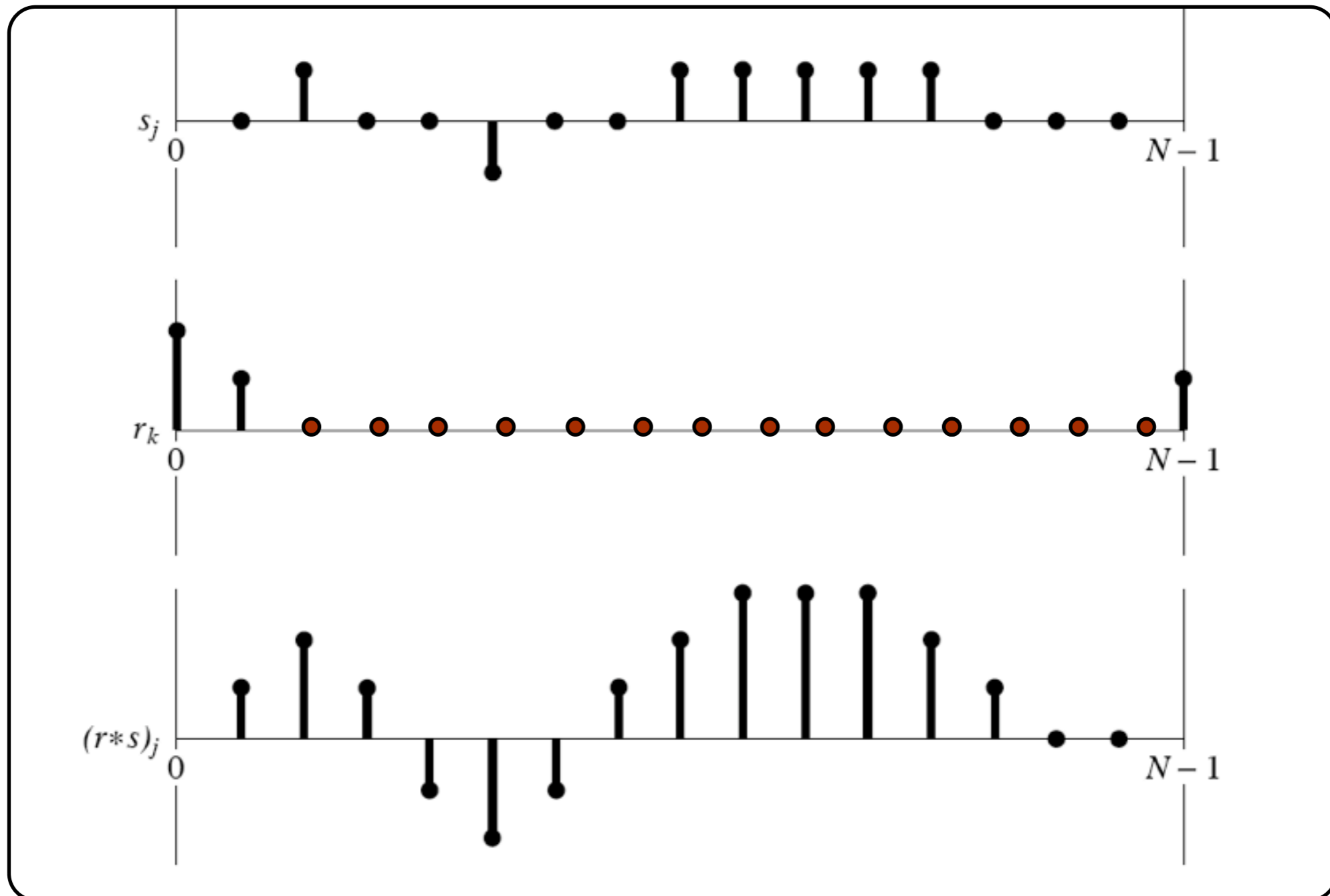
DECONVOLVE MEASURED SIGNAL AND
RESPONSE FUNCTION OF SAMPLED DATA



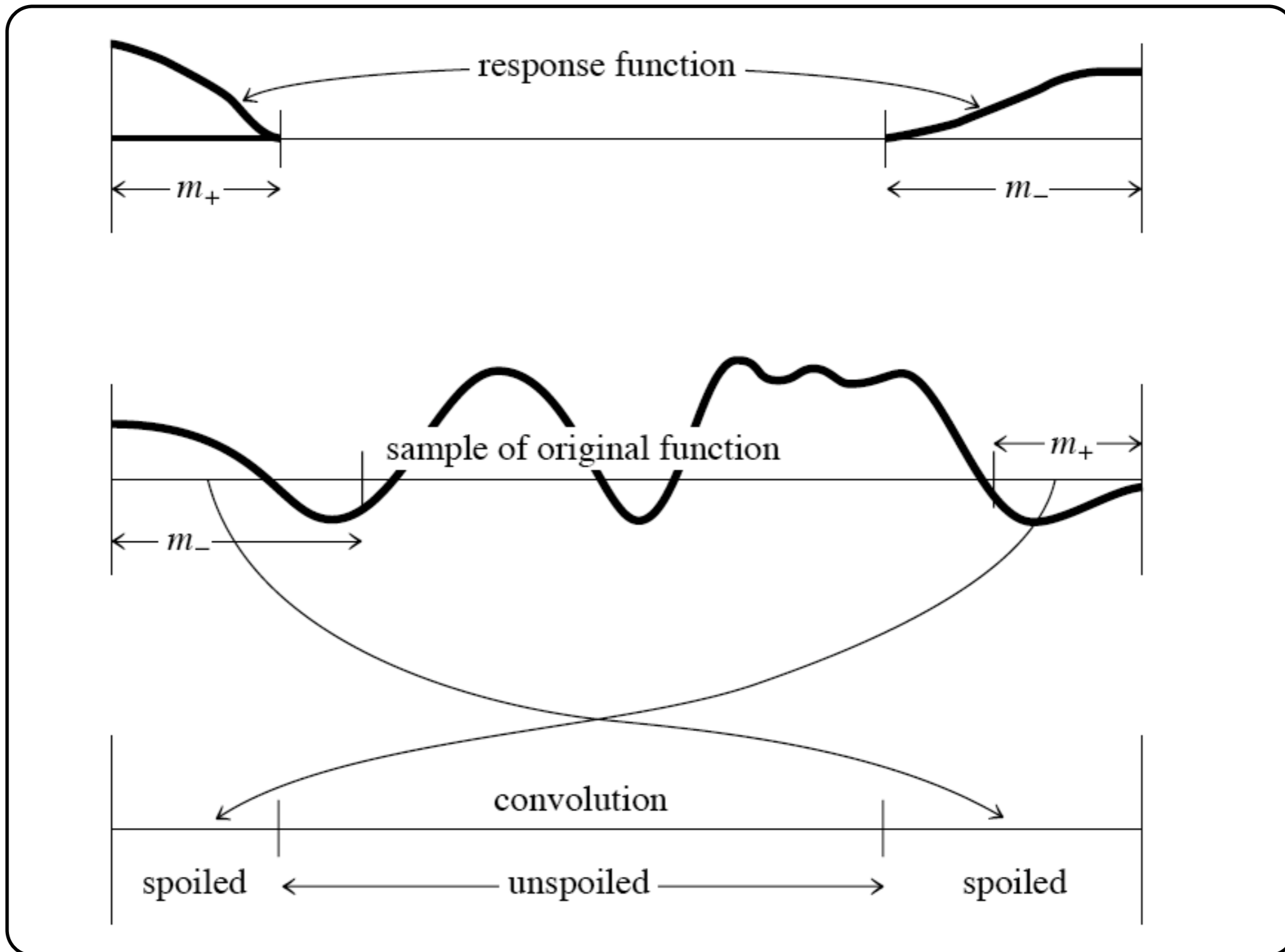
DISCRETE CONVOLUTION THEOREM

$$(r * s)_j \equiv \sum_{k=-M/2+1}^{M/2} s_{j-k} r_k \Leftrightarrow S_n R_n$$

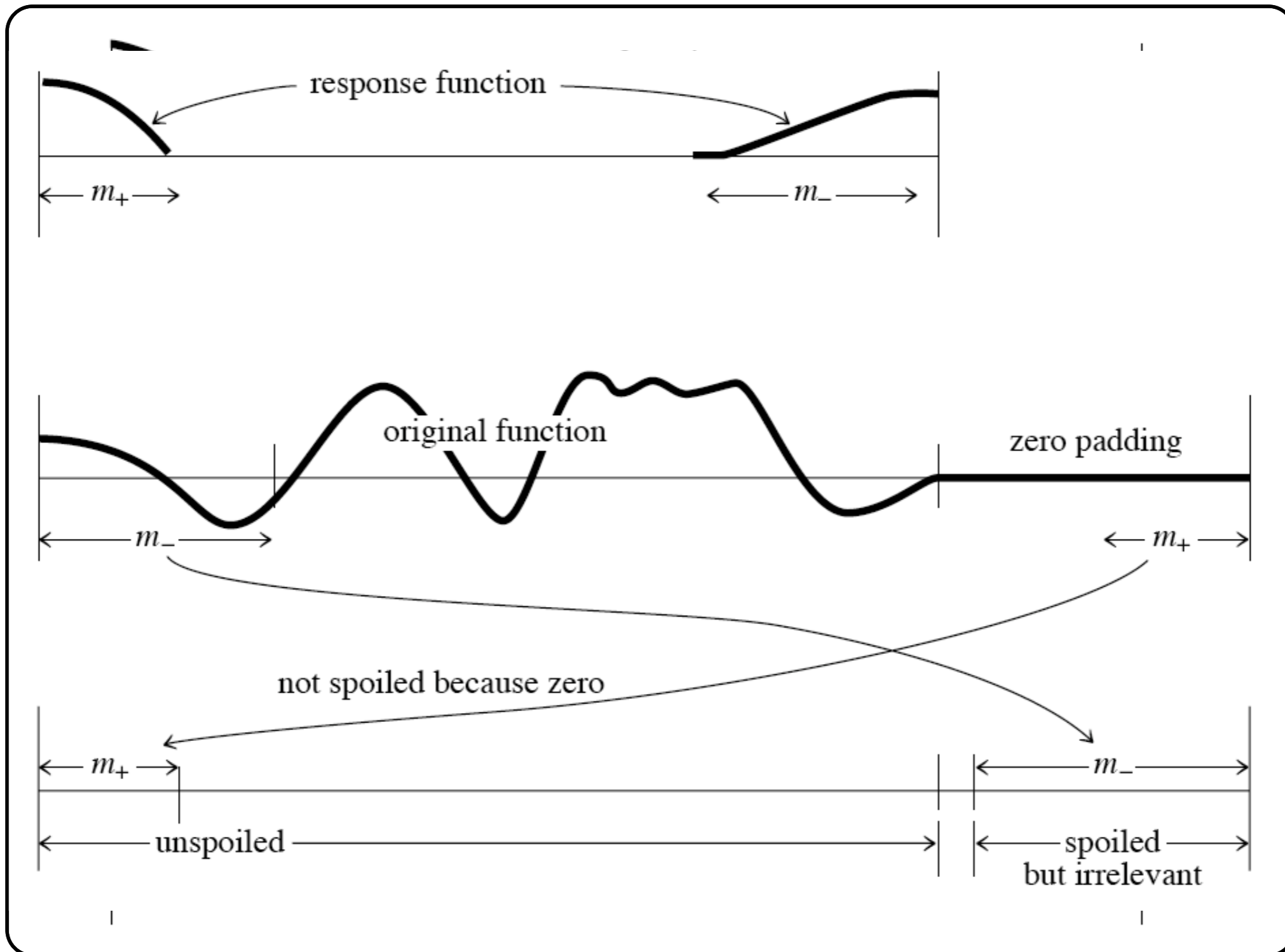
M ONLY NON-ZERO VALUES OF R_k



DISCRETE CONVOLUTION THEOREM



DISCRETE CONVOLUTION THEOREM



DISCRETE CONVOLUTION THEOREM

DISCRETE DECONVOLUTION

$$\frac{\tilde{F}(r * s)_j}{R_n} = S_n$$

HOWEVER NOISE AND UNCERTAINTIES IN
RESPONSE CAN MAKE THIS PROCESS
UNRELIABLE

DISCRETE CONVOLUTION THEOREM

DISCRETE DECONVOLUTION

$$\frac{\tilde{F}(r * s)_j}{R_n} = S_n$$

HOWEVER NOISE AND UNCERTAINTIES IN
RESPONSE CAN MAKE THIS PROCESS
UNRELIABLE

