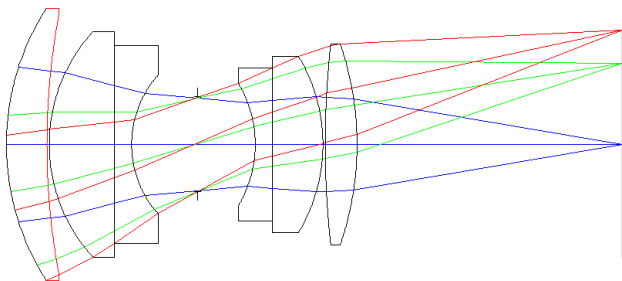


## Outline

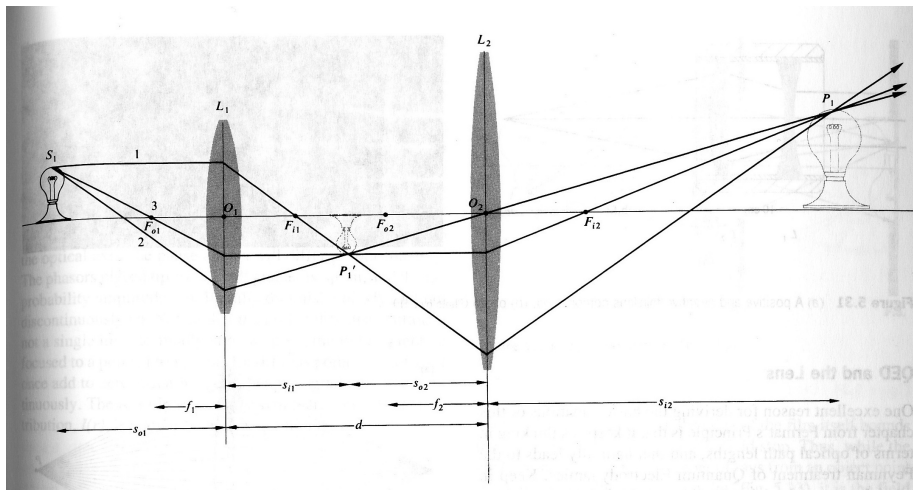
- 1 Optical Systems
- 2 Images and Pupils
- 3 Rays
- 4 Wavefronts
- 5 Aberrations

## Overview



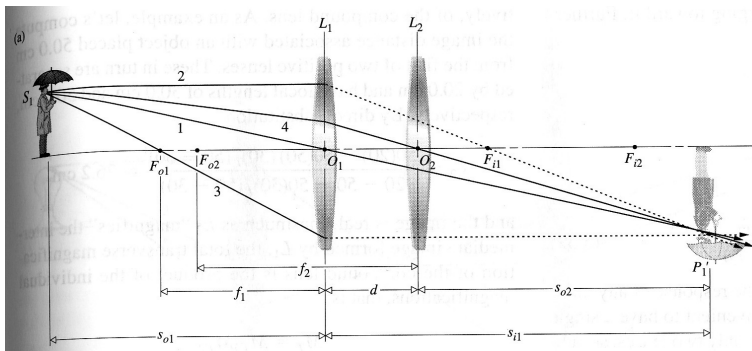
- combinations of several optical elements (lenses, mirrors, stops)
- examples: camera “lens”, microscope, telescopes, instruments
- thin-lens combinations can be treated analytically
- effective focal length:  $\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2}$

## Simple Thin-Lens Combinations



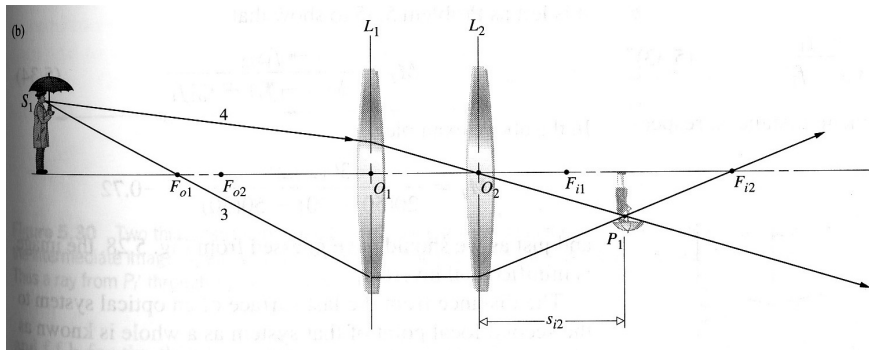
- distance  $>$  sum of focal lengths  $\Rightarrow$  real image between lenses
- apply single-lens equation successively

