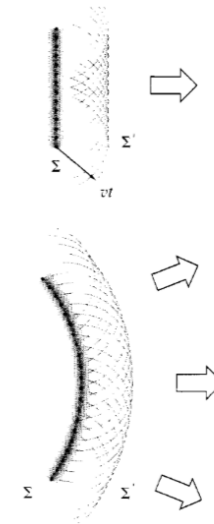


Wave Optics: Diffraction & Interference

- Huygens sources & wave propagation
 - Math: Green's function
- Fraunhofer vs. Fresnel interference
- Applications of diffraction & interference:
 - Diffraction limit of telescopes & microscopes
 - Fresnel diffraction behind halfscreen
 - Fresnel zone plates
 - Interference behind realistic double slit
 - Interference behind realistic N slit

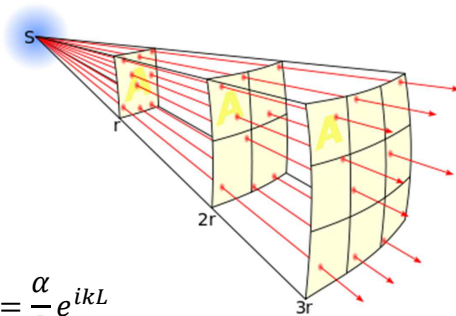
Huygens sources help to visualize field propagation

Rays & Phase fronts



Hecht, Fig. 4.26

Huygens sources: quantitative formulation

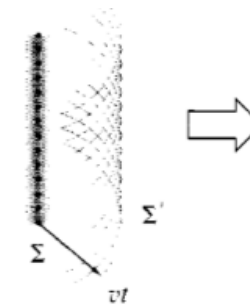


$$E(L) = \frac{\alpha}{L} e^{ikL}$$

Q: What is α ?

$$\alpha = E_s \cdot \frac{dx \cdot dy}{i\lambda}$$

Plane waves propagate undisturbed

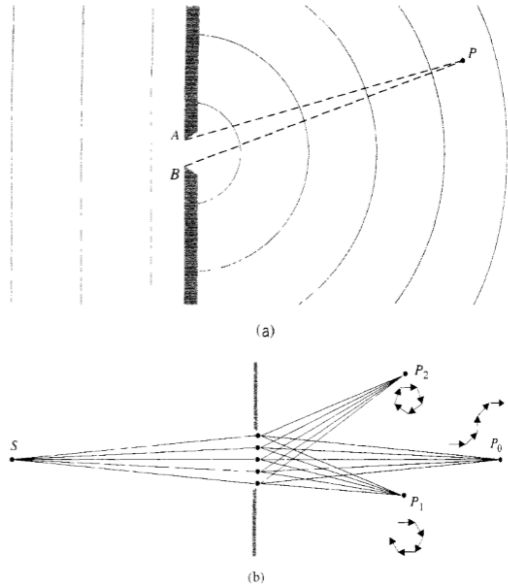


$$E(L_0) = E_s \int_0^\infty \frac{1}{i\lambda L} e^{ikL} 2\pi r dr$$

Q: Prove this equals E_s

$$\approx E_s \int_0^\infty \frac{1}{i\lambda L_0} e^{ik(L_0 + \frac{r^2}{2L_0})} 2\pi r dr = E_s e^{ikL_0} \int_0^\infty \frac{1}{i\lambda L_0} e^{i\pi \frac{r^2}{\lambda L_0}} \pi dr^2 = E_s e^{ikL_0}$$

Diffraction: integration over Huygens sources



$$E = \int_0^\infty \frac{E(x,y)}{i\lambda L} e^{ikL} dx dy$$

Q: When is this description valid?

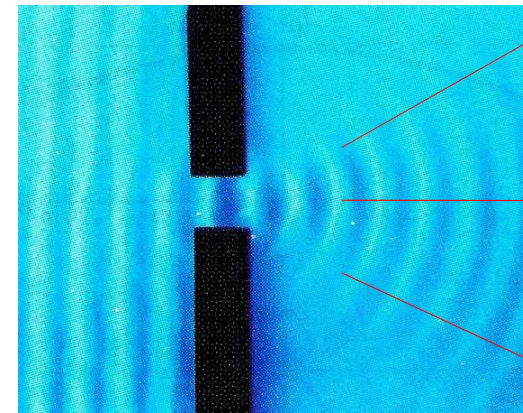
- Spectral coherence (monochromatic)
- Spatial coherence (one coherent source)

Q: How about backward light?