SIMILAR TO THE CONTINUOUS CASE!

NOISE REMOVAL BY OPTIMAL FILTERING

$$cs(t) = s(t) + n(t)$$

$s(t)$ is the smeared signal i.e. true \ast response

DESIGN AN OPTIMAL FILTER $\phi(t)$ THAT
PRODUCES A SIGNAL $\tilde{u}(t)$ AS CLOSE AS
POSSIBLE TO $u(t)$

$$\widetilde{U}(f) = \frac{C(f)\phi(f)}{R(f)}$$

CLOSE IN LEAST SQUARE SENSE

$$\int_{-\infty}^{\infty} |\widetilde{U}(f) - U(f)|^2 df \text{ IS MINIMISED}$$

NOISE REMOVAL BY OPTIMAL FILTERING

$$\int_{-\infty}^{\infty} \left| \frac{[S(f) + N(f)]\phi(f)}{R(f)} - \frac{S(f)}{R(f)} \right|^2 df$$

$$\int S(f)N(f)df \quad \text{TERMS ARE ZERO SINCE NOISE
AND SIGNAL ARE UNCORRELATED}$$

$$\int_{-\infty}^{\infty} |R(f)|^{-2} \underbrace{\{|S(f)|^2|1 - \phi(f)|^2 + |N(f)|^2|\phi(f)|^2\}}_{\Theta} df$$

