

# Cores/Cusp in Elliptical galaxies and SMBHs

*“Supermassive Black Holes and their host spheroids. I. Disassembling Galaxies”*

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*“The  $L \propto \sigma^8$  correlation for elliptical galaxies with cores: relation with black hole mass”*

Jhon Kormendy and Ralf Bender

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# The $L \propto \sigma^8$ correlation for elliptical galaxies with cores: relation with black hole mass

Jhon Kormendy and Ralf Bender (2013)

**Main idea:** Elliptical galaxies split into two types based on their central light profiles.

## (Giant) Core galaxies

The core is “missing light” with respect to a Sérsic profile.

## Coreless galaxies

They have extra light in the center (central power law).

**First question asked:** How does the Faber-Jackson relation change between the two galaxy types?

# The Faber-Jackson relation

Standard F-J relation:

$$L \propto \sigma^4$$

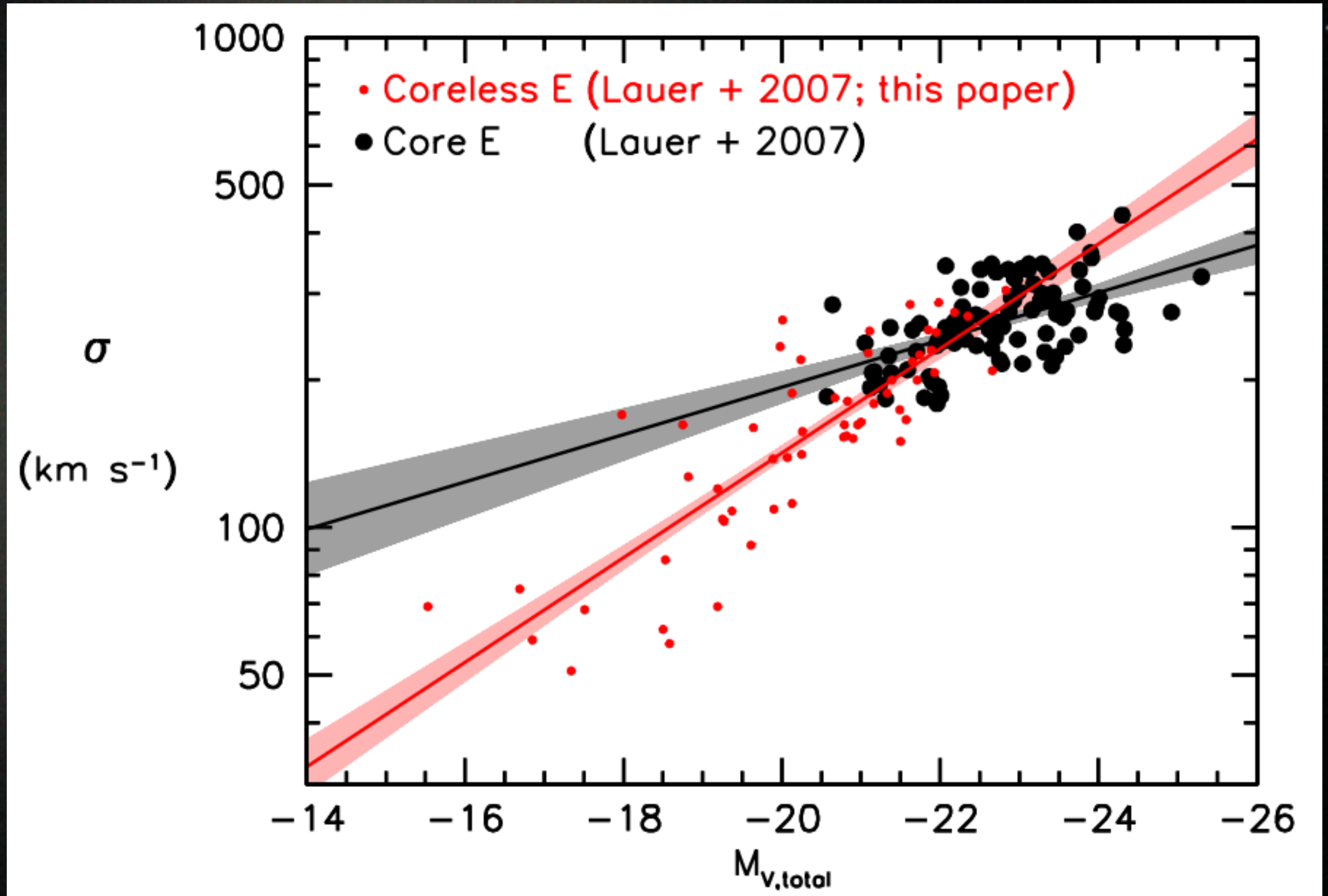
Observational results:

- Core galaxies

$$L \propto \sigma^{8.33}$$

- Coreless galaxies

$$L \propto \sigma^{3.74}$$



# The Faber-Jackson relation

Translated in terms of mass:

- Core galaxies  $\frac{d\log\sigma}{d\log M} = 0.091 \longrightarrow$  Large increases of mass produce only small changes in the velocity dispersion.
- Coreless galaxies  $\frac{d\log\sigma}{d\log M} = 0.203$

Comparison with the Virial theorem:  $M \propto \sigma^2 r_e$

This tells us that the structure of galaxies is not constant during their evolution.

# Formation Pathways: wet and dry mergers

For core galaxies:  $\frac{d\log\sigma}{d\log M} = 0.091$

Comparison with merger simulations:

• Minor wet mergers  $\frac{d\log\sigma}{d\log M} = -0.05$

• Major dry mergers  $\frac{d\log\sigma}{d\log M} = 0.15$

Massive elliptical galaxies grow mainly through dissipationless (dry) mergers.

# Core formation in elliptical galaxies

## The process:

- Two galaxies merge, their SMBHs form a binary.
- The binary "scours" the center by ejecting stars via 3-body interactions.
- This creates the "missing light" deficit.

## Evidence:

The mass deficit scales with the SMBH mass:  $M_{\text{def}} \propto M_{\bullet}$

