



Imaging extended objects with ALMA – “Data Combination” (Tutorial “3b”)

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Hands-on session

- a) Joint Deconvolution of interferometric datasets
- b) Preparation of SD images for combination with interferometric data
- c) Combination of interferometric with single-dish data – sdintimaging
- d) Using the `assess_ms` tool for uv coverage assessment.

TUTORIAL 3B - DATA COMBINATION

All sessions in room BW0.32, Total time: 4.5 h

11:00 - 12:30 h:

- Intro lecture: Imaging extended objects with ALMA
- Q&A

Slides: [Leiden-DataCombinationIntro-dpetry-Jan2026.pdf](#)

13:30 h - 15:00 h::

- Setup for hands-on tutorial
- Joint deconvolution of TM and 7M MOUSs
- Start of Combination of interferometric and SD data

Slides: [Leiden-DataCombinationHandsOn-dpetry-Jan2026.pdf](#)

[Leiden-DataCombination-sdintimaging-Jan2026.pdf](#)
(lecture during long exec times of sdintimaging/tclean)

Script: [leiden-tutorial3b.py](#)
(up to launching second sdintimaging)

15:30 h - 17:00 h:

- Combination of interferometric and SD data (continued)
- uv coverage assessment

Slides: [Leiden-DataCombination-assess_ms3-Jan2026.pdf](#)

Scripts: [leiden-tutorial3b.py](#)
(final steps: image analysis with imview, sdgain tests)

[leiden-tutorial3b-supplement.sh](#)
([assess_ms](#))

***ALMA delivers Groups of MOUSs which still need combination.
Possible combinations are (ordered by angular resolution):***

TM1	+	TM2			
TM1	+	TM2	+	7M	
TM1	+	TM2	+	7M	+
		TM1	+	7M	
		TM1	+	7M	+
				7M	+
					TP

If ALMA TP not available, use SD data from other observatories!

Two basic combination methods needed:

1) Combination of interferometric data from different arrays

Visibilities + Visibilities

2) Combination of interferometric data with single-dish data

Visibilities + Image

or Visibilities + Visibilities + Image (if there is more than one INT dataset)

Possible MOUS combinations in ALMA GOUSs:

TM1 + **TM2**
TM1 + **TM2** + **7M**
TM1 + **TM2** + **7M** + **TP**
 TM1 + **7M**
 TM1 + **7M** + **TP**
 7M + **TP**

We use as example for both hands-on sessions a GOUS of the most general case:

TM1 + TM2 + 7M + TP

2017.1.01355.L (ALMA Cycle 5), GOUS uid___A001_X1296_X1a2 consisting of

MOUS	SB Name
uid___A001_X1296_X1a3	G333.60_a_03_ TM1
uid___A001_X1296_X1a5	G333.60_a_03_ TM2
uid___A001_X1296_X1a7	G333.60_a_03_ 7M
uid___A001_X1296_X1a9	G333.60_a_03_ TP

Joint deconvolution is needed in all ALMA GOUS types except 7M+TP:

<i>TM1</i>	+	<i>TM2</i>			
<i>TM1</i>	+	<i>TM2</i>	+	<i>7M</i>	
<i>TM1</i>	+	<i>TM2</i>	+	<i>7M</i>	+ <i>TP</i>
		<i>TM1</i>	+	<i>7M</i>	
		<i>TM1</i>	+	<i>7M</i>	+ <i>TP</i>

Data Combination Hands-On Session: Joint Deconvolution

Joint deconvolution is needed in all ALMA GOUS types except 7M+TP:

TM1	+	TM2			
TM1	+	TM2	+	7M	
TM1	+	TM2	+	7M	+ TP
		TM1	+	7M	
		TM1	+	7M	+ TP
				7M	+ TP

If no TP data is involved, we speak of **pure joint deconvolution** and can use **task “tclean”** with **gridder “mosaic”**.

