

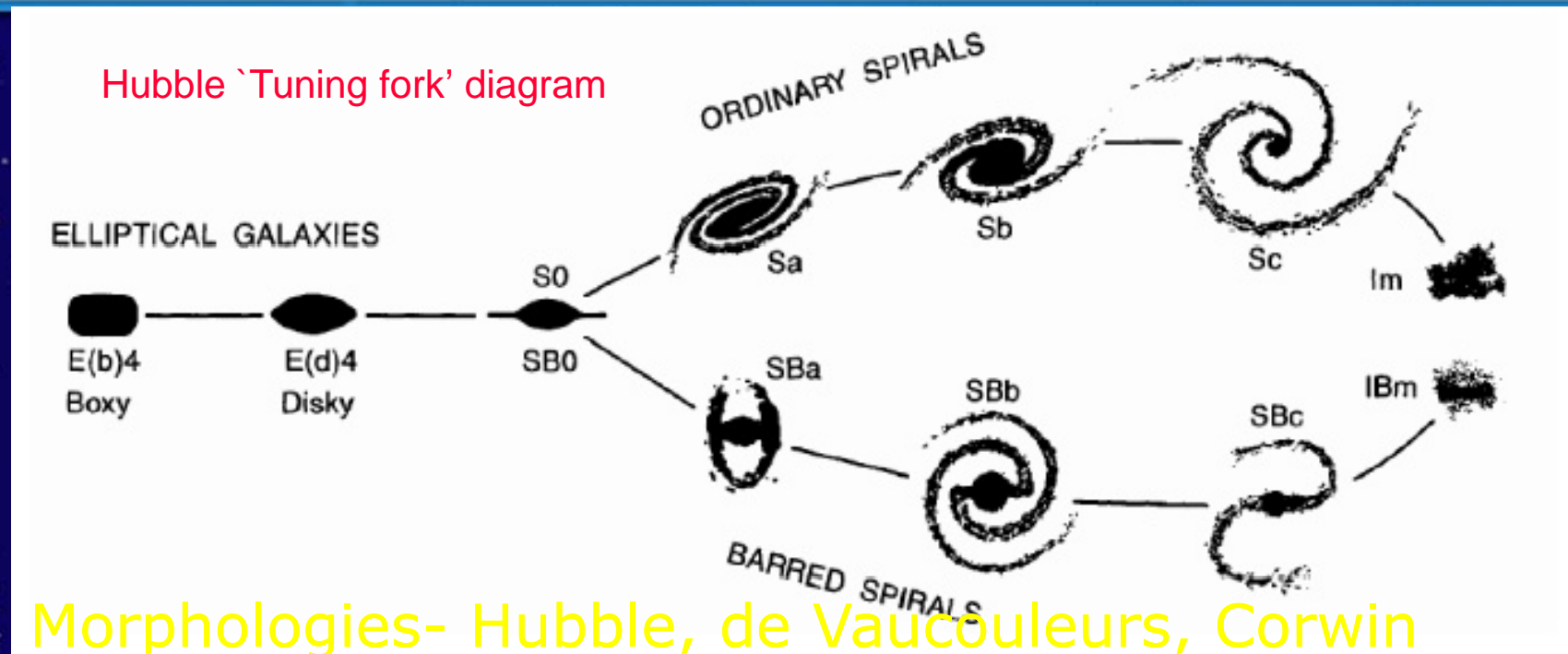
# Interacademic Lecture 8

Future directions in Astronomical data management

Kapteyn Astronomical Institute  
University of Groningen

Edwin A. Valentijn  
12 May 2010

# Case: morphologies parameters vs theory -degeneracy



## Morphologies- Hubble, de Vaucouleurs, Corwin

- U, B, V, R, I, Z, K
  - $m_{tot}$ ,  $m_{25}$ ,  $m_{26}$ ,  $D_{26}$ ,  $D_{25}$ ,  $D_{90\%}$ ,  $r_{eff}$ ,  $sb_0$ ,  $sb_{eff}$
- Structure/radial
  - $a/b$ ,  $pa$ ,  $B/D$ ,  $N$ , exp scale length,  $\delta$  (B-R)
- colours
  - Central (B-R),  $(B_R)_{tot}$

# What characterizes?

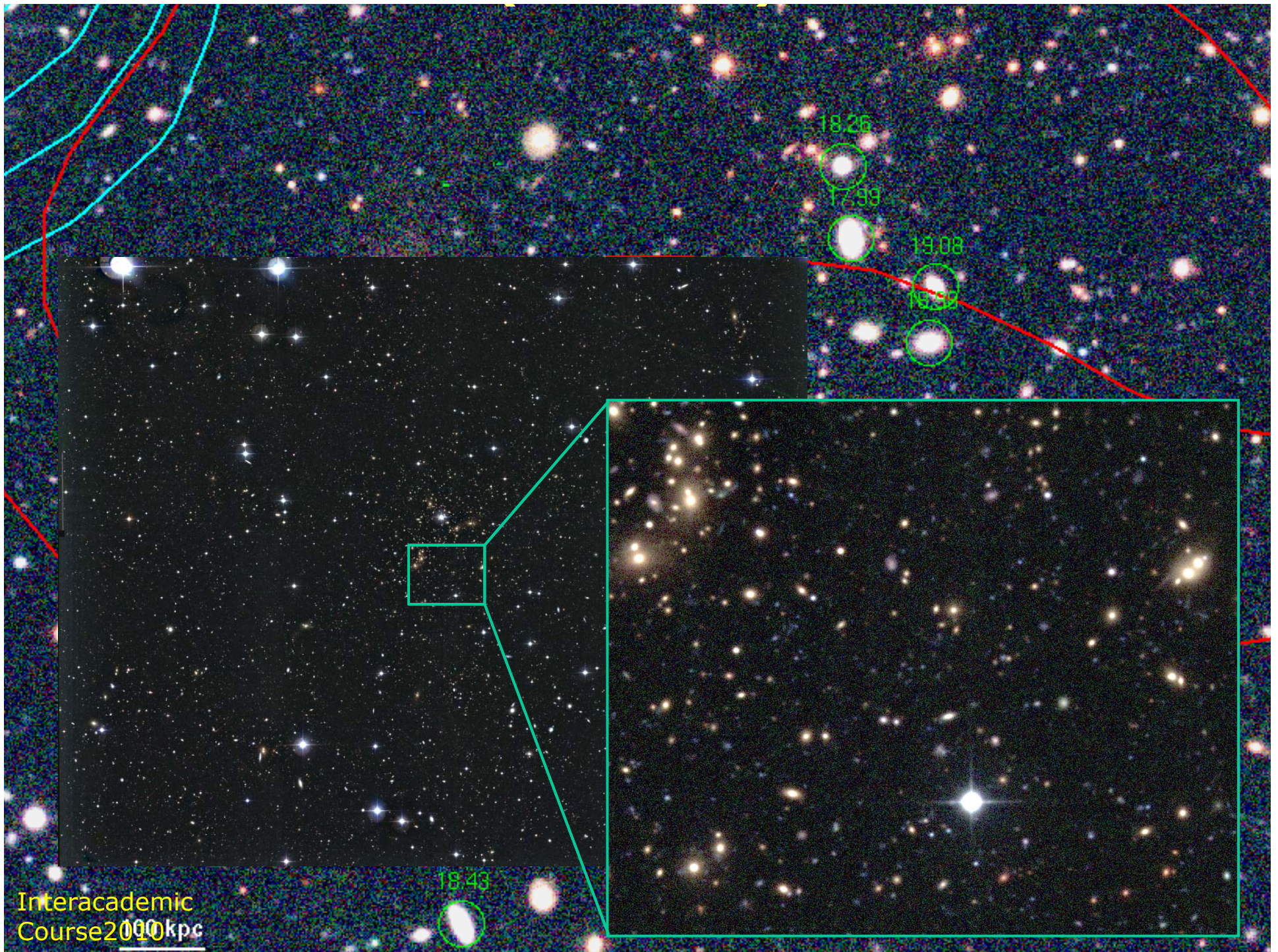
Phenomenology ← understanding → theory

Observations - interpretation

Primary relations - secondary

N parameters      N datapoints

Example Morphological transitions

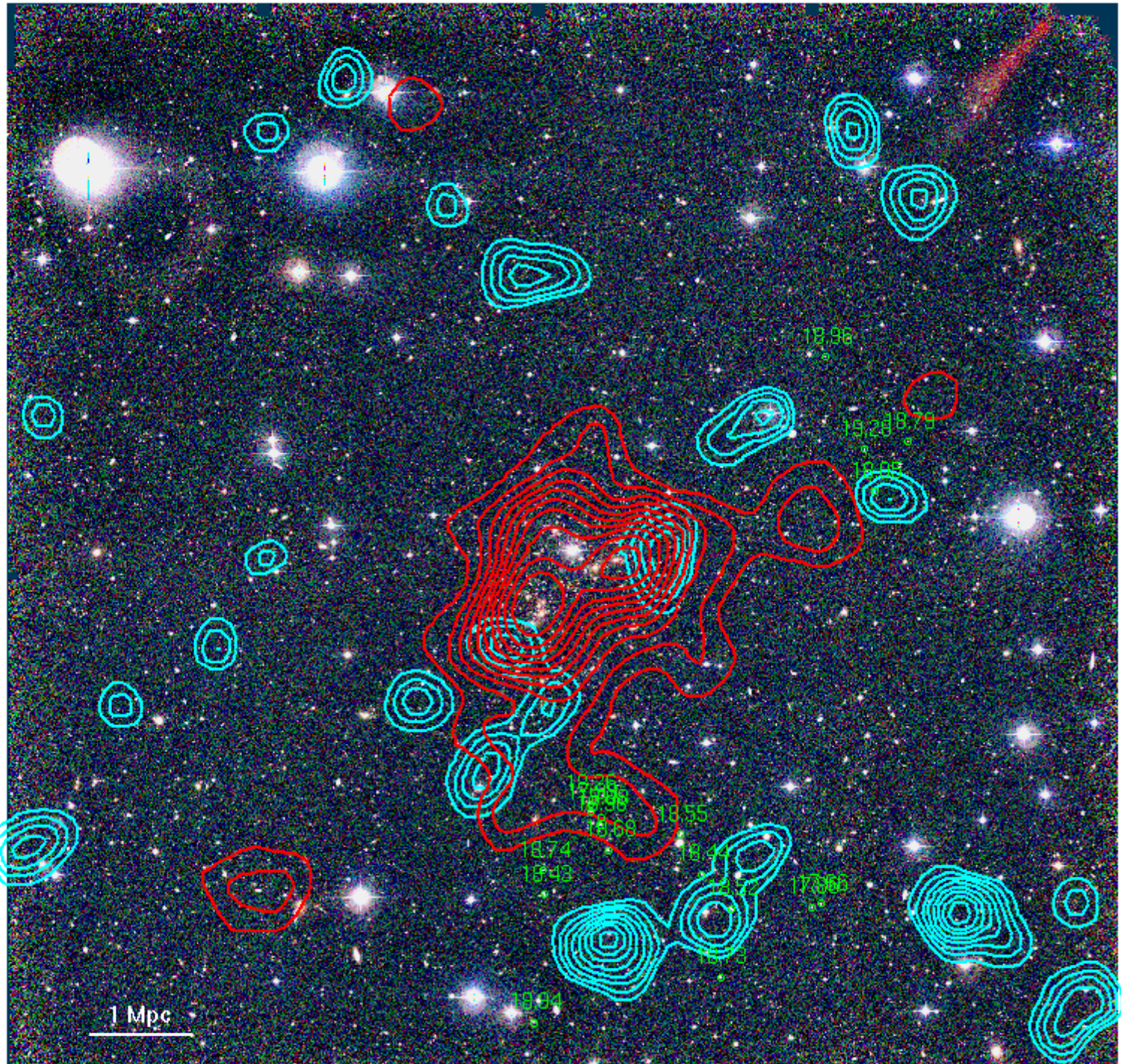


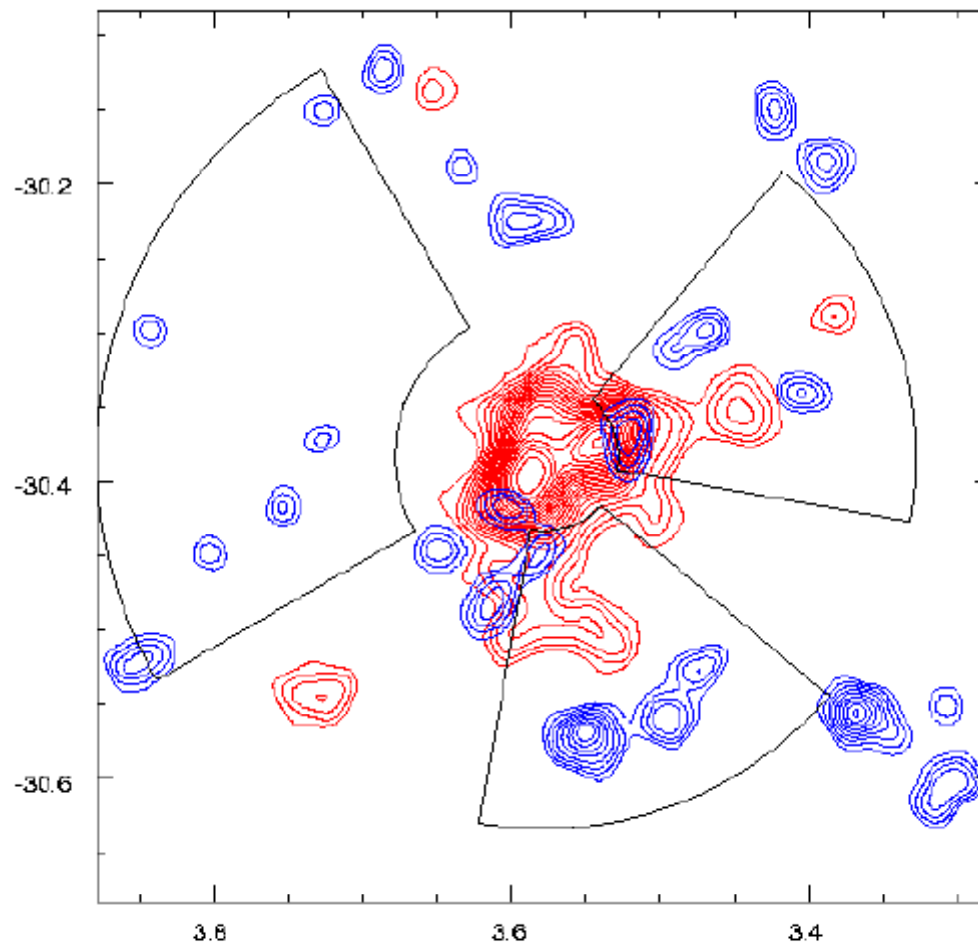
Braglia  
Pierini  
Bohringer  
AA L 2007