## Thin-Lens Combinations 1



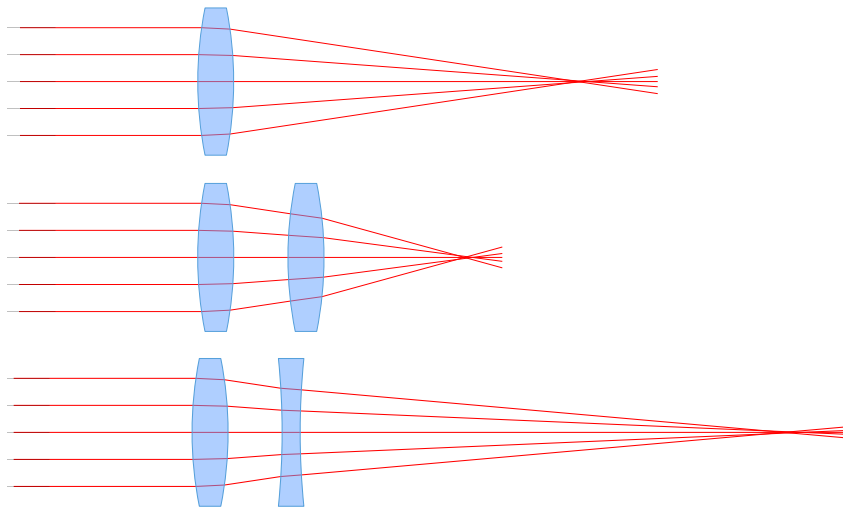
- construct image formed by lens 1 using rays 2 and 3
- ray 2 passes through focal point  $F_{i1}$
- ray 3 passes through focal point  $F_{o1}$
- ray 4 passes backwards through center of lens 2

## Thin-Lens Combinations 2

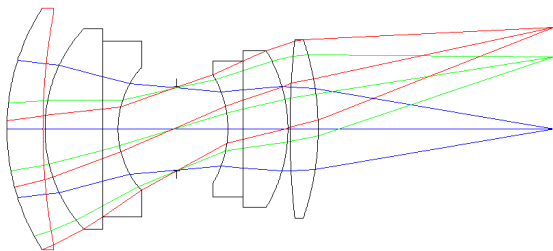


- adding lens 2 does not refract ray 4
- ray 3 is refracted to image focus  $F_{i2}$
- intersection of rays 3 and 4 determine image location
- lens 2 adds convergence or divergence

## Second Lens Adds Convergence or Divergence

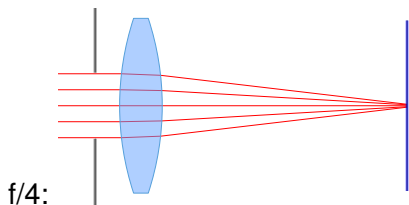
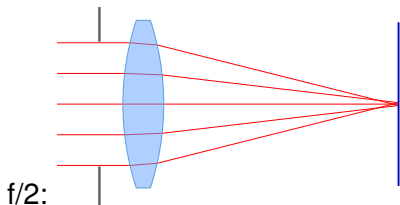


## Aperture



- all optical systems have a place where 'aperture' is limited
- main mirror of telescopes
- aperture stop in photographic lenses
- aperture typically has a maximum diameter
- aperture size is important for diffraction effects

## F-number



- describes the light-gathering ability of the lens
- f-number given by  $F = f/D$
- also called focal ratio or f-ratio, written as:  $f/F$
- the bigger  $F$ , the better the paraxial approximation works
- fast system for  $F < 2$ , slow system for  $F > 2$

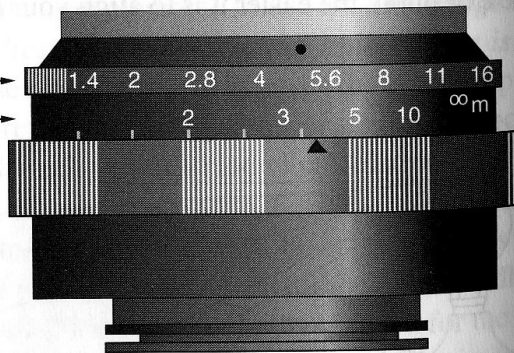


# F-number on Camera Lens

(b)

*f*-number scale

Distance scale



*f*-number 2.8

4

5.6

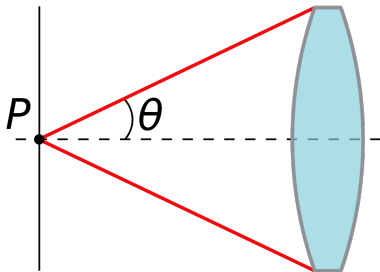
8

11

16



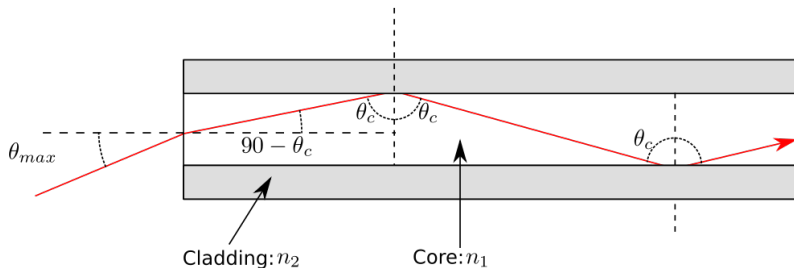
## Numerical Aperture



[en.wikipedia.org/wiki/File:Numerical\\_aperture.svg](https://en.wikipedia.org/wiki/File:Numerical_aperture.svg)

