

# TODAY'S COURSE

CHAPTER 1.3, 1.4 & 2.2 OAF-2  
NUMERICAL RECIPES CHAPTER 13.3

## TOPICS:

ALIASING & NYQUIST FREQUENCY

(OPTIMAL) FILTERING

MEASURING MOMENTS OF A S.P.

# RECAP LECTURE 1

DETECTION OF ASTRONOMICAL SIGNALS (S.P.) PLUS NOISE IS INFLUENCED BY INSTRUMENT TRANSFER FUNCTION AND DATA SAMPLING

STATISTICAL MOMENTS CHARACTERISE THE SIGNAL (PLUS NOISE)

NOISE CAN BE DUE TO THE DETECTOR, BACKGROUND, AND/OR INTRINSIC TO THE SIGNAL

ASSUME WSS S.P. (MEAN DOES NOT DEPEND ON TIME, OR MUCH SLOWER THAN MEASURING PROCESS, AUTO-CORRELATION DEPENDS ON OFFSET ONLY)

# ANOTHER MATH TOOL

## POWER SPECTRAL DENSITY

( $\infty$  AMPLITUDE OF INDIVIDUAL SINUSOIDS)

(WILL RETURN IN MORE DEPTH IN CHAPTER 6)

CONTINUOUS FT: 
$$F(f) = \int_{-\infty}^{\infty} f(t) e^{-2\pi i f t} dt$$

CONTINUOUS PSD: 
$$P(f) = F(\tilde{f})F(\tilde{f})^*$$

FOR WSS SIGNALS: 
$$P(f) = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i f \tau} d\tau$$

HENCE:

$$F(\tilde{f})F(f) = |F(f)|^2 = \int_{-\infty}^{\infty} R(\tau) e^{-2\pi i f \tau} d\tau$$

# DATA SAMPLING

TIME DOMAIN MULTIPLY S.P. WITH SHAH FUNCTION

$$m_{s\text{amp},n} = m_s(x) = m(x) \frac{1}{\tau} \text{III}\left(\frac{x}{\tau}\right) = \sum_n m(n\tau) \delta(x - n\tau)$$

DISCRETE FT:  $M_{s\text{amp},k} = \sum_{n=0}^{N-1} m_{s\text{amp},n} e^{2\pi i n k / N}$

DISCRETE PSD:  $P_j = \frac{2}{a_0} |a_j|^2$  POWER  $\propto$  AMPLITUDE SQUARED:

$$a_0 = M_{s\text{amp},k=0} = \sum_{n=0}^{N-1} m_{s\text{amp},n} \equiv N_0$$

$$a_k = M_{s\text{amp},k} = \sum_{n=0}^{N-1} m_{s\text{amp},n} e^{2\pi i n k / N}$$

# NYQUIST FREQUENCY

NYQUIST FREQUENCY = HALF THE  
FREQUENCY RELATED WITH THE  
SAMPLING RATE

SHARP NARROW SIGNAL REQUIRES MORE  
FREQUENCIES FOR ITS DESCRIPTION THAN  
BROAD SIGNAL

CF. THE NUMBER OF SIN+COS NECESSARY TO DESCRIBE THE SIGNAL

OPTICAL SPECTRA: RANGE FROM MIN  
TO MAXI FREQUENCY = BANDWIDTH SET  
BY THE WIDTH AND SHAPE OF THE  
SPECTRAL LINES

# NYQUIST FREQUENCY

if  $f_{nyq} = 2f_{max}$  in signal

SAMPLING AT THE NYQUIST FREQUENCY:  
NO LOSS OF INFORMATION

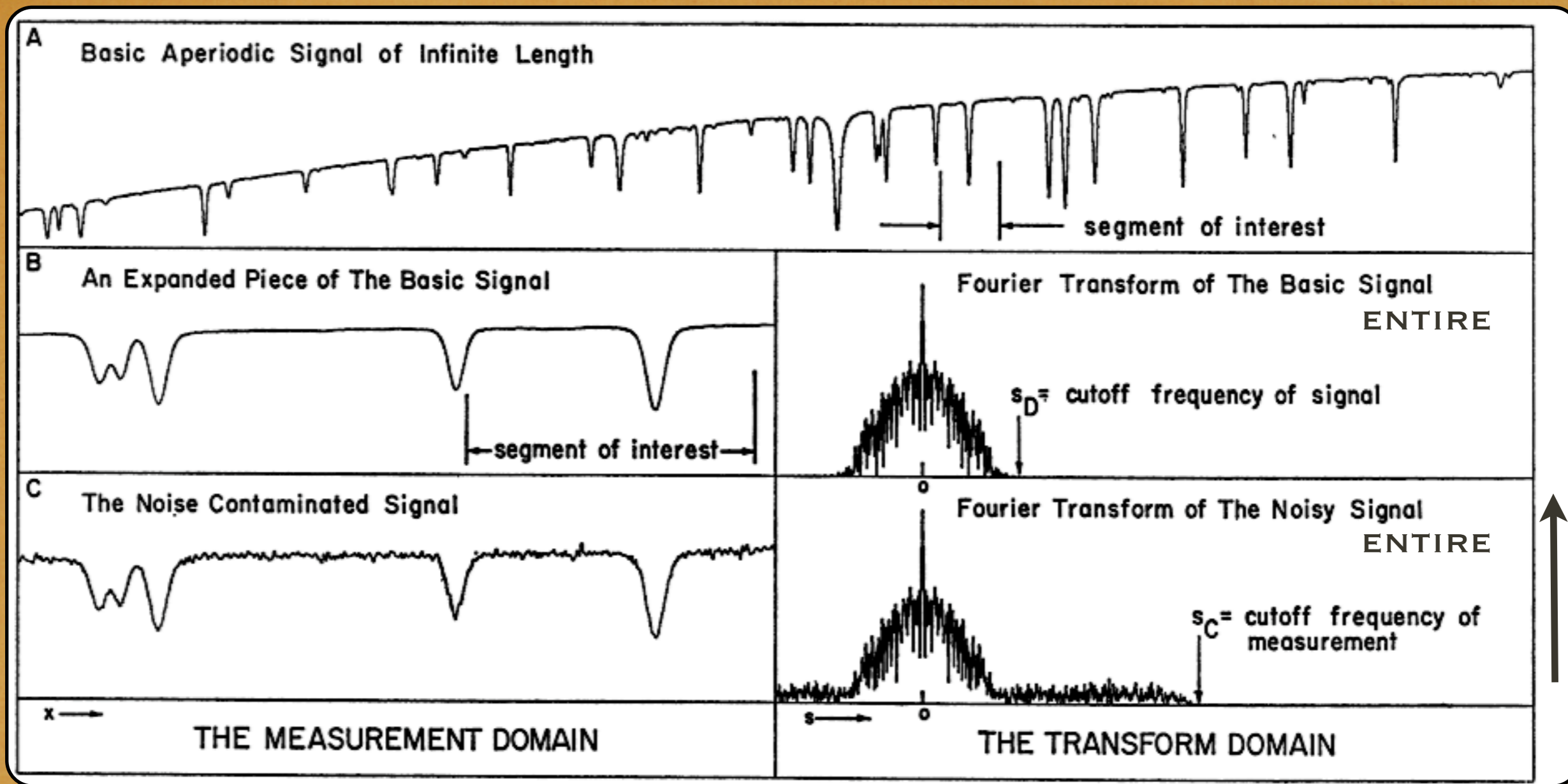
CONTINUOUS SIGNAL  $h(t)$  FULLY DESCRIBED  
BY THE SAMPLES

$$h(t) = \Delta t \sum_{n=-\infty}^{\infty} h_n \frac{\sin[2\pi f_{nyq}(t - n\Delta t)]}{\pi(t - n\Delta t)}$$