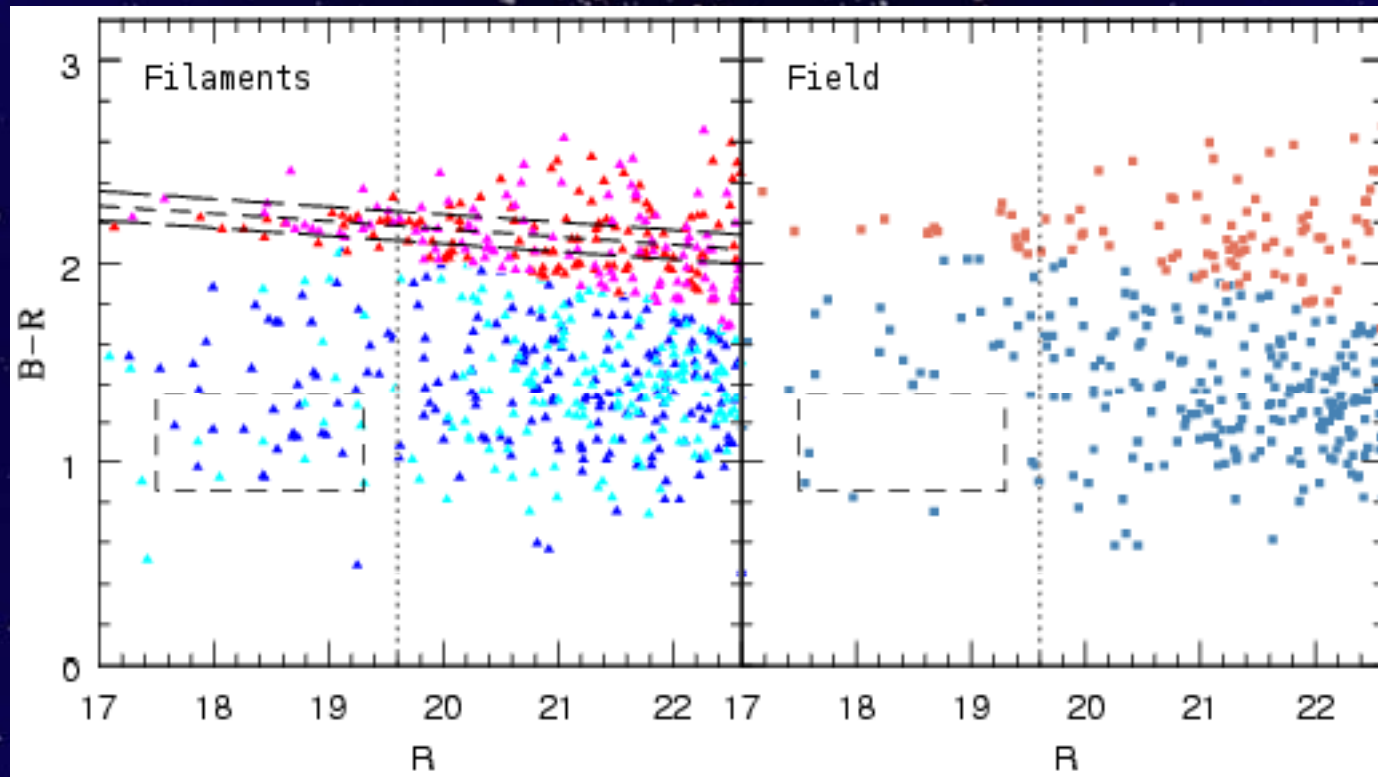
A2744  
 $Z \sim 0.3$

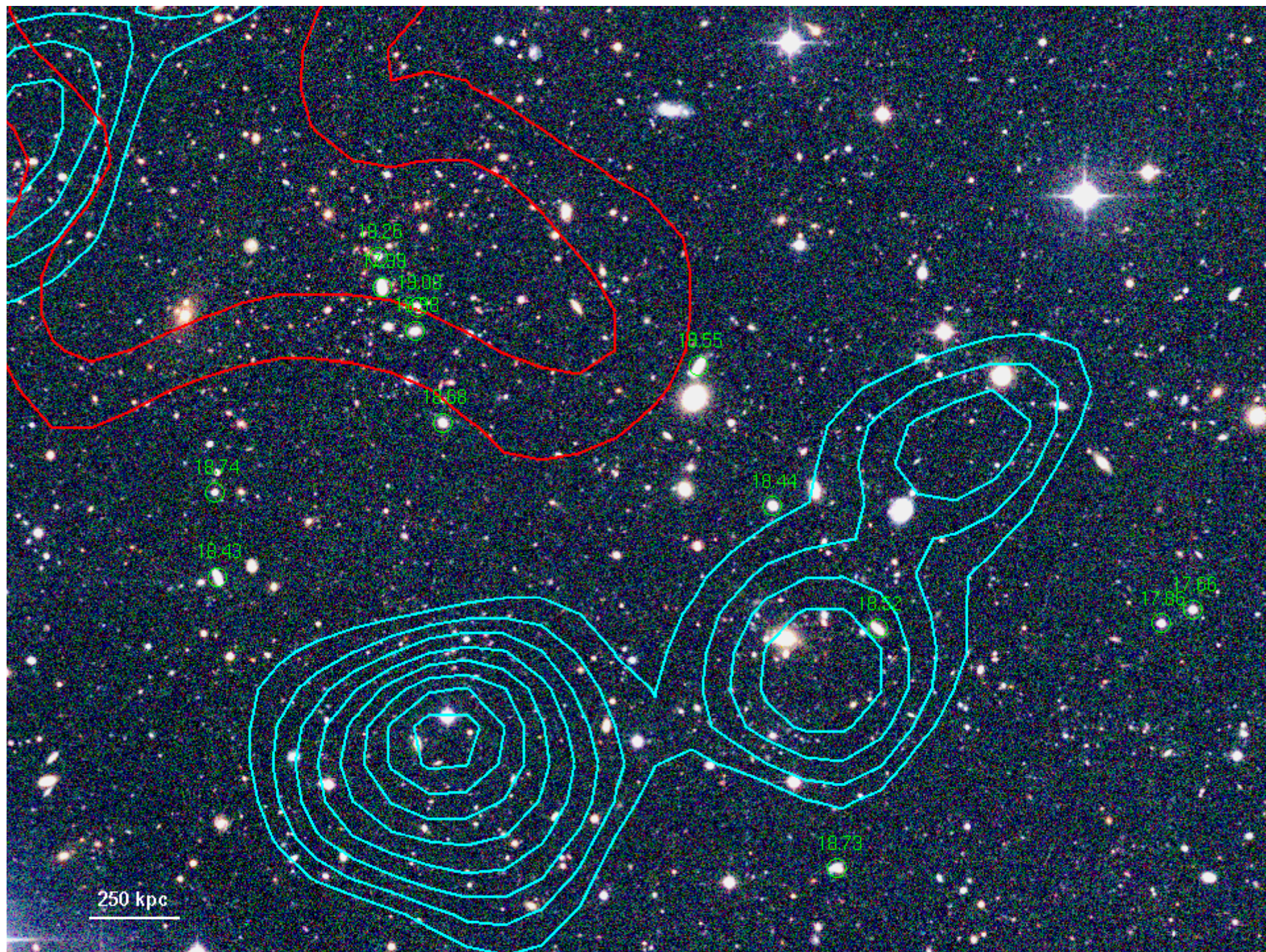
WFI@2.2m

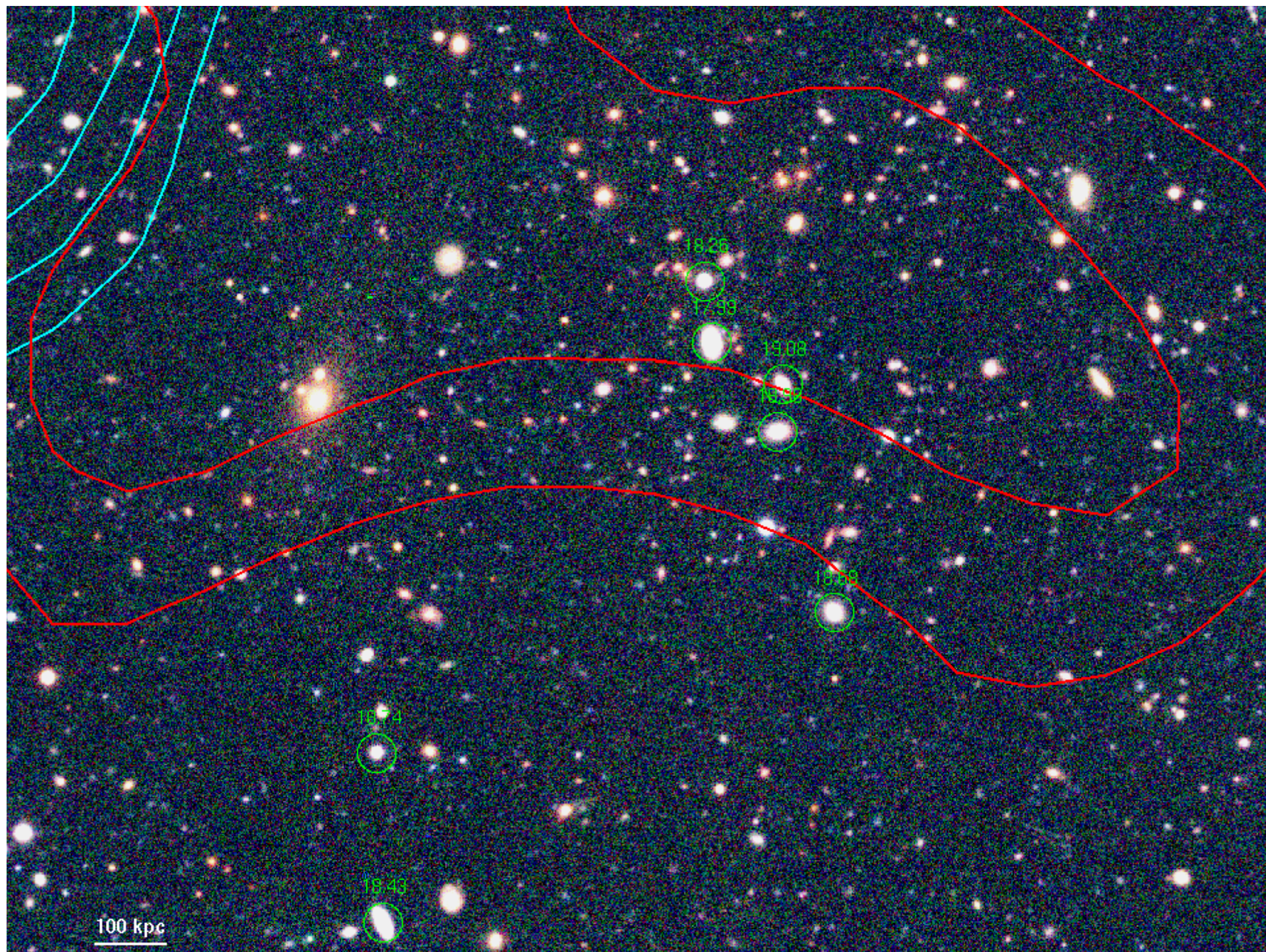
Interac  
Course





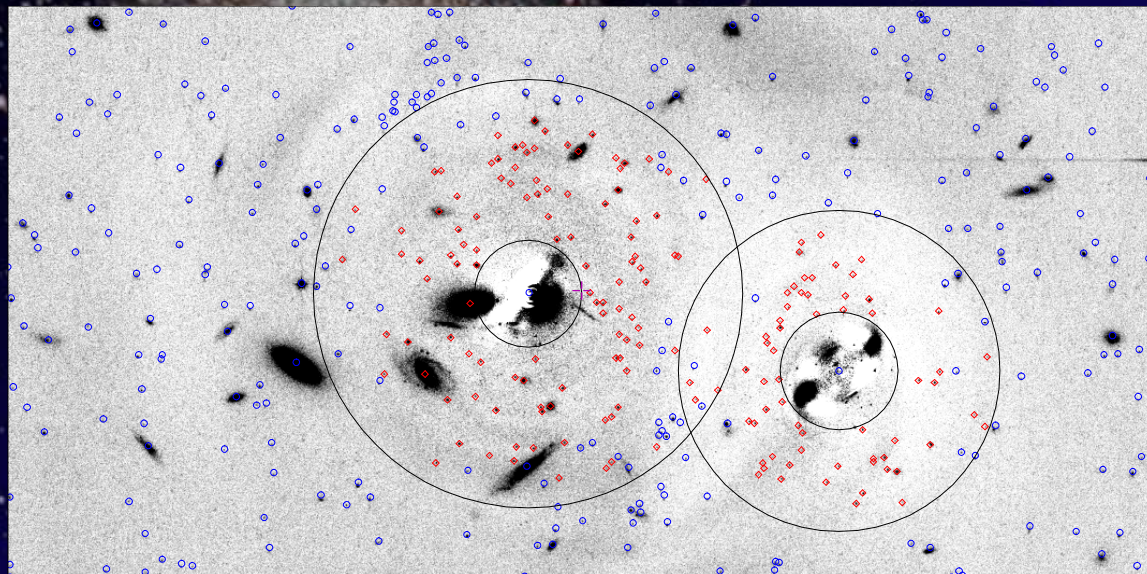
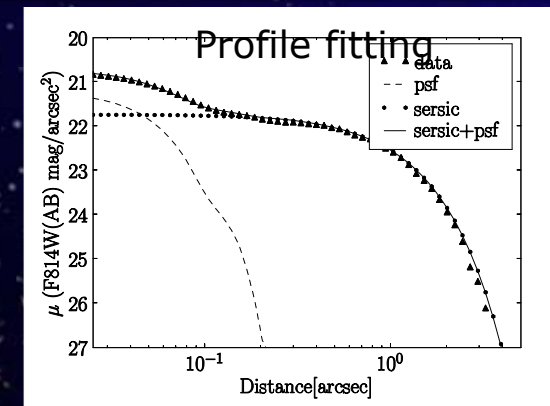