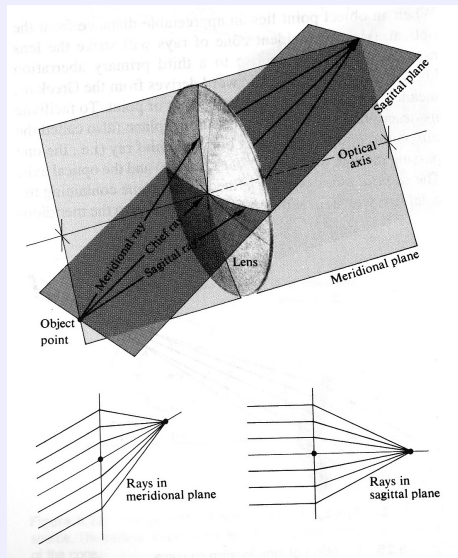
- numerical aperture (NA):  $n \sin \theta$
- $n$  index of refraction of working medium
- $\theta$  half-angle of maximum cone of light that can enter or exit lens
- important for microscope objectives ( $n$  often not 1)

## Numerical Aperture in Fibers



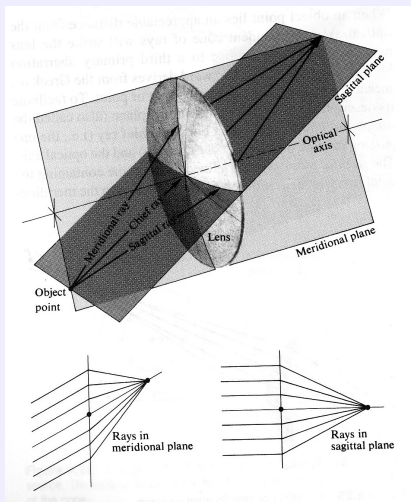
[en.wikipedia.org/wiki/File:OF-na.svg](http://en.wikipedia.org/wiki/File:OF-na.svg)

- acceptance cone of the fiber determined by materials
- $NA = n \sin \theta = \sqrt{n_1^2 - n_2^2}$
- $n$  index of refraction of working medium



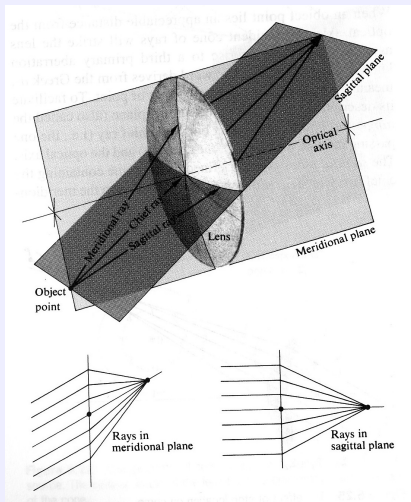
## Planes and Rays

- meridional plane defined by optical axis and chief ray going through center of optical system
- sagittal plane is perpendicular to it



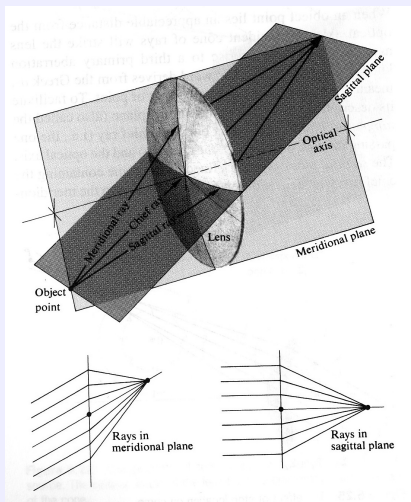
## Meridional (or Tangential) Ray

- confined to plane containing optical axis and object point from which ray originates



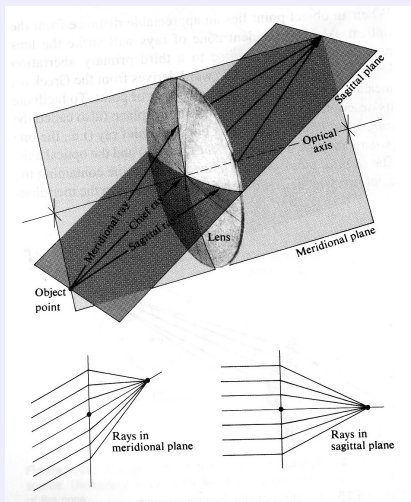
## Chief (or Principal) Ray

- goes through center of aperture
- meridional ray that starts at edge of object, and passes through center of aperture stop
- crosses optical axis at locations of pupils
- chief rays are equivalent to the rays in pinhole camera
- distance between chief ray and optical axis at an image location defines size of image



## Skew Ray

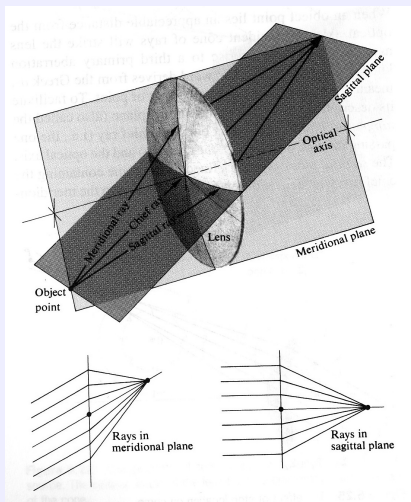
- does not propagate in plane that contains both object point and optical axis
- does not cross optical axis anywhere, and not parallel to it