# NYQUIST THEOREM: CONT'D

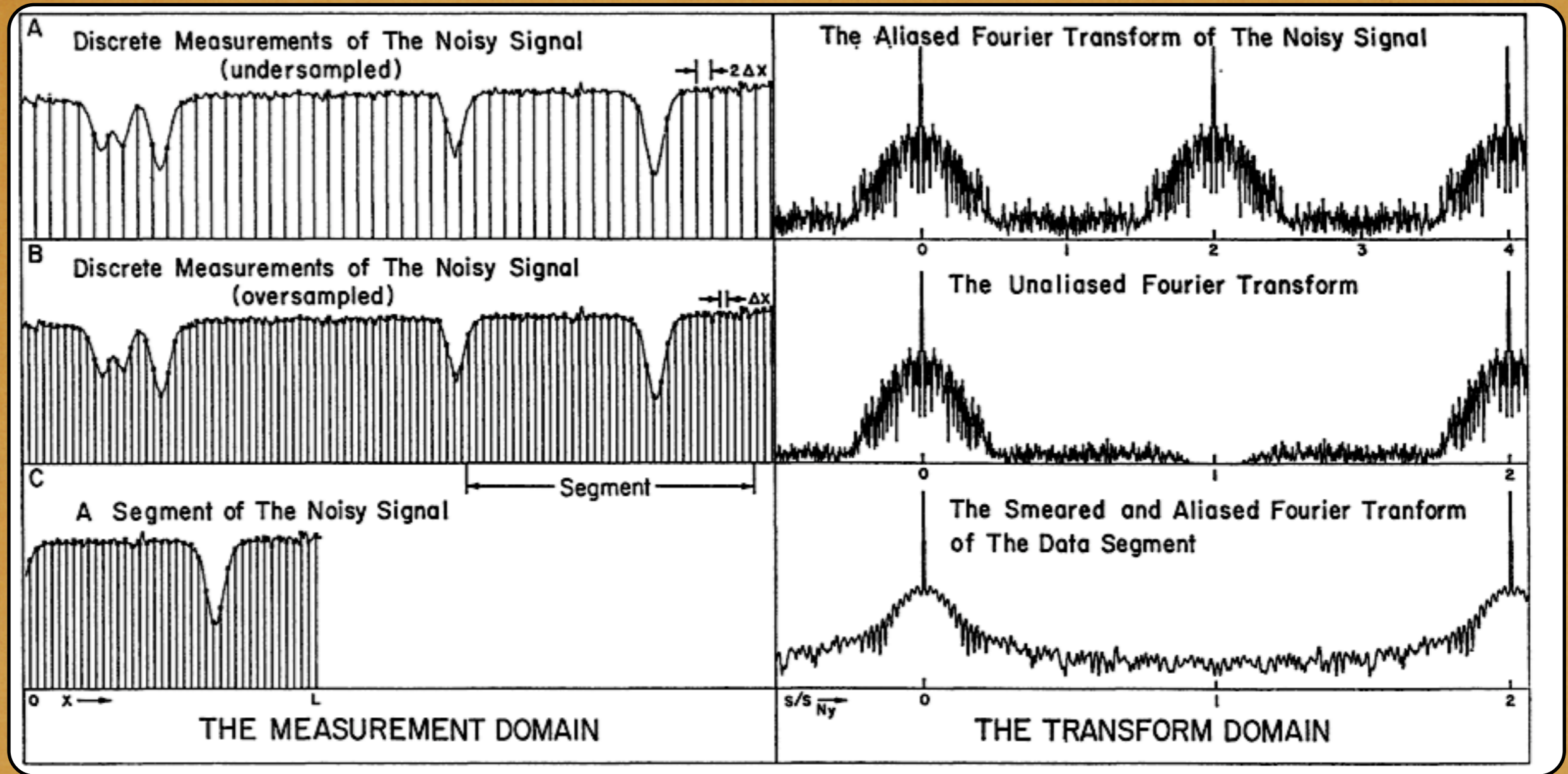
FLUX

LOG(PSD)



$\lambda$  → FREQUENCY (S)