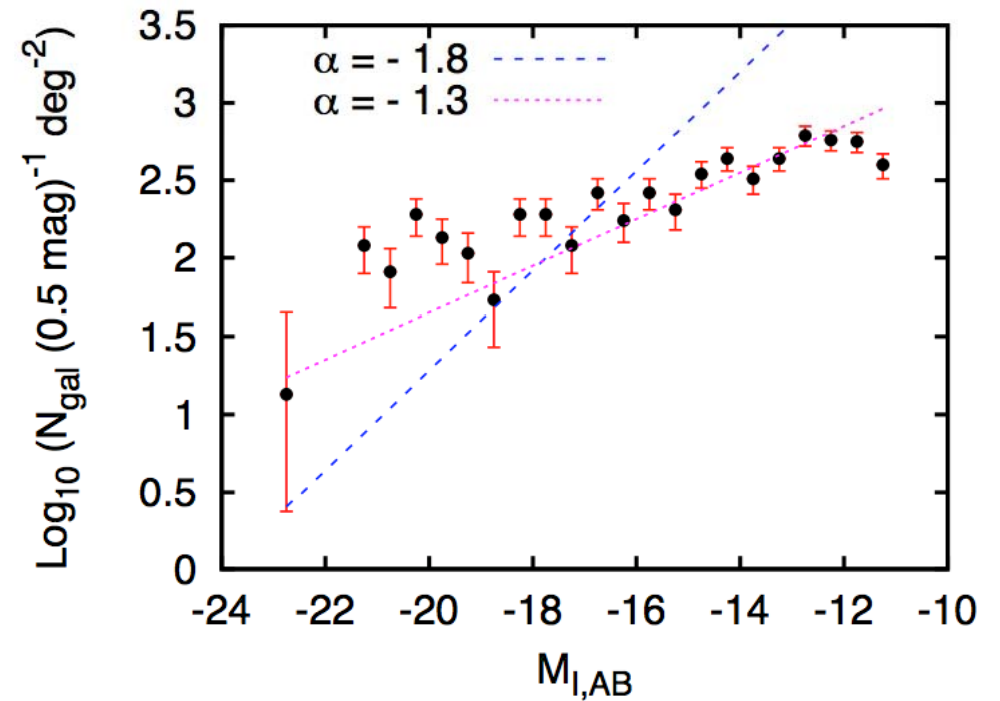
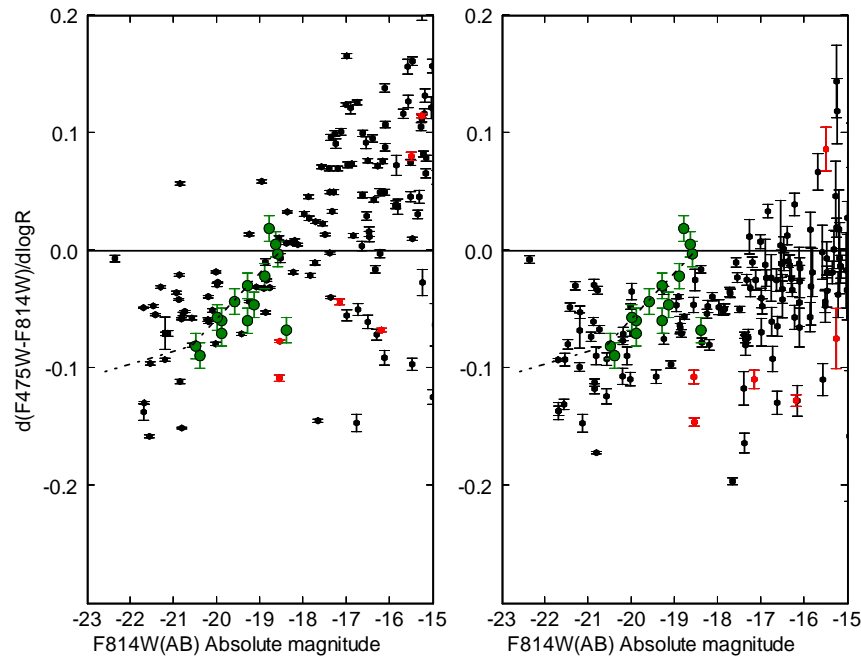


# ACS@HST Coma LS

- AstroWise enables COMA LS team
- 10,000 objects via galphot/ galfit parallel processing
  1. Photometric catalog full survey 70,000
  2. Structural parameters catalog full survey



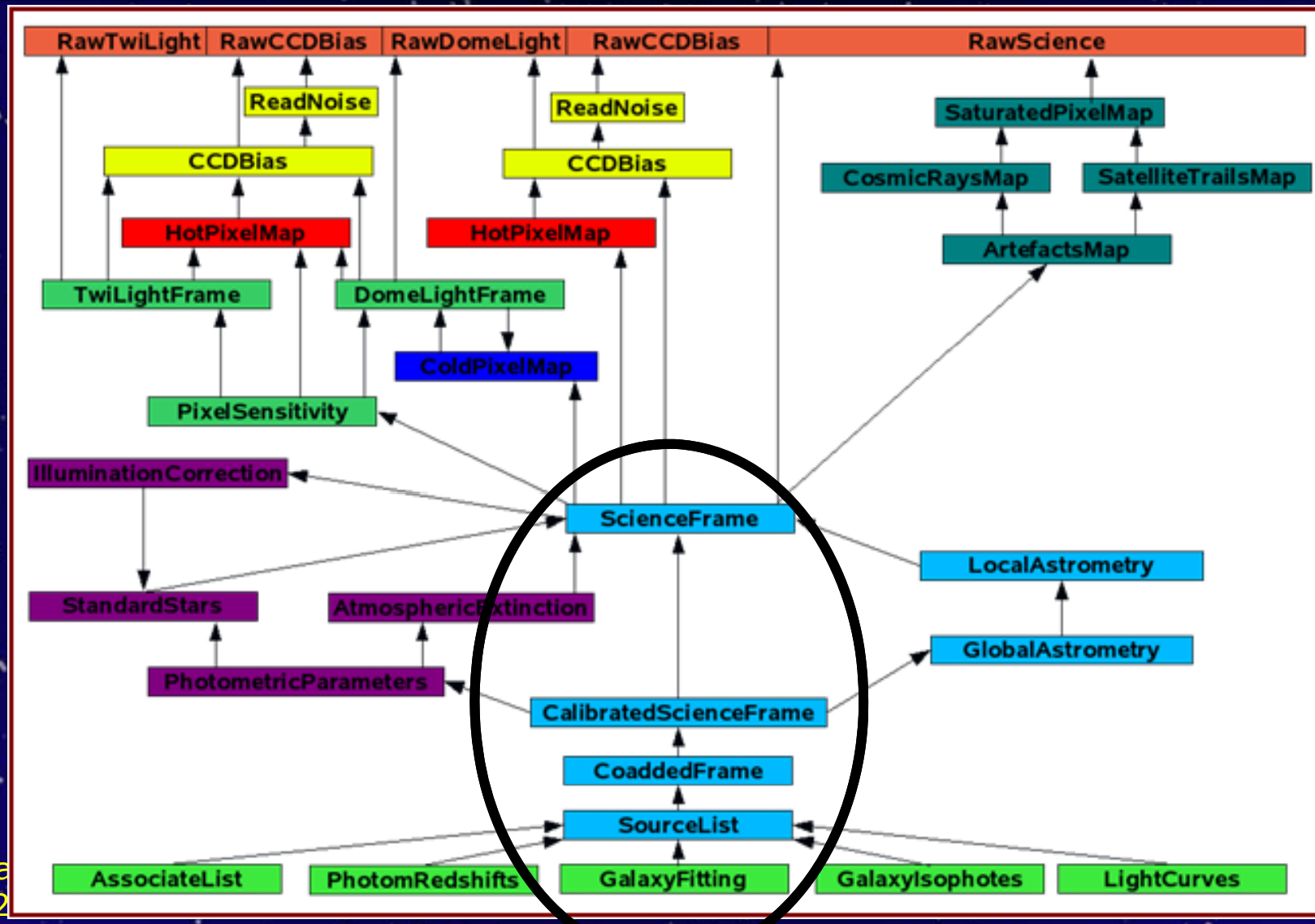
Carter et al. 2008; Hammer et al, 2010 submitted; Hoyos, et al., 2010, in prep



## PhD Den Brok 2010

- Extreme faint-end luminosity function Slope  $\alpha \approx -1.2$
- Much shallower than Cold dark Matter prediction  $\alpha \approx -1.8$
- clean radial colour gradients are much smaller
- Hint of extensive Intracluster globular cluster population

# TARGET diagram



# N params N data back to basics

- Joins – links
- ++ Inheritance – dependencies
- Everything in cs is addresses  
memory, ASCII, namespaces,  
registry
- Optimize , organize, index
- management

# Target processing: ++ the make metaphor

```
awe> targethot=HotPixelMap.get(date='2003-02-14', chip='A5382')
```

The processing chain is

ReadNoise <-- Bias <-- HotPixels

```
➤ > class HotPixelMap(ProcessTarget):
➤ > > def self.make()

➤ > class ProcessTarget():
➤ > > def get(date, chip) # if not exist/up-to-date then make()
➤ > > def exist()        # does the target exist?
➤ > > def uptodate()     # is each dependency up to date?
```

# Persistency dbobject

```
Class DBMeta                                # python<->db
    def __new__    # makes any derived Class persistent
    def __call__  # instantiate persistent object - attributes

Class DBObject:
    __metaclass__ = DBMeta
    object_id = persistent('The object identifier', oidtype) #unique

# make it
    example = DBObject()
    example.commit()

# get it
    oid = example.object_id
    result = DBObject(object_id = oid)
```

# Persistency dataobject

```
from astro.database.DBMain import DBObject, persistent
class DataObject(DBObject):
    filename = persistent('File part of this object',
str, "")
```

```
example = DataObject(pathname='example.txt')
example.store()
example.commit()
```

```
g = DataObject.filename.like('example*')
```

# Example 5LS

```
# Find ScienceFrames for a ccd named ccd53 and filter
```

```
Awe> q = (ReducedScienceFrame.chip.name == 'ccd53') and  
         (ReducedScienceFrame.filter == '#841')
```

```
# From the query result, get the rms of the sky in image
```

```
Awe> x = [k.imstat.stdev for k in q]
```

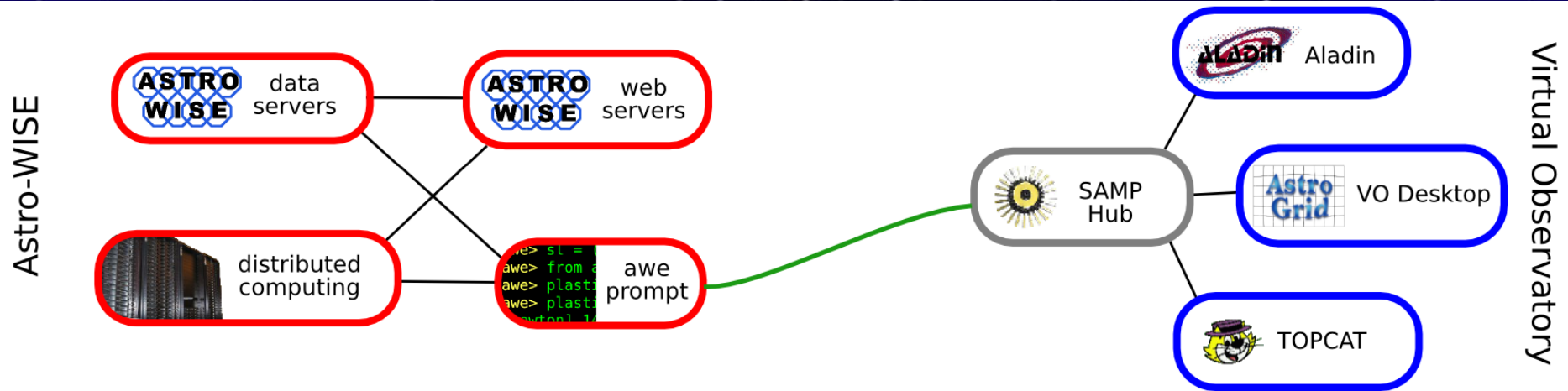
```
# get the rms of the used Masterflat
```

```
Awe> y = [k.flat.imstat.stdev for k in q]
```

```
# Make a plot
```

```
Awe> pylab.scatter(x,y)
```

# Query driven visualization



Ph D Buddelmeijer

# Query driven visualisation

**TOPCAT**

File Views Graphics Joins Windows YO Interop Help

Table List  
1: SourceList-135651

Current Table Properties  
Label: SourceList-135651  
Location: SL-135651-2df\_R\_17.votable  
Name:  
Rows: 4787  
Columns: 35  
Sort Order: ↑  
Row Subset: All  
Activation Action: (no action)  Broadcast Row

-SAMP-  
Messages: Clients:

60 / 485 M

**Aladin v6.0 \*\*\* PROTOTYPE VERSION (based on v6.000) \*\*\***

File Edit Image Catalog Overlay Tool View Interop Help

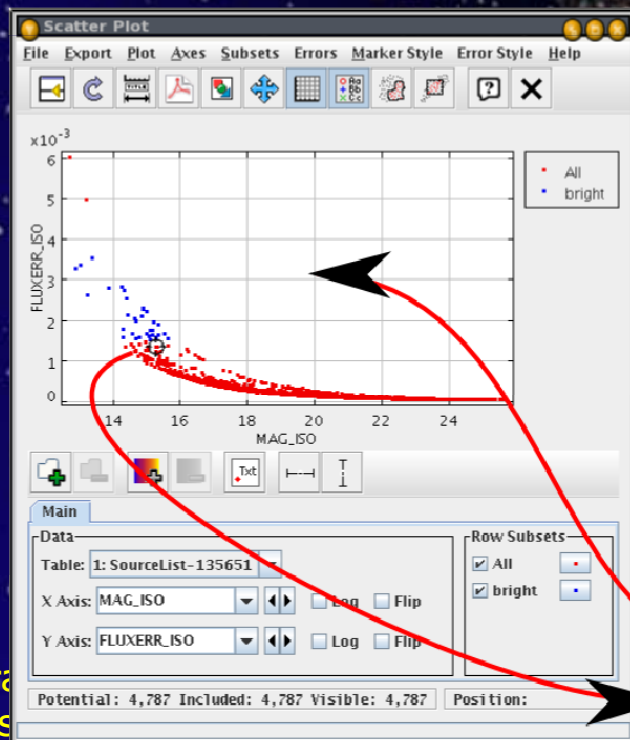
Location: ICRS

file:localhost:newton.data.users.buudelresearch1mp.Sci-6JJKEM

2 superimposed objects

BackGr	ERRA_IMAGE	SID	MAGERR_ISO
<input type="checkbox"/> 5.44149848878E-13	8678.15332031	2521	1202.79
<input type="checkbox"/> 3.09843269748E-13	9033.68652344	2550	1760.55
<input type="checkbox"/> 1.46398917295E-11	16169.5732422	2604	853.52
<input type="checkbox"/> 1.96915183812E-13	4169.22070312	2740	1987.10
<input type="checkbox"/> 2.30313733579E-13	7846.08886719	2828	1228.85

TIP: Write a script command directly into the "Command/Location" field 33 sel / 4787 src 112Mb



**SAMP Control**

File Connect Help

Sent Messages  
Received Messages  
Clients

Aladin   
Hub   
topcat    
Astro-WISE

```

Imp
kag
Initializing Distributed Pr
Current profile:
- username : AWHBUDELMEIJER
- database : db.astro.rug.astro-wise.org
- project : ALL
- current privileges : 1 (MyDB)

awe> sl = (SourceList.SLID == 135651)[0]
awe> from astro.services.samp.Samp import Samp
awe> samp = Samp()
awe> samp.broadcast(sl.frame)
awe> samp.broadcast(sl)
awe> s.highlighted(sl)
2474
awe>
    
```

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# Extreme data lineage

	RawFrame	ReducedFrame	RegriddedFrame	CoaddedRegriddedFrame	BiasFrame	ColdPixelMap	MasterFlatFrame	FringeFrame	HotPixelMap	IlluminationCorrecti
SLID=4147 SID=0 RA=11.3289 DEC=- 29.3984 X=1765 Y=84										
SLID=136151 SID=27 RA=9.5151 DEC=- 28.9031 X=883 Y=45								None		
SLID=136151 SID=29 RA=9.6949 DEC=- 28.9023 X=538 Y=126								None		
SLID=136151 SID=28 RA=9.8784 DEC=- 28.9041 X=247 Y=96								None		
SLID=4147 SID=40 RA=11.4650 DEC=- 29.3785 X=284 Y=187										

