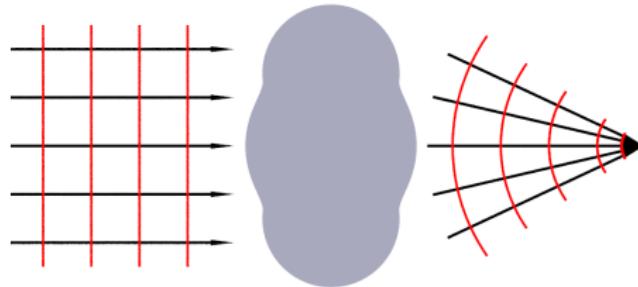


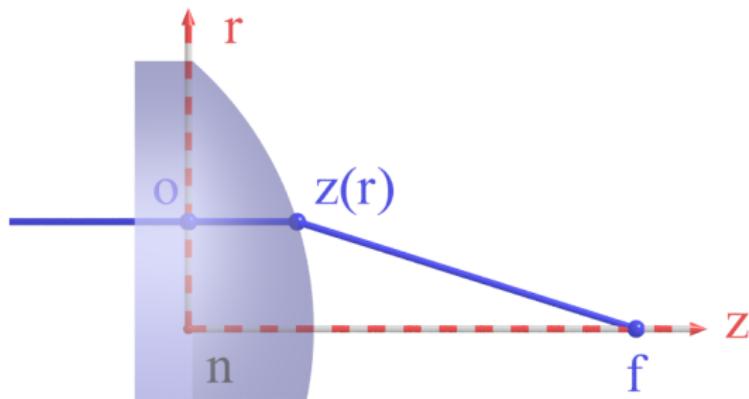
Outline

- ① Geometrical Optics
- ② Lenses
- ③ Mirrors
- ④ Optical Systems
- ⑤ Aberrations

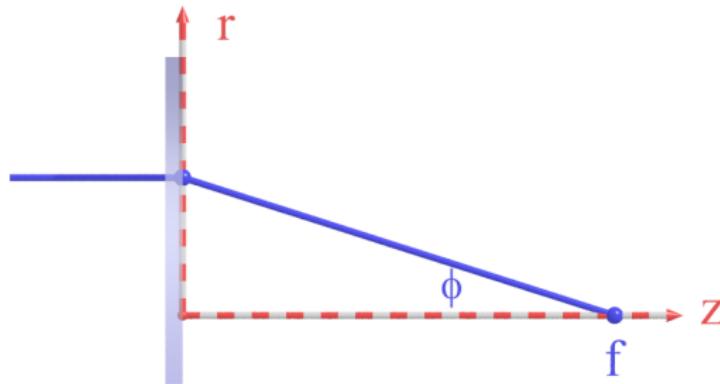


- rays are normal to locally flat wave
- rays are reflected and refracted according to Fresnel equations
- phase is neglected \Rightarrow incoherent sum
- rays can carry polarization information
- geometrical optics neglects diffraction effects: $\lambda \Rightarrow 0$
- *physical optics* $\lambda > 0$
- simplicity of geometrical optics mostly outweighs limitations

Surface Shape of Perfect Lens

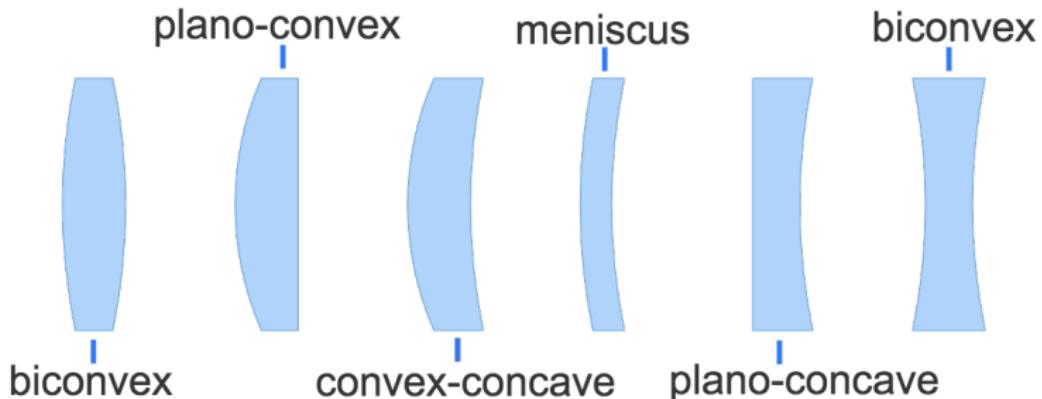


- lens material has index of refraction n
- $\overline{o z(r)} \cdot n + \overline{z(r) f} = \text{constant}$
- $n \cdot z(r) + \sqrt{r^2 + (f - z(r))^2} = \text{constant}$
- solution $z(r)$ is hyperbola with eccentricity $e = n > 1$



- Assumptions:
 - 1 Snell's law for small angles of incidence ($\sin \phi \approx \phi$)
 - 2 ray height r small so that optics curvature can be neglected (plane optics, $(\cos \phi \approx 1)$)
 - 3 $\tan \phi \approx \phi = r/f$
- paraxial approximation reasonable for $\phi \lesssim 10$ degrees

Spherical Lenses



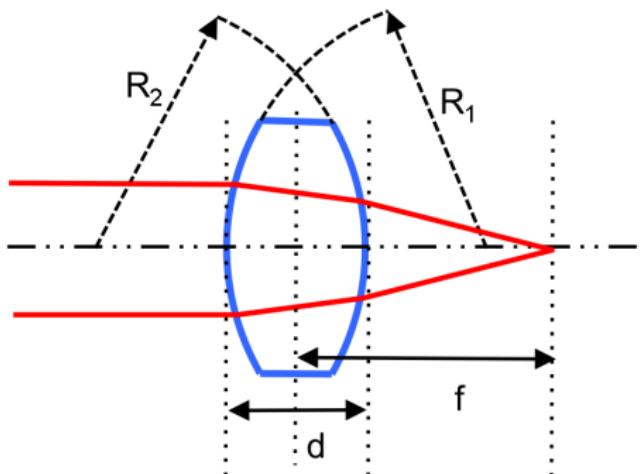
- surface error requirement less than $\lambda/10$
- most optical surfaces are spherical due to ease of manufacturing
- can easily join two spherical surfaces with equal radii

1.1-meter Singlet Lens of Swedish Solar Telescope



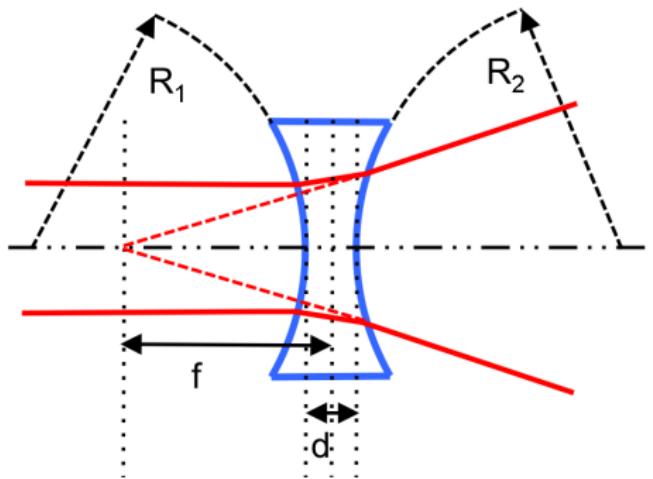
www.isf.astro.su.se/NatureNov2002/images/dan/DCP_5127.JPGg

