

Mass Scales and the CMBR

Observation and analysis of the glow of the Big Bang mark the beginning of 'precision cosmology' (and the end of classical astronomical methods)

Equations of Motion

- Einstein Equation
- Friedmann Equation
- Required:
 - Global Cosmic Parameters
 - Equation of State
 - Particle Processes

Einstein Equation

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} - \Lambda g_{\mu\nu} = 8\pi GT_{\mu\nu}$$

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Friedmann Equations

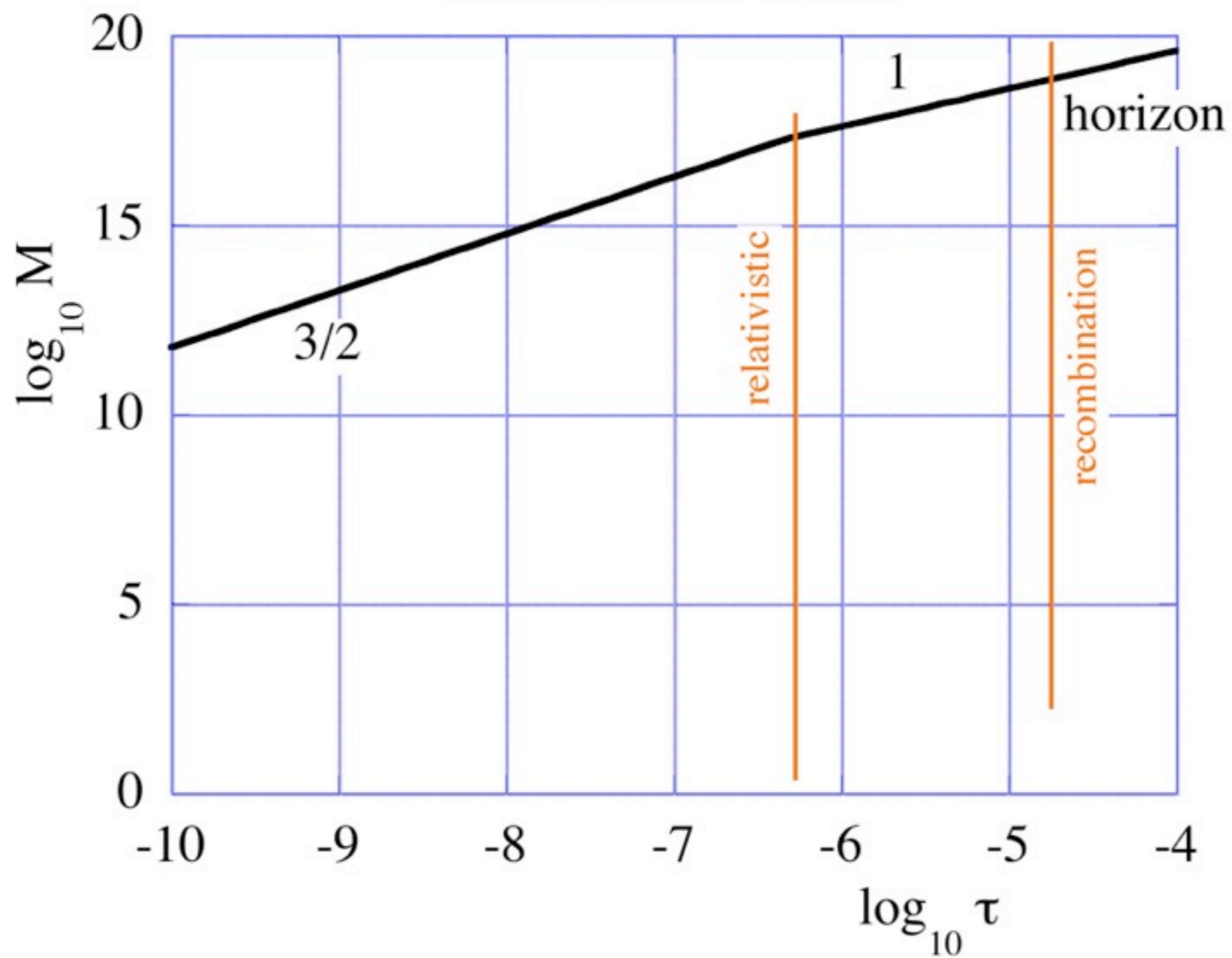
$$\left(\frac{da}{dt}\right)^2 = \frac{8\pi}{3}G\rho a^2 - kc^2 + \frac{\Lambda}{3}a^2$$