Comp. science journal PhD Mwebaze

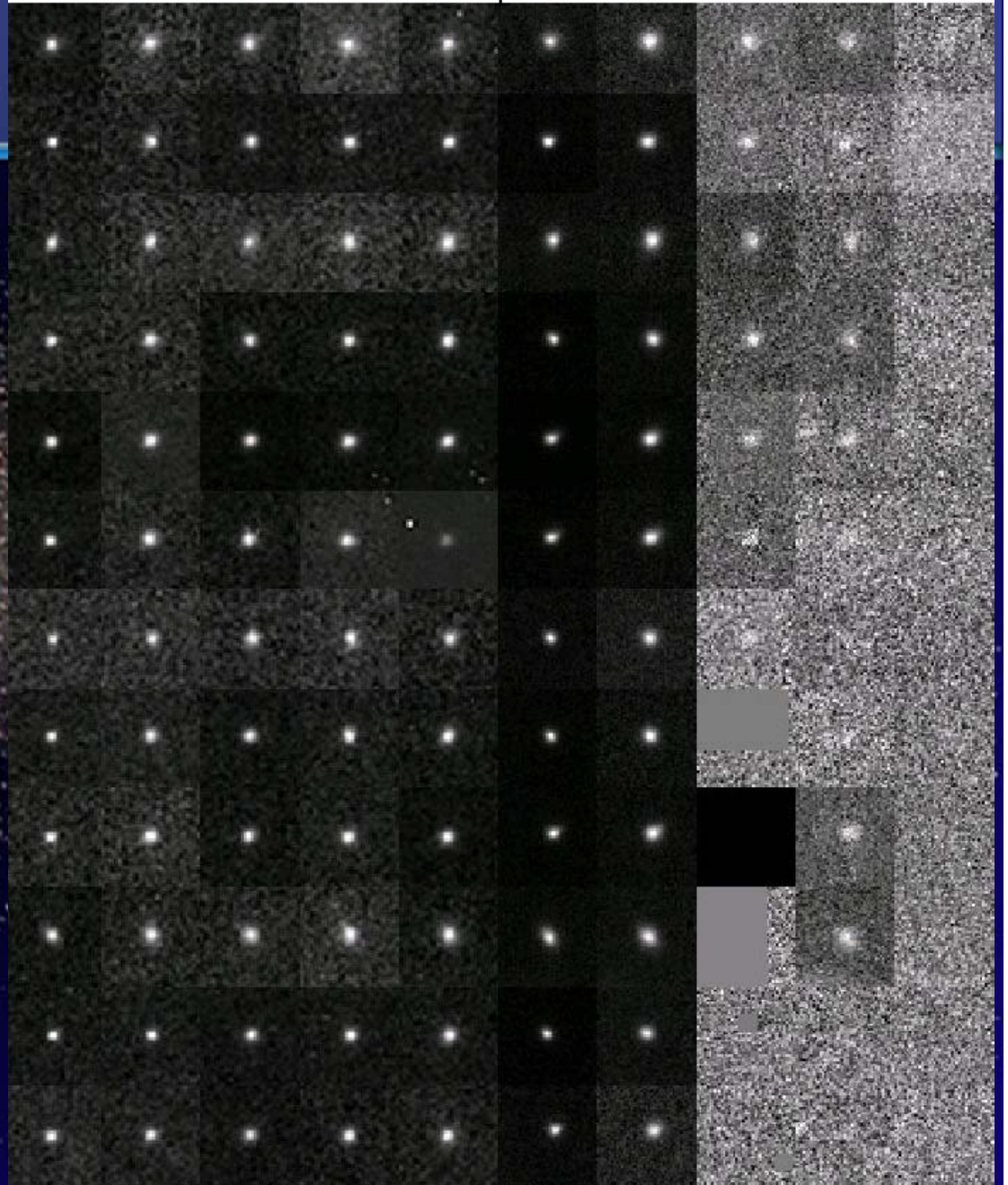
Interacademic  
Course2010

# Drop-outs

PhD Bout

Interacademic Course2010

<-- VIRCAM@VISTA --> | <-- MegaCam@CFHT -->  
K H J Y Z | z i r g u



# Astro-WISE information system – fully datacentric

All data beyond pixel data is Metadata

all pixel data  $\leftrightarrow$  data servers

all Metadata  $\leftrightarrow$  database

compute clusters / GRIDs all I/O to db

- all components scalable
- all components EU distributed

# time

Everything changes in time

- Physical changes
- Our inside in modeling
- Methods, code, bugs

### Quality of REGRIDDEDFRAME:

Sci-EVAL ENTYN-WF1---#842-ccd50-Regr---Sci-54566\_3131050-f56144d965b5e765b40bdec3d685fe595215d52b.fits

ASTROWISE DBView GetR Process

no previous comments

ObjectName: [awqvarloss](#) project: WFI@2.2m

is\_valid = 1: valid

#### Processing Details

creation_date	2008-04-10 07:31:02
is_valid	1
quality_flags	0
Privileges	4

#### Image Statistics Details

mean	+9.327e+01
median	+8.887e+01
stdev	+2.945e+03
min	-6.036e+06
max	+3.727e+06

#### Local Astrometry Details

creation_date	2008-04-10 07:30:42
is_valid	1
quality_flags	0
RMS	0.252
SEEING	0.856
NREF	317
SIG_DRA	0.209
SIG_DDEC	0.178
MEAN_DRA	-0.001
MEAN_DDEC	0.002

#### Photometry Details

creation_date	2008-03-29 20:15:06
is_valid	1
quality_flags	0
zeropoint	24.759
zp_error	0.000
zp_origin	derived
num_sources	173

#### Observational Details

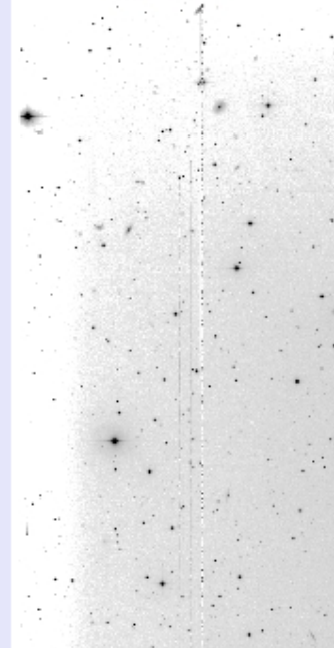
DATE_OBS	2002-03-18 03:38:21	OBSERVER	UNKNOWN
MJD_OBS	52351.1495509	EXPTIME	299.9176
OBJECT	SNW	AIRMSRT	1.246
R.A.	13:28:30.0000	AIRMSD	1.246
Dec.	-31:35:44.6308	Filter	#842
		mag_id	JohnsonB