Positive/Converging Spherical Lens Parameters



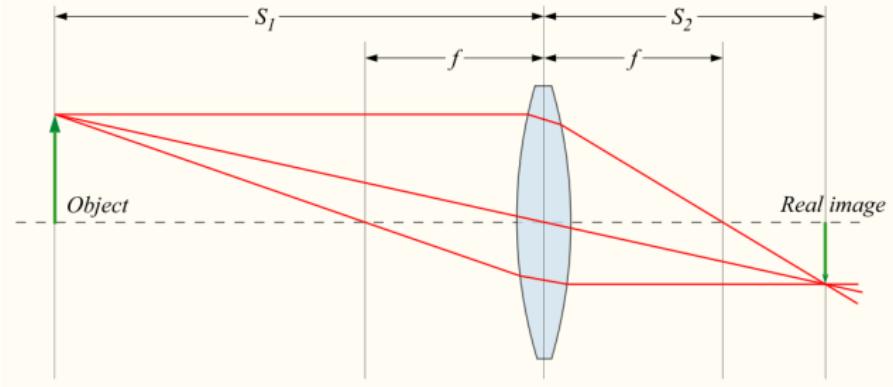
- center of curvature and radii with signs: $R_1 > 0, R_2 < 0$
- center thickness: d
- positive focal length $f > 0$

Negative/Diverging Spherical Lens Parameters



- note different signs of radii: $R_1 < 0, R_2 > 0$
- virtual focal point
- negative focal length ($f < 0$)

General Lens Setup



commons.wikimedia.org/wiki/File:Lens3.svg

- *object distance S_1 , object height h_1*
- *image distance S_2 , image height h_2*
- axis through two centers of curvature is *optical axis*
- surface point on optical axis is the *vertex*
- *chief ray* through center maintains direction

Thin Lens Approximation

- thick-lens equation:

$$\frac{1}{S_1} + \frac{1}{S_2} = (n - 1) \left(\frac{1}{R_1} - \frac{1}{R_2} + \frac{(n - 1)d}{nR_1 R_2} \right) = \frac{1}{f}$$

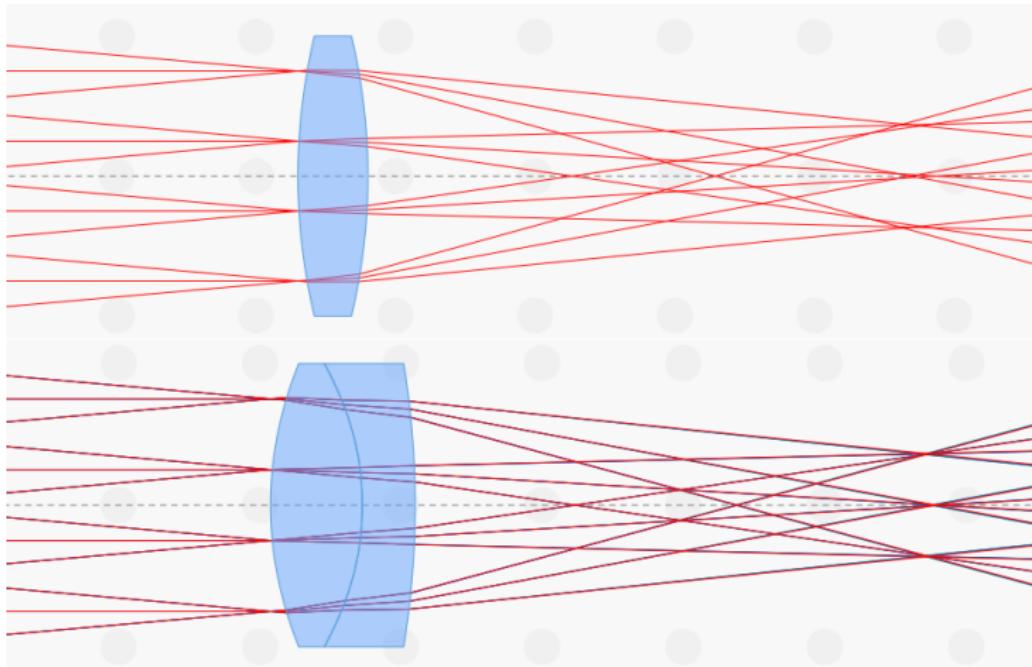
- thin-lens equation for $d \ll R_1 R_2$:

$$\frac{1}{S_1} + \frac{1}{S_2} = (n - 1) \left(\frac{1}{R_1} - \frac{1}{R_2} \right) = \frac{1}{f}$$

Finite Imaging

- rarely image point sources, but extended object
- object and image size are proportional
- orientation of object and image are inverted
- magnification perpendicular to optical axis: $M = h_2/h_1 = -S_2/S_1$

Achromatic Lens

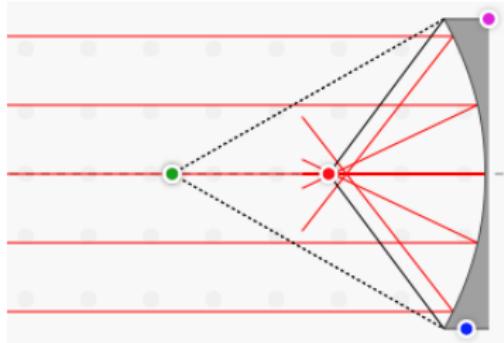
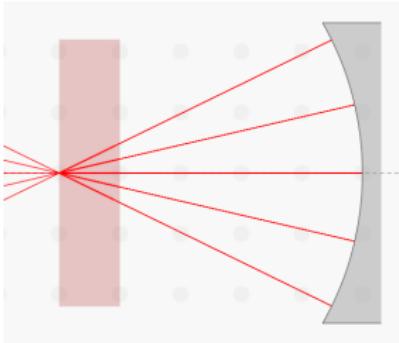


- combination of 2 lenses, different glass dispersion
- also less spherical aberration

Mirrors vs. Lenses

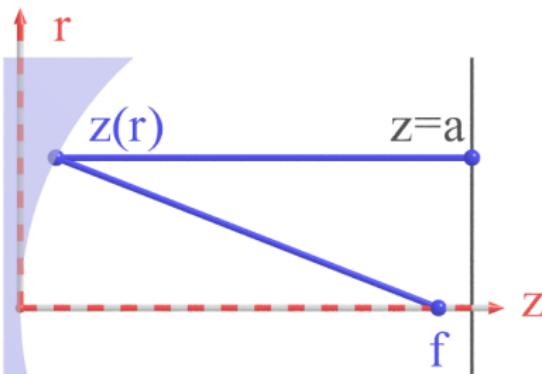
- mirrors are completely achromatic
- reflective over very large wavelength range (UV to radio)
- can be supported from the back
- can be segmented
- wavefront error is twice that of surface, lens is $(n-1)$ times surface
- only one surface to 'play' with