- Factor $(1+\cos\theta)/2$

Hecht, Fig. 10.1

Diffraction behind an aperture



Diffraction angle $\theta \approx \lambda/D$

Fresnel

Fraunhofer

NOTE: long-wavelength water wave is not representative for optics

Diffraction pattern changes with distance from slit

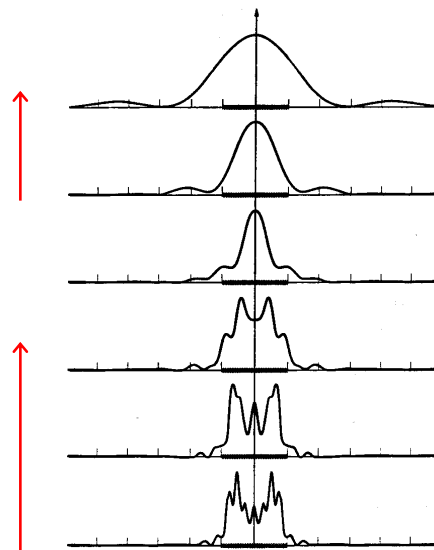
10.2 Fraunhofer diffraction

In **far field**:
- fixed shape
- only scaling $\propto L$

In **near field**:
Shape changes with distance from slit

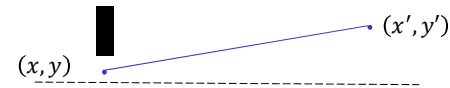
10.3 Fresnel diffraction

Q: What is far enough?



Hecht Fig. 10.5

Fraunhofer regime when $L \gg D^2/\lambda$



$$E(x', y') = \int_0^\infty \frac{E(x,y)}{i\lambda L} e^{ikL} dx dy$$

$$L \approx L_0 + \frac{(x-x')^2 + \dots}{2L_0} \approx L_0 + \frac{x^2 - 2xx' + x'^2}{2L_0}$$

$$E(x', \dots) \approx e^{ikL_0} e^{i\frac{kx'^2}{2L_0}} \int_0^\infty E(x, \dots) e^{-i2\pi\frac{xx'}{\lambda L_0}} e^{i\frac{kx^2}{2L_0}} \dots dx dy / i\lambda L_0$$

Fresnel term only for $L_0 \ll x^2/\lambda$

Fourier transform in Fraunhofer regime: $L_0 \gg x^2/\lambda$

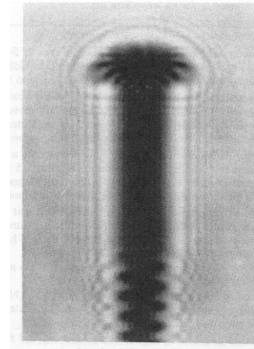
Many examples of Fraunhofer & Fresnel diffraction

- Fresnel diffraction behind slit & half screen
- Fresnel zone plate
- Fraunhofer diffraction behind slit
- Fraunhofer diffraction behind circular aperture
- Rayleigh criterium for resolution
- Diffraction limit of telescope & microscope
- Fraunhofer interference behind Young's double slit
- Interference behind N identical slits

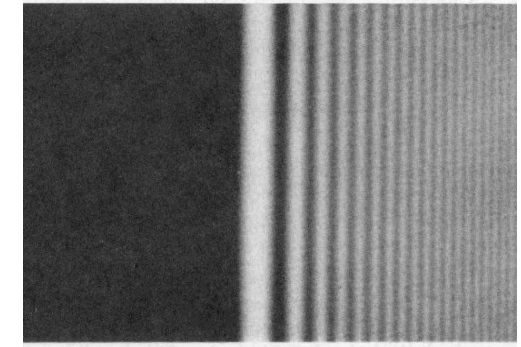
$$\sin \theta = \frac{1.22\lambda}{D} \Rightarrow \theta \approx \frac{1.22\lambda}{D}$$

Interference fringes in Fresnel regime

Pattern behind screw



Pattern behind half opaque screen



Fresnel diffraction in 2D and 3D

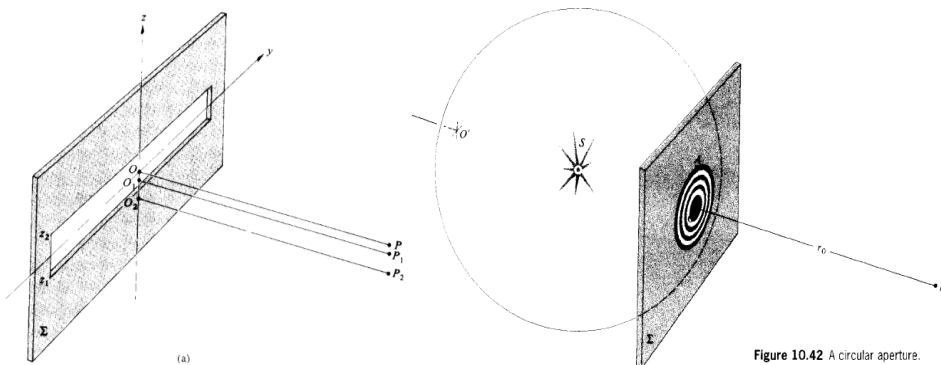


Figure 10.42 A circular aperture.