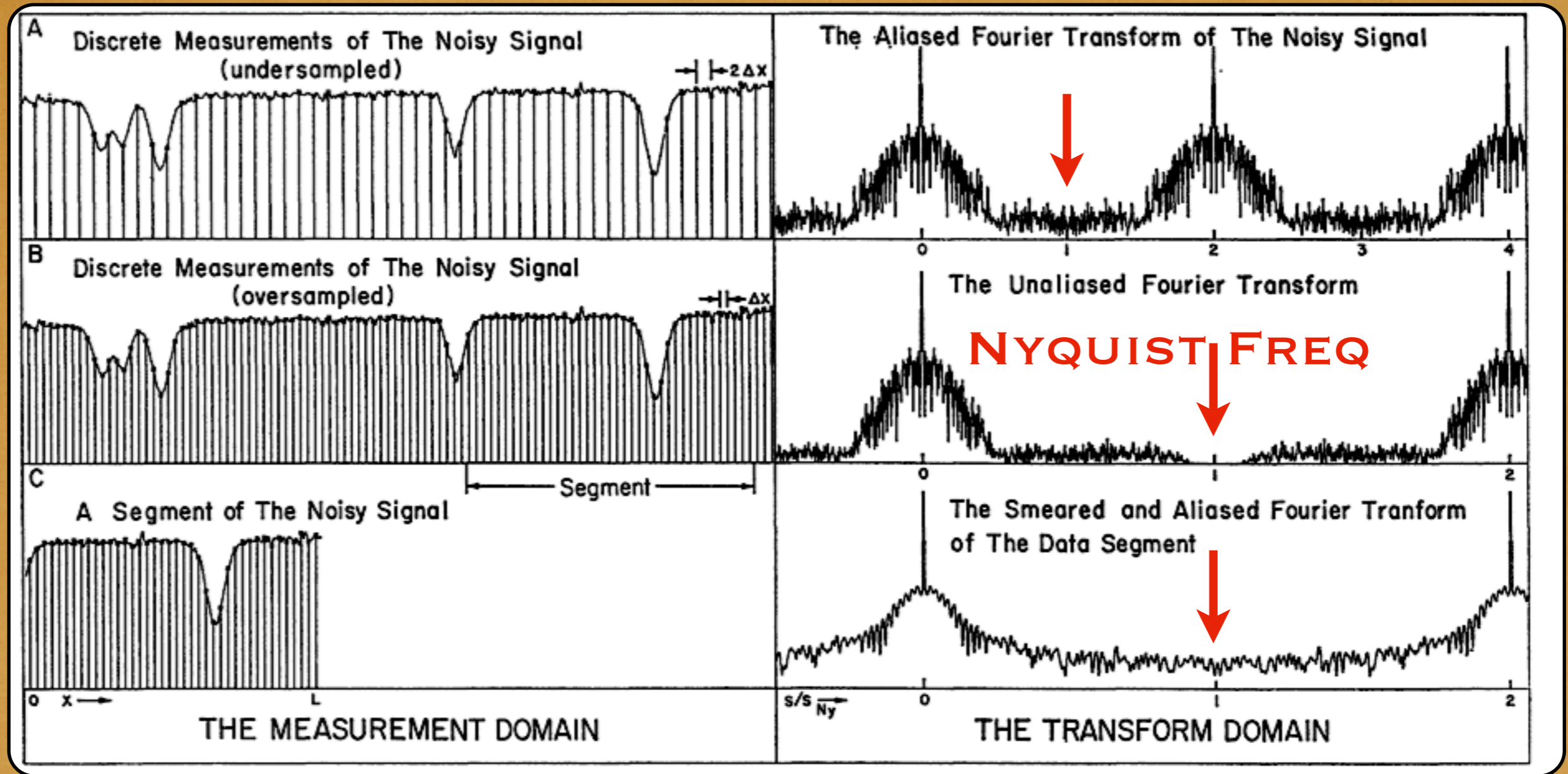
# NYQUIST THEOREM: CONT'D



SAMPLING; BRAULT & WHITE 1971, A&A



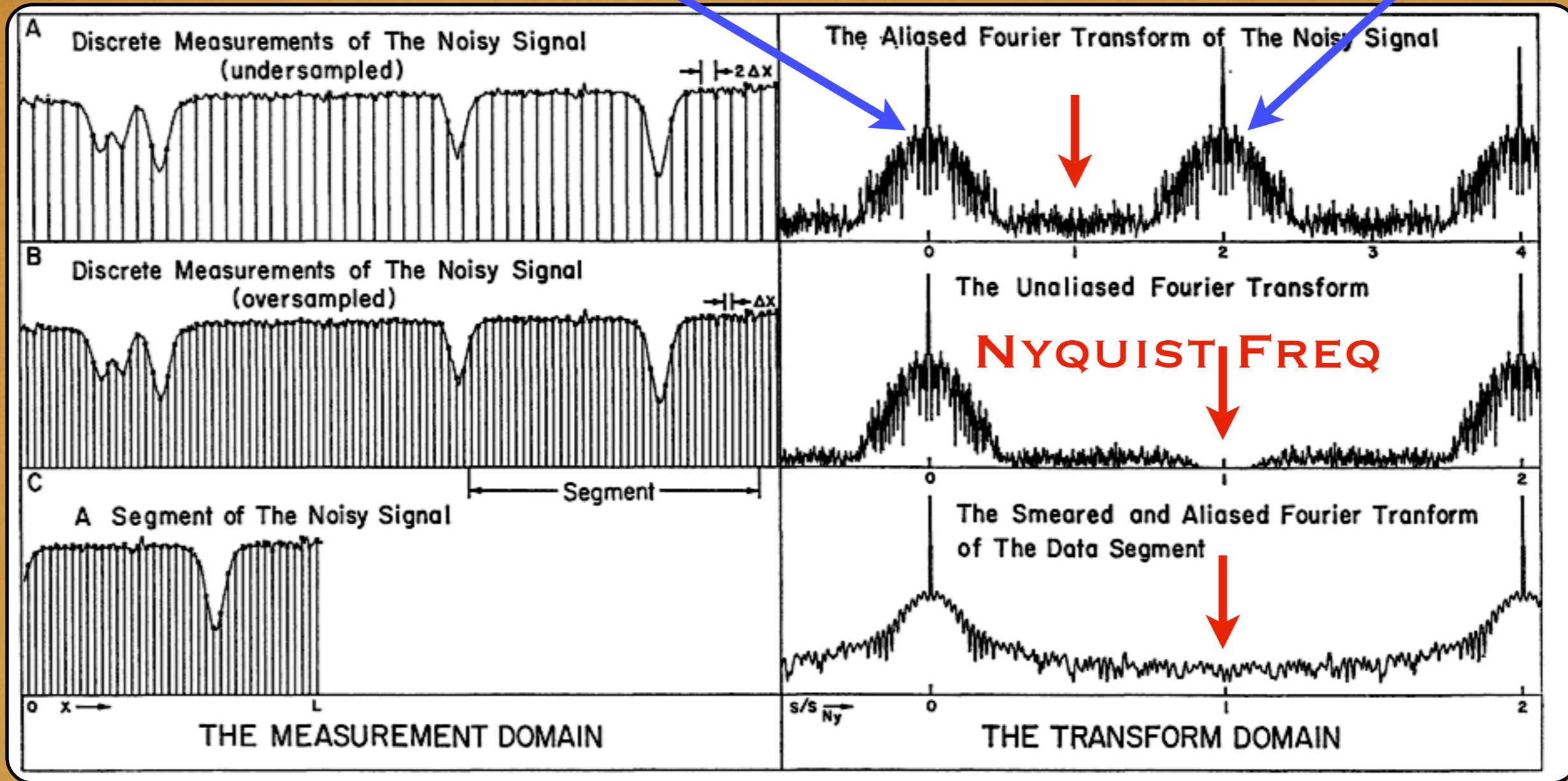
# NYQUIST THEOREM: CONT'D



SAMPLING; BRAULT & WHITE 1971, A&A

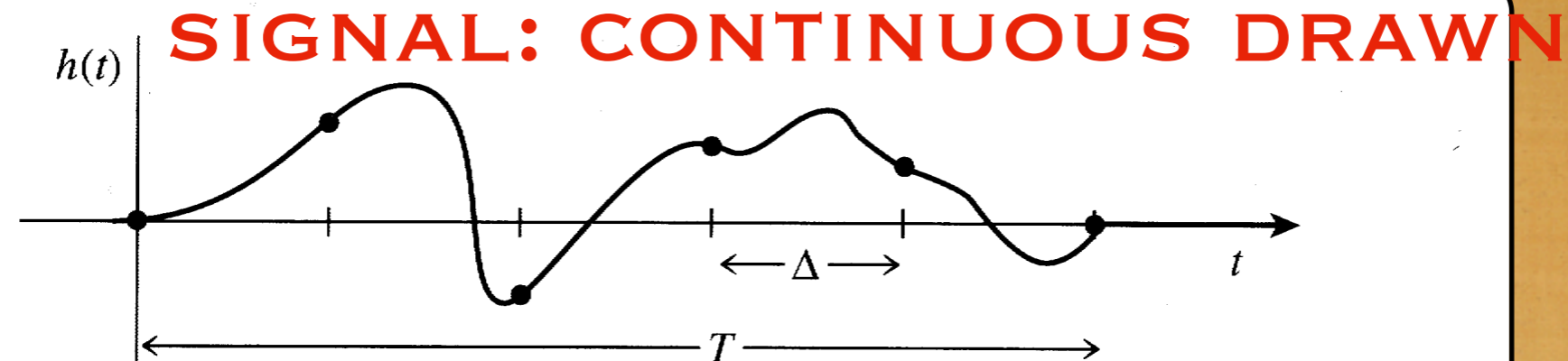
# NYQUIST THEOREM: CONT'D

SAMPLING CAUSES REPLICATION OF SIGNAL

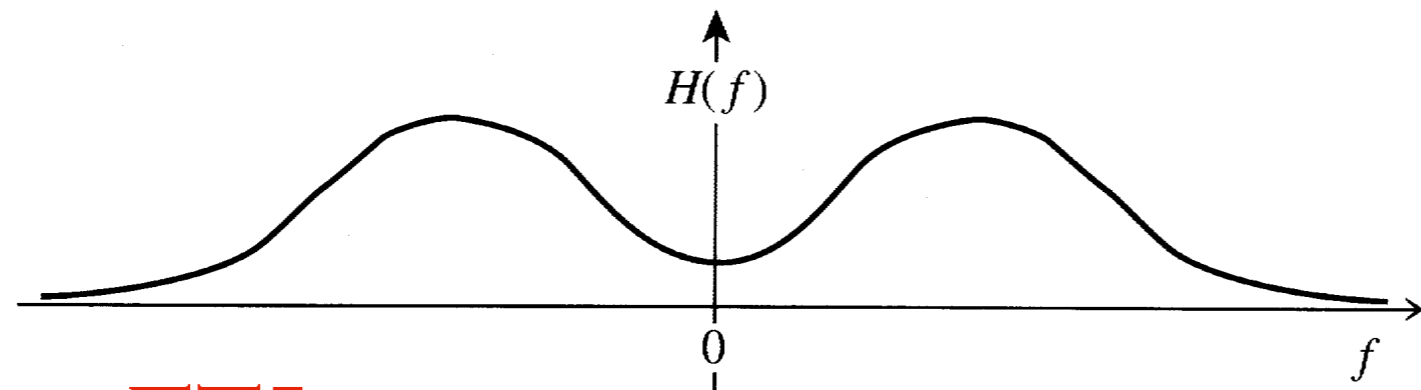


SAMPLING; BRAULT & WHITE 1971, A&A

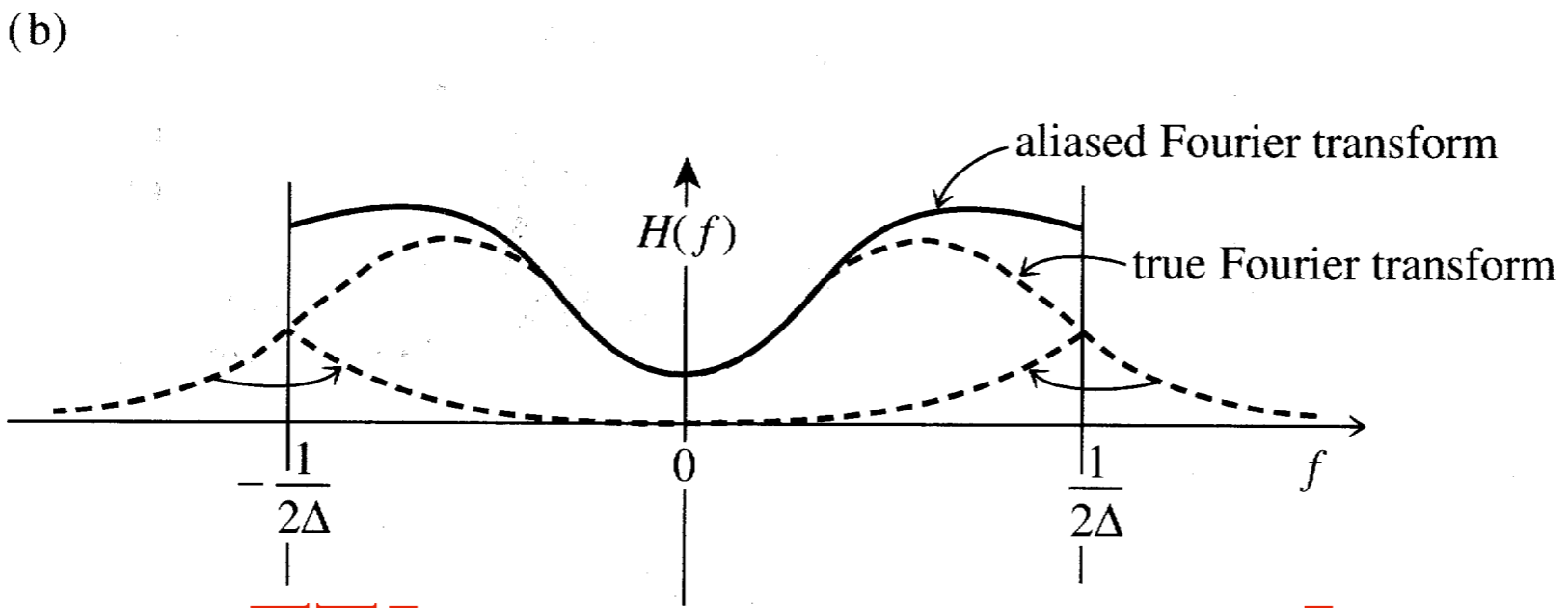
# ALIASING



(a) **SIGNAL: SAMPLES DOTS**



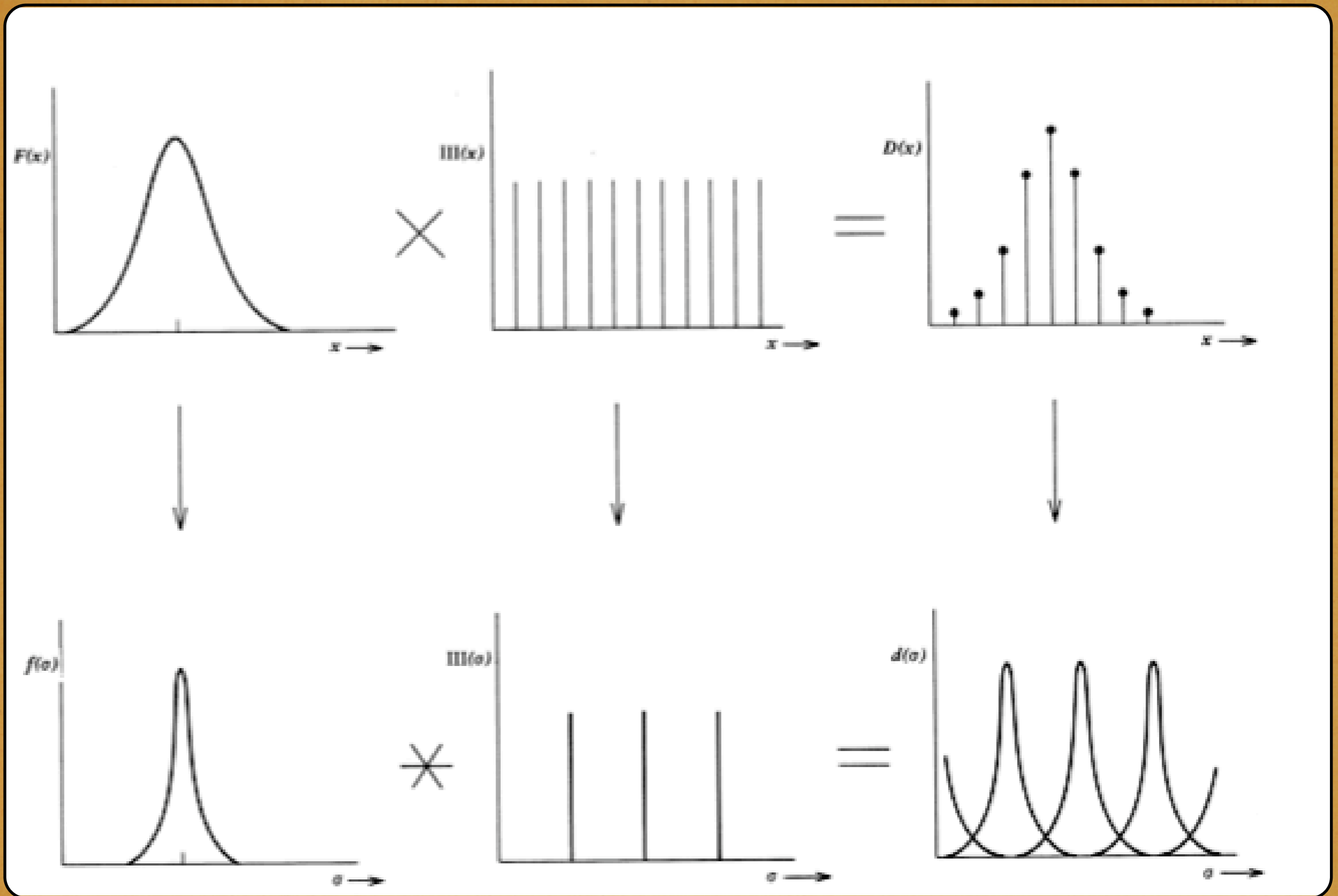
**FT[CONTINUOUS SIGNAL]**



**FT[SAMPLED SIGNAL]**

(c)