## Marginal Ray

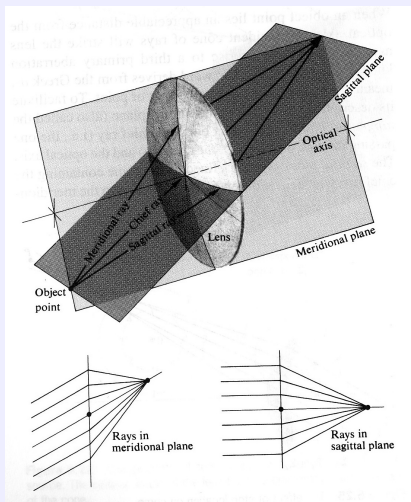
- is meridional ray that starts at point where object crosses optical axis and touches edge of aperture stop
- useful because it crosses optical axis again at locations where image is formed
- distance of marginal ray from optical axis at entrance and exit pupils defines their sizes





## Sagittal (or Transverse) Ray

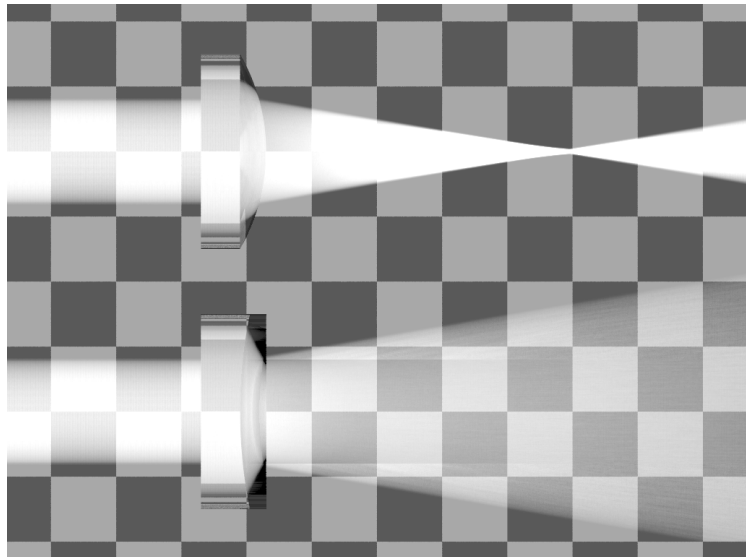
- comes from off-axis object point, propagates in plane perpendicular to meridional plane
- intersects the pupil along a line that is perpendicular to meridional plane
- chief ray is both sagittal and meridional
- all other sagittal rays are skew rays

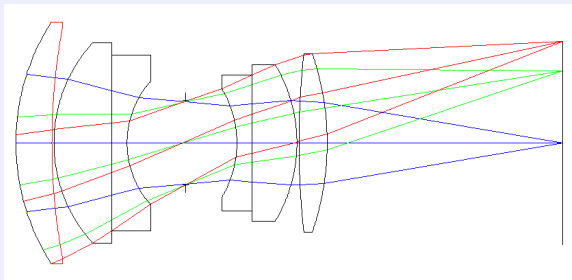


## Paraxial Ray

- makes a small angle to the optical axis of the system
- lies close to the axis throughout the system
- can be modeled reasonably well by using the paraxial approximation.

## Converging, Diverging and Collimated beams

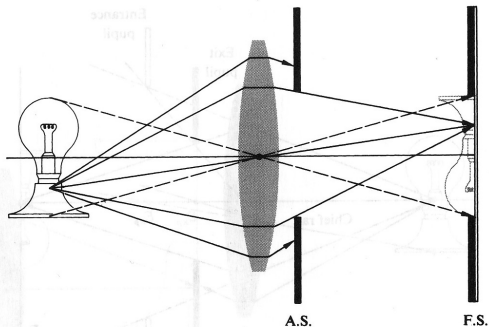




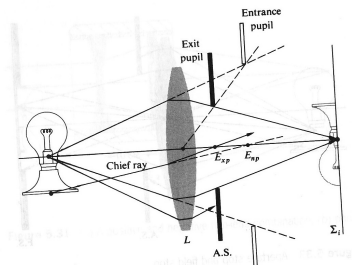
## Images and Pupils

- image
  - every object point comes to a focus in an image plane
  - light in one image point comes from pupil positions
  - object information is encoded in position, not in angle
- pupil
  - all object rays are smeared out over complete aperture
  - light in one pupil point comes from different object positions
  - object information is encoded in angle, not in position