#### Chip address of instrument WFI:



RegriddedFrame

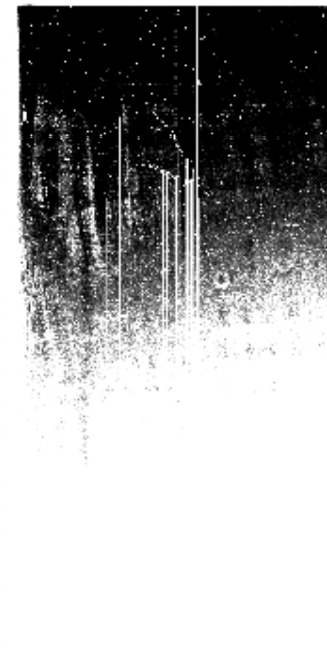
2439 X 4873 pixel  
8.13 X 16.24 arcmin



AstrometricParameters

WeightFrame

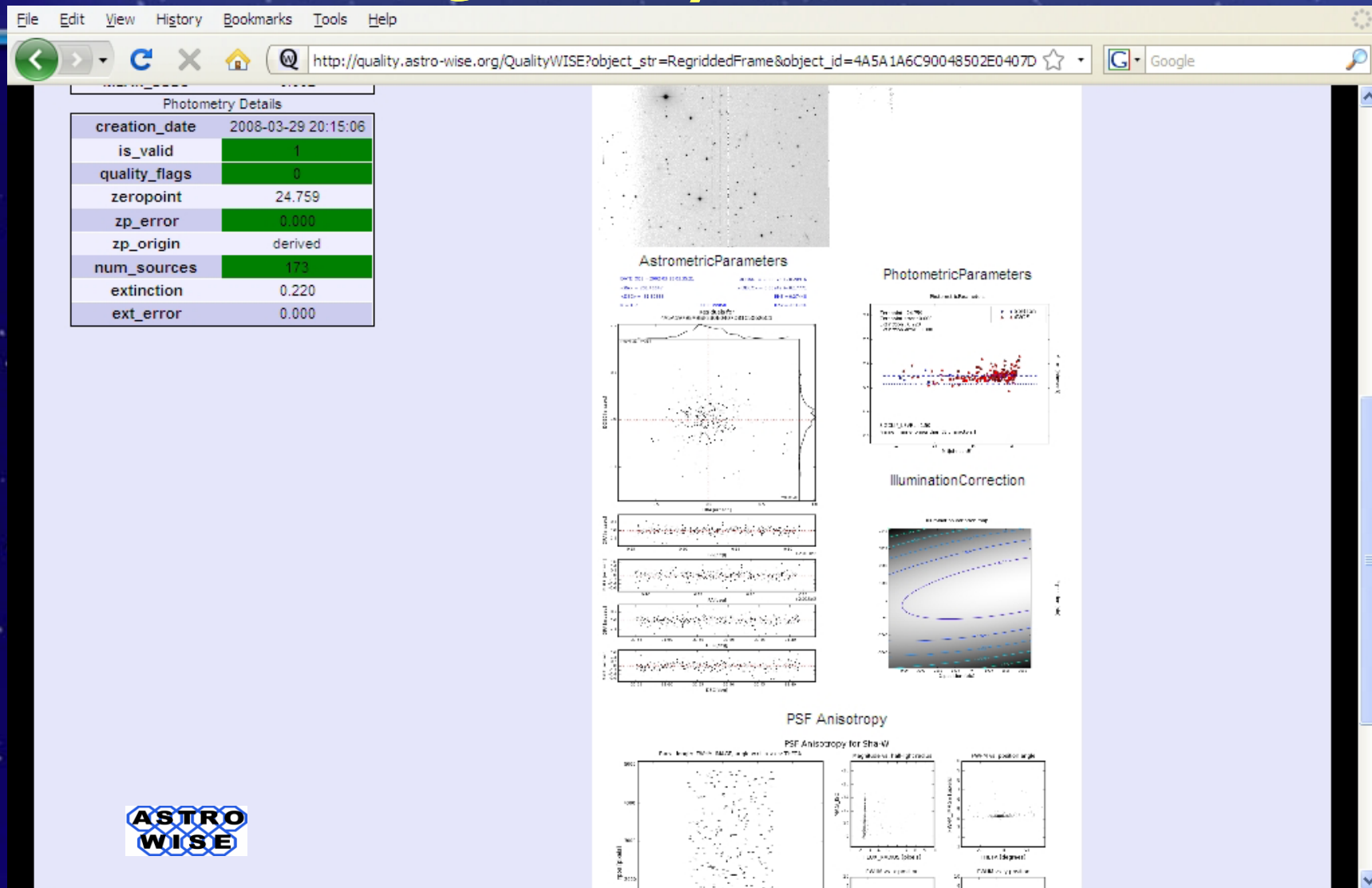
2439 X 4873 pixel  
8.13 X 16.24 arcmin



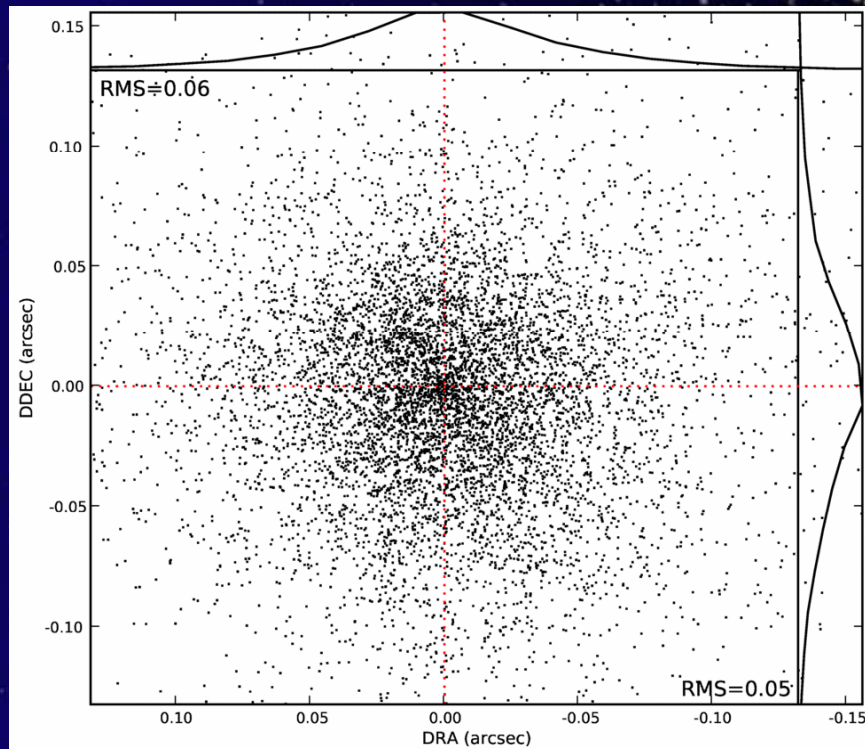
Done

# Post Doc Verdoes Kleijn

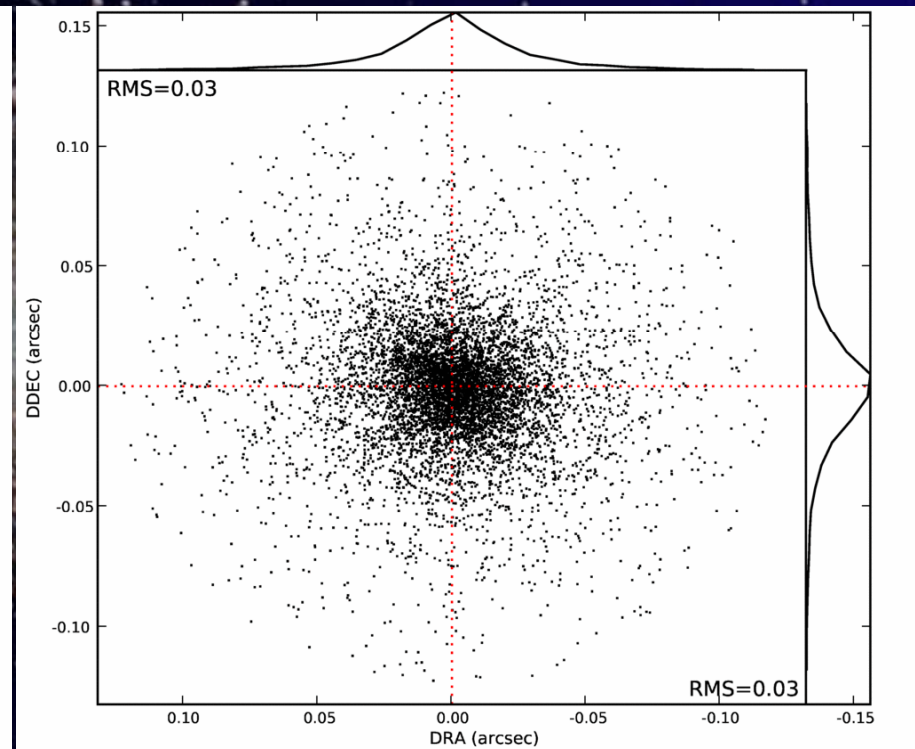
# Quality view



# Global Astrometry reproducibility



Local solution



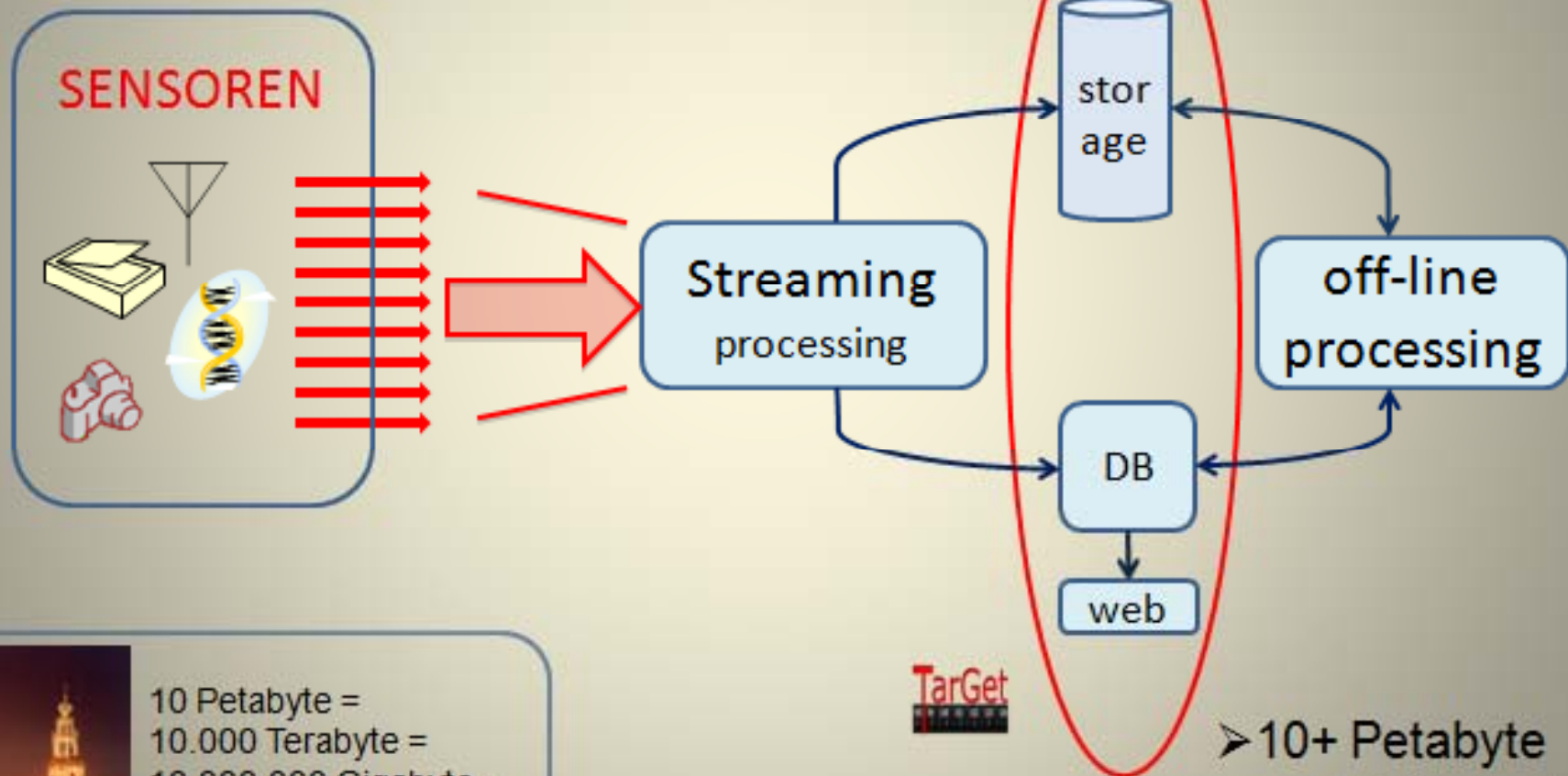
Global solution

(Internal regrid-to-regrid residuals  
of one WFI N=4 dither)

# future

- Information systems

## het probleem Sensor netwerken-data centric



10 Petabyte =  
10.000 Terabyte =  
10.000.000 Gigabyte =  
30 Martini torens DVD's

Date

- 10+ Petabyte
- gedistribueerd



## OmegaCAM

16k x 16k

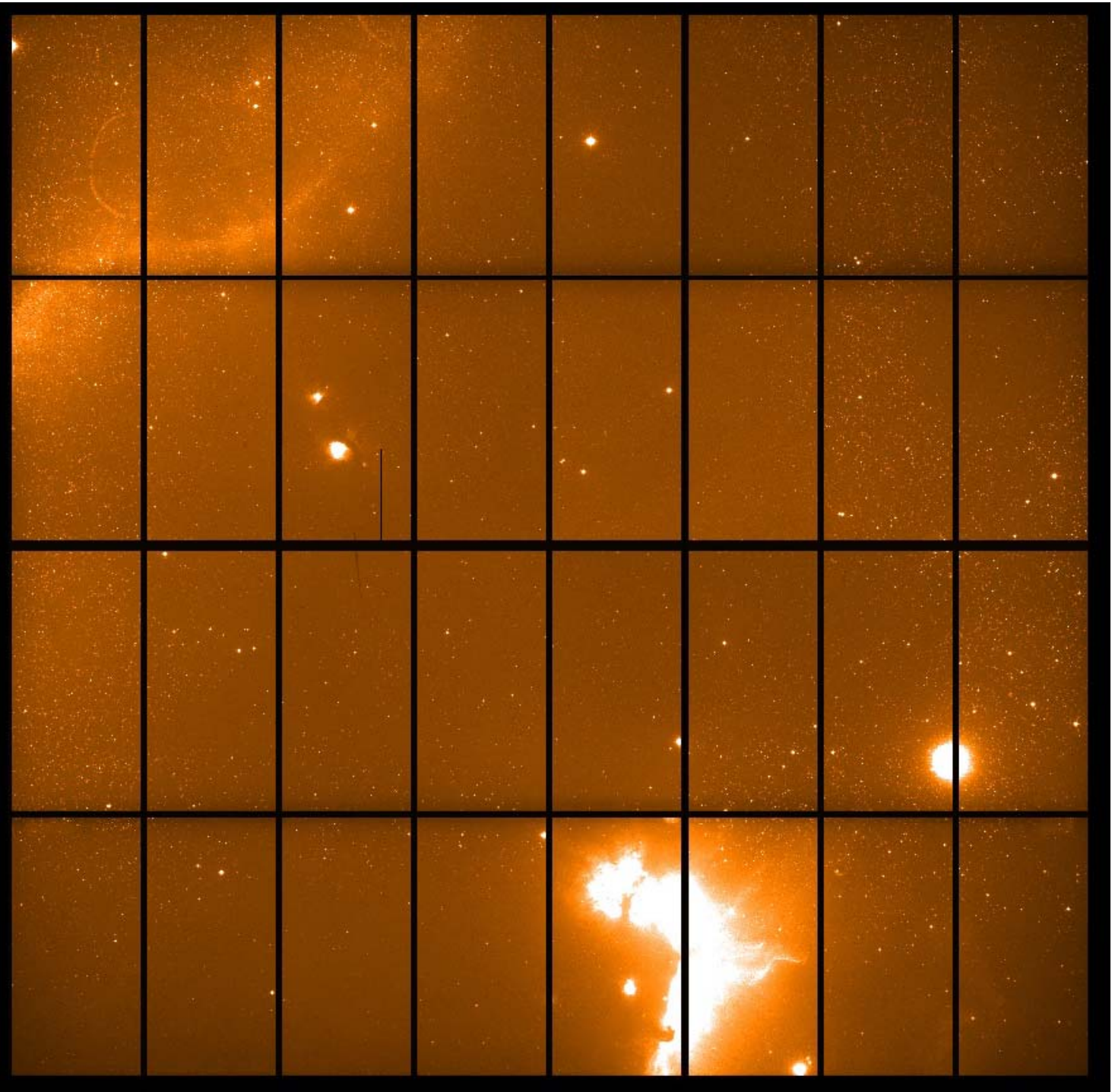
256Mpix

0.2" pixels

1 x 1 degree

2.6 m VST

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# KIDS vs. SDSS, CFHTLS

SDSS

CFHTLS

6 x area

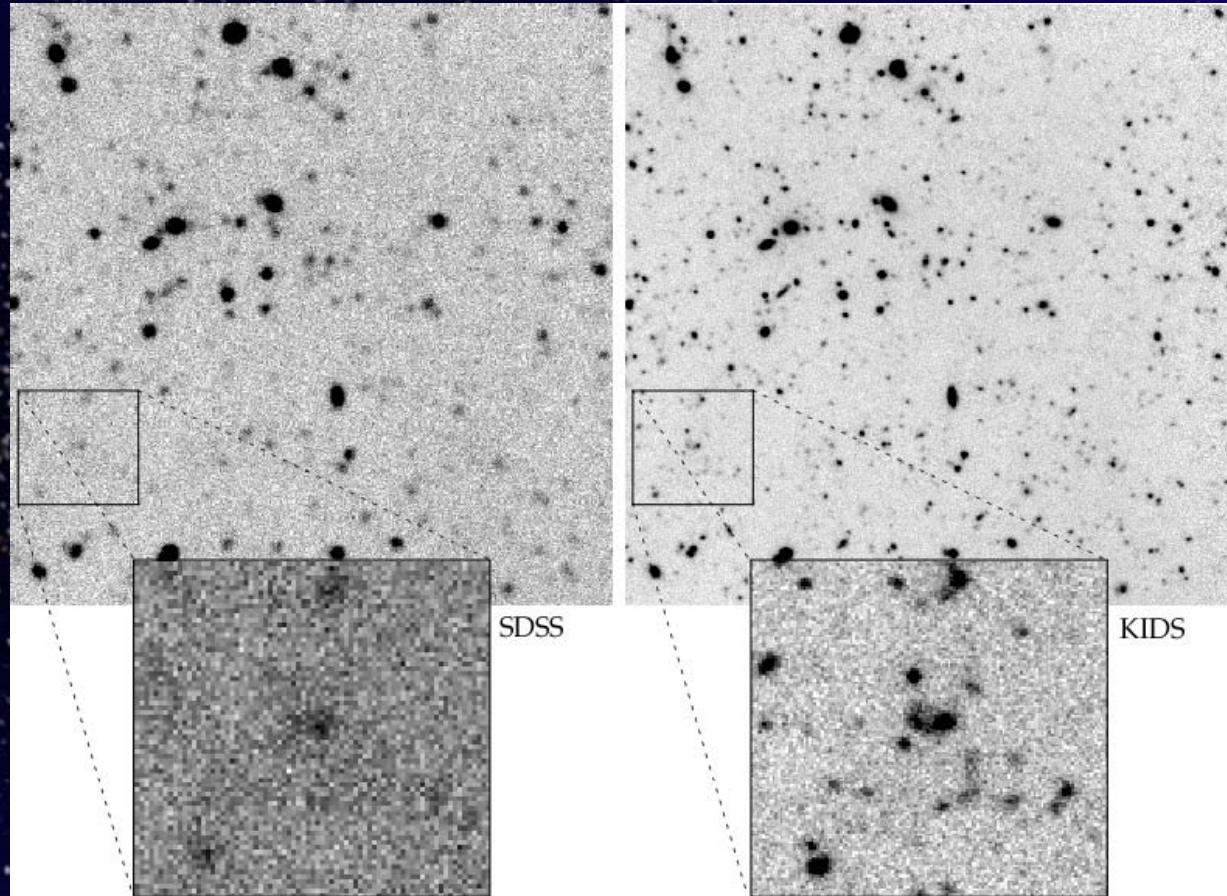
1/9th area

2 mag shallower

1 mag deeper

2x worse seeing

~same



(M.Neeser)



# KIDS – VIKING

250 nights

440 nights

VIKING

VISTA

4m telescope

0.6 sq.deg.  
InfraRed camera

16 2kx2k  
detectors

0.35" pixels

KIDS

VST

2.6m telescope

1 sq.deg. optical  
camera OmegaCAM

32 2kx4k  
detectors

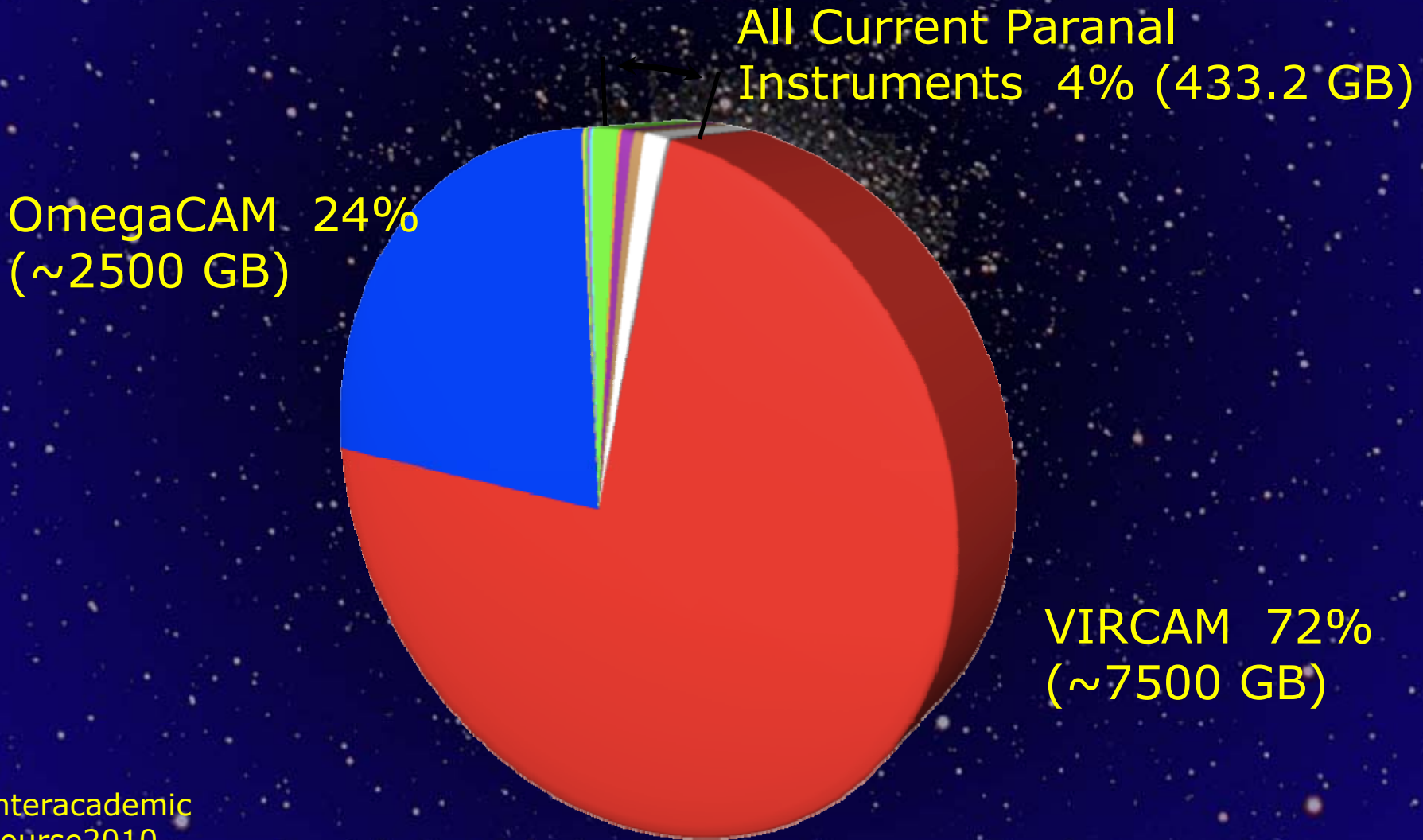
0.21" pixels



# Paranal Monthly Data Rates 2007 statistics



- |          |            |           |         |         |
|----------|------------|-----------|---------|---------|
| ● FORSI  | ● FORS2    | ● GIRAFFE | ● ISAAC | ● MIDI  |
| ● NACO   | ● SINFONI  | ● LIVES   | ● VIMOS | ● VISIR |
| ● VIRCAM | ● OmegaCAM |           |         |         |





