Astronomical Observing Techniques

Lecture 9: Silicon Eyes 1

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Overview

- 1. Solid State Physics
- 2. Intrinsic Photoconductors
- 3. Extrinsic Photoconductors
- 4. Readout & Operations
- 5. Detector Noise

Modern Detectors

1. Photon detectors

Respond directly to individual photons -> releases bound charge carriers. Used from X-ray to infrared.

Examples: photoconductors, photodiodes, photoemissive detectors

2. Thermal detectors

Absorb photons and thermalize their energy-> changes resistance -> modulates electrical current. Used mainly in IR and sub-mm detectors. *Examples: bolometers*

3. Coherent receivers

Respond directly to electrical field and preserve phase information (but need a reference phase "local oscillator"). Mainly used in the sub-mm and radio regime.

Examples: heterodyne receivers

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PERIODIC TABLE OF THE ELEMENTS

GROUP

	1 IA						/ /	///		+++	$\rightarrow \rightarrow$		http	://www.ktf-	split.hr/peri	iodni/en/		18 VIIIA
9	1 1.0079	RELATIVE ATOMIC MASS (1)			Me	Nonmetal										2 4.0026		
-ă 1	H	GROUP IUPAC GROUP CAS				Alkali matal										He		
E	HYDROGEN	2 11A 13 111A				Alkaline earth metal				. —	13 IIIA	14 IVA	15 VA	16 VIA	17 VIIA	HELIUM		
	3 6.941	4 9.0122	ATOMIC N	IUMBER - 5	10.811	_/_	Tra	insition metals	5	18 Noble	gas		5 10.811	6 12.011	7 14.007	8 15.999	9 18.998	10 20.180
1	Li	Be	s	YMBOL	B	/ /		Lanthanide	STAN	DARD STATE	(25 °C: 101)	kPa)	B	C	N	0	F	Ne
	LITHIUM	BERYLLIUM	/		BORON	_/_		Actinide	Ne	- gas	Fe - solid		BORON	CARBON	NITROGEN	OXYGEN	FLUORINE	NEON
_	11 22.990	12 24.305			A	/	/		Ga	- liquid	Tc - synthe	tic	13 26.982	14 28.086	15 30.974	16 32.065	17 35.453	18 39.948
3	Na	Μσ		ELE	MENT NAME		/			/			AL	Si	P	S	CI	Ar
	SODIUM	MAGNESIUM	3 1118	4 IVB	5 VB	6 VIB		8	VIIIB -	10	11 18	12 118	ALUMINIUM	SILICON	PHOSPHORUS	SULPHUR	CHLORINE	ARGON
/	19 39.098	20 40.078	21 44.956	22 47.867	23 50.942	24 51.996	25 54.938	26 55.845	27 58.933	28 58.693	29 63.546	30 65.39	31 69.723	32 72.64	33 74.922	34 78.96	35 79.904	36 83.80
4	K	Ca	Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cn	Zn	Ga	Ge	As	Se	Br	Kr
	DOTASSIUM	CALCIUM	SCANDILIM	TITANUM	VANADILIM	CHROMUM	MANCANESE	IRON	COBALT	NICKEL	COPPER	ZINC	GALLIUM	GERMANNUM	ARSENIC	SELENIUM	BROMINE	KEVETON
/	37 85.468	38 87.62	39 88.906	40 91.224	41 92.906	42 95.94	43 (98)	44 101.07	45 102.91	46 106.42	47 107.87	48 112.41	49 114.82	50 118.71	51 121.76	52 127.60	53 126.90	54 131.29
<u> </u>	Ph	Sr	v	7r	Nh	Mo	The	Ru	Rh	Pd	Aa	Cd	In	Sn	Sh	То	T	Vo
	PUPIDIUM	STRONTUNA	VITTORINA	ZIRCONIUM	NICOBILINA	IVIU	TECHNETIUM		DHODIUM	1 U	Ag	Cu	A A A	JIN	ANTHONY	TELLIPIUM		YENON
	55 132.91	56 137.33	57.71	72 178.49	73 180.95	74 183.84	75 186.21	76 190.23	77 192.22	78 195.08	79 196.97	80 200.59	81 204.38	82 207.2	83 208.98	84 (209)	85 (210)	86 (222)
	Co	Pa	La-Lu	Пf	То	W	Do	Os	Ir	Dt	A 11	Ha	TI	Dh	Di	Do	At	Dn
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	E.	Do	89-103	10.0	IDIA	S.	TDID		TATA	II Innoo	ПТорор	II Innlls		II Inn an				
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	/	LANTHANIDE						/ \				Copyright © 1998-2003 EniG. (eni@ktf-sp				(eni@ktf-split.hr)		
(1) Pu	ure Appl. Chem., 73, No. 4, 667-683 (2001) elative atomic mass is shown with five gnificant figures. For elements have no stable uclides, the value enclosed in brackets dicates the mass number of the longest-lived notice of the element.			57 138.91	58 140.12	59 140.91	60 144.24	61 (145)	62 150.36	63 151.96	64 157.25	65 158.93	66 162.50	67 164.93	68 167.26	69 168.93	70 173.04	71 174.97
si				La	Ce	Pr	Nd	IPm	Sm	Eu	Gd	Tb	Dv	Ho	Er	Tm	Yb	Lu
in is				LANTHANUM	CERIUM	PRASEODYMIUM	NEODYMIUM	PROMETHIUM	SAMARIUM	EUROPIUM	GADOLINIUM	TERBIUM	DYSPROSIUM	HOLMIUM	ERBIUM	THULIUM	YTTERBIUM	LUTETIUM
H	towever three such elements (Th, Pa, and U) o have a characteristic terrestrial isotopic composition, and for these an atomic weight is abulated.													·				
co ta				89 (227)	90 232.04	91 231.04	92 238.03	93 (237)	94 (244)	95 (243)	96 (247)	97 (247)	98 (251)	99 (252)	100 (257)	101 (258)	102 (259)	103 (262)
				Ac	Th	Pa	U	Np	Pu	Am	Cm	Bk	Cf	Es	Fm	MId	No	Lr
E	litor: Aditya Vardh	Vardhan (adivar@nettlinx.com)		ACTINIUM	THORIUM	PROTACTINIUM	URANIUM	NEPTUNIUM	PLUTONIUM	AMERICIUM	CURIUM	BERKELIUM	CALIFORNIUM	EINSTEINIUM	FERMIUM	MENDELEVIUM	NOBELIUM	LAWRENCIUM
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Diamond Lattice