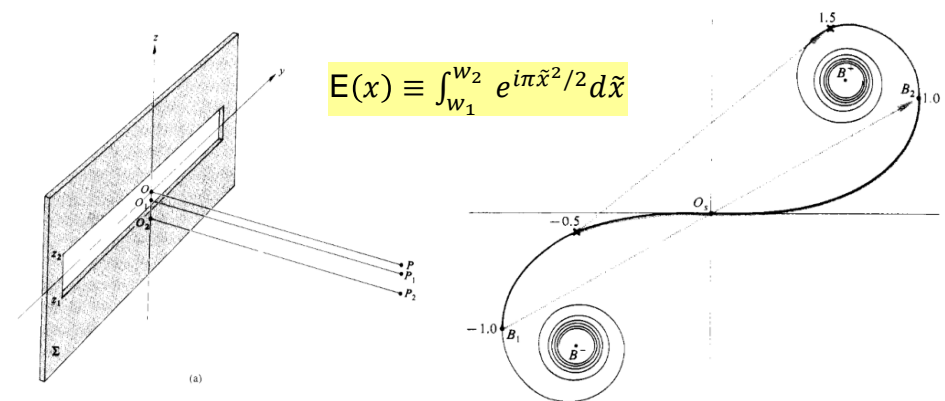
2D: Fresnel integral

$$F(w) \equiv \int_0^w e^{i\pi x^2/2} dx$$

3D: Trivial integral

$$\int_0^{x^2=r^2} e^{i\pi x^2/2} \pi dx^2$$

Fresnel diffraction in 2D: Fresnel number & Cornu spiral

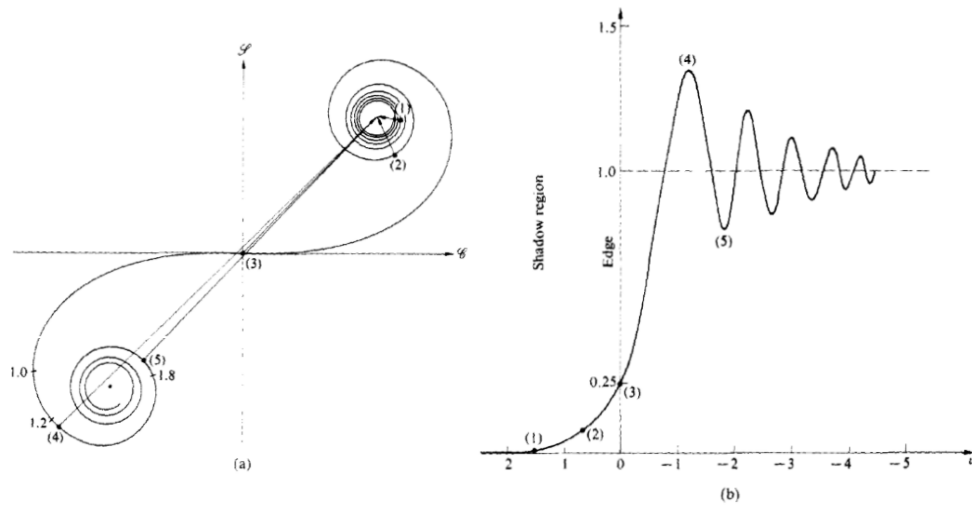


$$E(x) \equiv \int_{w_1}^{w_2} e^{i\pi \tilde{x}^2/2} d\tilde{x}$$

Figure 10.55 Cornu spiral for the slit.