Spherical Mirrors



- easy to manufacture
- focuses light from center of curvature onto itself
- focal length is half of curvature: $f = R/2$
- tip-tilt misalignment does not matter
- has no optical axis
- does not image light from infinity correctly (spherical aberration)

Parabolic Mirrors

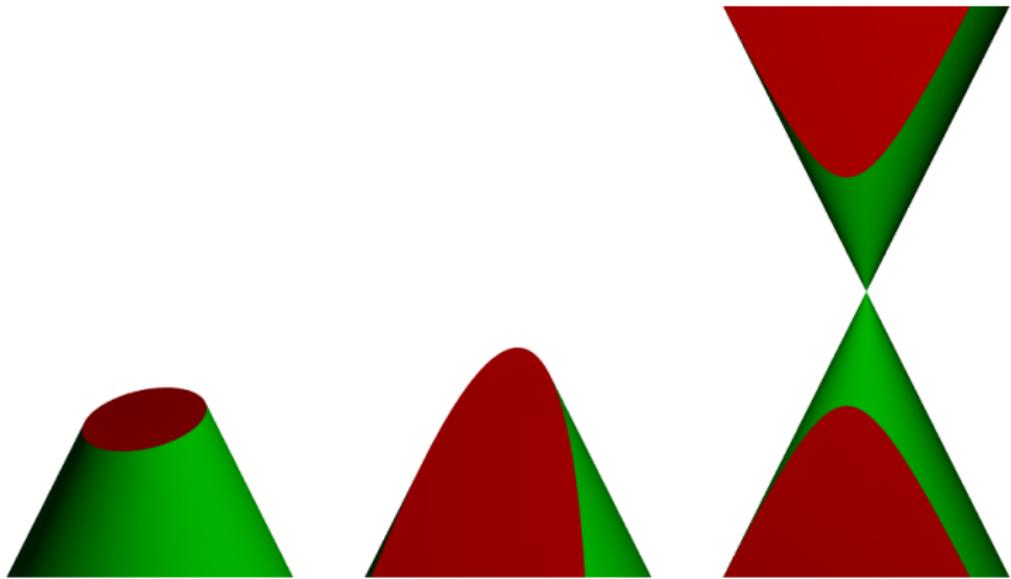


- want to make flat wavefront into spherical wavefront
- distance $\overline{az(r)} + \overline{z(r)f} = \text{const.}$
- $z(r) = r^2/2R$
- perfect image of objects at infinity
- has clear optical axis

8.4-meter Large Binocular Telescope Primary Mirror Nr. 2

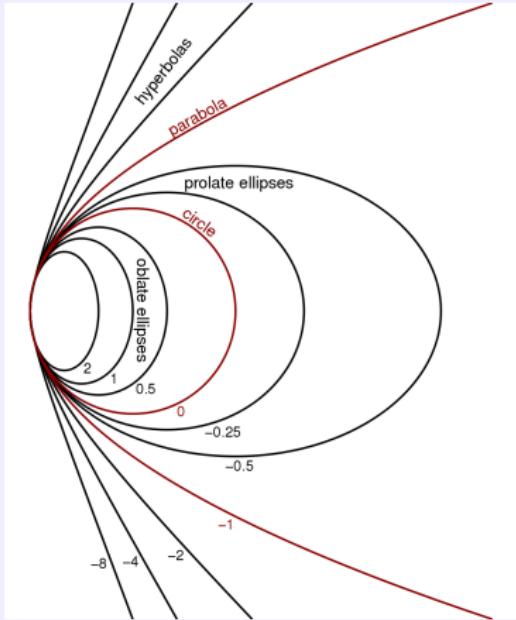


mirrorlab.as.arizona.edu/sites/mirrorlab.as.arizona.edu/files/LBT_2_DSCN1893.jpg



- circle and ellipses: cuts angle $<$ cone angle
- parabola: angle = cone angle
- hyperbola: cut along axis

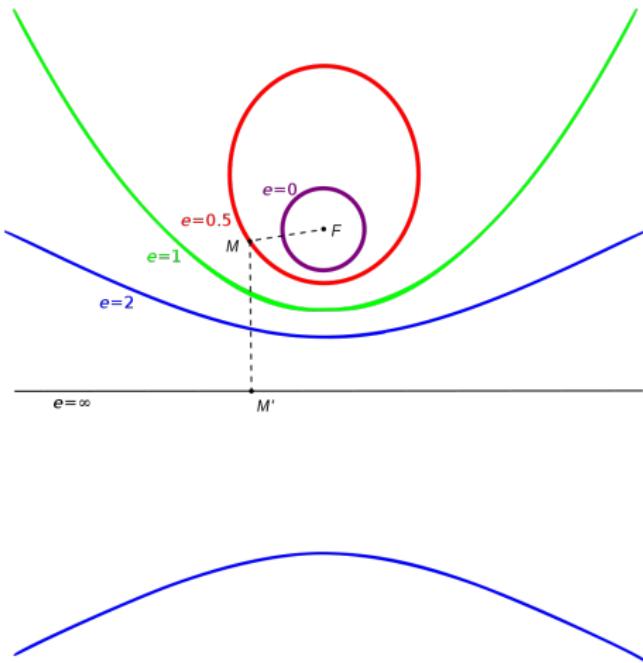
Conic Constant K



en.wikipedia.org/wiki/Conic_constant

- $r^2 - 2Rz + (1 + K)z^2 = 0$ for $z(r = 0) = 0$
- $z = \frac{r^2}{R} \frac{1}{1 + \sqrt{1 - (1+K)\frac{r^2}{R^2}}}$
- R radius of curvature
- $K = -e^2$, e eccentricity
- prolate ellipsoid ($K > 0$)
- sphere ($K = 0$)
- oblate ellipsoid ($0 > K > -1$)
- parabola ($K = -1$)
- hyperbola ($K < -1$)
- all conics are almost spherical close to origin
- analytical ray intersections

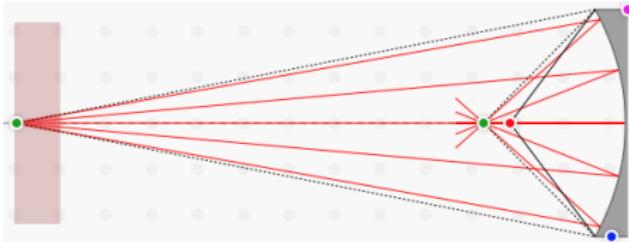
Foci of Conic Sections



en.wikipedia.org/wiki/File:Eccentricity.svg

- sphere has single focus
- ellipse has two foci
- parabola (ellipse with $e = 1$) has one focus (and another one at infinity)
- hyperbola ($e > 1$) has two focal points