# ALIASING: CONT'D



CONVOLUTION WITH SHAH FUNCTION: REPLICATION

SAMPLING: HIGH FREQUENCIES ARE FILTERED OUT  
WINDOW: LOW FREQUENCIES ARE FILTERED OUT



LEADS TO BAND LIMITED DATA

## FILTERING

→ FREQUENCY FILTERING  $Y(f) = X(f)H(f)$

$$y(t) = \int_{-\infty}^{\infty} x(t - \theta)h(\theta)d\theta$$

$$y(t) = x(t) * h(t)$$

FILTERING OF PROCESS X WITH FILTER H

→ TIME FILTERING

MEASURE A PROCESS  $x(t)$  OVER INTERVAL  $T$  ASSUMED  
ZERO OUTSIDE  $T$

$$\equiv y(t) = \Pi\left(\frac{t}{T}\right)x(t)$$

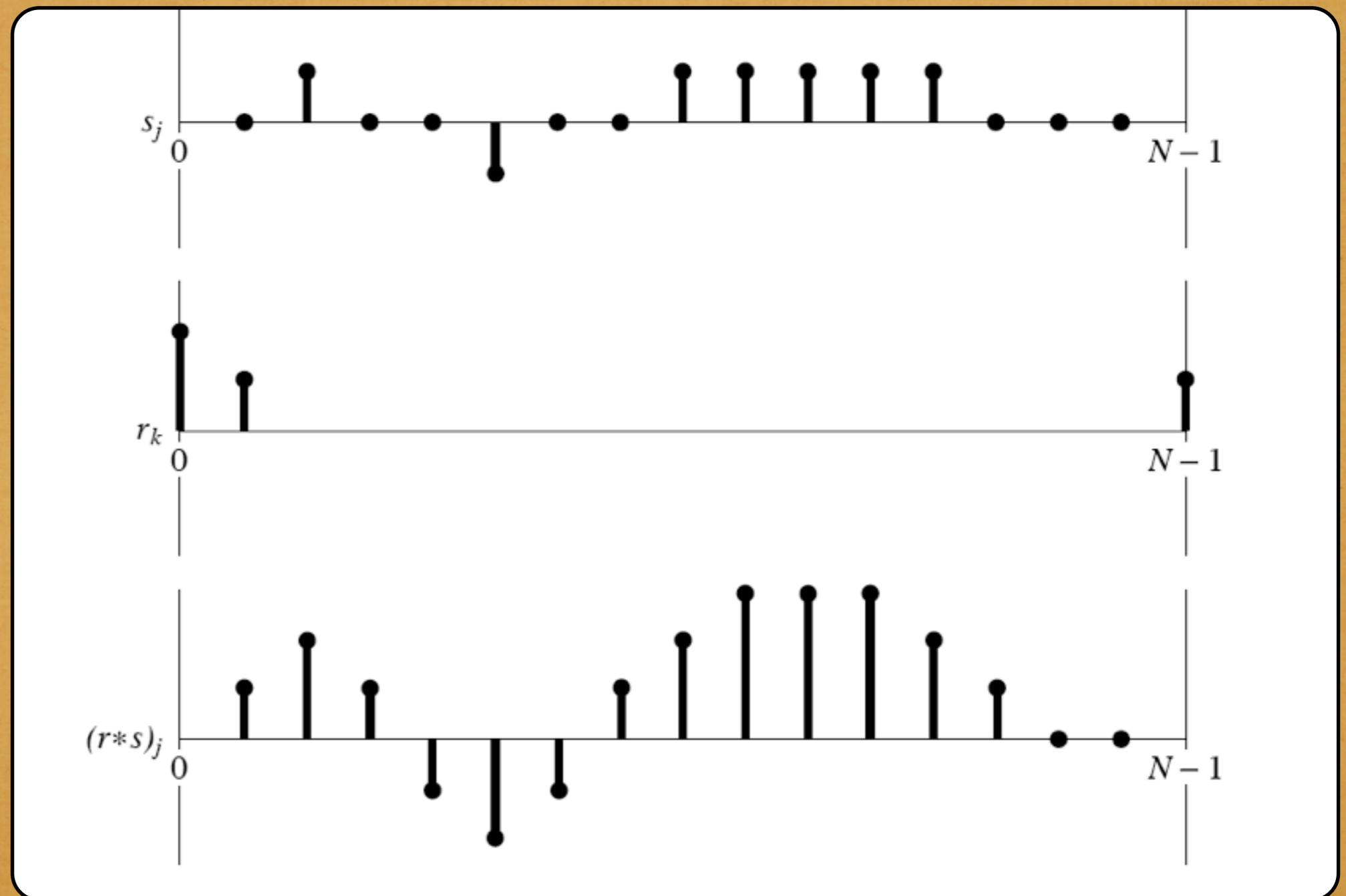
$$Y(f) = X(f) * T \text{sinc}(Tf)$$

ALL INFORMATION ABOUT FREQUENCIES  $< 1/T$  IS LOST!