Q: What is w_1 and w_2 ? A: $\frac{\pi \tilde{x}^2}{2} = k\Delta L = \frac{\pi x^2}{\lambda L}$

The semi-infinite opaque screen



Hecht Fig.10.59

Fresnel diffraction behind circular aperture

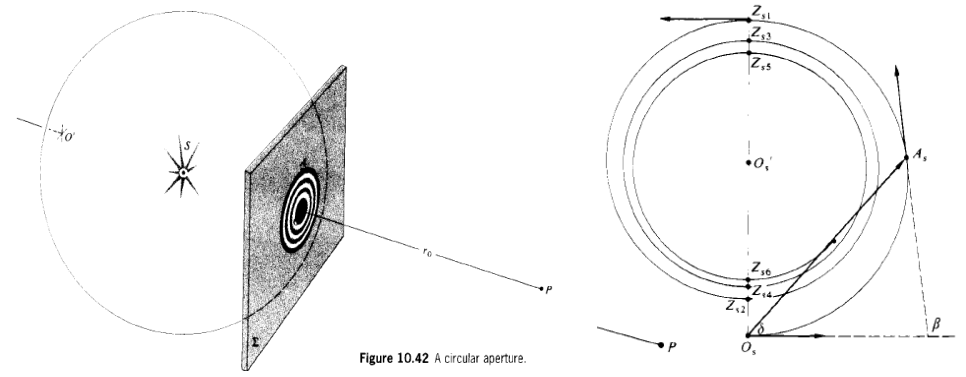


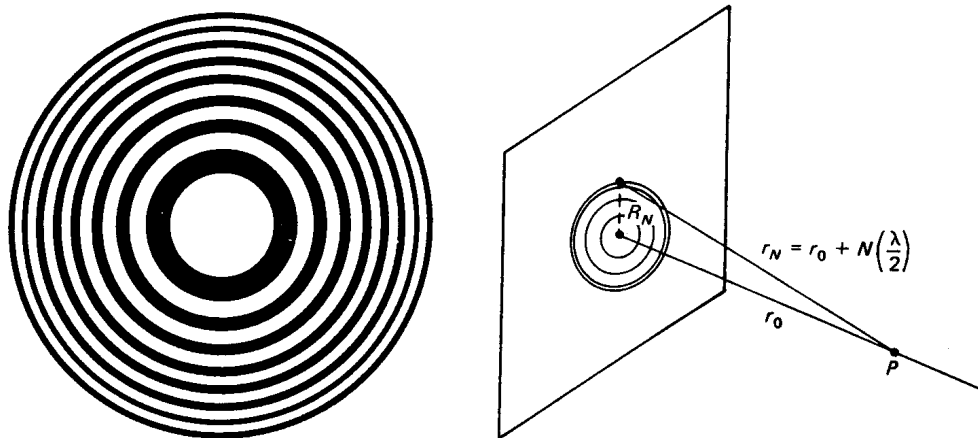
Figure 10.42 A circular aperture.

$$E_{\text{centre}} = \int_0^{x^2 = \frac{2r^2}{\lambda L}} e^{i\pi x^2/2} \pi dx^2$$

Hecht Fig.10.41 & 10.42

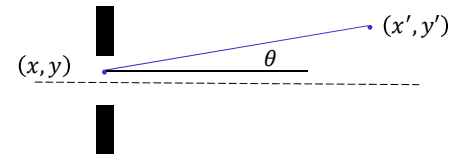
Fresnel lens: focusing with black-and-white pattern !!

How to increase the intensity in P by blocking Fresnel zones in the beam !



Fresnel number $N_F = b^2/L\lambda$ counts number of "Fresnel zones"

Fraunhofer diffraction behind slit of width w

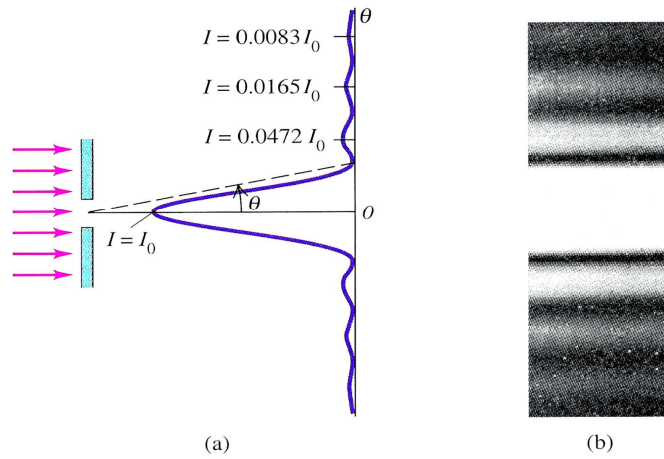


$$E(x') \propto \int_{-b/2}^{b/2} E(x) e^{ik\Delta L} dx = \int_{-b/2}^{b/2} E_0 e^{-ik \sin \theta x} dx$$

As graphical construction yields $\Delta L = -\sin \theta x$

$$E(\theta) \propto \int_{-b/2}^{b/2} E_0 e^{-ik \sin \theta x} dx = -E_0 b \frac{\sin \beta}{\beta} \text{ with } \beta = \frac{1}{2} k \sin \theta b$$

