TODAY'S COURSE

Chapter 1.3, 1.4 & 2.2 OAF-2 Numerical Recipes Chapter 13.3

Topics:

- Fourier transformations
- ALIASING & NYQUIST FREQUENCY
- A (OPTIMAL) FILTERING
- measuring moments of a stochastic process

RECAP LECTURE 1

DETECTION OF ASTRONOMICAL SIGNALS (S.P.) plus noise is convoluted by instrument transfer function and data sampling

STATISTICAL MOMENTS CHARACTERISE THE signal (plus noise)

NOISE CAN BE DUE TO THE DETECTOR, background, and/or intrinsic to the signal

ASSUME WSS S.P. (MEAN DOES NOT DEPEND ON time, or much slower than measuring process, auto-correlation depends on OFFSET ONLY)

Convolutions and cross-correlations

CONTINUOUS FOURIER TRANSFORMATIONS

$$
F(t) \Leftrightarrow f(x)
$$

$$
F(t) = \int_{-\infty}^{\infty} f(x)e^{-2\pi ixt} dx
$$

$$
f(x) = \int_{-\infty}^{\infty} F(t)e^{2\pi ixt} dt
$$

 $Euler's$ *relation* : $e^{ix} = cosx + isinx$ Used in restoration and/or spectral analysis of the signal

Convolution using FTs in practice

$f(x) * g(x) =$ \int_0^∞ $-\infty$ $f(x)g(x_1-x)dx$ **CONVOLUTION**

Convolution always broaden the input FUNCTION

figure from Gray

CENTRAL LIMIT THEOREM

many convolutions ➞smoothing until Gaussian PDF

$$
p_X(x) = p_{X_1}(x) * p_{X_2}(x) * p_{X_3}(x) * \cdots p_{X_n}(x)
$$

$$
\lim_{n \to \infty} p_X(x) = \frac{1}{\sqrt{2\pi}\sigma} \exp{-\frac{(x-\eta)^2}{2\sigma^2}}
$$

many physical processes/measurements yield a

Convolution using Fourier **TRANSFORMATIONS**

Convolution theorem $M(\lambda) = S(\lambda) * R(\lambda)$ $F(M(\lambda)) = F(S(\lambda)) \cdot F(R(\lambda))$

RECONSTRUCTION OF THE input=source spectrum

$$
M(\lambda) = S(\lambda) * R(\lambda)
$$

Convolution theorem

$$
F(M(\lambda)) = F(S(\lambda)) \cdot F(R(\lambda))
$$

$$
F(M(\lambda)) \equiv M(s) (etc)
$$

$$
M(s) = S(s) \cdot R(s)
$$

$$
S(s) = \frac{M(s)}{R(s)}
$$

$$
S(\lambda) = F^{-1} \left(\frac{M(s)}{R(s)}\right)
$$

SOME SPECIAL FUNCTIONS: Shah's function/Dirac comb

$$
III(t) = \sum_{n = -\infty}^{\infty} \delta(t - nT)
$$

Box/window function

$$
B(t) = 0 \text{ for } -\frac{W}{2} > t > \frac{W}{2}
$$

$$
B(t) = 1 \text{ for } -\frac{W}{2} < t < \frac{W}{2}
$$

Fourier transformations of these special functions

A sharp narrow signal needs more/ higher frequencies to be described in the Fourier Transform than broad SHALLOW SIGNAL

cf. the number of sin+cos necessary to describe the signal

Optical spectra: bandwidth set by the width of the spectral lines

Sampling: no loss of information if the input process has no frequencies > 1 2∆*tcrit* SAMPLING THEOREM

CONTINUOUS SIGNAL H(T) FULLY DESCRIBED by the samples

Optical spectra: bandwidth set by THE WIDTH OF THE SPECTRAL LINES

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Another math tool Power Spectral Density

$(\alpha$ AMPLITUDE OF INDIVIDUAL SINUSOIDS)

(will return in more depth in Chapter 6)

CONTINUOUS FT:

$$
F(f) = \int_{-\infty}^{\infty} f(t) e^{-2\pi i f t} dt
$$

CONTINUOUS PSD:

 $P(f) = F(f)F(f)^*$

 $P(f) =$ \int_0^∞ $-\infty$ FOR WSS SIGNALS: $P(f) = \int R(\tau) e^{-2\pi i f \tau} d\tau$

$$
\mathsf{HENCE}:
$$

$$
F(f)F(f) = |F(f)|^2 = \int_{-\infty}^{\infty} R(\tau)e^{-2\pi i f\tau}d\tau
$$

Important concept in PSD=Nyquist theorem

WINDOWING & NOISE, BRAULT & WHITE 1971, A&A

in time domain multiply s.p. with shah function DATA SAMPLING $m_s(x) = m(x)$ 1 τ $\prod(\frac{x}{x})$ τ $m_{samp,n} = m_s(x) = m(x)\frac{1}{\tau}\Pi(\frac{x}{\tau}) = \sum m(n\tau)\delta(x - n\tau)$ *n* D *SCRETE FT:* $M_{samp,k} =$ $N-1$ $\sum_{n=1}^{N-1}$ $m_{samp,n}$ $e^{2\pi i n k/N}$ DATA IS DISCRETE NOT CONTINUOUS

2

 $n=0$

 a_0

 D **SCRETE PSD:** $P_j = \frac{2}{a_s} |a_j|^2$ power α amplitude squared: $|a_j|^2$

$$
a_0 = M_{samp,k=0} = \sum_{n=0}^{N-1} m_{samp,n} \equiv N_0
$$

$$
a_k = M_{samp,k} = \sum_{n=0}^{N-1} m_{samp,n} e^{2\pi i n k/N}
$$

NYQUIST THEOREM: CONT'D

Sampling; Brault & White 1971, A&A, 13, 169 (in list of presentation papers!)

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Nyquist theorem: cont'd SAMPLING CAUSES REPLICATION OF SIGNAL

Sampling; Brault & White 1971, A&A, 13, 169 (in list of presentation papers!)

Aliasing

ALIASING: CONT'D

Convolution with shah function in freq space:

REPLICATION

Sampling: high frequencies are filtered out LEADS TO BAND LIMITED DATA WINDOW: LOW FREQUENCIES ARE FILTERED OUT

FILTERING FREQUENCY FILTERING $Y(f) = X(f)H(f)$

$$
y(t) = \int_{-\infty}^{\infty} x(t - \theta)h(\theta)d\theta
$$

$$
y(t) = x(t) * h(t)
$$

Filtering of process x with filter h

Sampling: high frequencies are filtered out LEADS TO BAND LIMITED DATA Window: low frequencies are filtered out

FILTERING FREQUENCY FILTERING $Y(f) = X(f)H(f)$

$$
y(t) = \int_{-\infty}^{\infty} x(t - \theta)h(\theta)d\theta
$$

$$
y(t) = x(t) * h(t)
$$

Filtering of process x with filter h

Remember inserting filter in optical imaging

GAUSSIAN RESPONSE FUNCTION

$$
R(\lambda)\frac{1}{\sqrt{(2\pi)\sigma}}exp-(\frac{\lambda^2}{2\sigma^2})
$$

Gaussian response function

Gaussian response function

FILTERING Sampling: high frequencies are filtered out LEADS TO BAND LIMITED DATA WINDOW: LOW FREQUENCIES ARE FILTERED OUT

measure a process x(t) over interval T assumed zero outside T

$$
\equiv y(t) = \Pi(\frac{t}{T})x(t)
$$

$$
Y(f) = X(f) * Tsinc(Tf)
$$

all information about frequencies <1/T is lost!