# (OPTIMAL) FILTERING

NUM RES CHAPTER 13.0-13.3

DECONVOLVE MEASURED SIGNAL AND  
RESPONSE FUNCTION OF SAMPLED DATA



NUM RES CHAPTER 13.1

# DISCRETE CONVOLUTION

## THEOREM

$$(r * s)_j \equiv \sum_{k=-M/2+1}^{M/2} s_{j-k} r_k \Leftrightarrow S_n R_n$$

## DISCRETE DECONVOLUTION

$$\frac{\tilde{F}(r * s)_j}{R_n} = S_n$$

HOWEVER NOISE AND UNCERTAINTIES IS  
RESPONSE CAN MAKE THIS PROCESS  
UNRELIABLE

# NOISE REMOVAL BY OPTIMAL FILTERING

$$cs(t) = s(t) + n(t)$$

$s(t)$  is the smeared signal i.e. true  $\times$  response

DESIGN AN OPTIMAL FILTER  $\phi(t)$  THAT  
PRODUCES A SIGNAL  $U(t)$  AS CLOSE AS  
POSSIBLE TO  $U(t)$

$$\widetilde{U}(f) = \frac{C(f)\phi(f)}{R(f)}$$

CLOSE IN LEAST SQUARE SENSE

$$\int_{-\infty}^{\infty} |\widetilde{U}(f) - U(f)|^2 df \text{ IS MINIMISED}$$



# NOISE REMOVAL BY OPTIMAL FILTERING

$$\int_{-\infty}^{\infty} \left| \frac{[S(f) + N(f)]\phi(f)}{R(f)} - \frac{S(f)}{R(f)} \right|^2 df$$

$\int S(f)N(f)df$  TERMS ARE ZERO SINCE NOISE  
AND SIGNAL ARE UNCORRELATED

$$\int_{-\infty}^{\infty} |R(f)|^{-2} \underbrace{\{|S(f)|^2 |1 - \phi(f)|^2 + |N(f)|^2 |\phi(f)|^2\}}_{\Theta} df$$