## Aperture and Field Stops

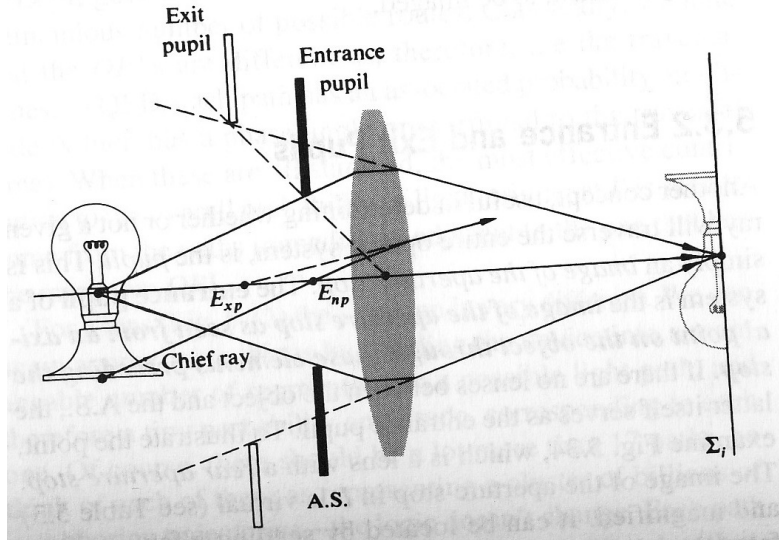


- aperture stop limits the amount of light reaching the image
- aperture stop determines light-gathering ability of optical system
- field stop limits the image size or angle

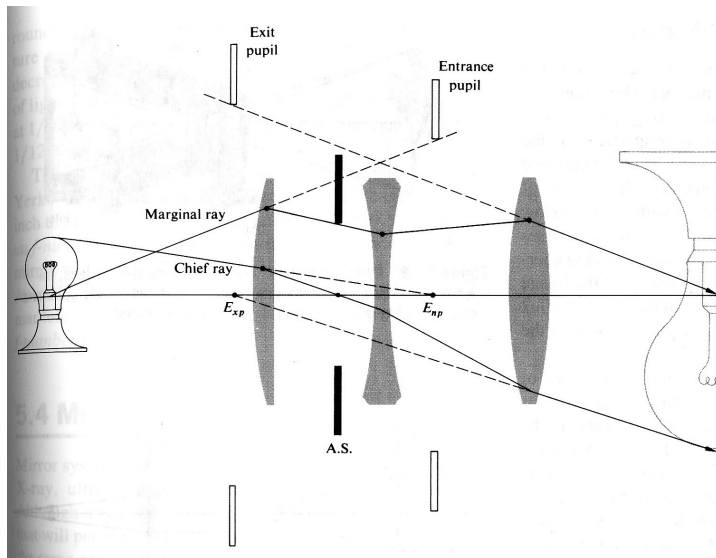


- pupil is an image of the aperture stop
- entrance pupil: image of the aperture stop as seen from a point on the optical axis and on the object through optical elements preceding the aperture stop
- exit pupil: image of the aperture stop as seen from a point on the optical axis and in the image through optical elements after the aperture stop

# Entrance and Exit Pupils

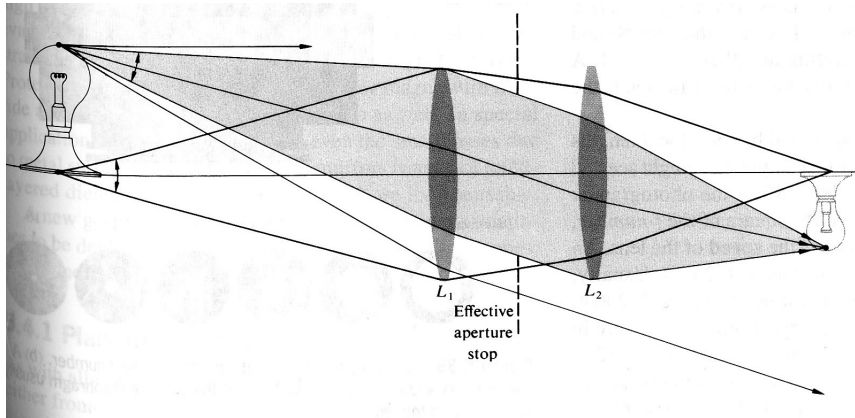


## Entrance and Exit Pupils



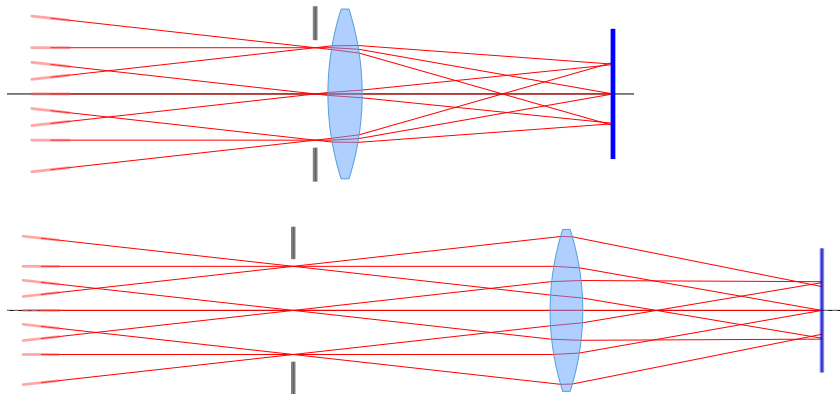


# Vignetting



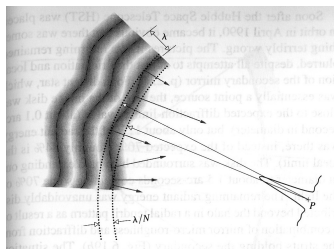
- effective aperture stop depends on position in object
- image fades toward its edges