Diffraction minima and maxima behind slit

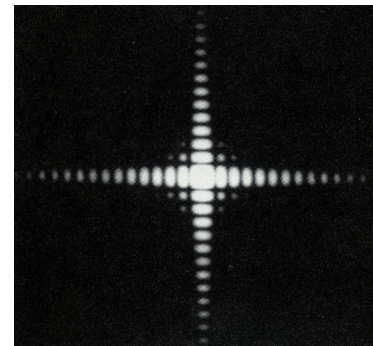


Opening angle of first minimum: $\sin(\theta_{\text{first min.}}) = \frac{\lambda}{b}$

Q: What is spacing maxima & minima A: Only minima equidistant

Fraunhofer diffraction behind rectangular hole

Diffraction behind square hole

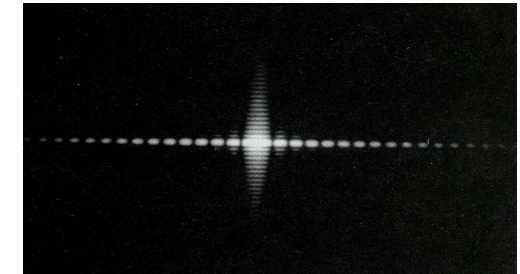


Field profile factorizes !!

$$E_P = E_0 \frac{\sin \beta_x}{\beta_x} \frac{\sin \beta_y}{\beta_y}$$

$$\beta_x = \frac{1}{2} k_0 b_x \sin \theta_x ; \beta_y = \frac{1}{2} k_0 b_y \sin \theta_y$$

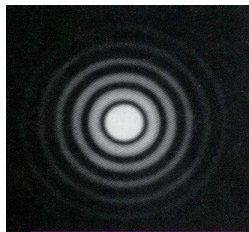
Diffraction behind rectangular hole



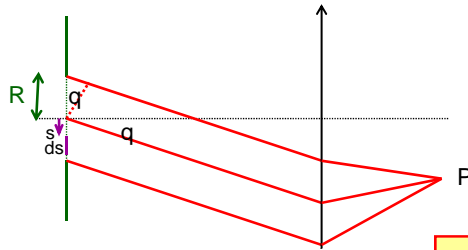
What is orientation of hole?



Fraunhofer diffraction behind round hole



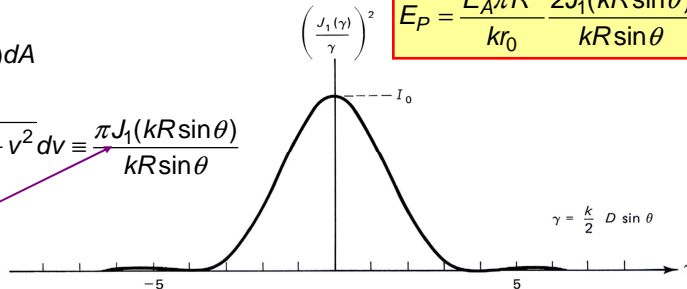
Field per unit Area



$$E_P \propto \frac{E_A}{r_0} \iint_{\text{Area}} \exp(isk \sin \theta) dA$$

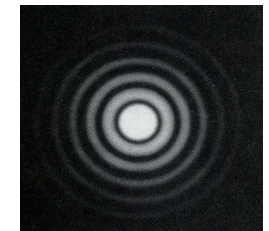
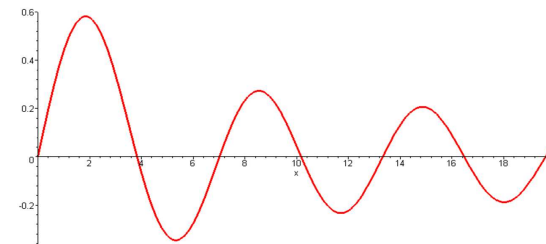
$$\propto \int_{-1}^1 \exp(ivkR \sin \theta) \sqrt{1-v^2} dv \equiv \frac{\pi J_1(kR \sin \theta)}{kR \sin \theta}$$