- Elements with 4 e⁻ in valence shell form crystals with diamond lattice structure (each atom bonds to four neighbors).
- These double-bonds between neighbours are due to "shared" electrons



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Diamond Lattice

Diamond lattice not only formed by group IV elements (C, Si, Ge) but also by III-V semiconductors (InSb, GaAs, AIP)



Electronic States and Bands

Atomic crystal

Wavefunctions Ψ overlap

Single atomic system

Example: H atom



Fermi Energy

- Fermi energy E_F determines concentration of thermally excited electrons in conduction band
- Energy valence band: E_v ; Energy conduction band: E_c
- Fermi function f(E): probability that state of energy E is occupied at temperature T.



Fermi energy = energy of the highest occupied quantum state in a system of fermions at T=0K

QM: fermions obey the Pauli exclusion principle \rightarrow two fermions cannot occupy the same quantum state. Fermions consecutively fill up the unoccupied quantum states starting with the lowest energy; when all the particles have been put in, the Fermi energy is the energy of the highest occupied state.

Fermi level = chemical potential

The Fermi level is the energy at which there is a 50% chance of finding an occupied energy state. The Fermi level can be calculated from the density of states in the conduction and valence bands. The Fermi level may increase, remain the same or decrease with increasing temperature, depending on the number of states in the conduction and valence bands.