## Telecentric Arrangement

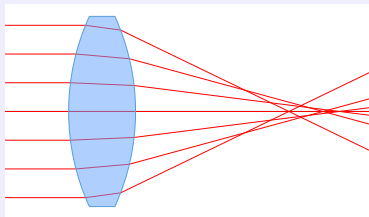


- as seen from image, pupil is at infinity
- easy: lens is its focal length away from pupil (image)
- magnification does not change with focus positions
- ray cones for all image points have the same orientation

## Spot Diagrams and Wavefronts



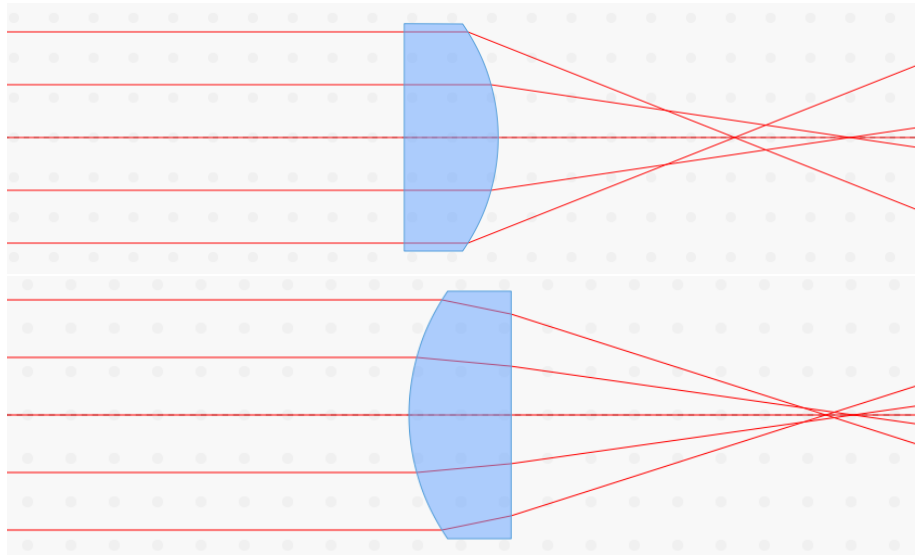
- plane of least confusion is location where image of point source has smallest diameter
- spot diagram: shows ray locations in plane of least confusion
- spot diagrams are closely connected with wavefronts
- aberrations are deviations from spherical wavefront



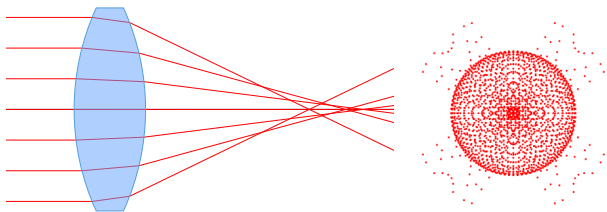
## Spherical Aberrations

- different focal lengths of paraxial and marginal rays
- longitudinal spherical aberration along optical axis
- transverse (or lateral) spherical aberration in image plane
- much more pronounced for short focal ratios

## Minimizing Spherical Aberrations

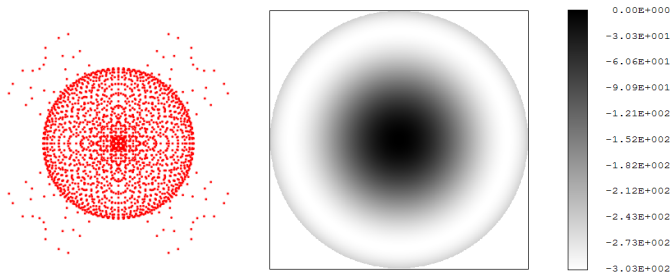


## Spherical Aberration of Spherical Lens



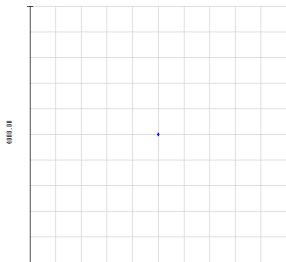
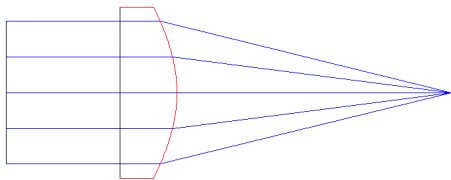
- foci from paraxial beams are further away than marginal rays
- spot diagram shows central area with fainter disk around it

## Spherical Aberration Spots and Waves



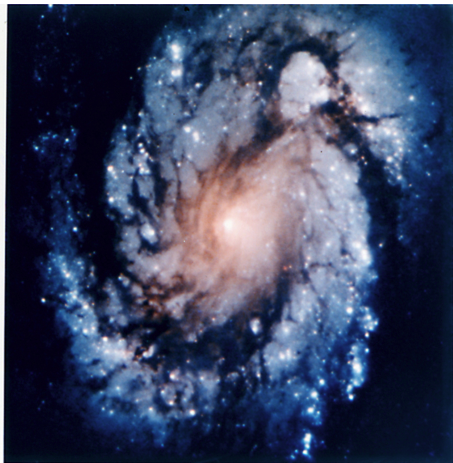
- spot diagram shows central area with fainter disk around it
- wavefront has peak and turned-up edges