In our example we just omit the TP data and jointly deconvolve TM1+TM2+7M

MOUS	SB Name
uid__A001_X1296_X1a3	G333.60_a_03_TM1
uid__A001_X1296_X1a5	G333.60_a_03_TM2
uid__A001_X1296_X1a7	G333.60_a_03_7M

Data Combination Hands-On Session

If you wanted to start our example from scratch,
you would need to obtain the following calibrated MSs:

uid___A002_Xc8ed16_X5c3d.ms.split.cal.tgz (TM1)

uid___A002_Xcd07af_X41d5.ms.split.cal.tgz (TM2)

uid___A002_Xc6ff69_X41e4.ms.split.cal.tgz (7M)

uid___A002_Xc7a409_X342e.ms.split.cal.tgz (7M)

uid___A002_Xc74b5b_X458a.ms.split.cal.tgz (7M)

uid___A002_Xc74b5b_X4102.ms.split.cal.tgz (7M)

uid___A002_Xc790bf_X4ea0.ms.split.cal.tgz (7M)

Our example: the N_2H^+ ($v=0$, $J=1-0$) line at 93.173397 GHz, in a galactic star forming region

To make the problem tractable in the short time of this tutorial,
we only look at a small spectral region around the line.

The line is at -47 km/s. We image 30 channels from -50 km/s to -45 km/s.

The following analysis steps *have already been performed* for you on the 12M and 7M data:

- determination of the continuum channels
- uvcontsub
- split out the SPW and field of interest
- concat of the split-out MSs into one MS containing all relevant 12M and 7M data

Next step (our first step): Joint Deconvolution of the concatenated MS

Joint Deconvolution of the concatenated MS

Three purposes:

- 1) demonstrate the joint deconvolution
- 2) obtain and *verify* the deconvolution (tclean) parameters for the later run of sdintimaging (shares most parameters with tclean)
- 3) obtain a dirty image to measure RMS and have a template image for the imregrid of the SD image

All necessary commands for the datacombination session are collected in the file

```
/almastorage/allegro/data/projects/ztuim1kb/data_products/tutorial3b-data_combination/  
leiden-tutorial3b.py
```

Copy this file into your working directory:

```
cd /almastorage/allegro/data/projects/ztuim1kb/analysis/$USER/tutorial3b-data_combination
```

```
cp ../../../../data_products/tutorial3b-data_combination/leiden-tutorial3b.py .
```

and open it in an editor (e.g. emacs).

(Beginning of the hands-on part)

(During waiting periods, we will learn more about sdintimaging and uv coverage assessment.)

```
emacsv@almanode1.strw.leidenuniv.nl
File Edit Options Buffers Tools Python Help
Save Undo
##### European ALMA school at Leiden Observatory
##### 26.01.2026 to 30.01.2026
#####
##### Material for Tutorial 3b
#####
##### DATA COMBINATION
#####
##### D. Petry, ESO, Jan 2026

##### NOTE: this script is not meant as an executable to be run
##### but as a collection of commands which you should
##### copy into your terminal / CASA session prompt as we go
##### through the procedure.

## The original raw data for GOUSS uid__A001_X1296_X198 (pro
## uid__A002_Xc8ed16_X5c3d.ms.split.cal (12M TM1)
## uid__A002_Xcd07af_X41d5.ms.split.cal (12M TM2)
## uid__A002_Xc74b5b_X4102.ms.split.cal (7M)
## uid__A002_Xc790bf_X4ea0.ms.split.cal (7M)
## uid__A002_Xc7a409_X342e.ms.split.cal (7M)
## uid__A002_Xc74b5b_X458a.ms.split.cal (7M)
## uid__A002_Xc6ff69_X41e4.ms.split.cal (7M)

### continuum subtraction
### 12M
## uvcontsub( vis='uid__A002_Xc8ed16_X5c3d.ms.split.cal',
##            outputvis='uid__A002_Xc8ed16_X5c3d.ms.split.c
$ L_Sou6333.60_Dim12m.contsub',
##            field='',
##            spw='25,27,29,31',
##            datacolumn='data',
##            fitspec='25:93.138120-93.190976GHz;93.197080-9
-:-- leiden-tutorial3b.py Top L1 (Python EDoc Abbre
Can't guess python-indent-offset, using defaults: 4
```