Θ MINIMISED WITH RESPECT TO ϕ

NOISE REMOVAL BY OPTIMAL FILTERING

$$\frac{d\theta}{d\phi} = 0$$

$$-2S^2(1 - \phi) + 2N^2\phi = 0$$

OPTIMAL FILTER

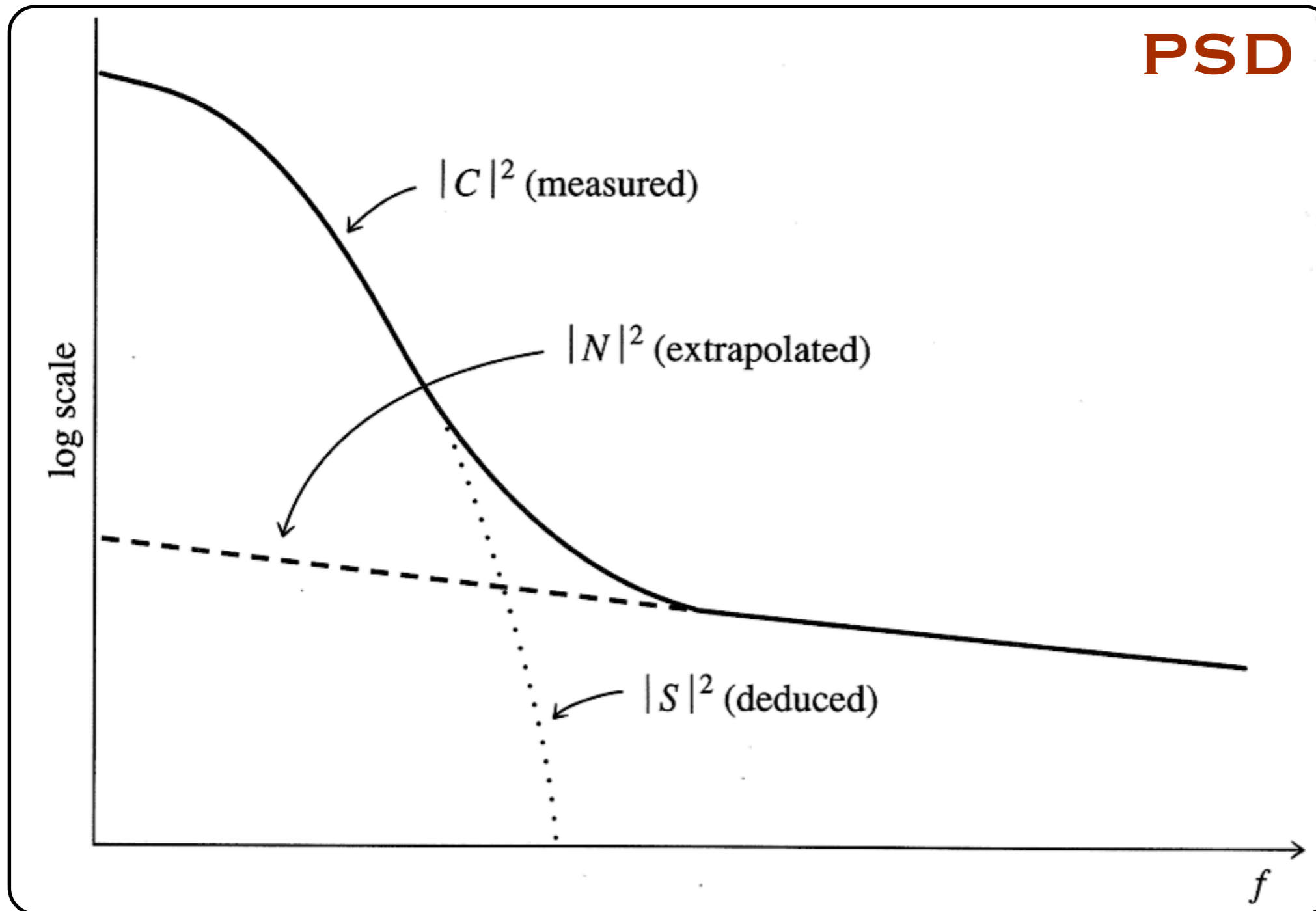
$$\phi = \frac{S^2}{S^2 + N^2}$$

DOES NOT CONTAIN TRUE SIGNAL DIRECTLY!