# Astro-WISE – LOFAR LTA



IBM- Blue Gene/P

5 Petabyte/yr

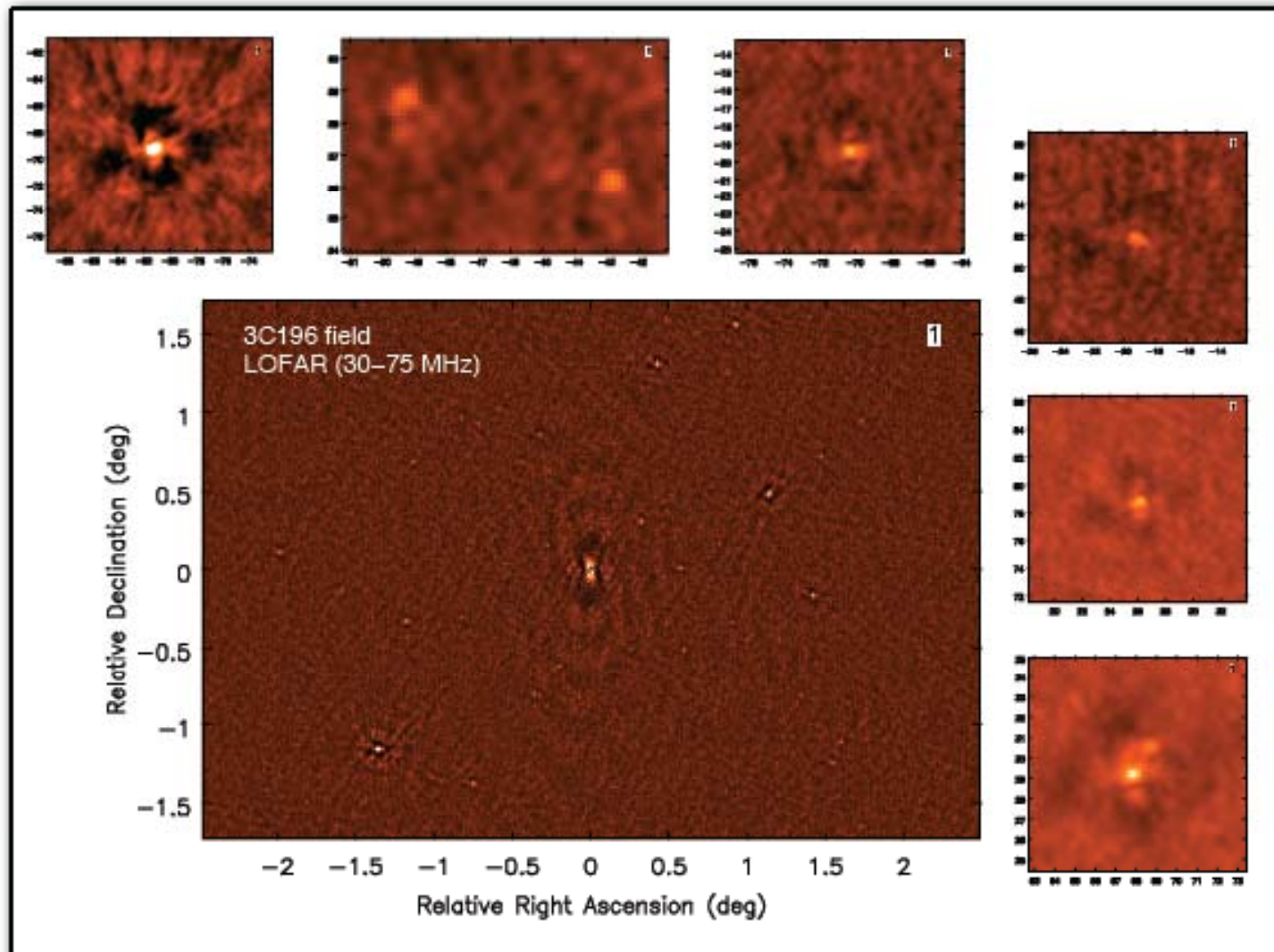
- EoR

- Survey

# LOFAR Superterp



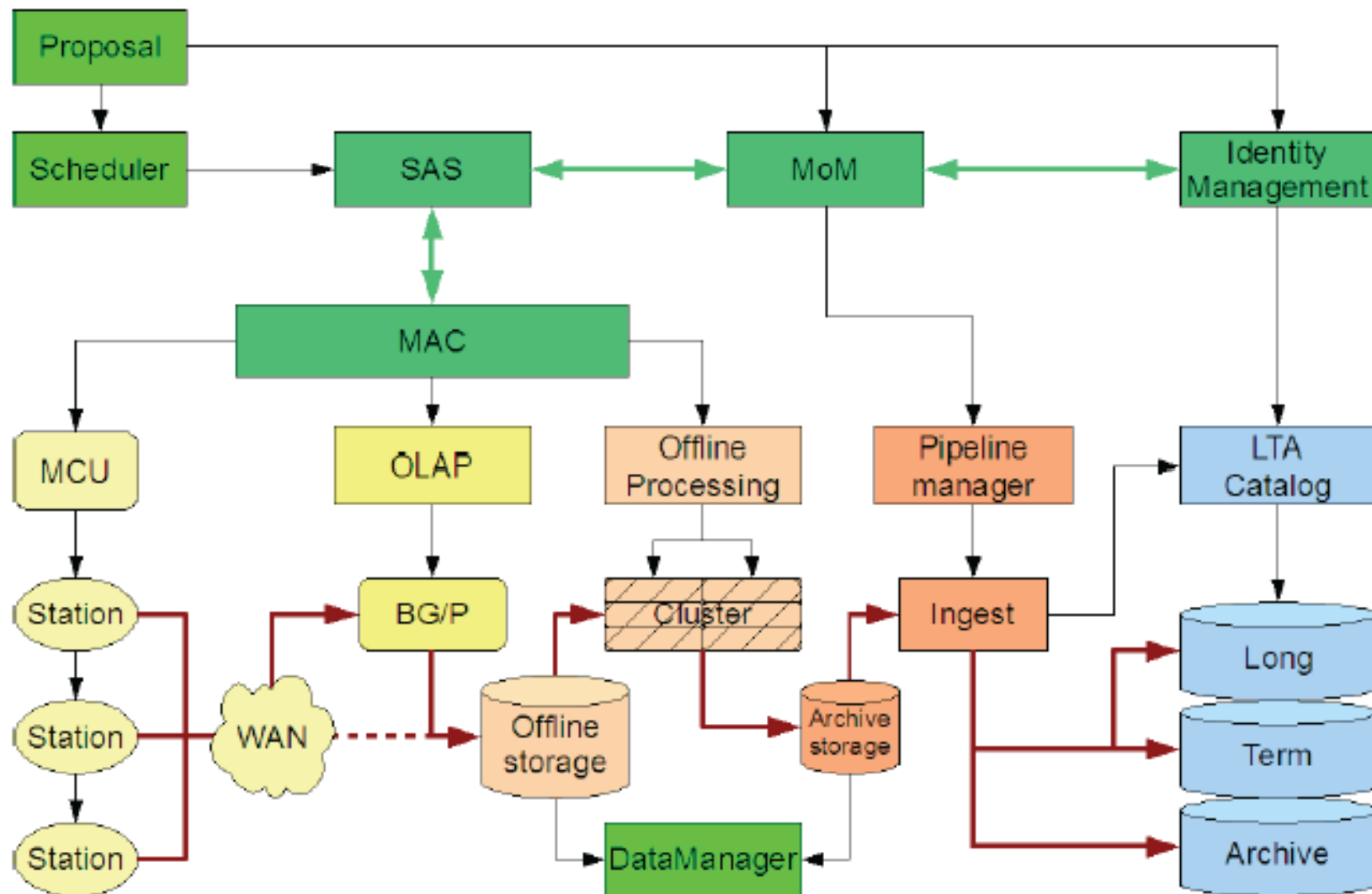
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*Fully automated processing, 72 sub-bands, ~4 hrs*

*(courtesy J. McKean)*

*Included DPPP, additional flagging pass, BBS, solution flagging, imaging*



*(courtesy H. Holtes)*

# Lofar Long Term Archive

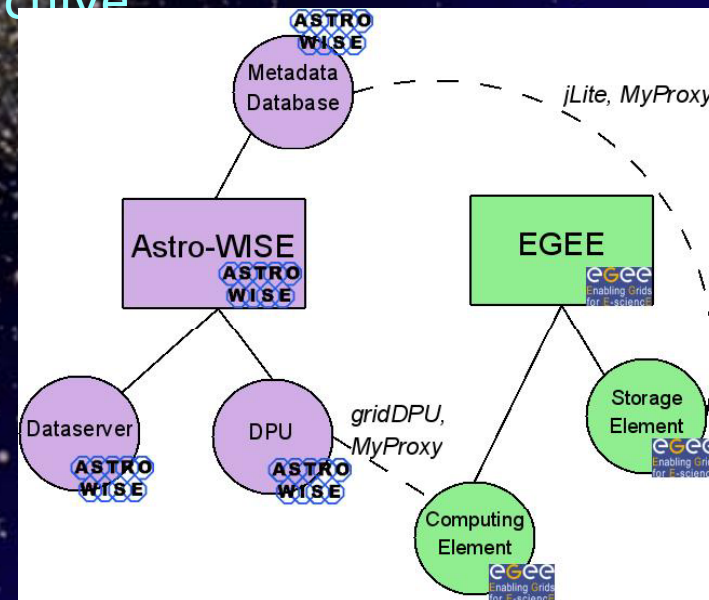
- Astro-WISE -> EGEE full integration
  - Astro-WISE stores on EGEE
  - launches Astro-WISE applications on EGEE
  - Incorporated 3<sup>rd</sup> party software gLite, MyProxy
- Designed+implemented LOFAR LTA
  - Ready for >10 Petabyte
  - Next step: integration pipeline+archive

SpringerLink - Journal Article | Lofar Archive

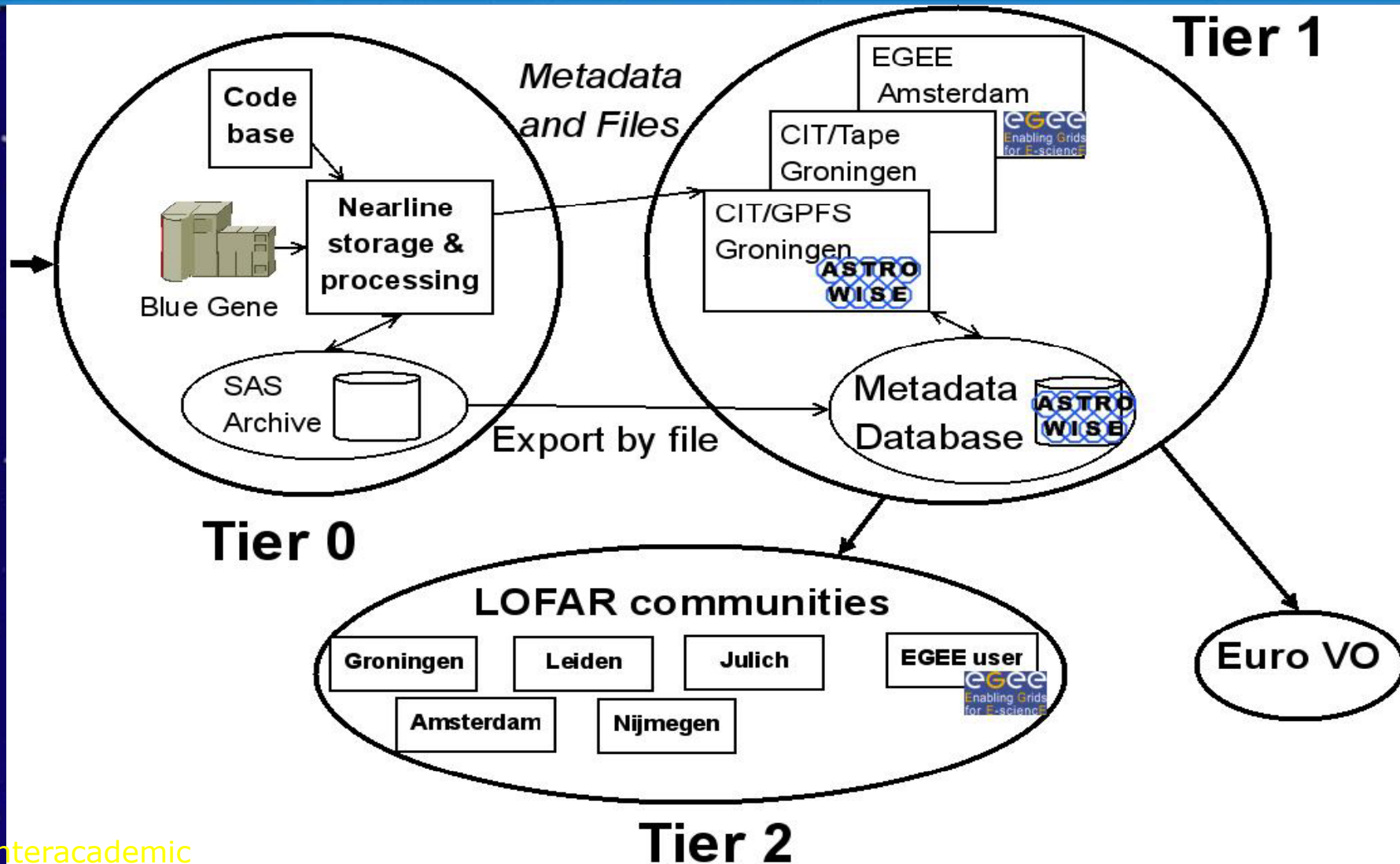
Home | Help | project test-lofar | user AWANONTHOUS | Search | Show Latest | Overview