Elliptical Mirrors



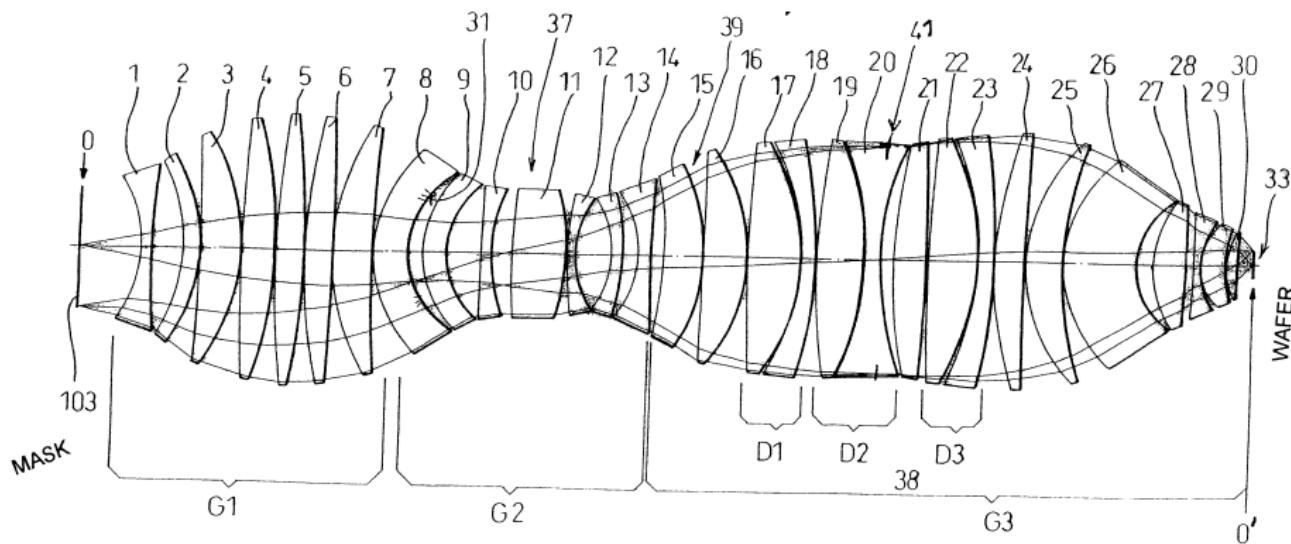
- have two foci at finite distances
- perfectly reimagine one focal point into another

Hyperbolic Mirrors



- have a real focus and a virtual focus (behind mirror)
- perfectly reimagine one focal point into another

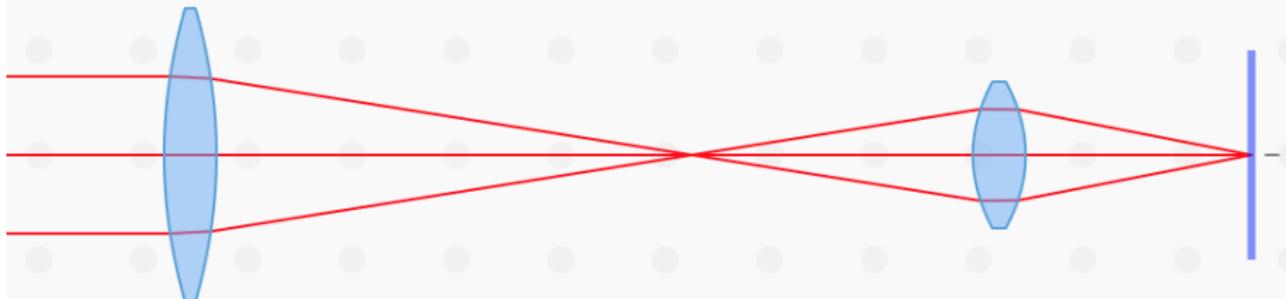
Overview



www.google.com/patents/US6560031

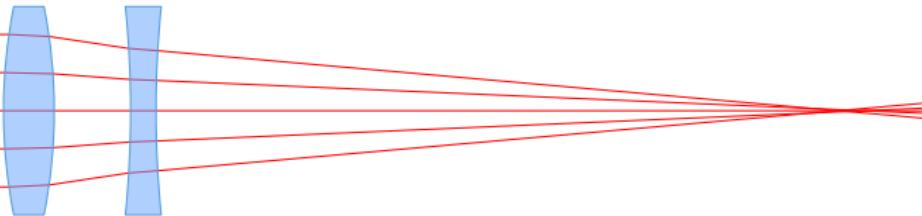
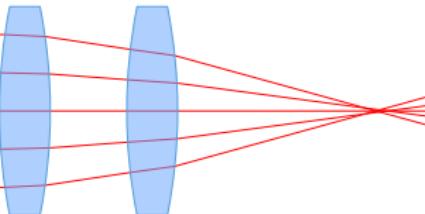
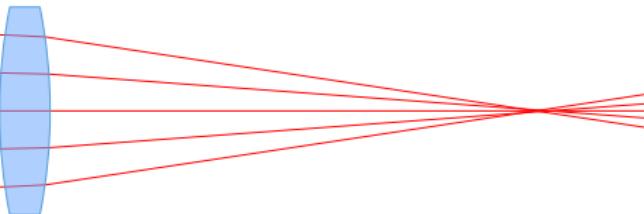
- combinations of several optical elements (lenses, mirrors, stops)
- examples: camera “lens”, microscope, telescopes, instruments

Simple Thin-Lens Combinations

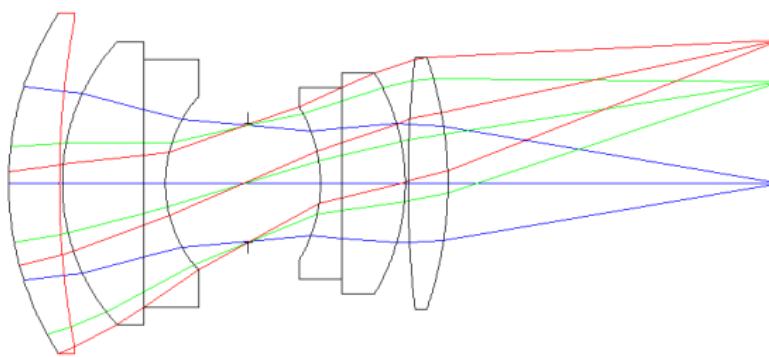


- thin-lens combinations can be treated analytically
- distance > sum of focal lengths \Rightarrow real image between lenses \Rightarrow apply single-lens equation successively
- focal length of general thin lens combination
 - focal lengths: f_1, f_2
 - distance between thin lenses: d_{12}
 - effective focal length: $\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2} - \frac{d_{12}}{f_1 f_2}$