Bessel function



$$E_P = \frac{E_A \pi R^2}{kr_0} \frac{2J_1(kR \sin \theta)}{kR \sin \theta}$$

Properties of Bessel function $J_1(x)$



Central disk = Airy disk
Rings = Airy rings

$$\lim_{x \downarrow 0} \frac{J_1(x)}{x} = \frac{1}{2}$$

$J_1(x) = 0$ als $x = 3.832, 7.016, 10.173$

$$I_P = I_0 \left(\frac{2J_1(kR \sin \theta)}{kR \sin \theta} \right)^2$$

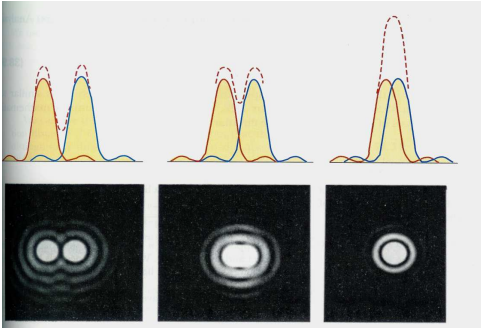
$$I = 0 \text{ als } kR \sin \theta = 3.832 \Rightarrow \sin \theta = \frac{1.22\lambda}{2R} = \frac{1.22\lambda}{D}$$

Compare with diffraction from slit of rectangle:

$$I = 0 \text{ als } \sin \theta = \frac{\lambda}{b}$$

Why is diffraction angle of round hole larger?

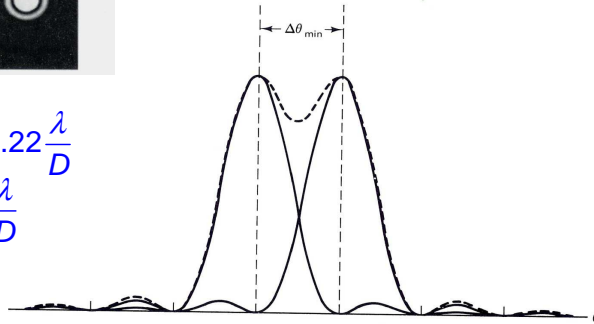
Rayleigh criterion for resolution



Q: When can you resolve two maxima?

Lord Rayleigh:

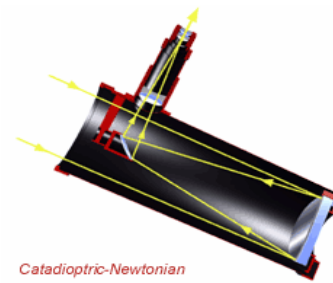
- If intensity between peaks drops 19%
- If one diffraction max. coincides with other diffraction minimum



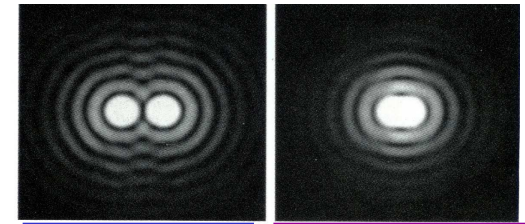
For round holes: $\Delta\theta > 1.22 \frac{\lambda}{D}$

For rectangular holes: $\Delta\theta > \frac{\lambda}{D}$

Resolution of telescope



Catadioptric-Newtonian



Well distinguishable

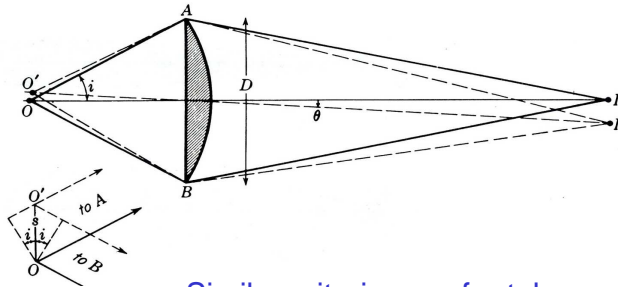
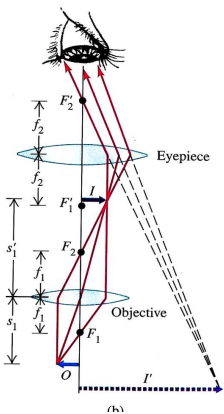
Barely distinguishable

What is angular resolution of telescope with $D = 4$ cm (for $\lambda = 500$ nm)?

