# TODAY'S COURSE

## CHAPTER 1.3, 1.4 & 2.2 OAF-2 NUMERICAL RECIPES CHAPTER 13.3

# TOPICS:

- FOURIER TRANSFORMATIONS
- ALIASING & NYQUIST FREQUENCY
- (Optimal) Filtering
- MEASURING MOMENTS OF A STOCHASTIC PROCESS

# RECAP LECTURE 1

DETECTION OF ASTRONOMICAL SIGNALS (S.P.) PLUS NOISE IS CONVOLUTED BY INSTRUMENT TRANSFER FUNCTION AND DATA SAMPLING

STATISTICAL MOMENTS CHARACTERISE THE SIGNAL (PLUS NOISE)

NOISE CAN BE DUE TO THE DETECTOR, BACKGROUND, AND/OR INTRINSIC TO THE SIGNAL

ASSUME WSS S.P. (MEAN DOES NOT DEPEND ON TIME, OR MUCH SLOWER THAN MEASURING PROCESS, AUTO-CORRELATION DEPENDS ON OFFSET ONLY)

CONVOLUTIONS AND CROSS-CORRELATIONS

### CONTINUOUS FOURIER TRANSFORMATIONS



$$F(t) \Leftrightarrow f(x)$$

$$F(t) = \int_{-\infty}^{\infty} f(x)e^{-2\pi i x t} dx$$

$$f(x) = \int_{-\infty}^{\infty} F(t)e^{2\pi i x t} dt$$

Euler's relation :  $e^{ix} = cosx + isinx$ Used in restoration and/or spectral ANALYSIS OF THE SIGNAL

#### CONVOLUTION USING FTS IN PRACTICE

# **CONVOLUTION** $f(x) * g(x) = \int_{-\infty}^{\infty} f(x)g(x_1 - x)dx$

## CONVOLUTION ALWAYS BROADEN THE INPUT FUNCTION



FIGURE FROM GRAY

# CENTRAL LIMIT THEOREM

MANY CONVOLUTIONS -SMOOTHING UNTIL GAUSSIAN PDF

$$p_X(x) = p_{X_1}(x) * p_{X_2}(x) * p_{X_3}(x) * \dots p_{X_n}(x)$$
$$\lim_{n \to \infty} p_X(x) = \frac{1}{\sqrt{(2\pi)\sigma}} exp - \frac{(x-\eta)^2}{2\sigma^2}$$



MANY PHYSICAL PROCESSES/MEASUREMENTS YIELD A GAUSSIAN PROBABILITY DENSITY FUNCTION

## CONVOLUTION USING FOURIER TRANSFORMATIONS

**CONVOLUTION THEOREM**  $M(\lambda) = S(\lambda) * R(\lambda)$  $F(M(\lambda)) = F(S(\lambda)) \cdot F(R(\lambda))$ 



# RECONSTRUCTION OF THE INPUT=SOURCE SPECTRUM

$$\begin{split} M(\lambda) &= S(\lambda) * R(\lambda) \\ \text{Convolution theorem} \\ F(M(\lambda)) &= F(S(\lambda)) \cdot F(R(\lambda)) \\ F(M(\lambda)) &\equiv M(s) \ (etc) \\ M(s) &= S(s) \cdot R(s) \\ S(s) &= \frac{M(s)}{R(s)} \\ S(\lambda) &= F^{-1} \left(\frac{M(s)}{R(s)}\right) \end{split}$$

# Some special functions: Shah's function/Dirac comb

$$III(t) = \sum_{n=-\infty}^{\infty} \delta(t - nT)$$



#### **BOX/WINDOW FUNCTION**



$$B(t) = 0 \ for \ -\frac{W}{2} > t > \frac{W}{2}$$
$$B(t) = 1 \ for \ -\frac{W}{2} < t < \frac{W}{2}$$

## FOURIER TRANSFORMATIONS OF THESE SPECIAL FUNCTIONS



# A SHARP NARROW SIGNAL NEEDS MORE/ HIGHER FREQUENCIES TO BE DESCRIBED IN THE FOURIER TRANSFORM THAN BROAD SHALLOW SIGNAL