Example Thin-Lens Combinations

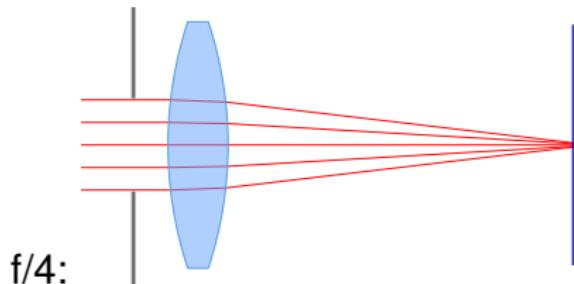
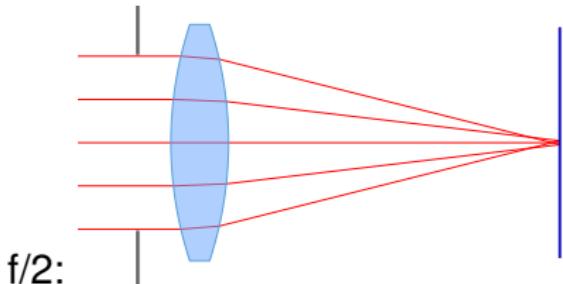


Aperture



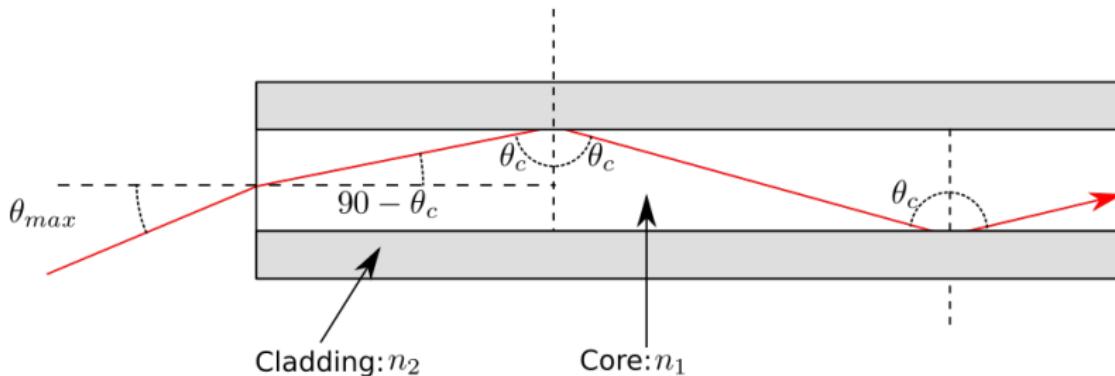
- all optical systems have a place where 'aperture' is limited
- main mirror of telescopes
- aperture stop in photographic lenses
- aperture typically has a maximum diameter
- aperture size is important for diffraction effects

F-number



- describes the light-gathering ability of the optical system
- f-number given by $F = f/D$
- also called focal ratio or f-ratio, written as: f/F
- the bigger F , the better the paraxial approximation works
- fast system for $F < 2$, slow system for $F > 2$

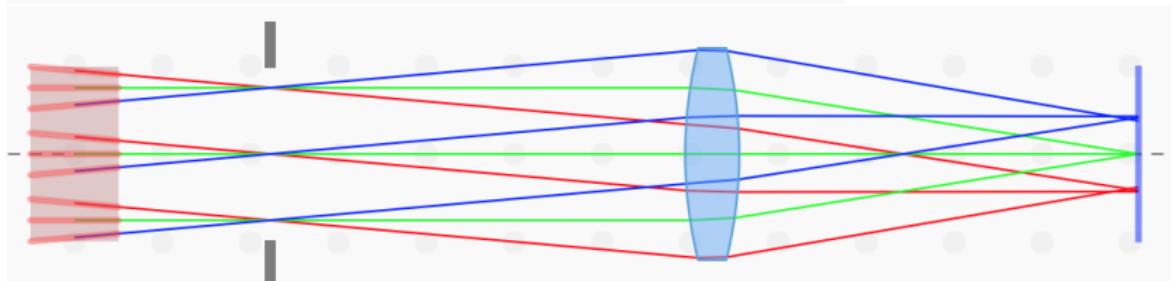
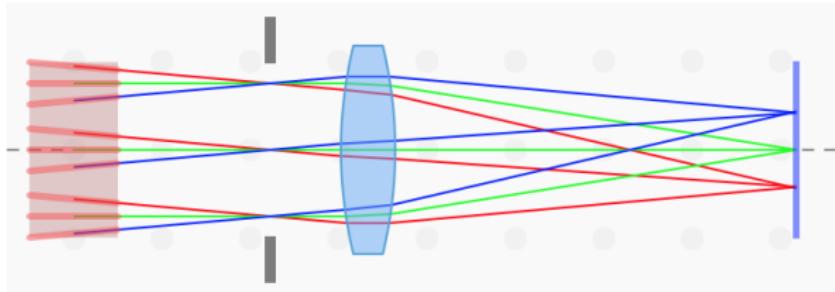
Numerical Aperture in Fibers



en.wikipedia.org/wiki/File:OF-na.svg

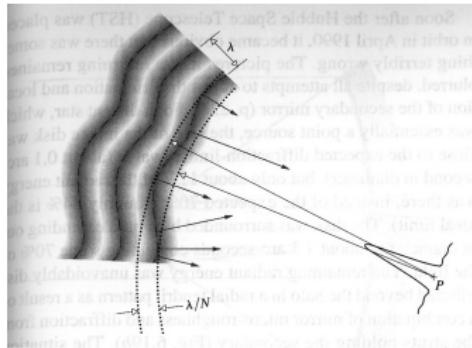
- acceptance cone of the fiber determined by materials
- $NA = n \sin \theta = \sqrt{n_1^2 - n_2^2}$
- n index of refraction of external medium

Telecentric Arrangement



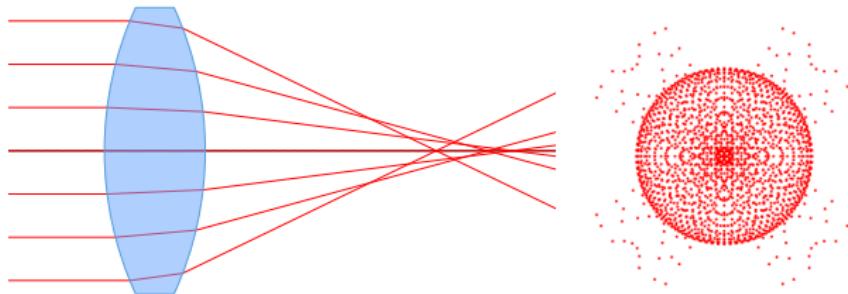
- as seen from image, pupil is at infinity
- easy: lens is its focal length away from pupil (image)
- magnification does not change with focus positions
- ray cones for all image points have the same orientation