Answer: $\theta_{res} \approx 15 \mu\text{rad}$ (diameter moon $\approx 0.5^\circ \approx 8$ mrad)

Note: Fluctuations in earth atmosphere ("seeing") typically limit resolution for $D > 20$ cm

Resolution of microscope



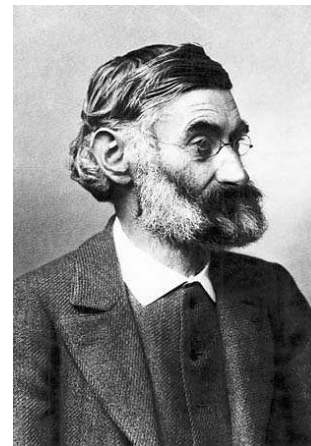
Similar criterium as for telescope:

$$\theta_{res} \approx 1.22 \frac{\lambda}{D} \Rightarrow x_{res} \approx f \theta_{res} \approx 1.22 \frac{\lambda}{D/f} \approx \frac{1.22 \lambda}{2 N.A.}$$

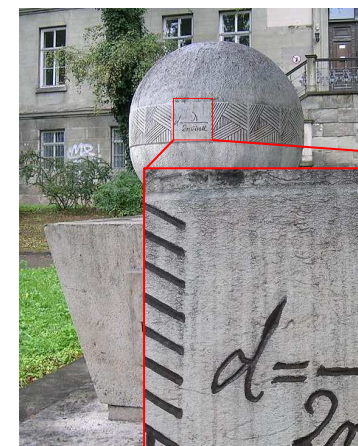
$$s_{min} = \frac{1.22 \lambda}{2n \sin i} = \frac{1.22 \lambda}{2 N.A.}$$

Note: also valid for large angles & immersion objectives

Abbe diffraction limit for microscopes



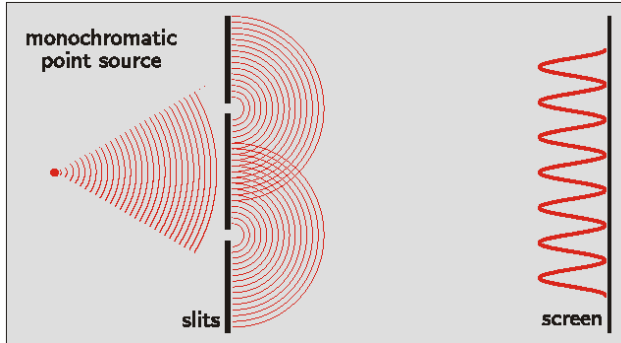
Ernst Abbe (1840-1905)



$$d = (1.22 \times) \frac{\lambda}{2n \sin \theta_{in}}$$

$$d = \frac{\lambda}{2n \sin \alpha}$$

Young's experiment with a double slit



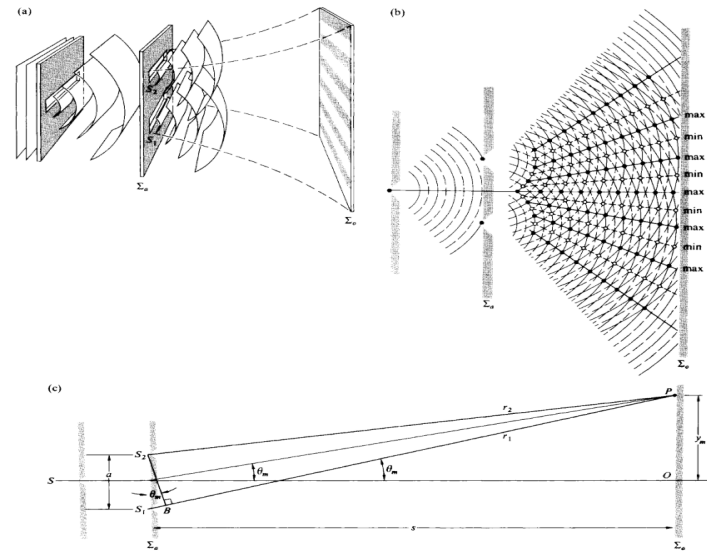
Interference of two diffracted waves

Q: Why did it take until 1803?