$\Theta$  MINIMISED WITH RESPECT TO  $\phi$

# NOISE REMOVAL BY OPTIMAL FILTERING

$$\frac{d\theta}{d\phi} = 0$$

$$-2S^2(1 - \phi) + 2N^2\phi = 0$$

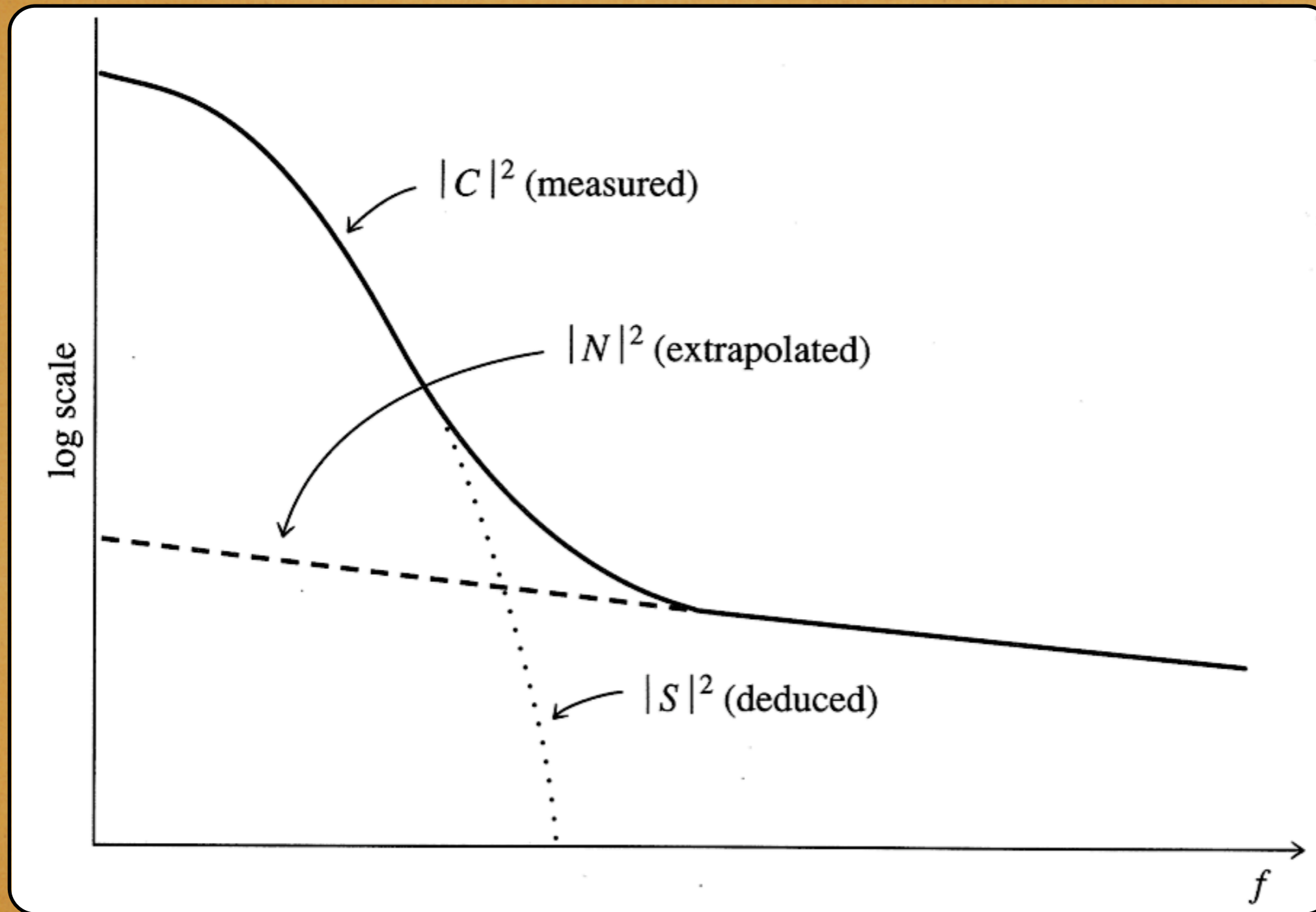
OPTIMAL FILTER

$$\phi = \frac{S^2}{S^2 + N^2}$$

**DOES NOT CONTAIN TRUE SIGNAL DIRECTLY!**

$$|S(f)|^2 + |N(f)|^2 = PDS(f) = |CS(f)|^2$$

# NOISE REMOVAL BY OPTIMAL FILTERING



PAGE 542 NUM RES

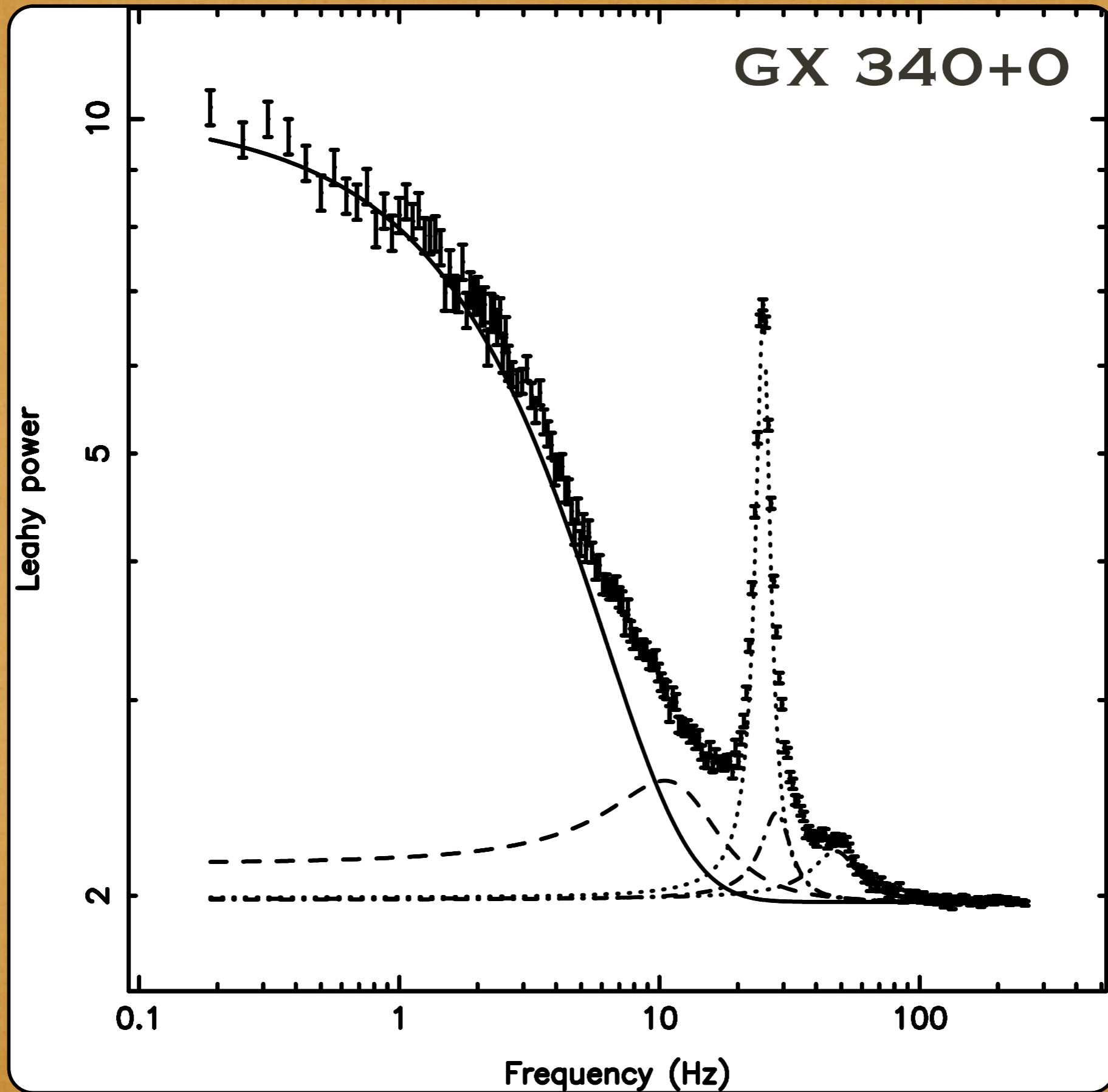
# SOME APPLICATIONS

ON OPTIMAL DETECTION OF POINT SOURCES IN CMB MAPS  
VIO ET AL, 2002, A&A, 391, 789

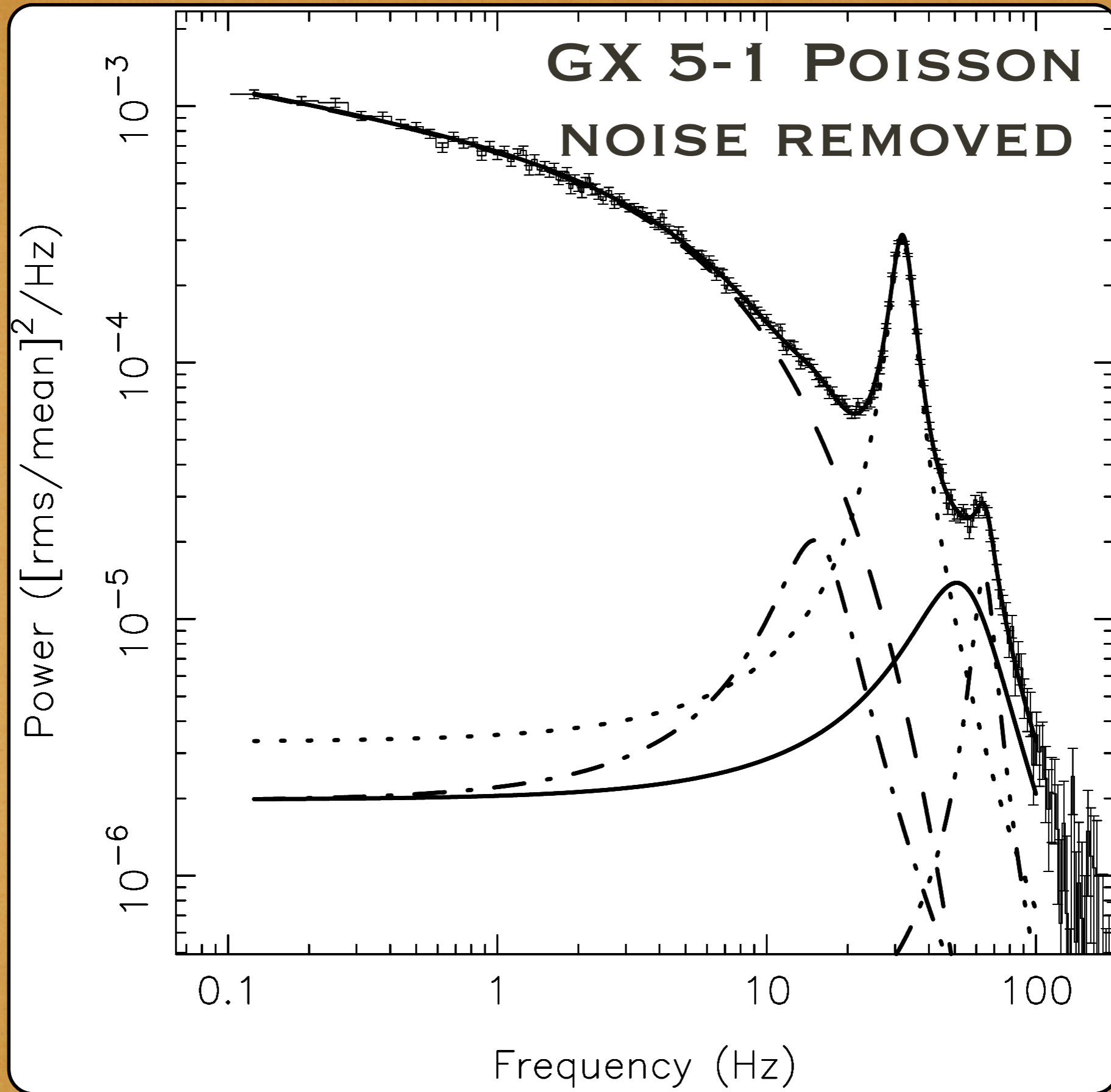
AN OPTIMAL FILTER FOR THE DETECTION OF GALAXY  
CLUSTERS THROUGH WEAK LENSING  
MATURI, ET AL. 2005, A&A, 442, 851

THE LARGEST SCALE PERTURBATIONS: A WINDOW ON THE  
PHYSICS OF THE BEGINNING  
WANDELT, NEW ASTRONOMY REVIEW, 2006, 11, 900