# Supermassive Black Holes and their host spheroids. I. Disassembling Galaxies

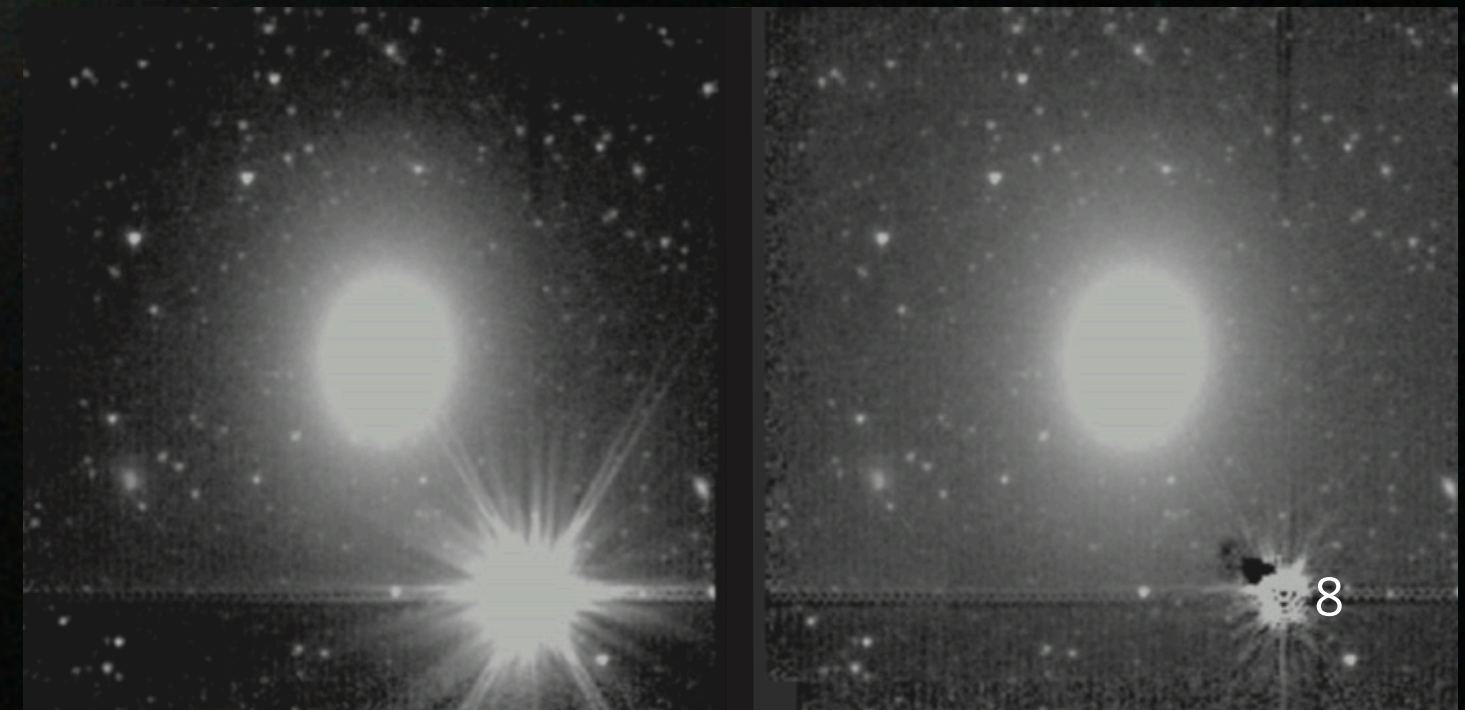
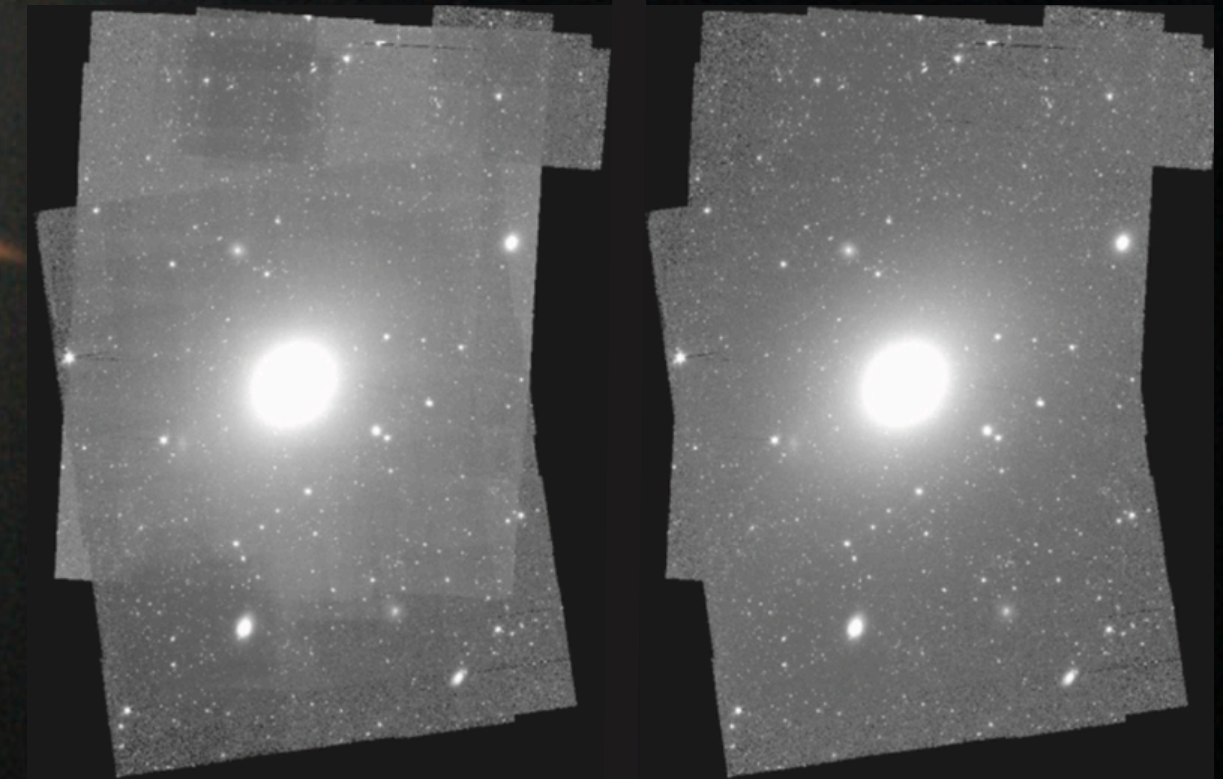
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## Disassembling Galaxies

### Data reduction

- Data acquisition from Spitzer Space telescope
  - Mosaicking and Overlapping
  - Sigma Mosaics
  - Aesthetic corrections and imaging masking
  - 2D PSF
- 
- Kinematics movements
  - AGNs and Dust
  - Core and Coreless