- Spectral coherence
- Spatial coherence

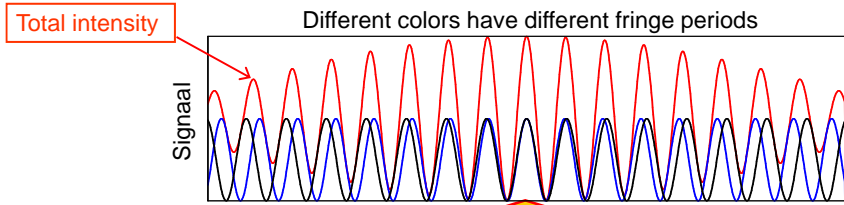
Young's double slit experiment

Path length differences are crucial

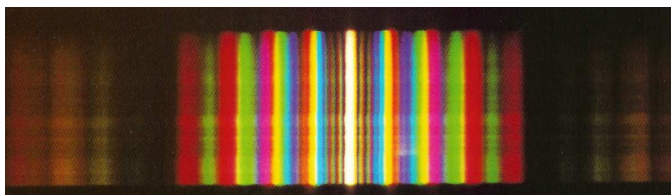


Hecht, Fig. 9.8

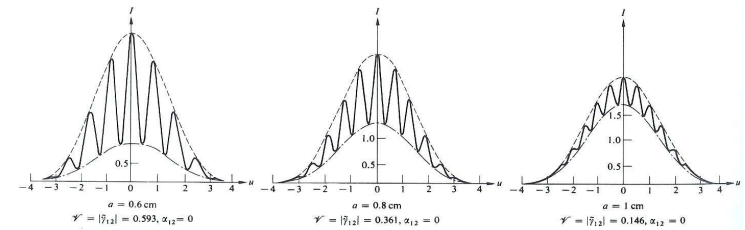
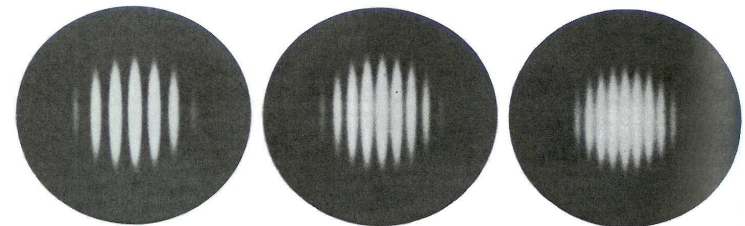
1. Young's experiment with limited spectral coherence



Constructive interference for all colors only in the center ($\theta=0$)



2. Young's experiment with limited spatial coherence



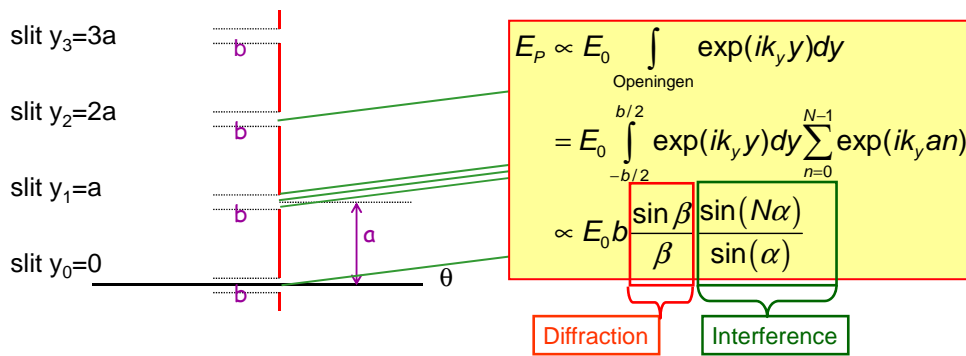
A=0.6 cm, V=0.59

A=0.8 cm, V=0.36

A=1.0 cm, V=0.15

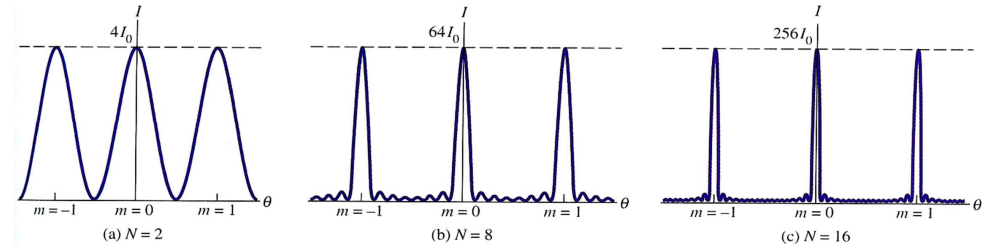
distance between slits increases

Fraunhofer diffraction behind N slits



Again: factorization of diffraction from single slit & mutual interference

Fraunhofer diffraction behind N small slits ($b \rightarrow 0$)



$$E(\theta) \propto \frac{\sin(N\alpha)}{\sin(\alpha)} ; \quad \alpha = \frac{1}{2} k_y a$$

Q: What happens at $\alpha = m\pi$?

A: Diffraction maximum (orde m with $a \cdot \sin\theta = m\lambda$)

Q: How does peak intensity scale with N?

A: Intensity $\propto N^2$ (coherent sources)

$$\lim_{\alpha \rightarrow m\pi} \frac{\sin(N\alpha)}{\sin(\alpha)} = N$$

Better angular resolution with multiple sources/detectors



Westerbork telescope array: $\lambda \approx 21$ cm ($\nu \approx 1.4$ GHz)

Angular resolution of array