Spot Diagrams and Wavefronts



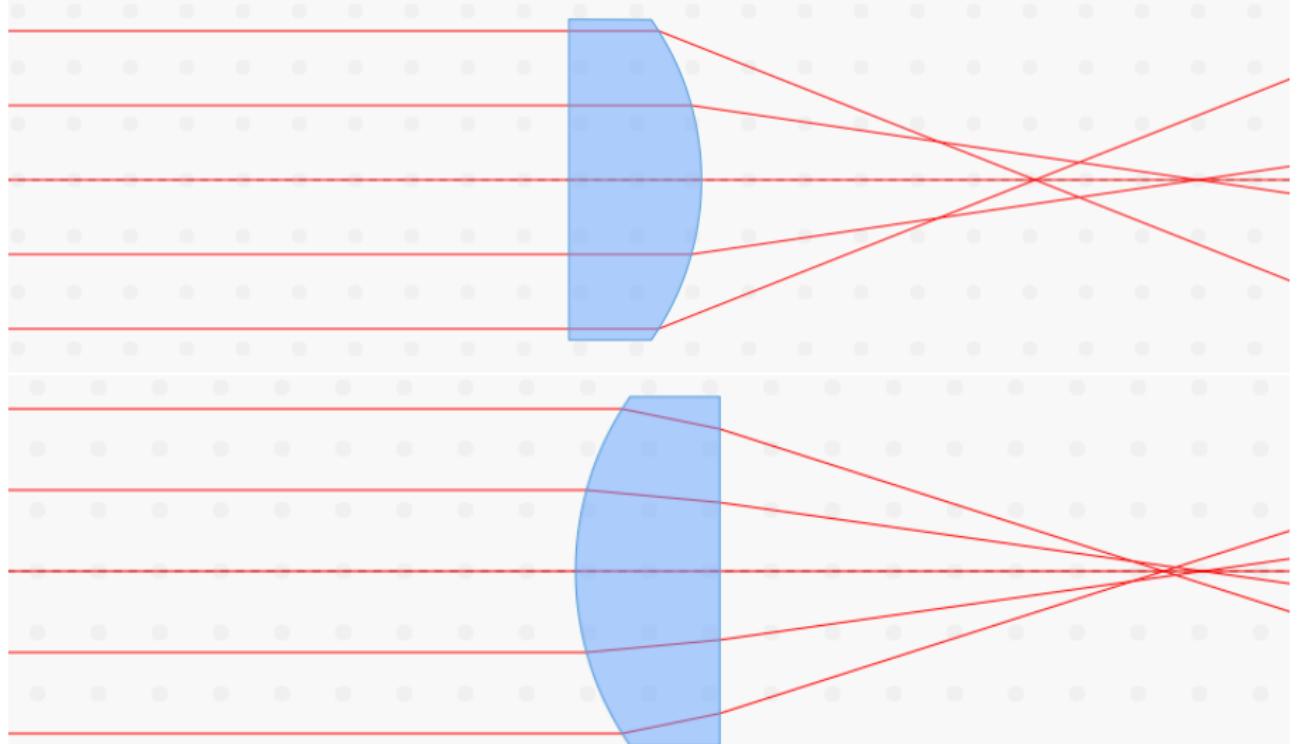
- plane of least confusion is location where image of point source has smallest diameter
- spot diagram: shows ray locations in plane of least confusion
- spot diagrams are closely connected with wavefronts
- aberrations are deviations from spherical wavefront

Spherical Aberration of Spherical Lens

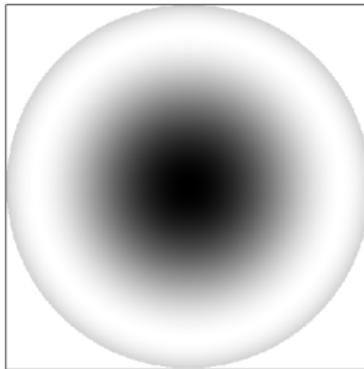
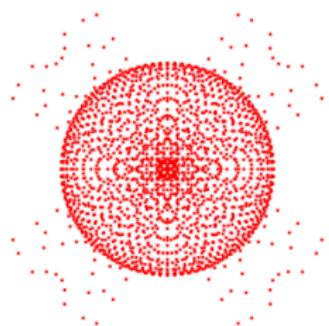


- different focal lengths of paraxial and marginal rays
- longitudinal spherical aberration along optical axis
- transverse (or lateral) spherical aberration in image plane
- much more pronounced for short focal ratios
- foci from paraxial beams are further away than marginal rays
- spot diagram shows central area with fainter disk around it

Minimizing Spherical Aberrations

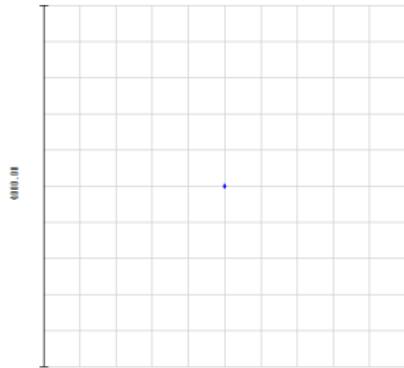
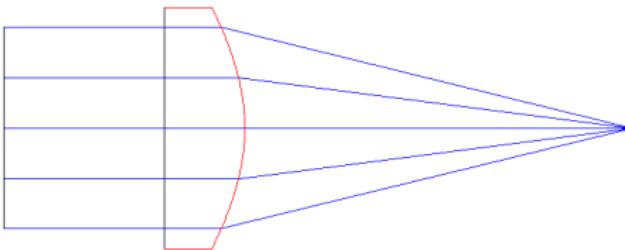


Spherical Aberration Spots and Waves



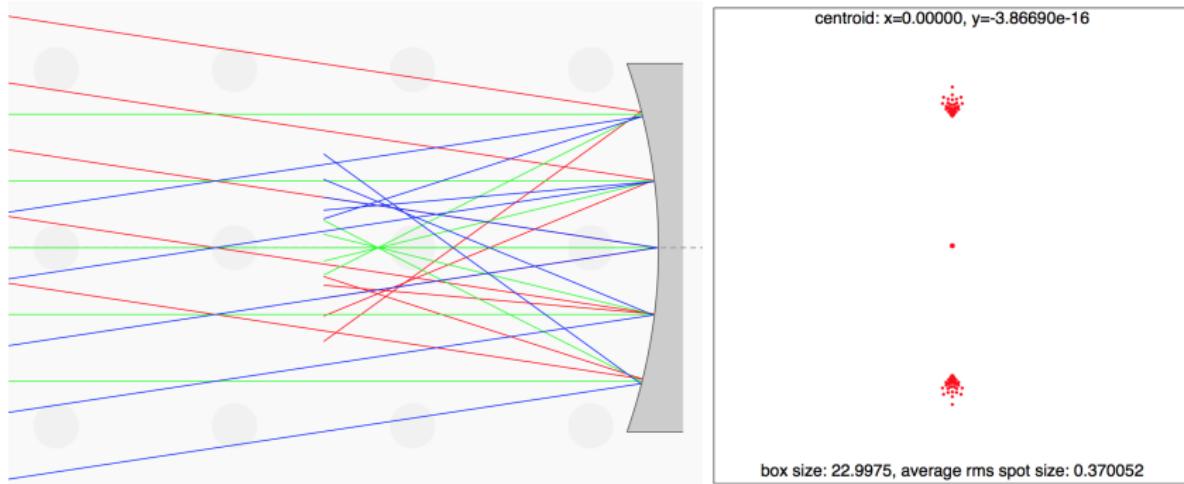
- spot diagram shows central area with fainter disk around it
- wavefront has peak and turned-up edges