$$|S(f)|^2 + |N(f)|^2 = PDS(f) = |CS(f)|^2$$

NOISE REMOVAL BY OPTIMAL FILTERING

PSD



SOME APPLICATIONS OF FILTERING

ON OPTIMAL DETECTION OF POINT SOURCES IN CMB MAPS

VIO ET AL, 2002, A&A, 391, 789

AN OPTIMAL FILTER FOR THE DETECTION OF GALAXY

CLUSTERS THROUGH WEAK LENSING

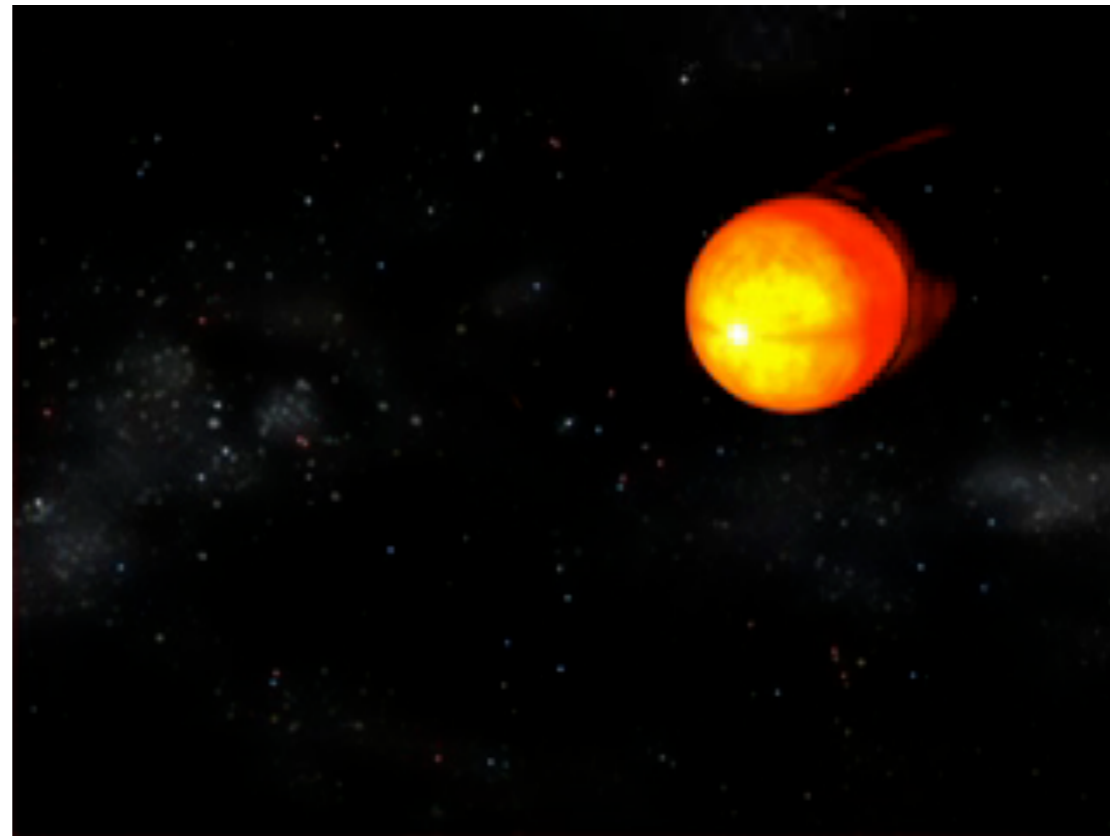
MATURI, ET AL. 2005, A&A, 442, 851

THE LARGEST SCALE PERTURBATIONS: A WINDOW ON THE