## Aspheric Lens

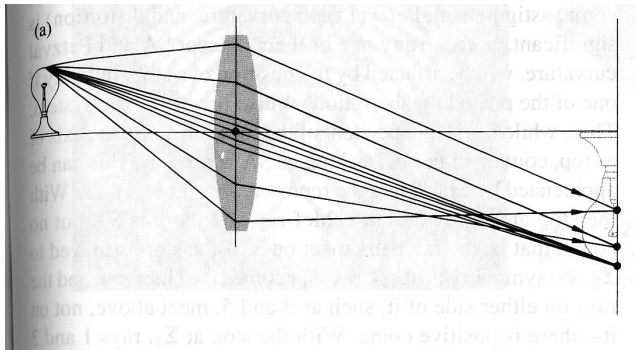


- conic constant  $K = -1 - \sqrt{n}$  makes perfect lens
- difficult to manufacture
- but possible these days



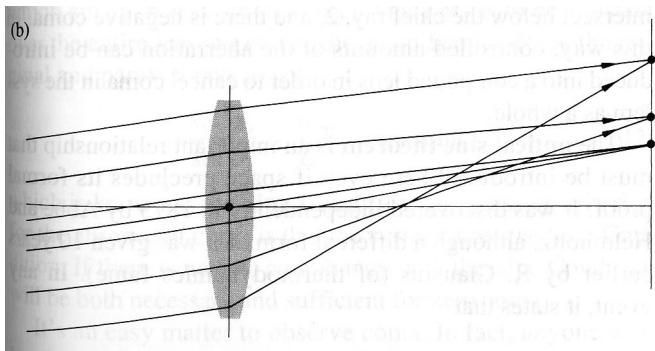


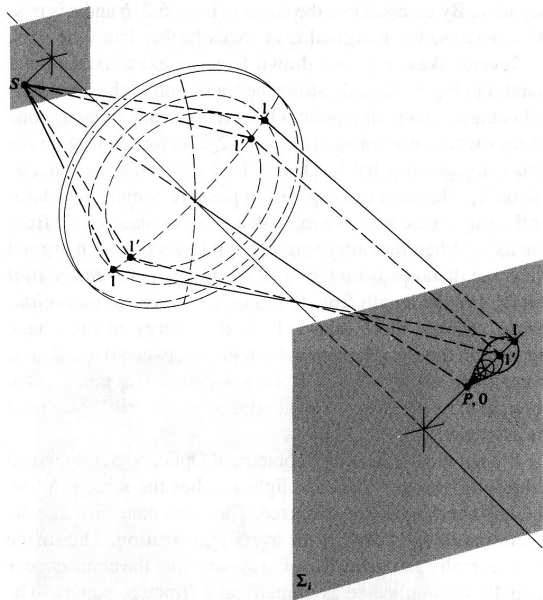
## Coma



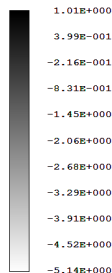
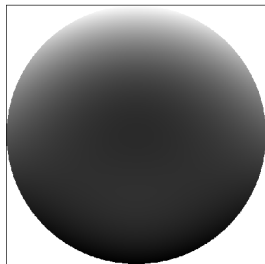
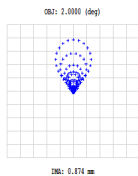
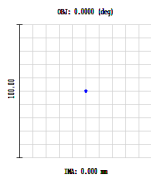
- typically seen for object points away from optical axis
- leads to 'tails' on stars

## Positive Coma



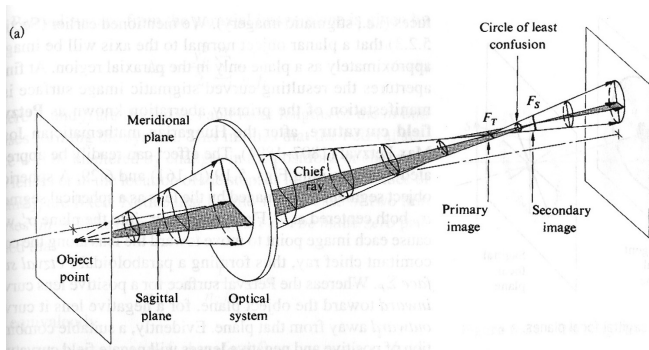


## Coma Spots and Waves



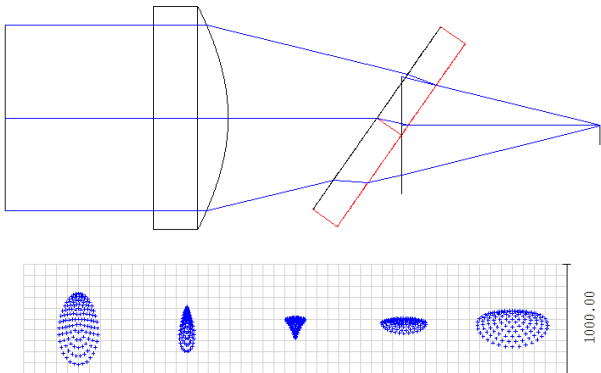
- parabolic mirror with perfect on-axis performance
- spots and wavefront for off-axis image points
- wavefront is tilted in inner part