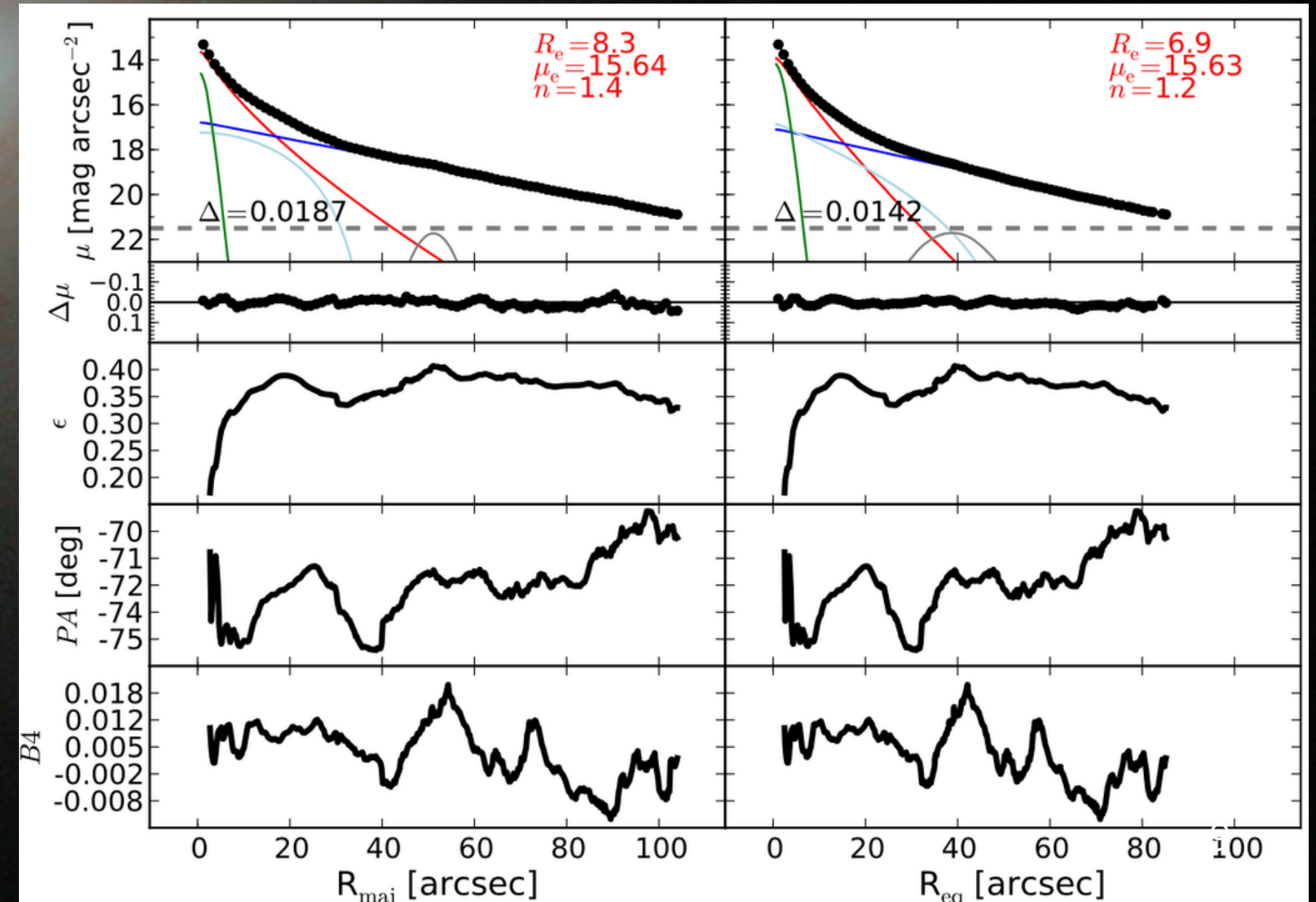


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## Disassembling Galaxies

### Analysis and Results

- 1D Isophotal Analysis
- Upsharp Masking
- Mathematical Modeling



Galaxy NGC 2974

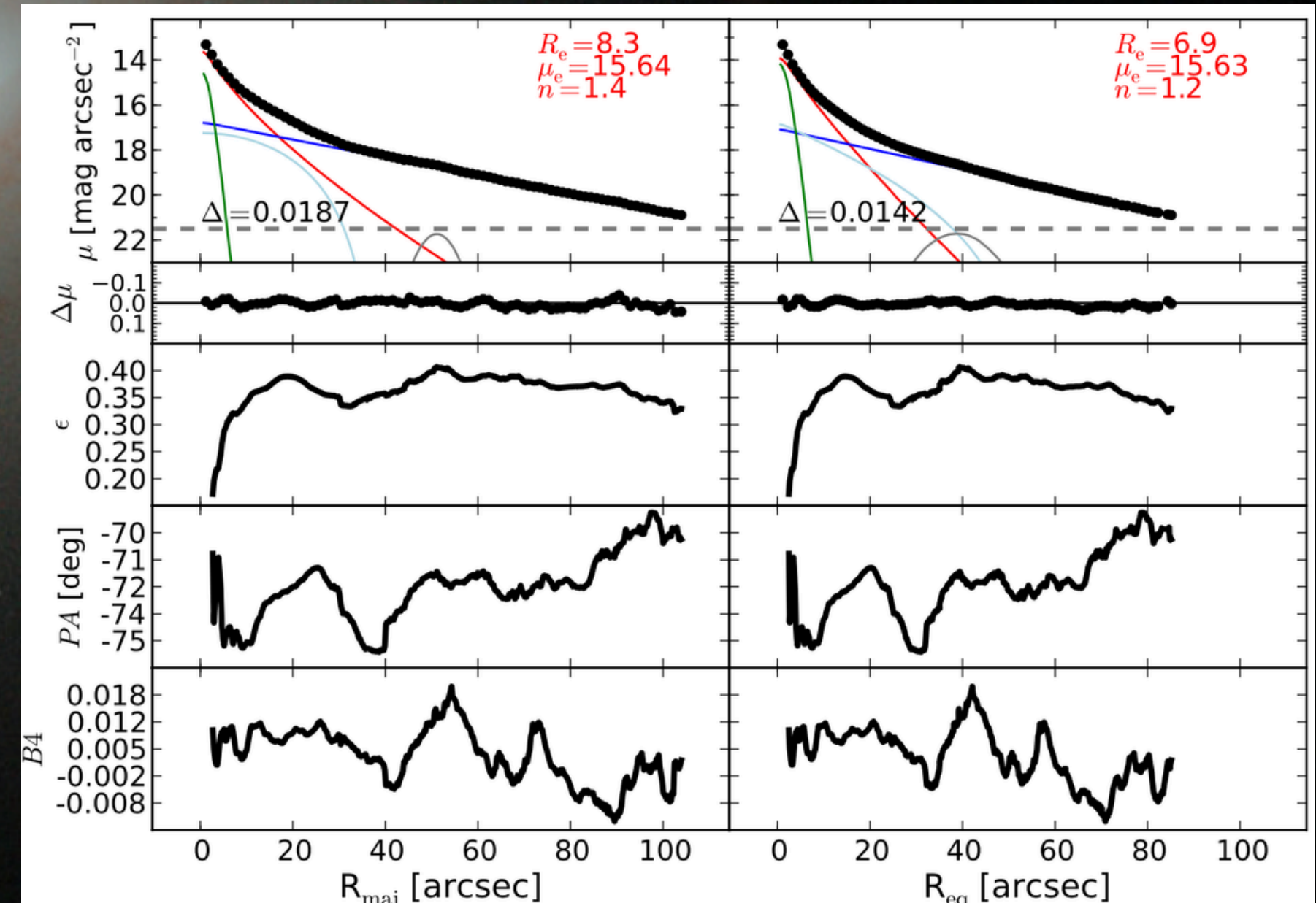
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## Disassembling Galaxies

### Analysis and Results

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| Work    | Model                           | $R_{e,\text{sph}}$<br>(arcsec) | $n_{\text{sph}}$ |
|---------|---------------------------------|--------------------------------|------------------|
| 1D maj. | S-bul + e-d + F-bar + G-n + G-r | 8.3                            | 1.4              |
| 1D eq.  | S-bul + e-d + F-bar + G-n + G-r | 6.9                            | 1.2              |
| 2D      | S-bul + e-d + G-bar + m-n       | 10.6                           | 1.3              |
| S+11 2D | S-bul + G-n                     | 27.2                           | 3.0              |



Galaxy NGC 2974