Fermi energy and Fermi level are only the same at absolute zero. At absolute zero temperature the Fermi level can be thought of as the energy up to which available electron states are occupied. At higher temperatures, the Fermi level is the energy at which the probability of a state being occupied has fallen to 0.5.

The Fermi function f(E) gives the probability that a given available electron energy state will be occupied at a given temperature. Typically, most of the levels up to the Fermi level E_{F} are filled, and relatively few electrons have energies above the Fermi level.

The population of states depends upon the *product* of the Fermi function and the electron density of states:

- In the gap there are no electrons because the density of states is zero.
- In the conduction band at OK, there are no electrons even though there are plenty of available states, but the Fermi function is zero.
- At high temperatures, both the density of states and the Fermi function have finite values in the conduction band, so there is a finite conducting population.

Electric Conductivity



Overcome bandgap E_g to lift e^- into conduction band:

- 1. external excitation, e.g. via a photon ← photon detector
- 2. thermal excitation
- 3. impurities

Intrinisic Photo-Conductors: Basic Principle

- semi-conductor: few charge carriers \rightarrow high resistance
- charge carriers = electron-hole pairs
- photon lifts e⁻ into conduction band
- applied electric field drives charges to electrodes





- $\eta = quantum efficiency$
- τ = mean lifetime before recombination
- μ_n = electron mobility ~ mean time between collisions.

drift velocity $v=\mu_n E$, current density $j=n_0 qv$

where:

Important Quantities and Definitions

Quantum efficiency $\eta = \frac{\# \text{ absorbed photons}}{\# \text{ incoming photons}}$

Responsivity $S = \frac{\text{electrical output signal}}{\text{input photon power}}$

Wavelength cutoff:
$$\lambda_c = \frac{hc}{E_g} = \frac{1.24\,\mu m}{E_g[eV]}$$

Photo-current: $I_{ph} = q \varphi \eta G$

Photoconductive gain G:
$$G = \frac{I_{ph}}{q \varphi \eta} = \frac{\tau}{\tau_t} = \frac{\text{carrier lifetime}}{\text{transit time}}$$

The product ηG describes the probability that an incoming photon will produce an electric charge that will reach an electrode.

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Limitations of Intrinsic Semiconductors

long-wavelength cutoffs

	→ Germanium	1.85µm
$\lambda_c = \frac{hc}{-}$	\rightarrow Silicon:	1.12µm
E_g	→ GaAs: 0	.87µm

- non-uniformity of material
- problems to make good electrical contacts to pure Si
- difficult to avoid impurities and minimize thermal (Johnson) noise

Extrinsic Semiconductors

- extrinsic semiconductors: charge carriers = electrons (n-type) or holes (p-type)
- achieved by addition of impurities at low concentration to provide excess electrons or holes
- → much reduced bandgap -> longer wavelength cutoff



Example: addition of boron to silicon in the ratio 1:100,000 increases its conductivity by a factor of 1000!

Extrinsic Semiconductor Band Gaps

		Ge	Si
Impurity	Туре	Cutoff wavelength λ _c (μm)	Cutoff wavelength λ_c (µm)
Al	р		18.5 ^a
В	p	119 ^b	28^a
Be	P	52 ^b	8.3 ^a
Ga	р	115 ^b	17.2^{a}
In	р	111 ^b	7.9^{a}
As	n	98 ^b	23^a
Cu	р	31 ^b	5.2^{a}
Р	n	103 ^b	27 ^a
Sb	n	1 29 ^b	29 ^a

Problem: absorption coefficients much less than for intrinsic photoconductors \rightarrow low QE \rightarrow active volumes (pixels) must be large

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Blocked Impurity Band (BIB) Detectors

Solution: use separate layers to optimize the optical and electrical properties independently:

- IR-active layer: heavily doped
- Blocking layer: thin layer of high purity (intrinsic photoconductor)
- Typical species are Si:As or Si:Sb BIBs



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Photodiodes

- junction between two oppositely doped zones
- Two adjacent zones create a depletion region with high impedance



- 1. Photon gets absorbed e.g. in the p-type part
- 2. Absorption creates an e⁻-hole pair
- 3. The e^- diffuses through the material
- 4. Voltage drives the e[−] across the depletion region → photo-current April 2016 Astronomical Observing Techniques: Detectors 1 18

Charge Coupled Devices (CCDs)

CCDs = array of integrating capacitors.