# Astigmatism



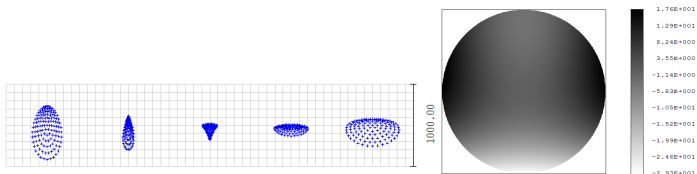
- image of a point forms focal lines at the sagittal and tangential foci
- in between an elliptical shape

## Tilted Glass Plate in Converging Beam



- astigmatism and spherical aberration
- note beam shift
- tilted plates: beam shifters, filters, beamsplitters

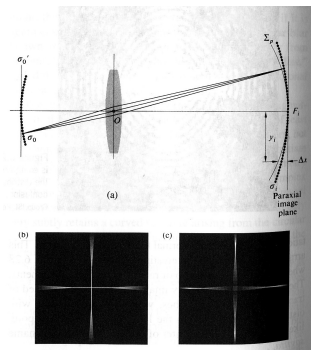
## Astigmatism Spots and Waves



- focus in two orthogonal directions, but not in both at the same time
- difference of two parabolae with different curvatures
- wavefront has saddle shape

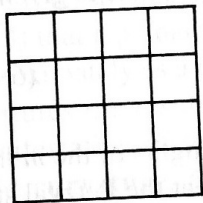


## Field Curvature

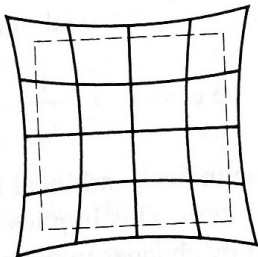


- field (Petzval) curvature: image lies on curved surface
- problems with flat detectors (e.g. CCDs)
- solution: field flattening lens close to focus

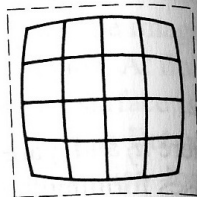
## Distortion



(a)



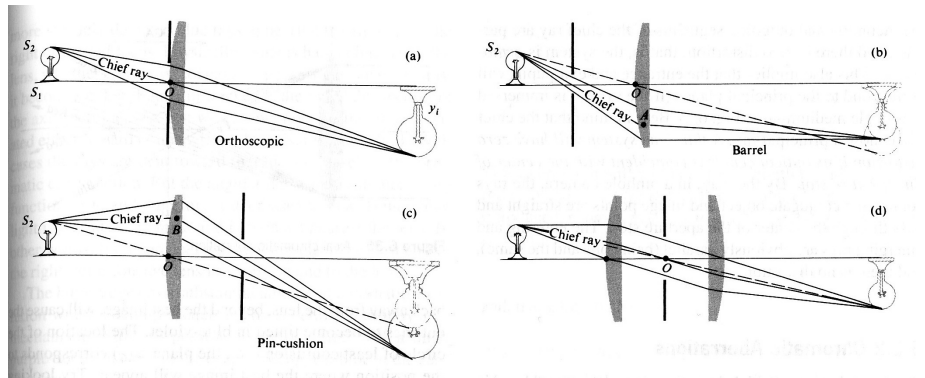
(b)



(c)

- image is sharp but geometrically distorted
- (a) object
- (b) positive (or pincushion) distortion
- (c) negative (or barrel) distortion

## Aperture Stop Creates Distortion

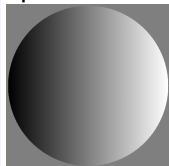


## Seidel Aberrations

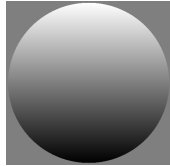
- Ludwig von Seidel (1857)
- Taylor expansion of  $\sin \phi$
- $\sin \phi = \phi - \frac{\phi^3}{3!} + \frac{\phi^5}{5!} - \dots$
- paraxial: first-order optics
- Seidel optics: third-order optics
- Seidel aberrations: spherical, astigmatism, coma, field curvature, distortion

# Zernike Polynomials

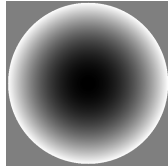
tip



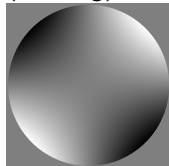
tilt



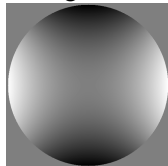
focus



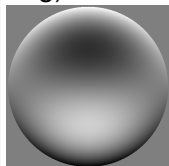
astigmatism  
(45 deg)



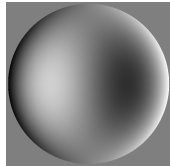
astigmatism  
0 deg



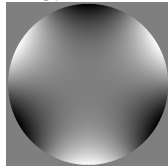
coma (0  
deg)



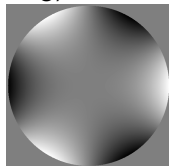
coma (90  
deg)



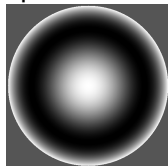
trefoil (0  
deg)



trefoil (30  
deg)



third-order  
spherical



- low orders equal Seidel aberrations
- form orthonormal basis on unit circle