Your analysis session at the beginning of the hands-on work (after copying and opening leiden-tutorial3b.py): Now follow along in the script file copying and pasting into your terminal ...

```
Terminal -
File Edit View Terminal Tabs Help
almanode1 [77] $ cd /almastorage/allegro/data/projects/ztuimlkb/analysis/allegroquest9/tutorial3b-data_combination
almanode1 [78] $ pwd
/almastorage/allegro/data/projects/ztuimlkb/analysis/allegroquest9/tutorial3b-data_combination
almanode1 [79] $ cp ../../../../data_products/tutorial3b-data_combination/leiden-tutorial3b.py .
almanode1 [79] $ emacs leiden-tutorial3b.py &
[1] 2569545
almanode1 [80] $
```

Applications | Log Messages (almano... | emacs@almanode1.str... | Terminal - IPython: alle... | 2026-01-22 15:04 | Guest 9 Allegro

emacs@almanode1.strw.leidenuniv.nl

Log Messages (almanode1.strw.leidenuniv.nl:/almastorage/allegro/data/projects/ztuim1kb/analysis/allegroquest9/tutorial3b-data_combination/casa-20260122-140401.log)

File Edit View

Search Message: Filter: Time

Time	Priority	Origin	Message
2026-01-22 14:04:04	INFO	::casa	nologger = False
2026-01-22 14:04:04	INFO	::casa	pipeline = False
2026-01-22 14:04:04	INFO	::casa	startupfile = '/home/allegroquest9/.casa/startup.py'
2026-01-22 14:04:04	INFO	::casa	user_site = False
2026-01-22 14:04:04	INFO	::casa	
2026-01-22 14:04:04	INFO	::casa	casarundata version : casarundata-2025.08.07-1.tar.gz
2026-01-22 14:04:04	INFO	::casa	measures version : WSRT_Measures_20251130-160001.ztar
2026-01-22 14:04:04	INFO	::casa	
2026-01-22 14:04:04	INFO	::casa	Checking Measures tables in data repository sub-directory /almastorage/software/casa-6.6.6-17-pipeline-2025.1.0.35-py3.10...
2026-01-22 14:04:04	INFO	::casa	IERSeop2000 (version date, last date in table (UTC)): 2025/11/10/15:15, 2025/10/11/00:00:00
2026-01-22 14:04:05	INFO	::casa	IERSeop97 (version date, last date in table (UTC)): 2025/11/10/15:15, 2025/10/11/00:00:00
2026-01-22 14:04:05	INFO	::casa	IERSpredict (version date, last date in table (UTC)): 2025/11/30/15:15, 2026/02/28/00:00:00
2026-01-22 14:04:05	INFO	::casa	TAI.UTC (version date, last date in table (UTC)): 2025/11/19/15:15, 2017/01/01/00:00:00

Insert Message: Lock scroll

Your analysis session after creating the links to the input data and starting casapy-666p:

(all further commands in the script are meant to run under CASA)

```
[1] 2569545
almanode1 [80] $ ln -sf ../../data_products/tutorial3b-data_combination/GOUS2017.1.01355.L_SouG333.60_ALMA12m+7m_93.17
376GHz_allvis.ms
almanode1 [81] $ ln -sf ../../data_products/tutorial3b-data_combination/member.uid__A001_X1296_X1a9.G333.60.spw17.I.s
d.im.fits
almanode1 [82] $ casapy-666p

optional configuration file not found, continuing CASA startup without it

IPython 8.26.0 -- An enhanced Interactive Python.

Using matplotlib backend: tkagg
CASA 6.6.6.17 -- Common Astronomy Software Applications [6.6.6.17]

CASA <1>: QStandardPaths: XDG_RUNTIME_DIR not set, defaulting to '/tmp/allegroquest9/runtime-allegroquest9'
```