Latest 9 UVMeasurement sets

ID	Creator	Privileges	Project	antennaSelection	beamNumber	clock	creationDate	instrumentFilter	instrumentType	IntegrationInterval	measurementDescription	measurementIdentifier
1	AWTTERO	2	test-lofar	LBA Sparse Even	0	200.0	2010-09-15 16:09:41	10-90 MHz	Interferometer	1.0	Single Beam vanCottin L50921	1009
2	AWTTERO	2	test-lofar	LBA Sparse Even	0	200.0	2010-03-12 16:57:10	10-90 MHz	Interferometer	1.0	Single subband beam	31
3	AWTTERO	2	test-lofar	LBA Both	0	160.0	2010-02-12 15:50:19	170-230 MHz	Interferometer	1.0	New Measurement because the old tests have disappeared	79
4	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-06 17:34:12	10-60 MHz	Interferometer	4.0	None	450
5	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-26 16:48:39	10-60 MHz	Interferometer	5.0	None	447
6	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-26 16:39:13	10-60 MHz	Interferometer	4.0	None	443
7	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-26 15:32:33	10-60 MHz	Interferometer	4.0	None	438
8	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-26 15:05:07	10-60 MHz	Interferometer	4.0	None	435
9	AWTTERO	2	test-lofar	LBA Outer	0	160.0	2009-11-06 14:54:53	10-60 MHz	Interferometer	4.0	None	418



# Lofar - design



	60 MHz	150 MHz
Total # fields (2 pi steradian)	609	3346
Total observing time (100% eff., using 4 beams)	456.75 hr	836.5 hr
Total # sources	Tbd	Tbd
Total uv data size	466 Tbyte	2.27 Pbyte
Total post DP <sup>3</sup> uv data size	~ 2.9 Tbyte	~ 7.1 Tbyte
Total image data size	Tbd	Tbd

*(R. Nijboer, March 2009)*

## Current archive resources

- Currently using ~50 TBytes
- 300 TBytes available @ SARA in Amsterdam
- 60 TBytes @ Astrowise server in Groningen
- 500 TBytes (to 1 PBytes) via TARGET in Groningen
- 1 PBytes tape storage @ Jülich
- Additional 2-3 PBytes to be added @ BiGGRID in Amsterdam



- Lofar 2010      [www.lofar.org](http://www.lofar.org)
- ALMA 2013      [www.eso.org/sci/facilities/alma](http://www.eso.org/sci/facilities/alma)
- EUCLID 2018
- LSST      30Tb/night      [www.lsst.org](http://www.lsst.org)
- SKA 2022      [www.skatelescope.org](http://www.skatelescope.org)

- CLOUD – computing
- Web 2.0
- Crowd sourcing
  - Galaxy zoo
  - annotation

# AstroInformatics

- [www.astroinformatics2010.org](http://www.astroinformatics2010.org)

## AstroInformatics2010

Pasadena, June 16 - 19, 2010

Sunday May 09 2010 13:06 PST

38 days until the conference

### AstroInformatics 2010

[Introduction](#)

[42 Participants](#)

[Conference Agenda](#)

[Invited Speakers](#)

[Proceedings Info](#)

[Registration](#)

[Pay by CreditCard](#)

### More Information

[Accommodation](#)

[Getting to Caltech](#)

[Local Information](#)

[Conference Poster](#)

[Conference Photos](#)

[Organizing Committee](#)

[Contact the LOC](#)

## AstroInformatics 2010

Caltech, June 16-19 (Wed-Sat), 2010  
Cahill Center for Astronomy & Astrophysics, Caltech, Pasadena, CA, USA

### Conference motivation and goals:

Science is being fundamentally transformed and empowered by computation and information technology, and it turn, it stimulates further technological developments. Astronomy, like many other fields, is becoming exponentially data-rich, and the tasks of data management, data exploration, and knowledge discovery become central to our research enterprise, bringing along many technical and methodological challenges. Information technology also provides the stage where we collaborate and interact, and publish, preserve, and disseminate knowledge.

The general philosophy behind the conference is to be future-oriented, and essentially define the emerging discipline of AstroInformatics. One way to think about it is, what is next for, and beyond the VO, scientifically, technologically, institutionally? AstroInformatics can be viewed as a broader intellectual, organizational, and funding environment, within which VOs serve as particular institutions and provide some fundamental functionalities and infrastructure.

There will be a modest number of invited review talks to serve as a basis for the discussion, and a lot of discussion, some of it led by panels. Contributed papers will be accomodated as posters.

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# ADASS

## Astronomical Data Analysis Software & Systems

The **Astronomical Data Analysis Software and Systems (ADASS)** conference is held each year at a different hosting astronomical institution. The conference provides a forum for scientists and programmers concerned with algorithms, software and software systems employed in the acquisition, reduction, analysis, and dissemination of astronomical data. An important element of the program is to foster communication between developers and users with a range of expertise in the production and use of software and systems. The program consists of invited talks, contributed oral and display papers, tutorials, user group meetings and special interest group meetings (called BOFs).

## ADASS XX



## ADASS XX

- The 20th annual ADASS conference will be in Boston, Massachusetts, USA, 7-11 November 2010.

# EURO-VO IVAO

## Welcome to the Euro-VO AIDA Astronomical Infrastructure for Data Access

This is the web-based collaboration area of the Euro-VO Astronomical Infrastructure for Data Access project. This project is supported by EU in the framework of the FP7 [einfrastructure Scientific Research Repositories](#) initiative (project RI2121104). It started on 1 February 2008, for a duration of 30 months.

### Highlights

- [Fifth Euro-VO Technology Forum](#): March 16 -18, 2010 in Heidelberg
- [Second Community feedback Workshop](#) : 25-28 January 2010, Strasbourg
- **The EuroVO-AIDA Project also realizes an [update of the DCA Census of European astronomical Data Centres](#).**
- EURO-VO in the ICT results: [e-Infrastructures give real boost to virtual observatories](#) - issued on 8 october 2009.
- [Fourth Euro-VO Technology Forum](#): 22 - 24 September 2009, Trieste
- [Second EuroVO-AIDA Research Initiative](#). Deadline 15 July 2009.
- [Data Centre Workshop on how to publish data in the VO](#) 22-26 June 2009, ESAC, Villafranca del Castillo
- [EuroVO-AIDA Hands-on workshop](#) : 30 March - 2 April 2009, ESO, Garching
- [Third Euro-VO Technology Forum](#) : 16- 18 March 2009, Strasbourg *Due to a strike warning on March 19, the meeting will be held March 16 -18*
- [Full Harvestable EuroVO Registry of Resources](#) *has been released (13 Mar 2009)*
- [EuroVO-AIDA Workshop MultiWavelength astronomy and the Virtual Observatory](#) : December 1-3 2008, ESAC, Villafranca del Castillo
- [Second Euro-VO Technology Forum \(with VO-TECH\)](#) : 29 September - 2 October 2008, Cambridge, UK
- [First EuroVO-AIDA Research Initiative](#). Deadline 15 June 2008.
- [First Euro-VO Technology Forum \(with VO-TECH\)](#): March 17 - 19 2008, Strasbourg

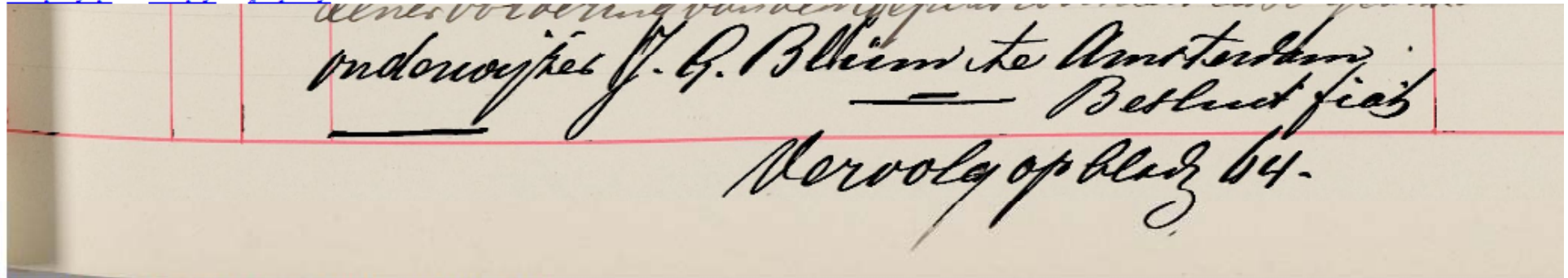


MONK



There are 182 results. Showing results 1 - 10 for amsterdam

[show paragraph](#) [show page](#) [Query image](#)



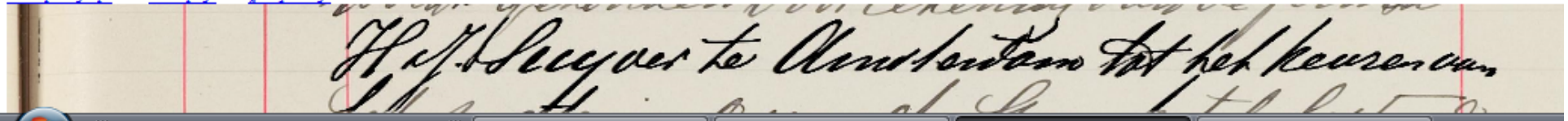
onderwijzer J.G. Bluum te Amsterdam. -- Besluit fiat -- Vervolg op blad 64.

[show paragraph](#) [show page](#) [Query image](#)



loniaal Etablissement te Amsterdam

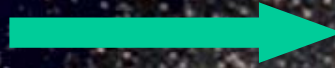
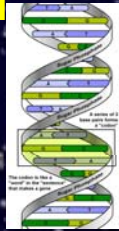
[show paragraph](#) [show page](#) [Query image](#)



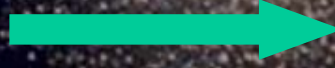
# LifeLines

## 165.000 patients 30 yr

### Molecular



### Physiology



### Care

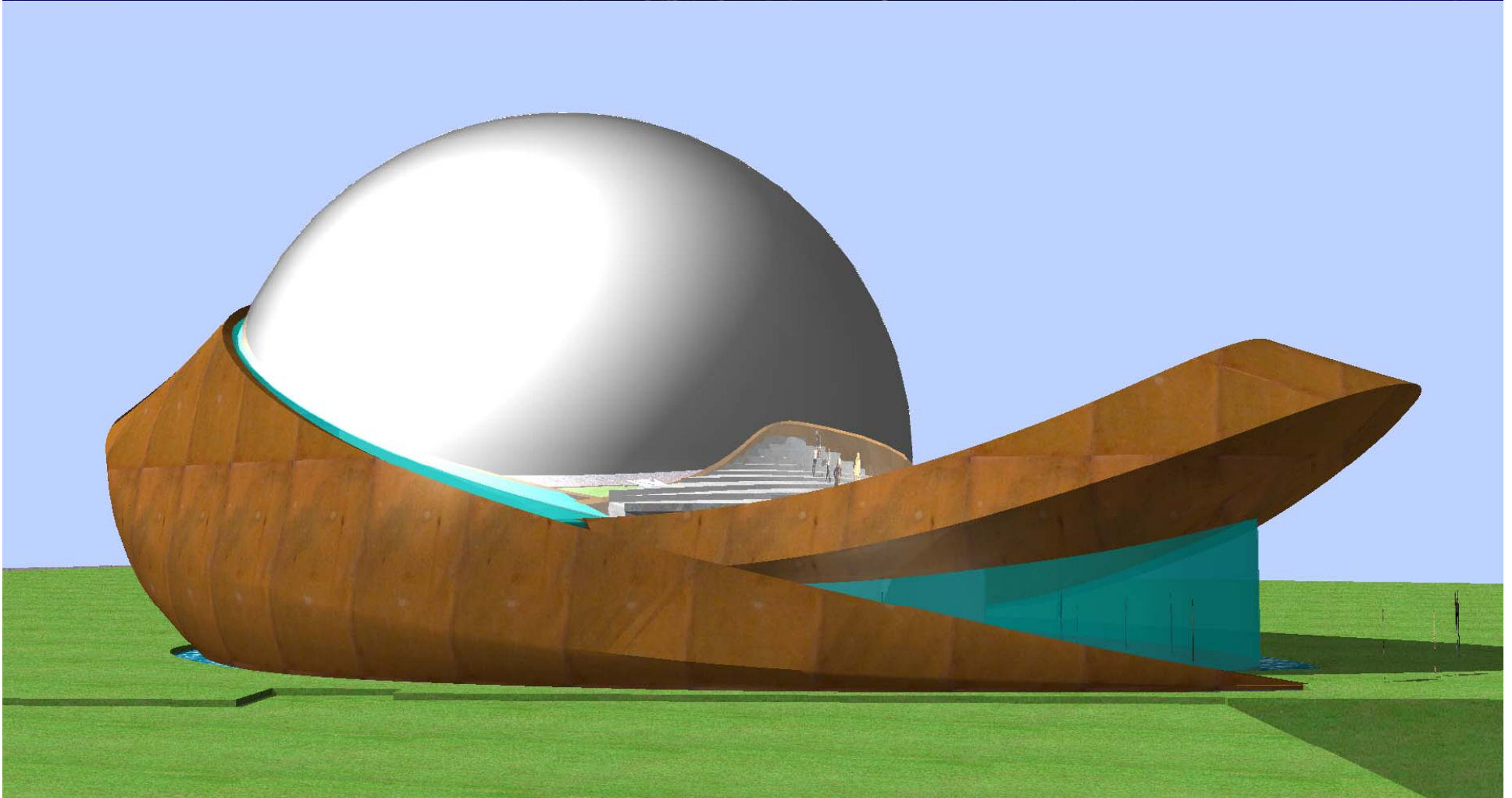


Integration

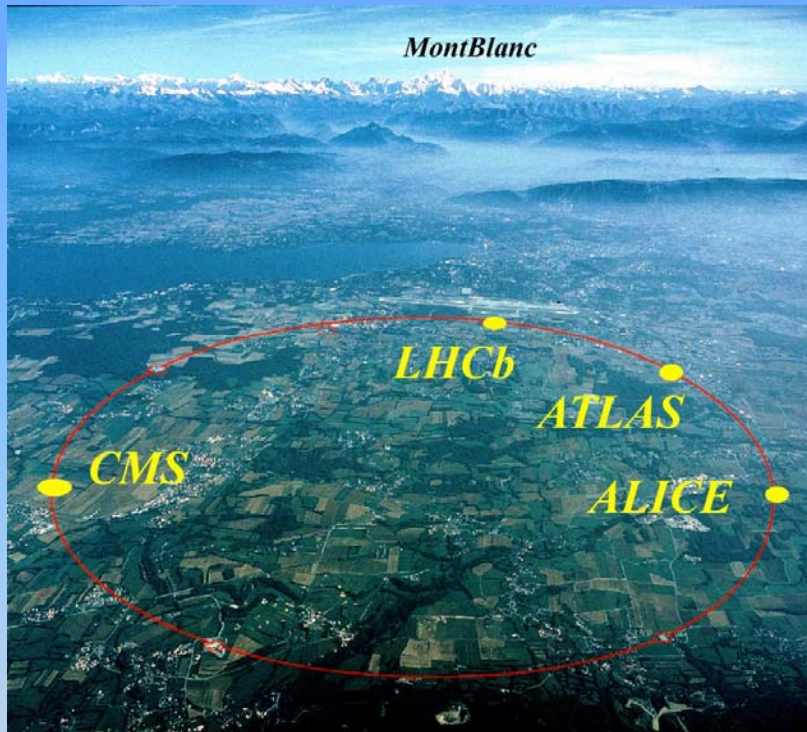
ASTRO  
WISE

# IJkdijk





# LHC & EGEE



- 15 PB data per year
- grid computing

EGEE: ~250 sites, >45000 CPU

eGee

Enabling Grids  
for E-science

Interacademic Course 2010

[www.eu-egee.org](http://www.eu-egee.org)