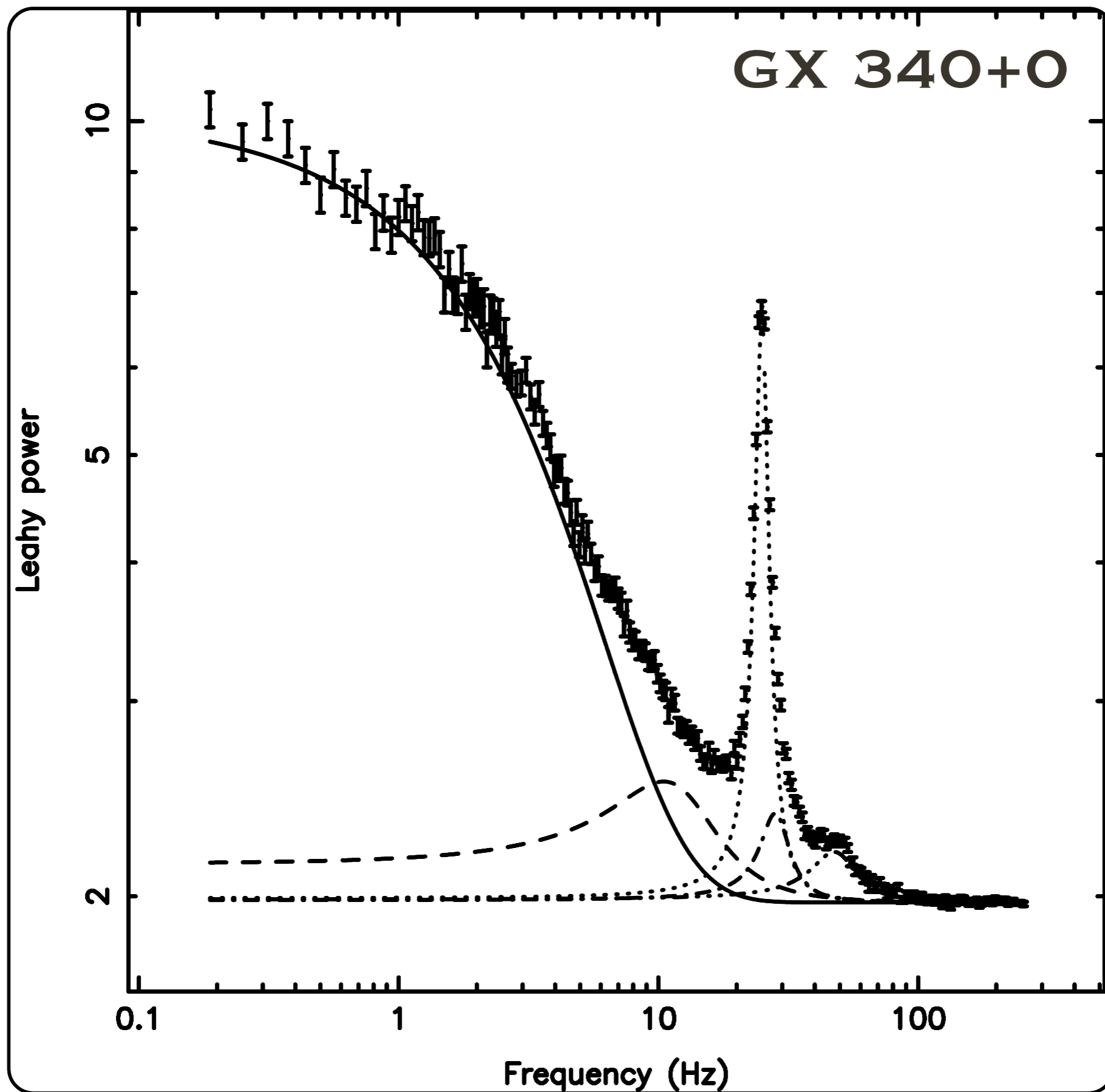
PHYSICS OF THE BEGINNING

WANDELT, NEW ASTRONOMY REVIEW, 2006, 11, 900

LOW-MASS X-RAY BINARIES

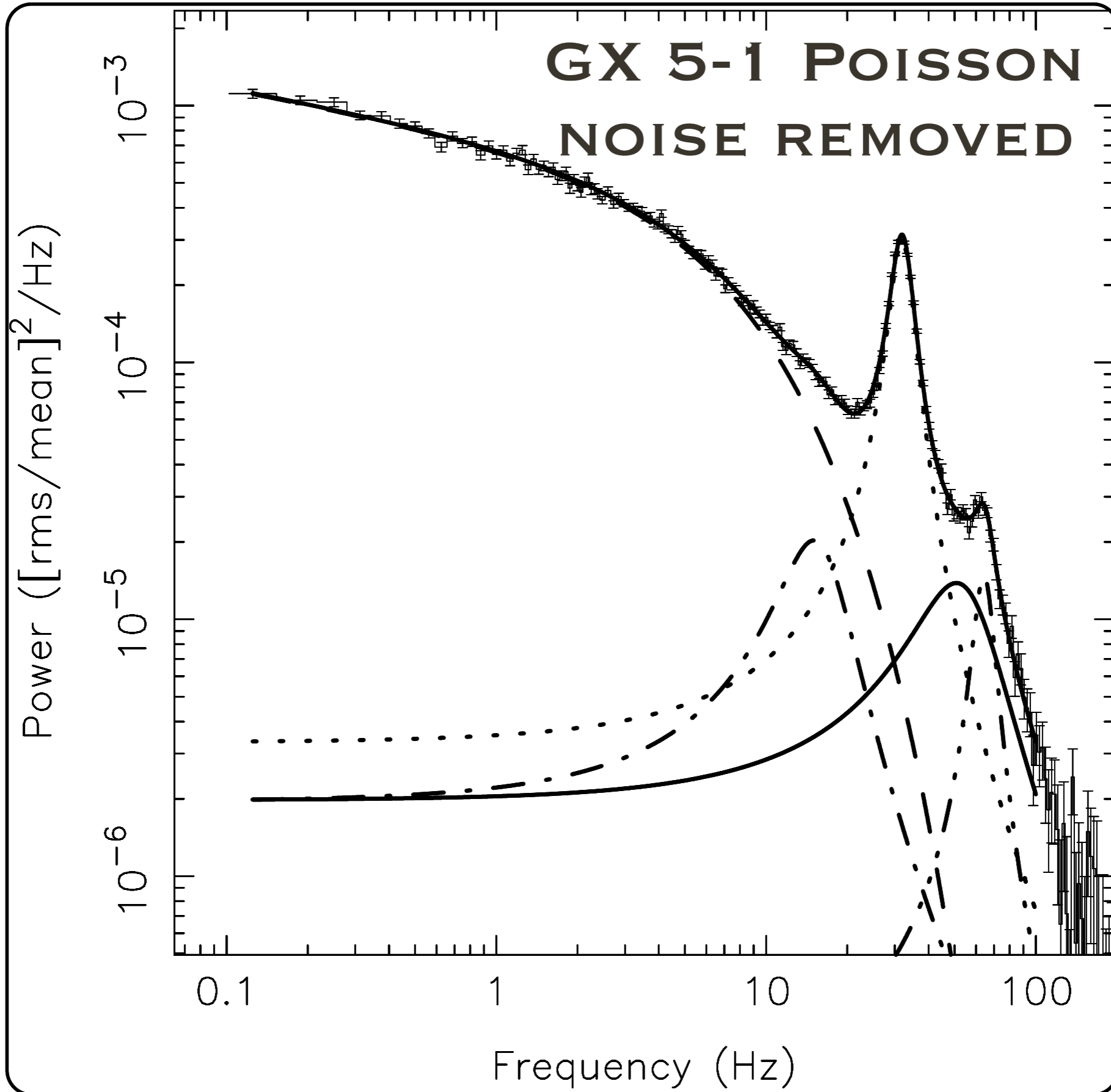


C.F. X-RAY TIMING EXPERIMENTS

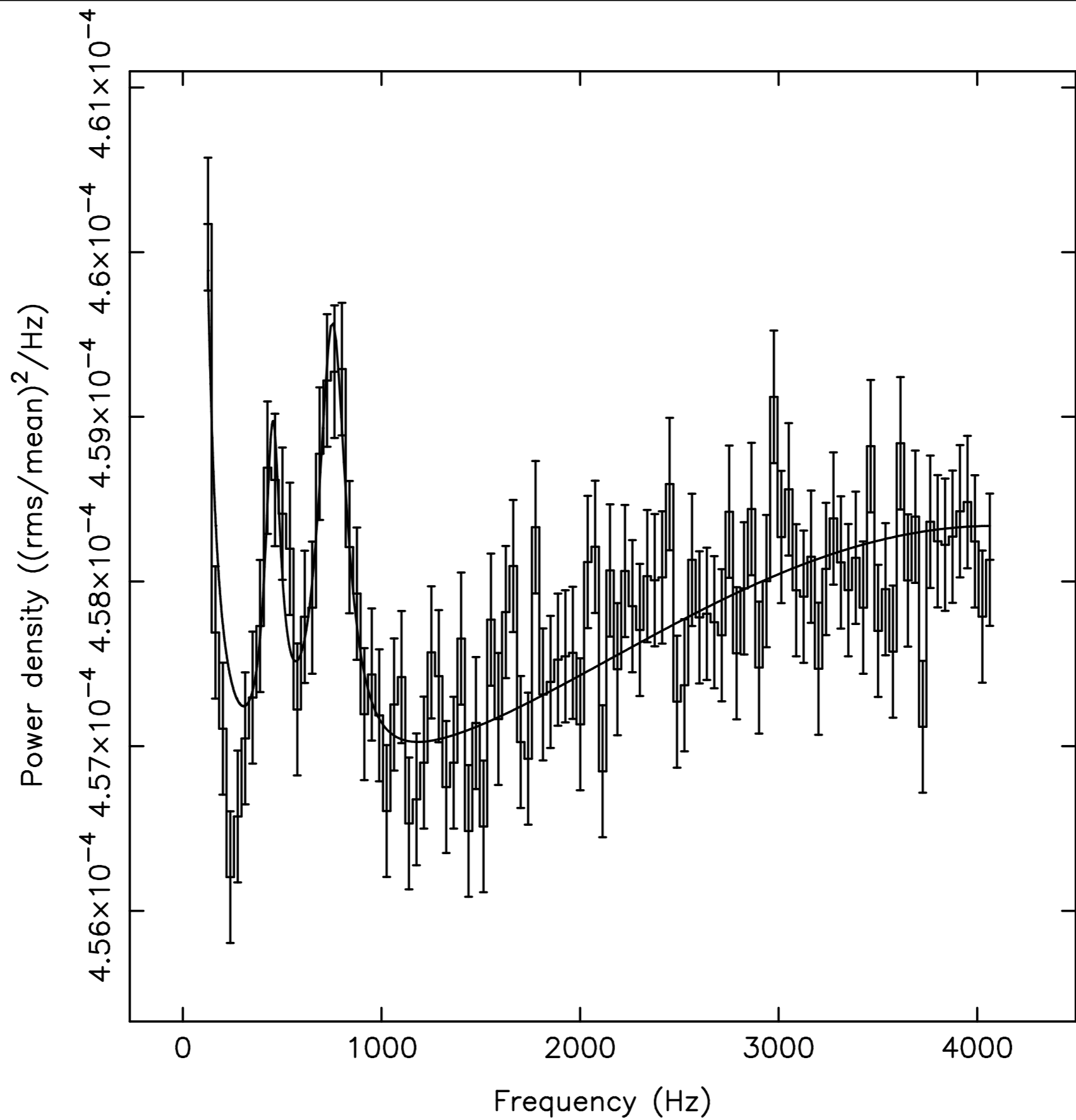


X-RAY TIMING EXPERIMENTS

**GX 5-1 POISSON
NOISE REMOVED**



X-RAY TIMING EXPERIMENTS



RECAP FILTERING:

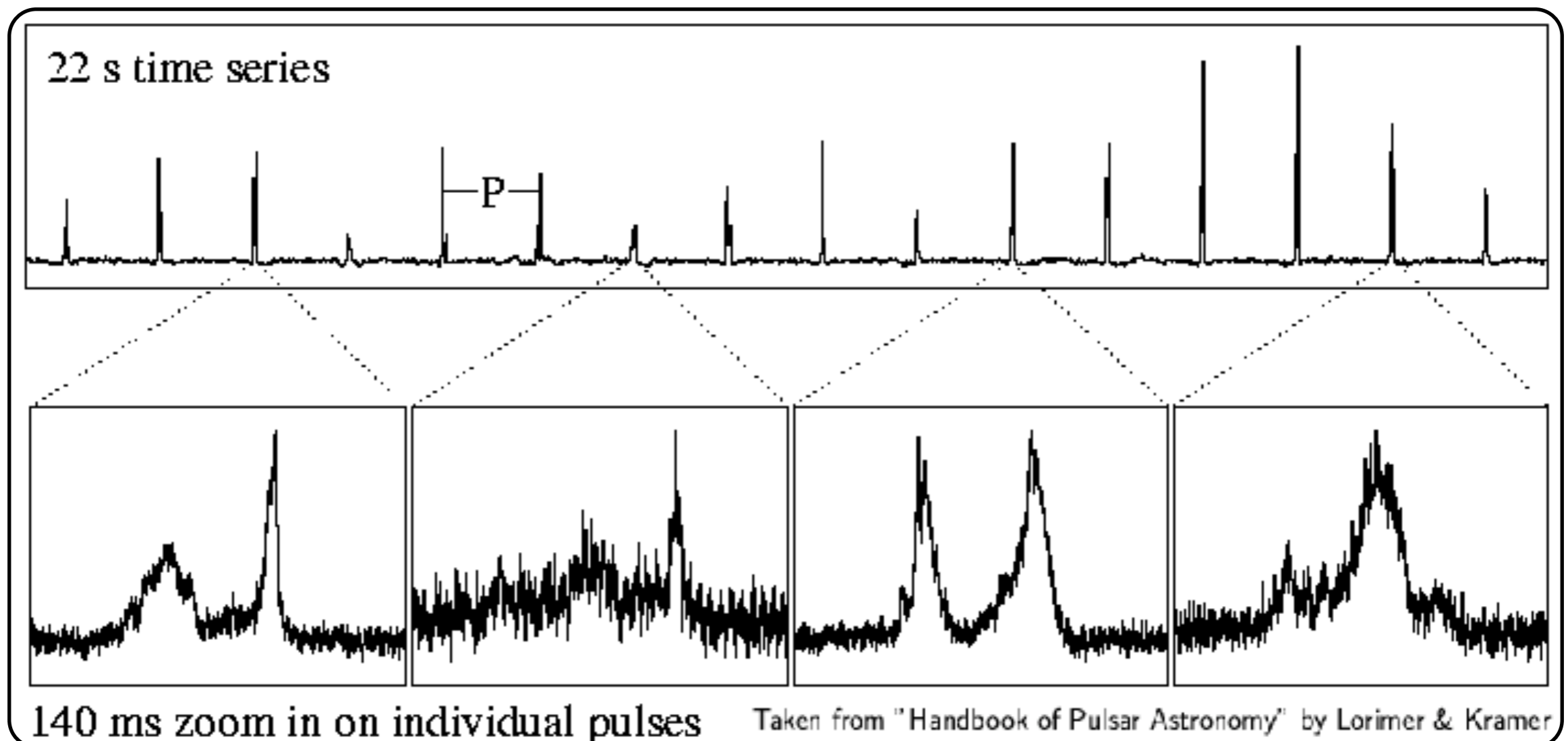
ONE CAN DESIGN AN OPTIMAL FILTER SUCH THAT THE FILTERED MEASURED DATA-SET IS AS CLOSE AS POSSIBLE (IN LEAST-SQUARE SENSE) TO THE UNCORRUPTED SIGNAL

ESTIMATING THE MOMENTS OF A STOCHASTIC PROCESS

CHAPTER 2.2.2

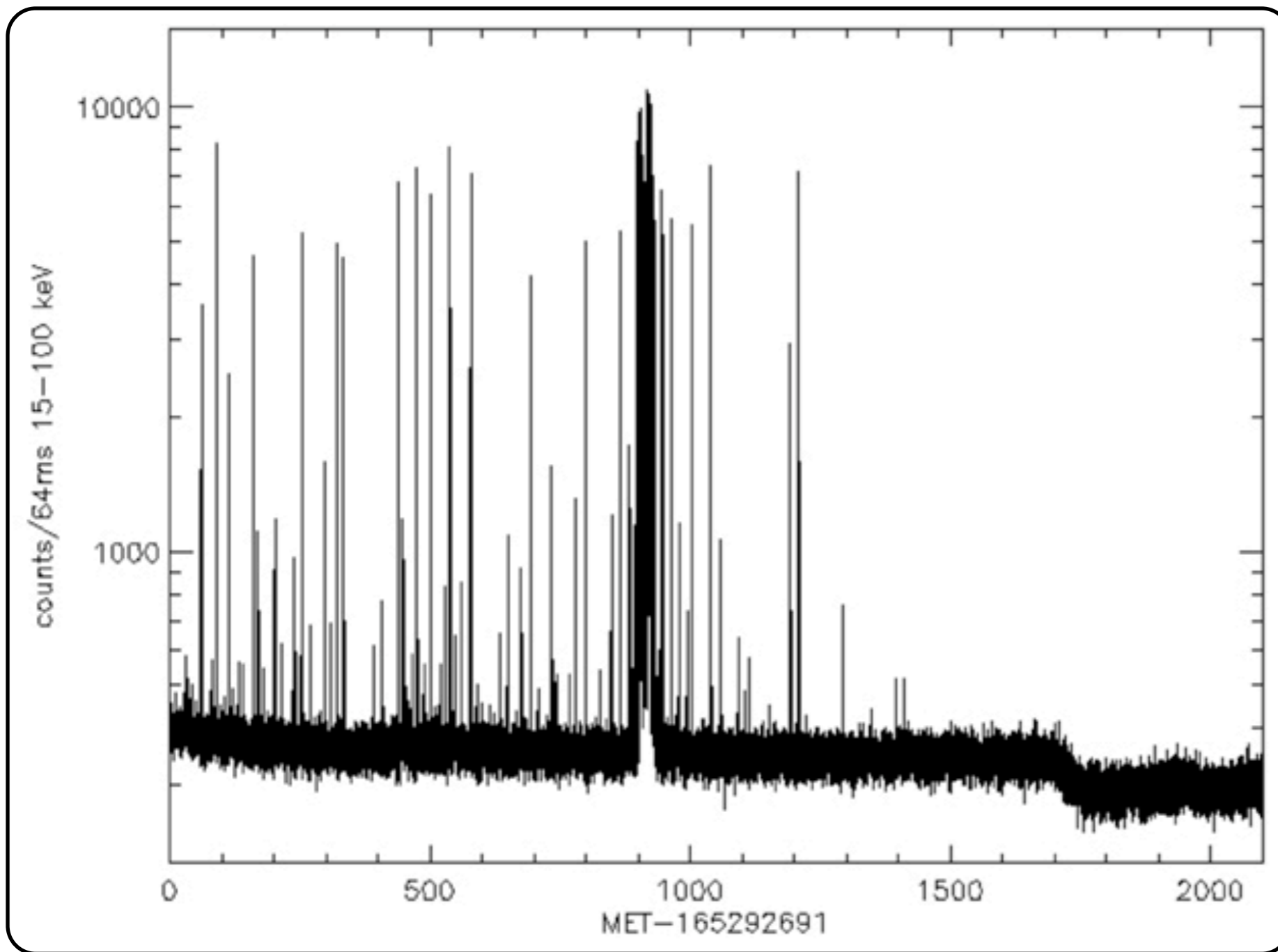
SEE ALSO APPENDIX B3.2, LENA EA.

HOW REPRESENTATIVE IS A MEASUREMENT OF A S.P.?