Aspheric Lens



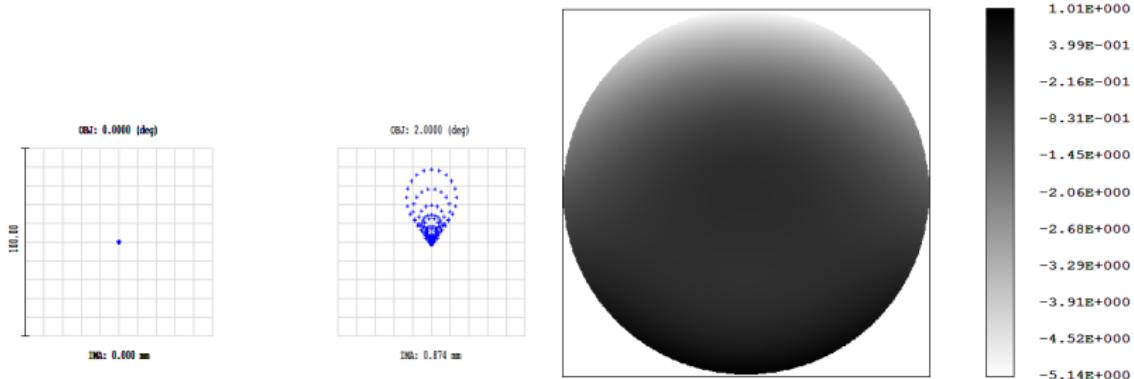
- conic constant $K = -1 - \sqrt{n}$ makes perfect lens
- difficult to manufacture
- but possible these days

Coma



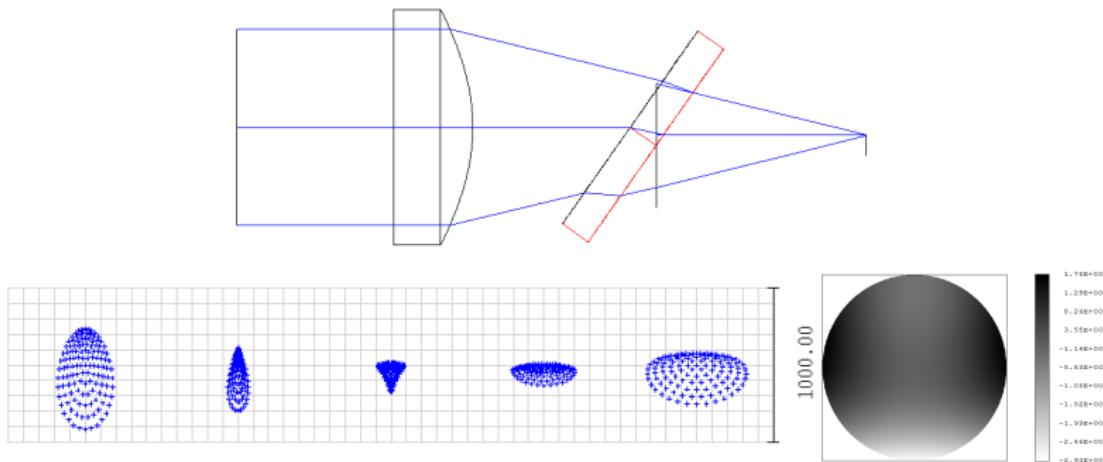
- typically seen for object points away from optical axis
- leads to 'tails' on stars

Coma Spots and Waves



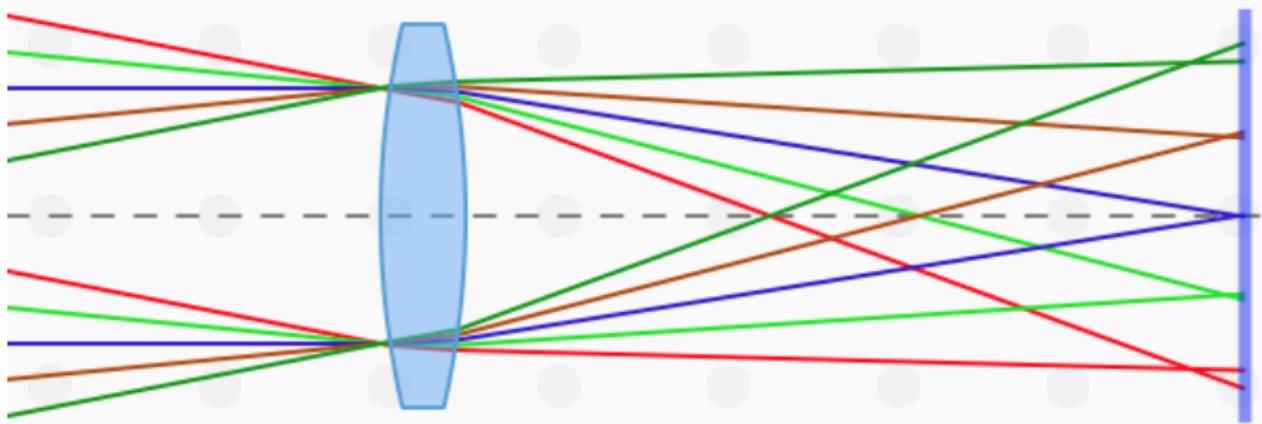
- parabolic mirror with perfect on-axis performance
- spots and wavefront for off-axis image points
- wavefront is tilted in inner part

Astigmatism due to Tilted Glass Plate in Converging Beam



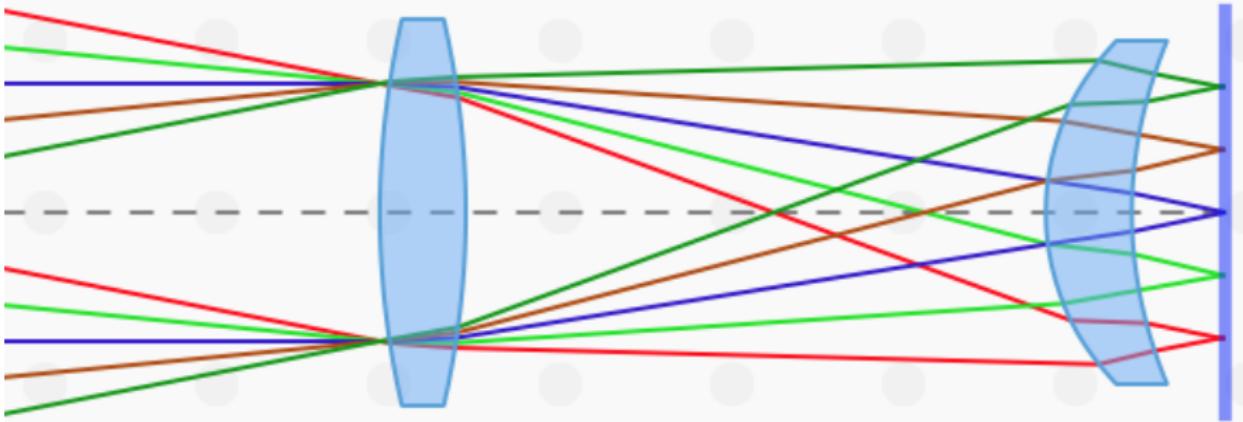
- astigmatism: focus in two orthogonal directions, but not in both at the same time
- tilted glass-plate: astigmatism, spherical aberration, beam shift
- tilted plates: beam shifters, filters, beamsplitters
- difference of two parabolae with different curvatures