EGEE-III

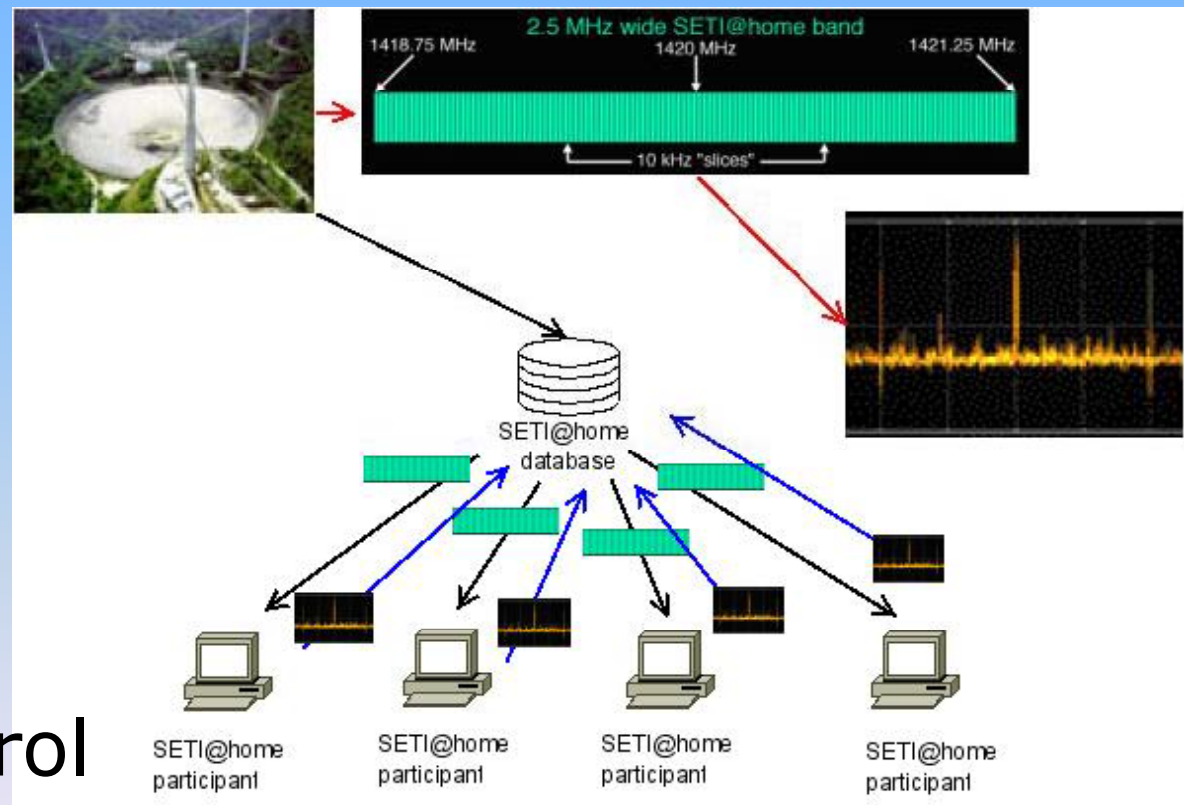


- First scientific grid computing

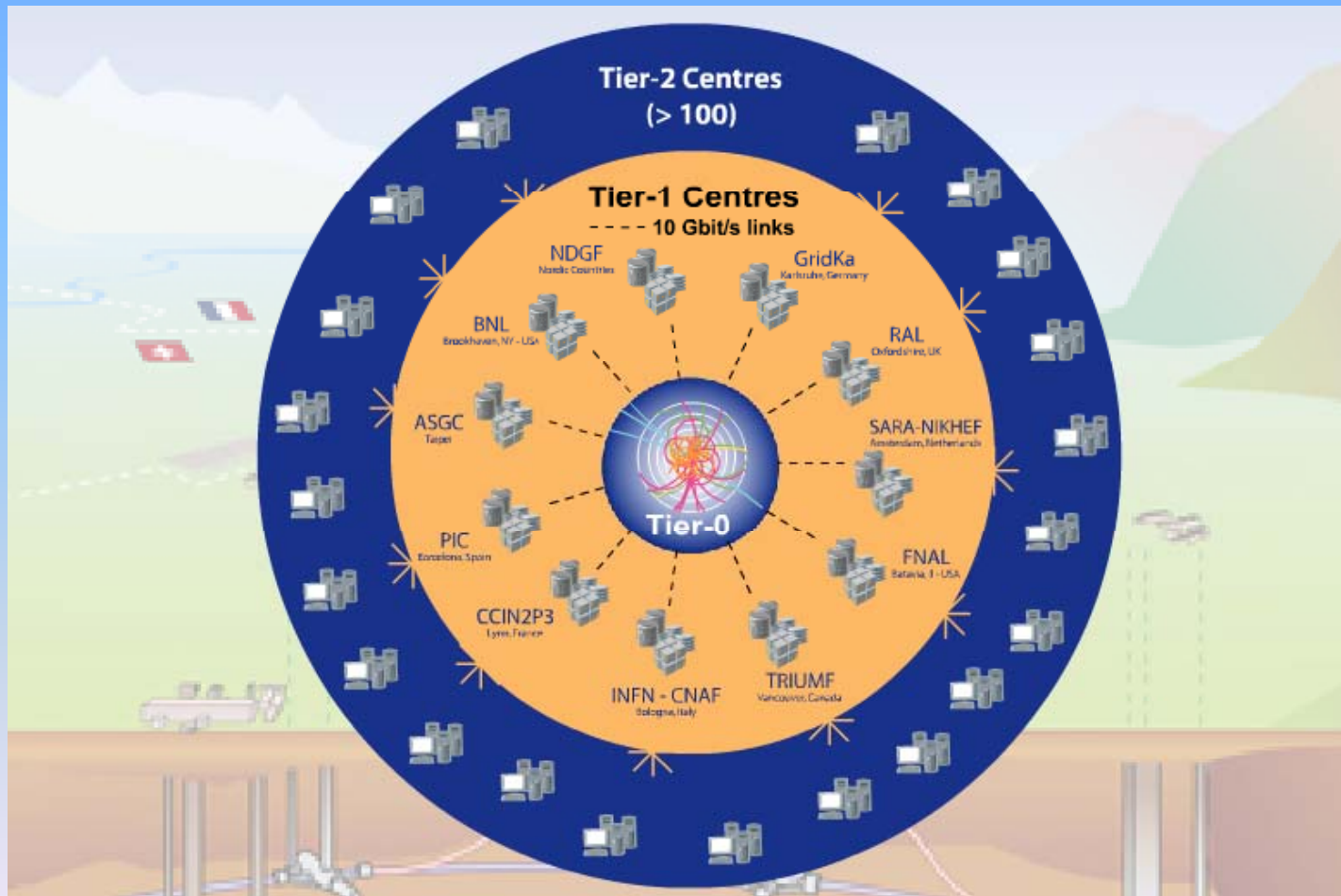
<http://setiathome.berkeley.edu/>

<http://boinc.berkeley.edu>

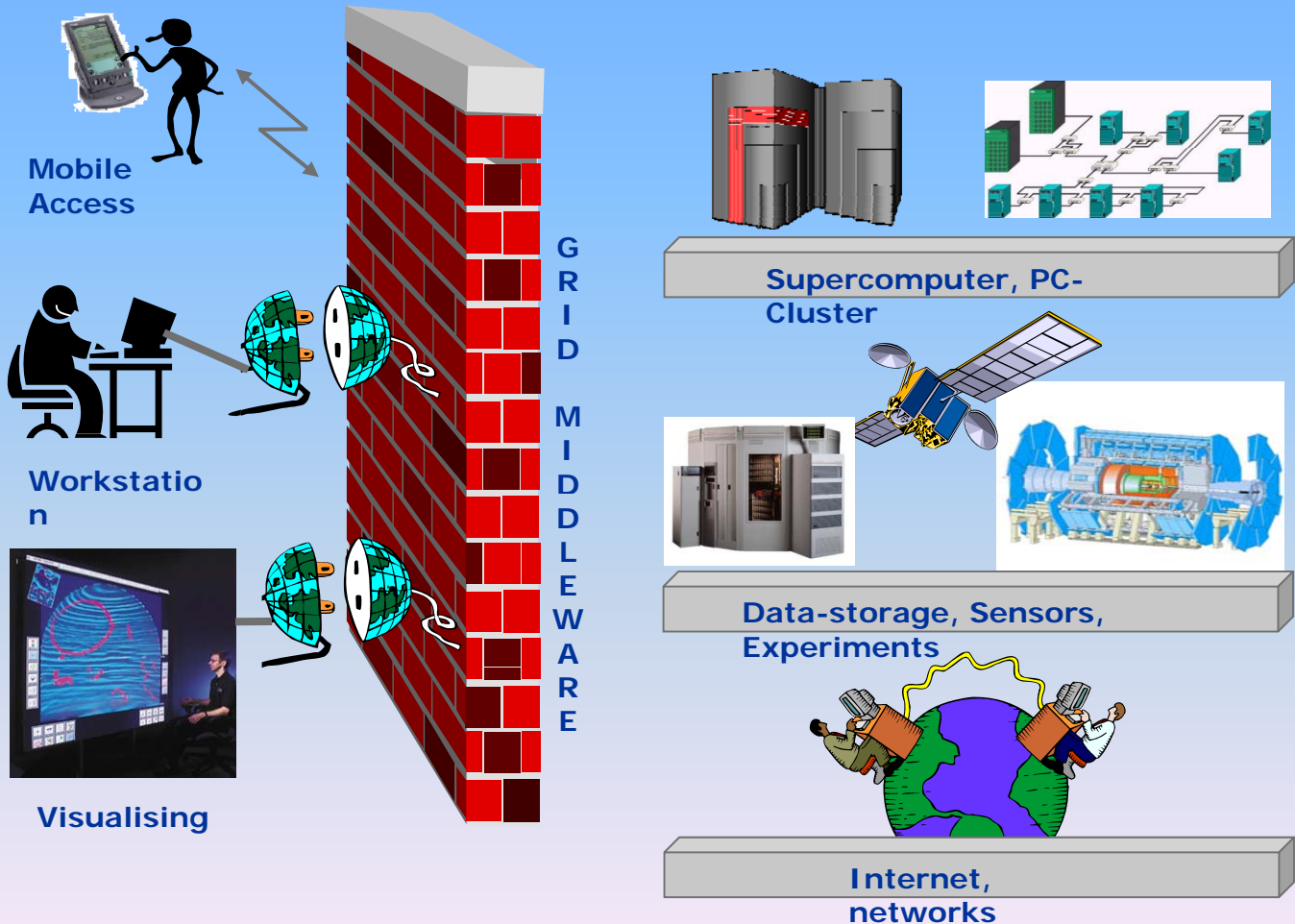
- Distributed resources
- No central control



# EGEE: Tiers approach



# Grid



# EGEE infrastructure

egEE  
Enabling Grids  
for E-science



Region	Countries	Sites	CPU
CERN	5	12	6400
UK/I	2	25	8384
Fr	2	12	7238
De/C/H	2	15	4413
It	1	34	4341
NE	9	30	3289
SEE	8	38	2727
CE	7	24	2588
SWE	2	18	1938
A-P	8	20	1884
Ru	2	15	738
Totals	48	243	44040

# Grid in Netherlands: BigGrid

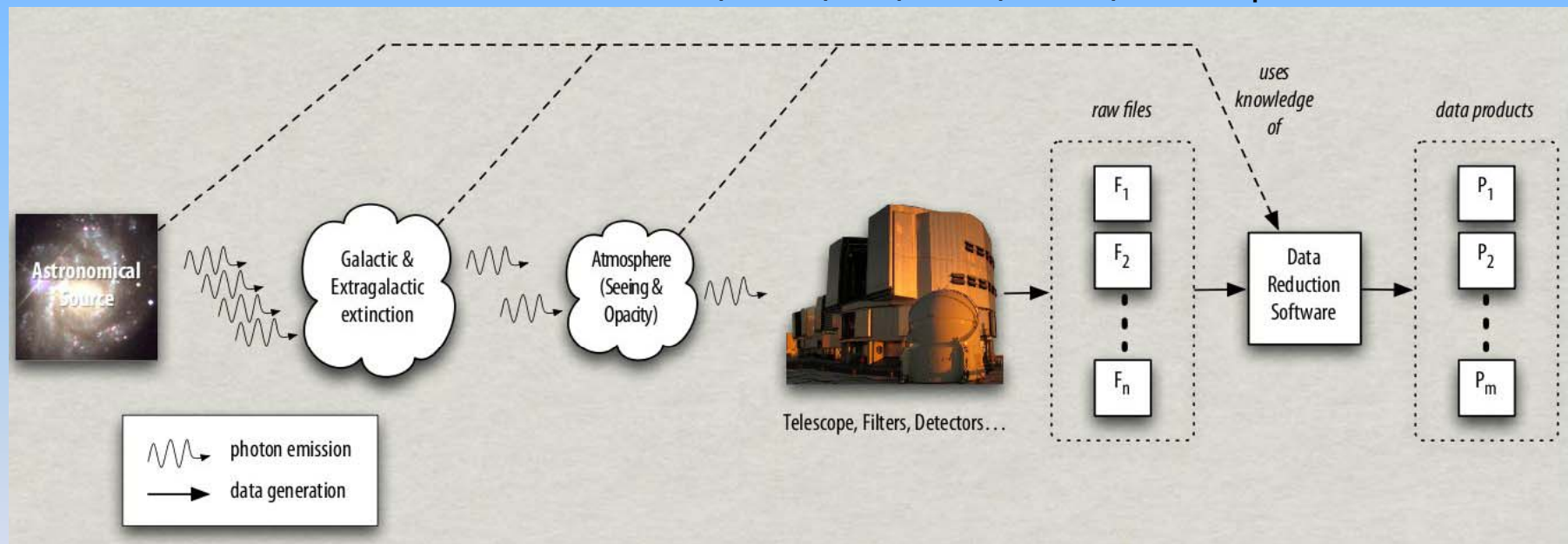
- 3 general sites
  - SARA
  - Nikhef
  - RuG
- 6 Life science sites
- About 3000 CPUs
- Several PB storage



# VO: Data Provenance

J.Santander, A.Delgado, The case for Data Provenance  
IVOA Interop 2009

<http://www.ivoa.net/cgi-bin/twiki/bin/view/IVOA/InterOpNov2009DM>



Connect data reduction facilities (like Astro-WISE) with telescopes (raw data archive on telescope) and end-user (VO interfaces) by providing complete description of raw data (general data model for raw data).