CF. THE NUMBER OF SIN+COS NECESSARY TO DESCRIBE THE SIGNAL

OPTICAL SPECTRA: BANDWIDTH SET BY THE WIDTH OF THE SPECTRAL LINES

# $\begin{array}{l} \mbox{Sampling theorem} \\ \mbox{Sampling: no loss of information} \\ \mbox{if the input process has no} \\ \mbox{frequencies} > \frac{1}{2\Delta t_{crit}} \end{array}$

CONTINUOUS SIGNAL H(T) FULLY DESCRIBED BY THE SAMPLES

## OPTICAL SPECTRA: BANDWIDTH SET BY THE WIDTH OF THE SPECTRAL LINES



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# ANOTHER MATH TOOL POWER SPECTRAL DENSITY

#### ( AMPLITUDE OF INDIVIDUAL SINUSOIDS)

(WILL RETURN IN MORE DEPTH IN CHAPTER 6)

CONTINUOUS FT:

$$F(f) = \int_{-\infty}^{\infty} f(t) \ e^{-2\pi i f t} dt$$

CONTINUOUS PSD:

 $P(f) = F(\tilde{f})F(\tilde{f})^*$ 

FOR WSS SIGNALS:  $P(f) = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i f \tau} d\tau$ 

$$\widetilde{F(f)}\widetilde{F(f)} = |F(f)|^2 = \int_{-\infty}^{\infty} R(\tau)e^{-2\pi i f \tau} d\tau$$

#### IMPORTANT CONCEPT IN PSD=NYQUIST THEOREM



WINDOWING & NOISE, BRAULT & WHITE 1971, A&A

# DATA SAMPLING DATA IS DISCRETE NOT CONTINUOUS TIME DOMAIN MULTIPLY S.P. WITH SHAH FUNCTION $m_{samp,n} = m_s(x) = m(x)\frac{1}{\tau}\Pi I(\frac{x}{\tau}) = \sum m(n\tau)\delta(x-n\tau)$ N-1DISCRETE FT: $M_{samp,k} = \sum m_{samp,n} e^{2\pi i n k/N}$ n=0

DISCRETE PSD:  $P_j = \frac{2}{a_0} |a_j|^2$  power  $\propto$  amplitude squared:

$$a_{0} = M_{samp,k=0} = \sum_{n=0}^{N-1} m_{samp,n} \equiv N_{0}$$
$$a_{k} = M_{samp,k} = \sum_{n=0}^{N-1} m_{samp,n} e^{2\pi i n k/N}$$

# NYQUIST THEOREM: CONT'D



SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

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SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

#### NYQUIST THEOREM: CONT'D SAMPLING CAUSES REPLICATION OF SIGNAL



SAMPLING; BRAULT & WHITE 1971, A&A, 13, 169 (IN LIST OF PRESENTATION PAPERS!)

# ALIASING



## ALIASING: CONT'D



CONVOLUTION WITH SHAH FUNCTION IN FREQ SPACE:

#### REPLICATION

# SAMPLING: HIGH FREQUENCIES ARE FILTERED OUT WINDOW: LOW FREQUENCIES ARE FILTERED OUT LEADS TO BAND LIMITED DATA

# FILTERING FREQUENCY FILTERING Y(f) = X(f)H(f)

$$y(t) = \int_{-\infty}^{\infty} x(t - \theta)h(\theta)d\theta$$
$$y(t) = x(t) * h(t)$$

FILTERING OF PROCESS X WITH FILTER H

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FILTERING OF PROCESS X WITH FILTER H

#### REMEMBER INSERTING FILTER IN OPTICAL IMAGING

## GAUSSIAN RESPONSE FUNCTION



$$R(\lambda)\frac{1}{\sqrt{(2\pi)\sigma}}exp-(\frac{\lambda^2}{2\sigma^2})$$

# GAUSSIAN RESPONSE FUNCTION



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SAMPLING: HIGH FREQUENCIES ARE FILTERED OUT WINDOW: LOW FREQUENCIES ARE FILTERED OUT LEADS TO BAND LIMITED DATA FILTERING

## -> TIME FILTERING

MEASURE A PROCESS X(T) OVER INTERVAL T ASSUMED ZERO OUTSIDE T

$$\equiv y(t) = \Pi(\frac{t}{T})x(t)$$

$$Y(f) = X(f) * Tsinc(Tf)$$

ALL INFORMATION ABOUT FREQUENCIES <1/T IS LOST!