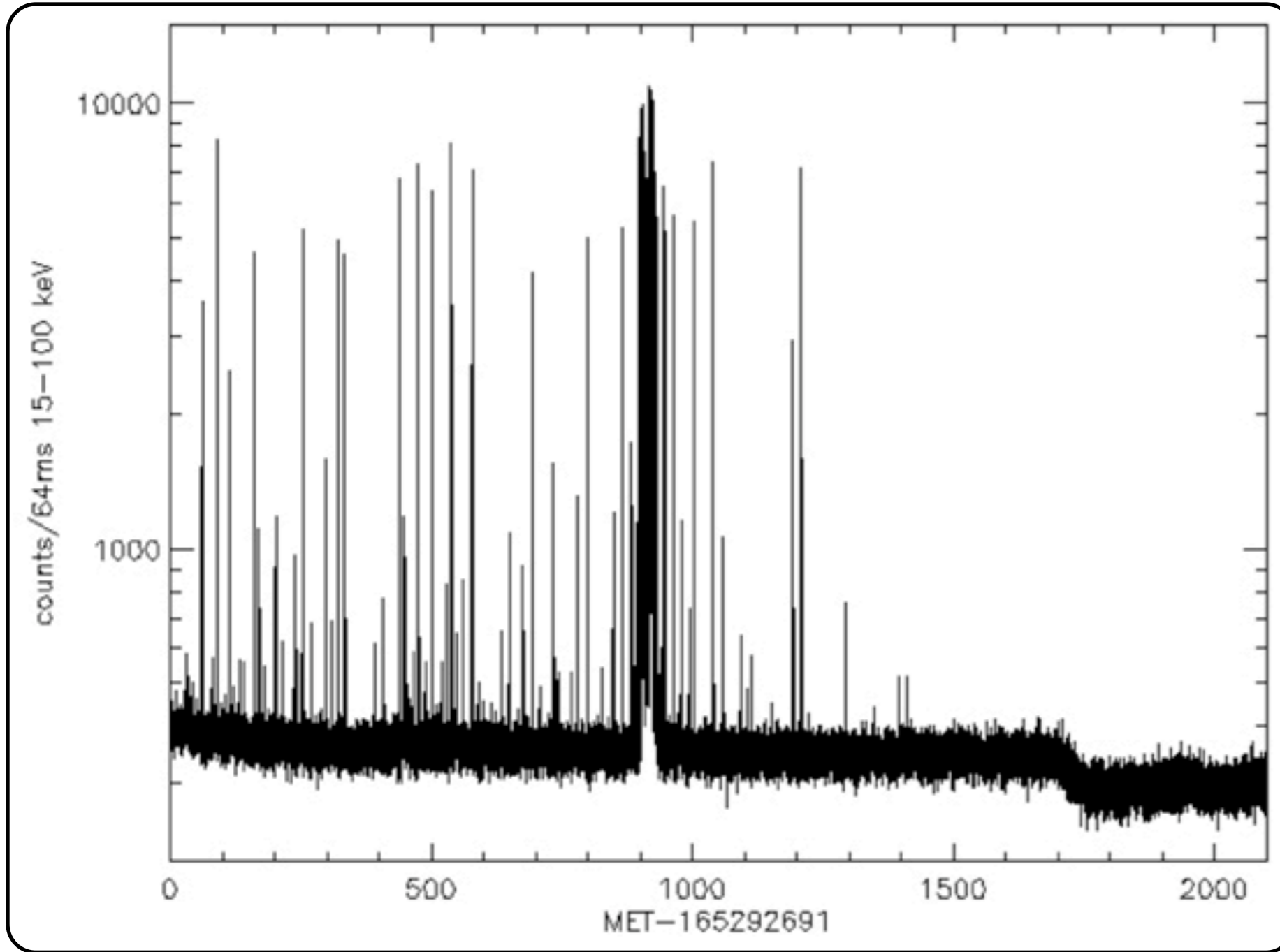
HOW REPRESENTATIVE IS A MEASUREMENT?

SGR 1900+14

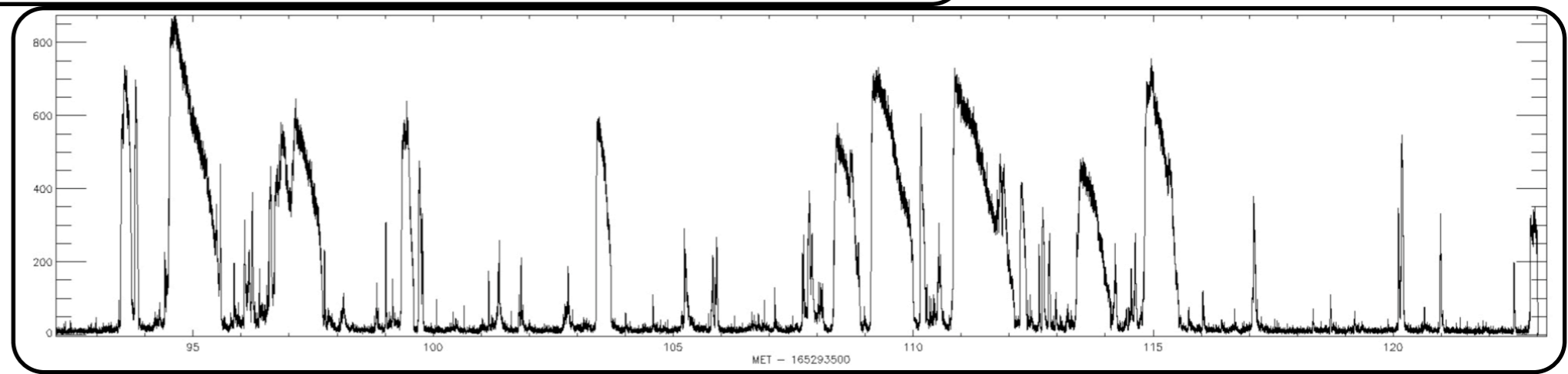


HOW REPRESENTATIVE IS A MEASUREMENT?