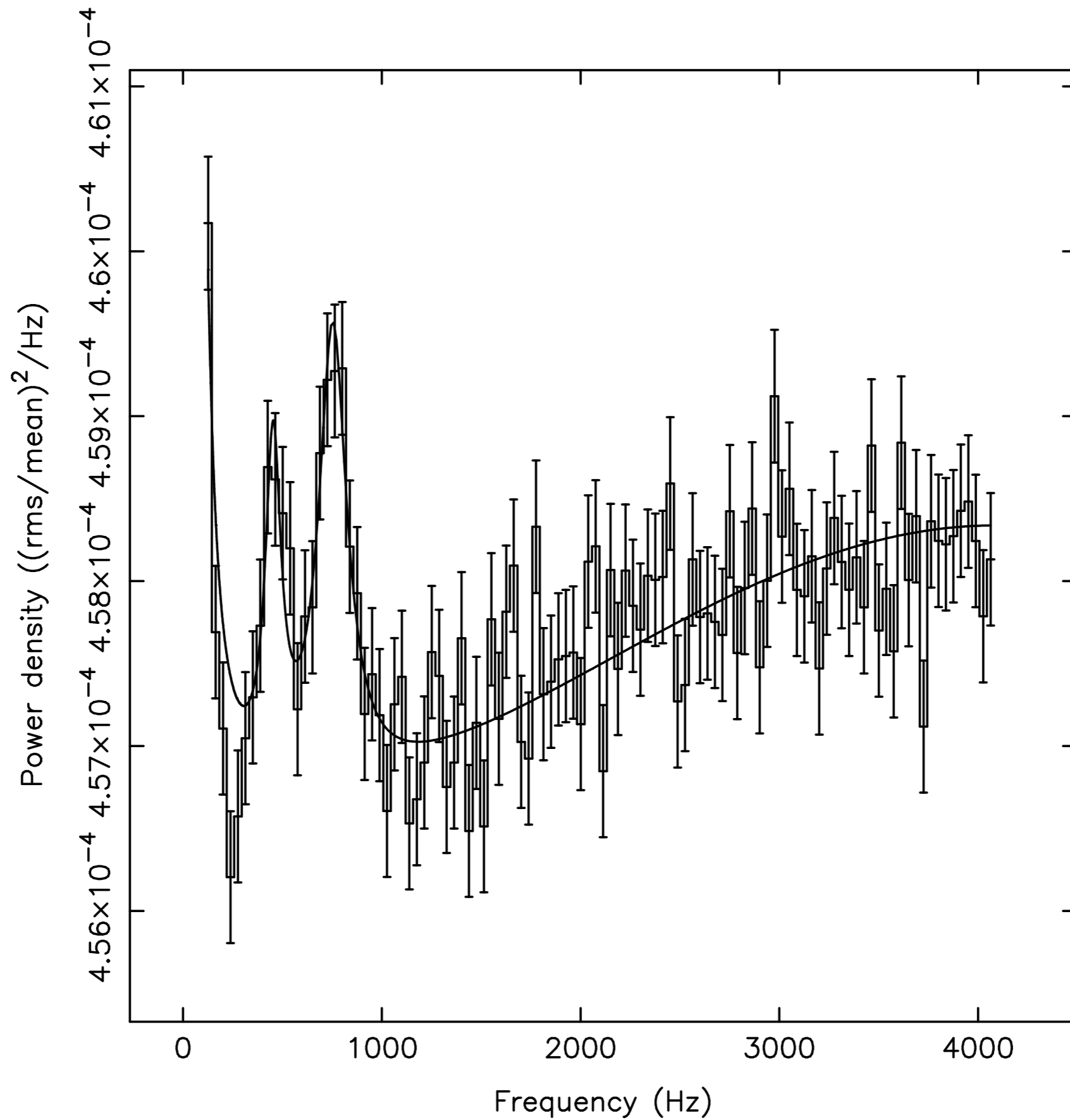
# C.F. X-RAY TIMING EXPERIMENTS



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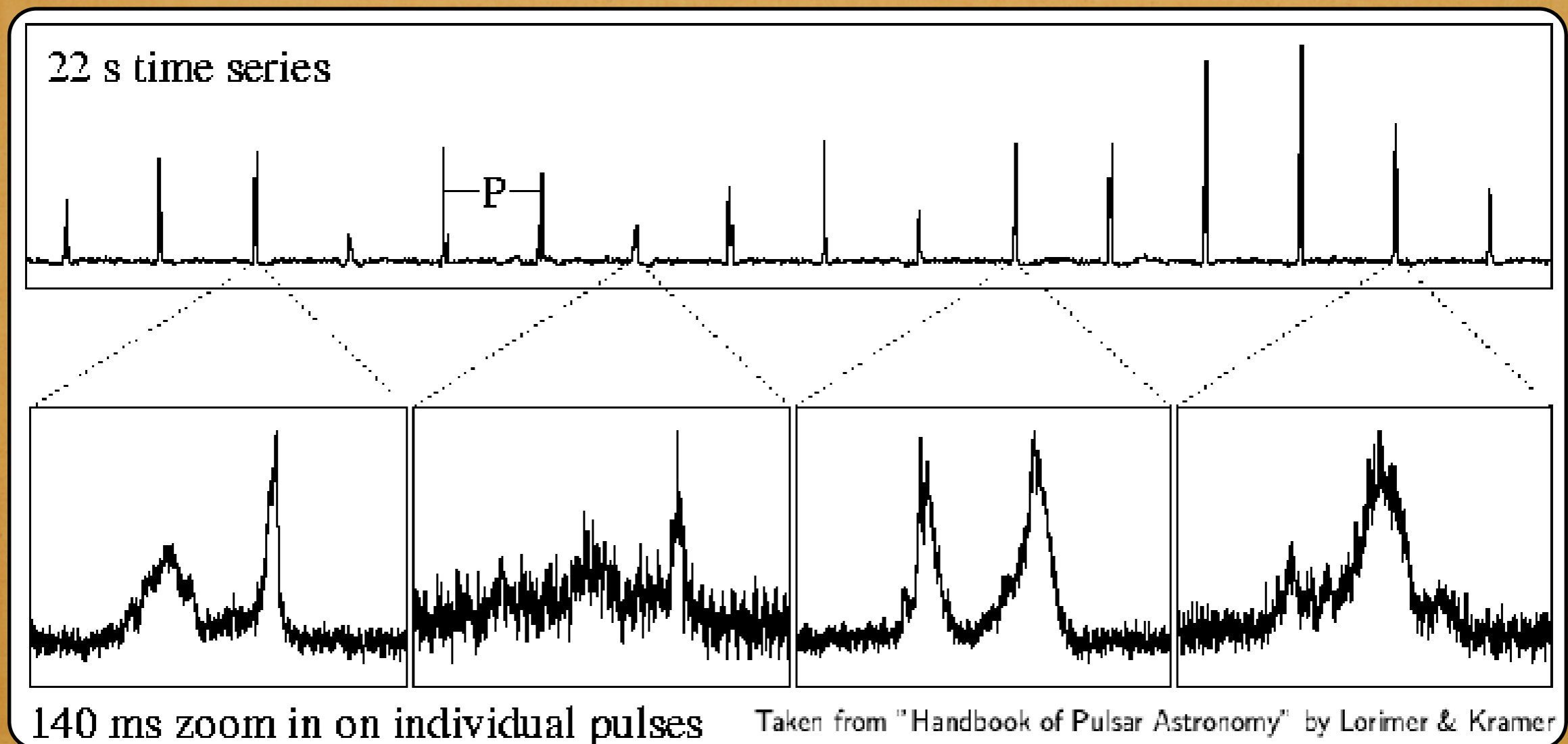


# ESTIMATING THE MOMENTS OF A STOCHASTIC PROCESS

CHAPTER 2.2.2

SEE ALSO APPENDIX B3.2, LENA EA.

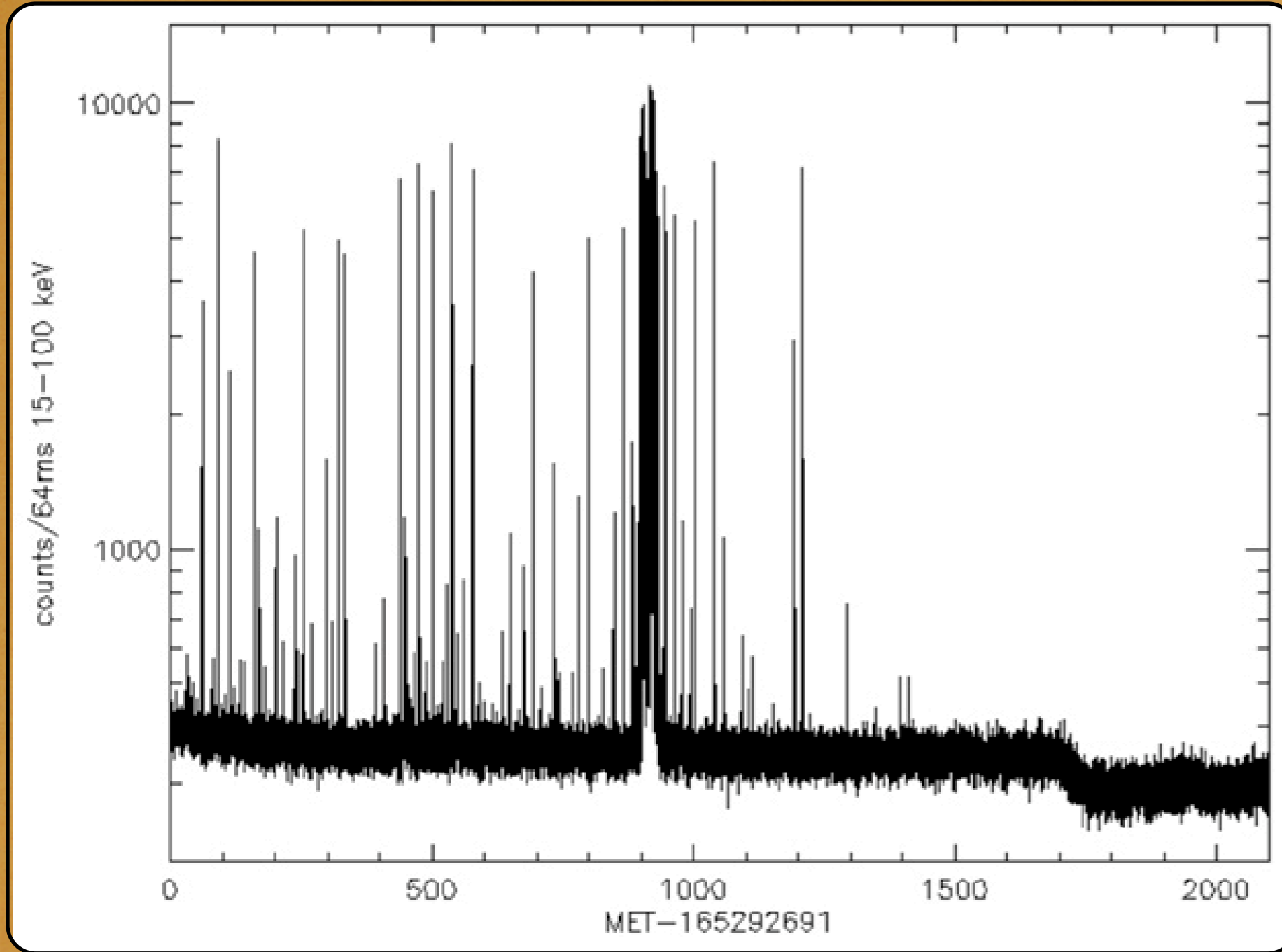
## HOW REPRESENTATIVE IS A MEASUREMENT OF A S.P.?