Pixel structure: metal "gate" evaporated onto SiO_2 (isolator) on silicon = MOS



- 1. photons create free e⁻ in the photoconductor
- 2. e^- drift toward the electrode but cannot penetrate the SiO₂ layer
- 3. e^- accumulate at the Si-SiO₂ interface
- 4. total charge collected at interface measures number of photons during the exposure
- 5. \rightarrow read out the number of e⁻ A April 2016 Astronomical Observing Techniques: Detectors 1

Charge Coupled Readouts

Collected charges are passed along the columns to the edge of the array to the output amplifier.



Be aware of charge transfer (in-)efficiencies (CTEs) due to electrostatic repulsion, thermal diffusion and fringing ufields tors 1 20



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Charge Transfer Efficiency (CTE)

Time-dependent mechanisms that influence the CTE:

- 1. Electrostatic repulsion causes electrons to drift to the neighbouring electrode with time constant for charge transfer τ_{sl} .
- 2. Thermal diffusion drives electrons across the storage well at τ_{th} .
- 3. "Fringing fields" due to dependency of the well on the voltages of neighbouring electrodes (τ_{ff}).

Approximation for the CTE of a CCD with m phases: C_{1}^{2}

$$CTE = \left(1 - e^{-t/\tau}\right)^m$$

Noise from charge transfer inefficiency: $\varepsilon = (1-CTE)$

CCD Color Sensors

- 1. Take three exposures through three filters subsequently only works for fixed targets (standard for astronomy).
- 2. Split the input beam in three channels, each with a separate and optimized CCD (professional video cameras).
- 3. Bayer mask over CCD each subset of 4 pixels has one filtered red, one blue, and two green.



Main Detector Noise Components

G-R noise

$$\left\langle I_{G-R}^{2}\right\rangle = 4q^{2}\varphi\eta G^{2}\Delta f$$

fundamental statistical noise due to the Poisson statistics of the photon arrival \rightarrow transferred into the statistics of the generated and recombined holes and electrons.

Johnson or kTC noise

$$\left\langle I_J^2 \right\rangle = \frac{4kT}{R} \Delta f$$

fundamental thermodynamic noise due to the thermal motion of the charge carriers. Consider a photo-conductor as an RC circuit. Since $\langle Q^2 \rangle = kTC$, the charge noise is also called kTC noise or reset noise.

1/f noise
$$\langle I_{1/f}^2 \rangle \propto \frac{I^2}{f} \Delta f$$

increased noise at low frequencies, due to bad electrical contacts, temperature fluctuations, surface effects (damage), crystal defects, and JFETs, ...

The total noise in the system is:
$$\langle I_N^2 \rangle = \langle I_{G-R}^2 \rangle + \langle I_J^2 \rangle + \langle I_{1/f}^2 \rangle_{24}$$

BLIP and NEP

Operationally, background-limited performance (BLIP) is always preferred: $\langle I_{G-R}^2 \rangle \gg \langle I_J^2 \rangle + \langle I_{1/f}^2 \rangle$

The noise equivalent power (NEP) is the signal power that yields an RMS S/N of unity in a system of $\Delta f = 1$ Hz: $NEP_{G-R} = \frac{2hc}{\lambda} \left(\frac{\varphi}{\eta}\right)^{1/2}$

In BLIP the NEP can only be improved by increasing the quantum efficiency η .