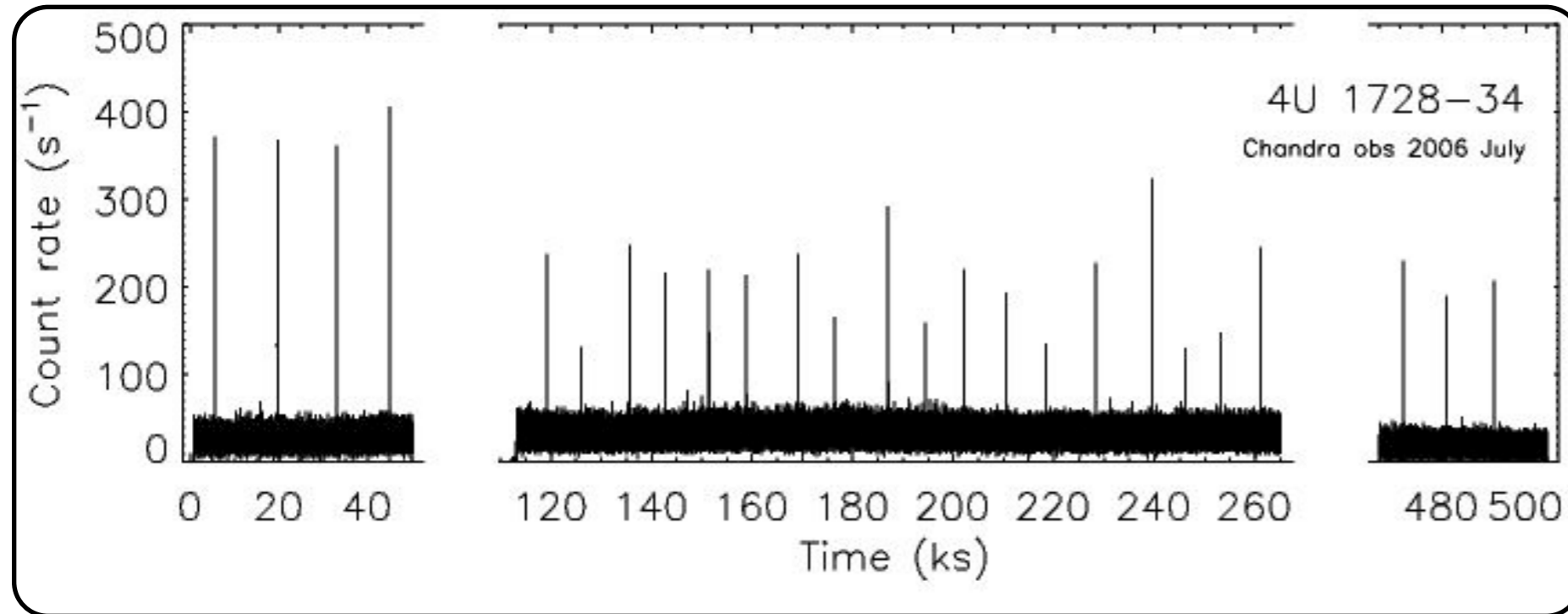
SGR 1900+14



ZOOM



THE CASE OF TYPE I X-RAY BURSTS



GALLOWAY PRIVATE COMMUNICATION

1 MEASUREMENT OF $x(t)$ IN A TIME $T \Rightarrow$ WINDOWING AND
AVERAGING OVER TIME ΔT

WINDOWING: $y(t) = \Pi\left(\frac{t}{T}\right)x(t)$

AVERAGING:

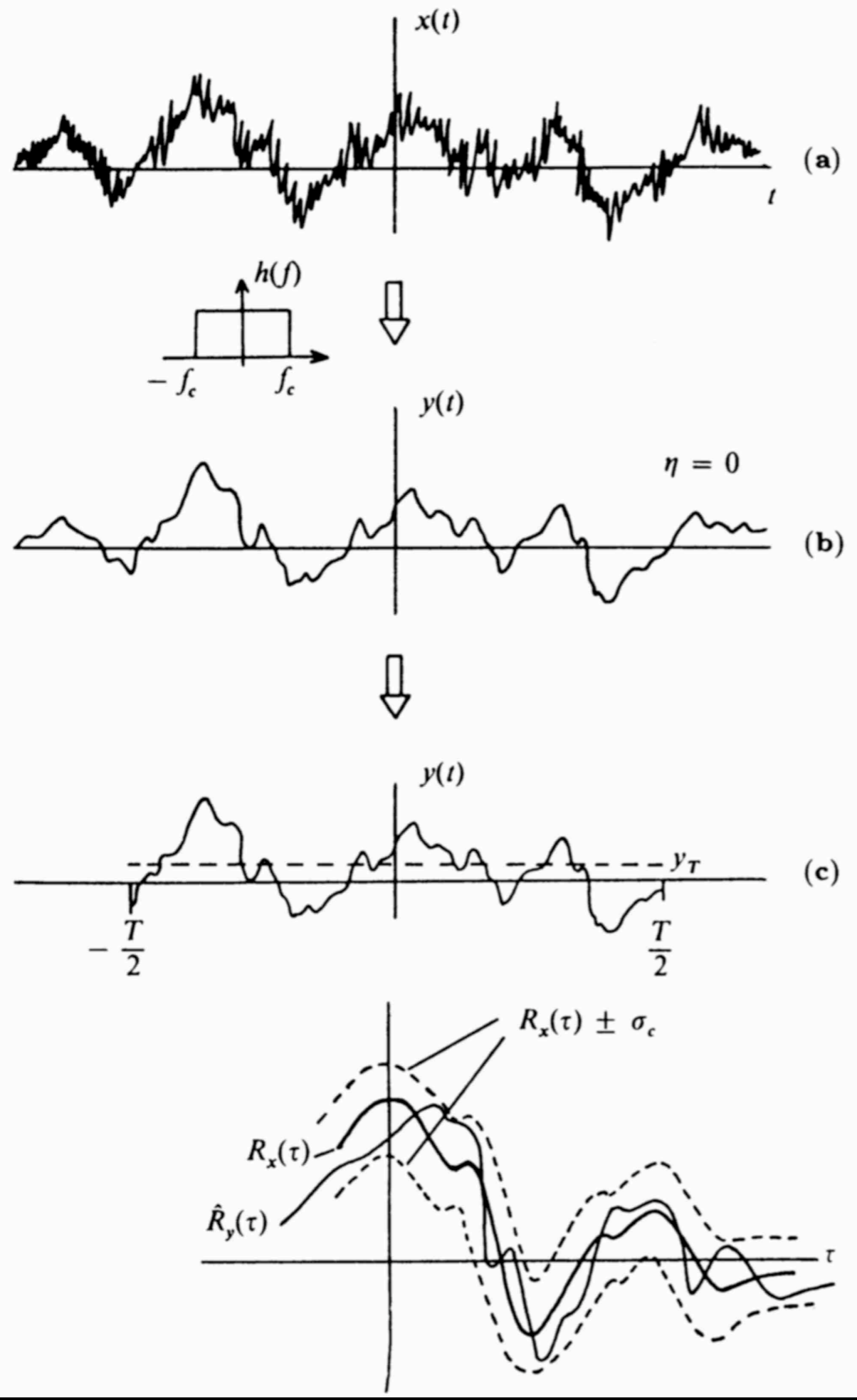
$$z(t) \equiv y_{\Delta T}(t) = \frac{1}{\Delta T} \int_{t-\Delta T/2}^{t+\Delta T/2} y(t') dt' = \frac{1}{\Delta T} \int_{-\infty}^{\infty} \Pi\left(\frac{t-t'}{\Delta T}\right) y(t') dt'$$

= LOW-PASS FILTER, REMEMBER NYQUIST THEOREM

RESPONSE →
 REMOVE HIGH FREQUENCIES
 DUE TO FINITE RESPONSE OF
 INSTRUMENT AND DETECTOR

FINITE DURATION →
 TIME DOMAIN,
 MULTIPLICATION WITH BOX

1 SAMPLE → **AVERAGING** →
 TIME -> CONVOLVING, HENCE
 SMOOTHING, WITH SINC IN
 FREQ DOMAIN



DERIVATION ON BLACK BOARD PAGE 25,26,27
LECTURE NOTES OAF2

$$\sigma_{x_T}^2 = \sigma^2 / N \text{ for } T \gg \tau_0$$

τ_0 DEPENDS ON TRANSFER
FUNCTION

WHICH IN ORDER TO AVOID ALIASING
SHOULD BE SUITED FOR THE
SYSTEM UNDER STUDY

ESTIMATING σ IF THE PROBABILITY DENSITY
FUNCTION OF THE S.P. IS KNOWN



PROPAGATION OF ERRORS **CHAPTER 5.2**

ESTIMATING σ IF THE PROBABILITY DENSITY
FUNCTION OF THE S.P. IS NOT KNOWN



E.G.

BOOTSTRAP METHOD

JACKKNIFE METHOD

MORE ON THIS LATER

STOCHASTIC DESCRIPTION OF RADIATION FIELDS

WHAT IS THE SIZE OF THE FLUCTUATIONS
IN THE RADIATION FIELD?