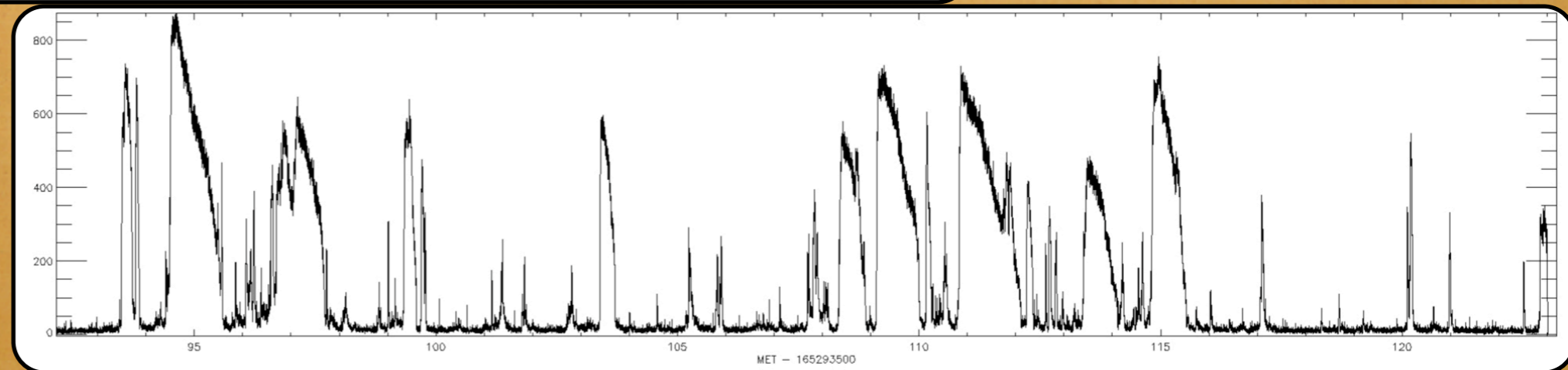
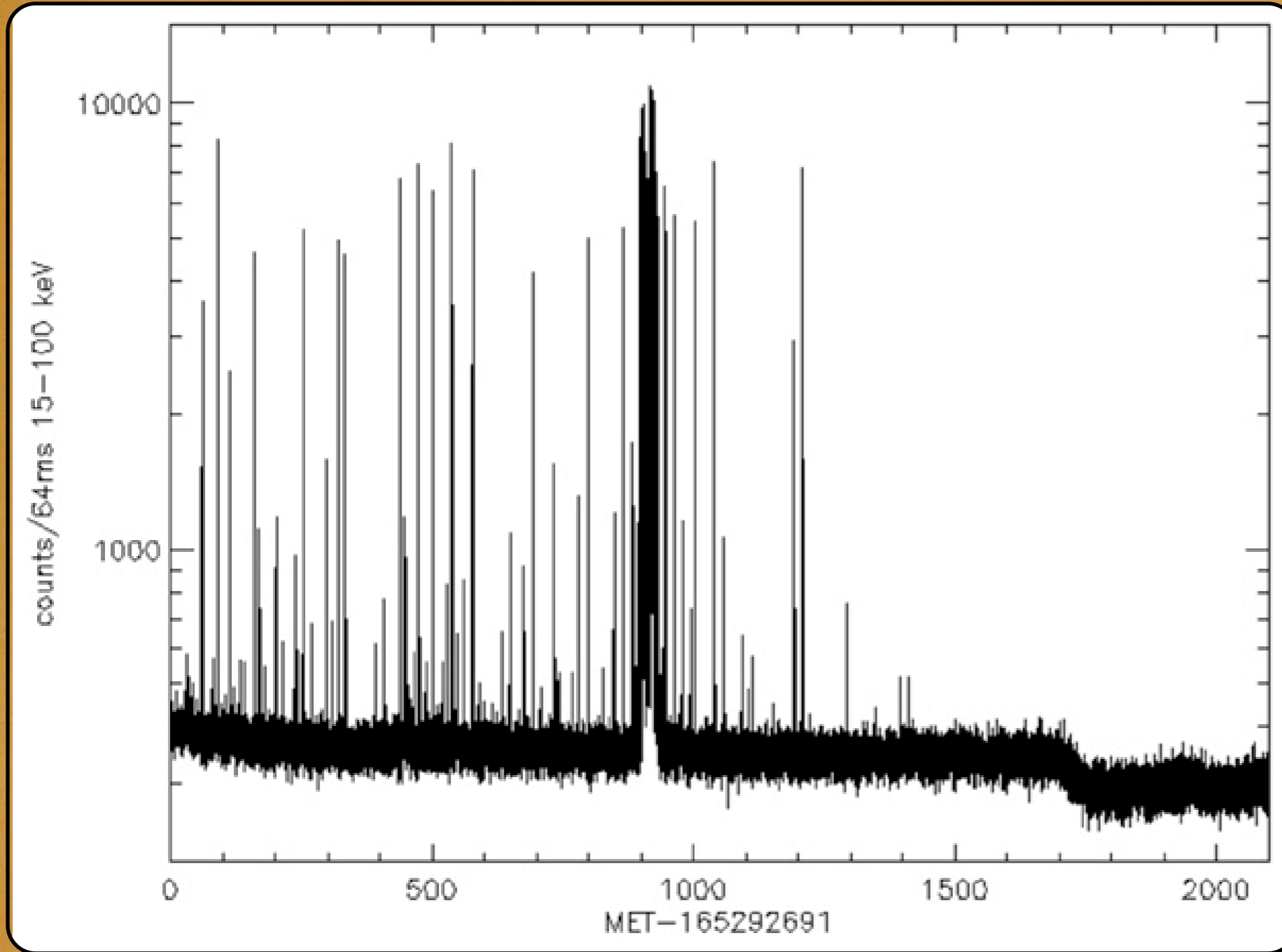
# HOW REPRESENTATIVE IS A MEASUREMENT?

SGR 1900+14

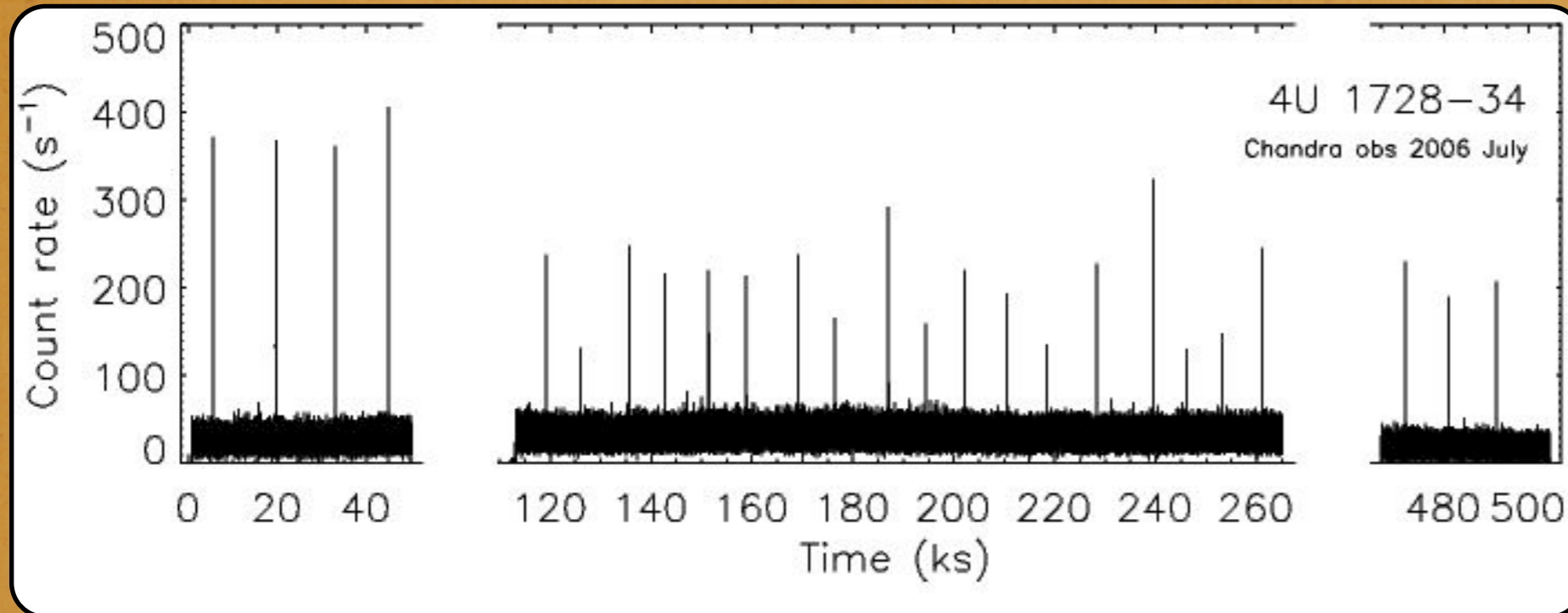


# HOW REPRESENTATIVE IS A MEASUREMENT?

SGR 1900+14



# DIFFERENT FOR THE CASE OF TYPE I X-RAY BURSTS



GALLOWAY PRIVATE COMMUNICATION

1 MEASUREMENT OF  $x(t)$  IN A TIME  $T \Rightarrow$  WINDOWING AND  
AVERAGING OVER TIME  $\Delta T$

WINDOWING:  $y(t) = \Pi\left(\frac{t}{T}\right)x(t)$

AVERAGING:

$$z(t) \equiv y_{\Delta T}(t) = \frac{1}{\Delta T} \int_{t-\Delta T/2}^{t+\Delta T/2} y(t') dt' = \frac{1}{\Delta T} \int_{-\infty}^{\infty} \Pi\left(\frac{t-t'}{\Delta T}\right) y(t') dt'$$

= LOW-PASS FILTER, REMEMBER NYQUIST THEOREM