Field Curvature



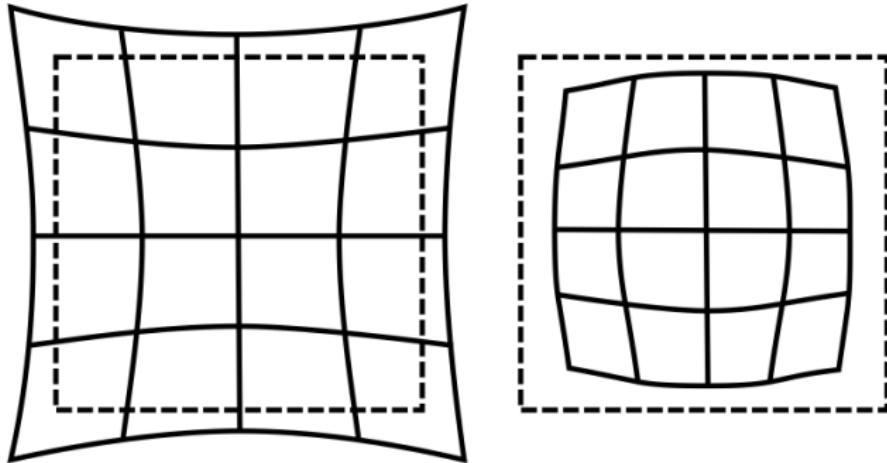
- field (Petzval) curvature: image lies on curved surface
- curvature comes from lens thickness variation across aperture
- problems with flat detectors (e.g. CCDs)
- potential solution: field flattening lens close to focus

Petzval Field Flattening



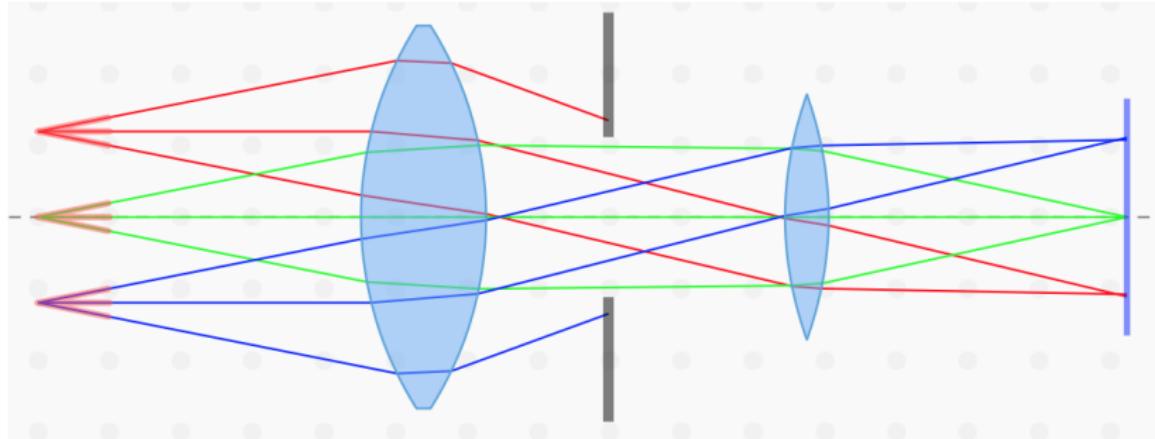
- Petzval curvature only depends on index of refraction (n_i) and focal length (f_i) of lenses
- Petzval curvature is independent of lens position!
- Petzval curvature vanishes for $\sum_i \frac{1}{n_i f_i} = 0$
- field flattener also makes image much more telecentric

Distortion



- image is sharp but geometrically distorted
- (a) object
- (b) positive (or pincushion) distortion
- (c) negative (or barrel) distortion

Vignetting



- effective aperture stop depends on position in object
- image fades toward its edges
- only influences transmission, not image sharpness

Aberration Descriptions

Seidel Aberrations

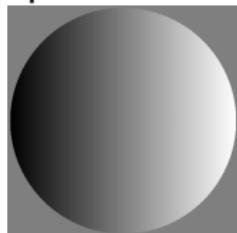
- introduced by Ludwig von Seidel (1857)
- Taylor expansion of $\sin \phi = \phi - \frac{\phi^3}{3!} + \frac{\phi^5}{5!} - \dots$
- paraxial approximation: first-order optics, $\sin \phi \sim \phi$
- Seidel aberrations: third-order optics, $\sin \phi \sim \phi - \frac{\phi^3}{3!}$
- describes spherical, astigmatism, coma, field curvature, distortion

Zernike Polynomials

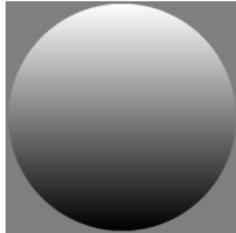
- orthonormal basis on unit circle
- lowest orders correspond to Seidel aberrations
- $Z_n^m(r, \phi) = R_n^m(r) \cos m\phi, Z_n^{-m}(r, \phi) = R_n^m(r) \sin m\phi$
- $n \geq m \geq 0$
- $R_n^m(r) = \sum_{k=0}^{\frac{n-m}{2}} (-1)^k \binom{n-k}{k} \binom{n-2k}{\frac{n-m}{2}-k} r^{n-2k}$

Zernike Polynomials

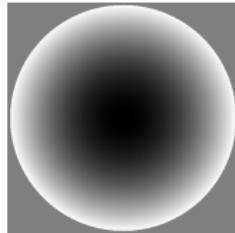
tip



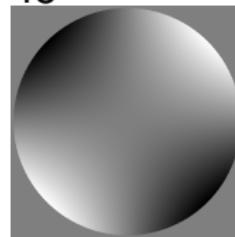
tilt



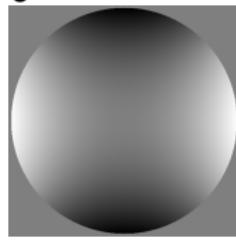
focus



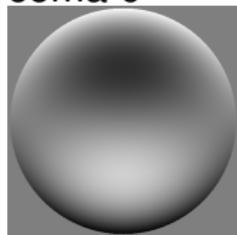
astigmatism
 45°



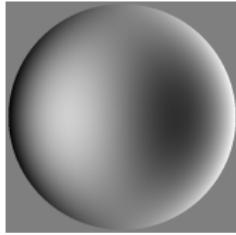
astigmatism
 0°



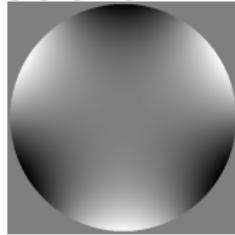
coma 0°



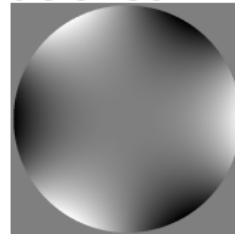
coma 90°



trefoil 0°



trefoil 30°



third-order
spherical