$$\frac{d}{dt}(\rho a^3) + \frac{P}{c^2} \frac{da^3}{dt} = 0$$

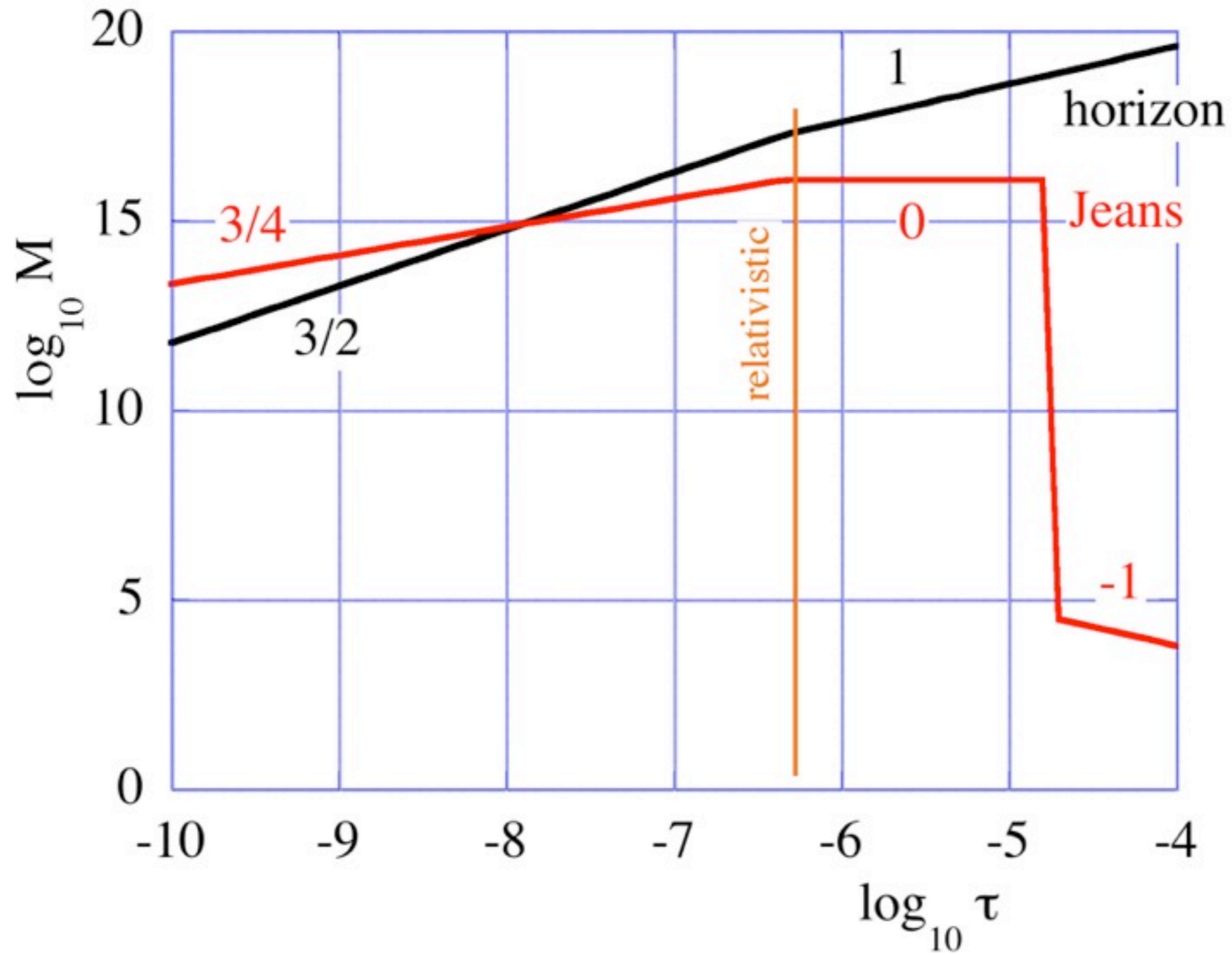
Mass Scales

- Analytic Examples
- Einstein-De Sitter Models
- (Non)relativistic
- Thermal Processes

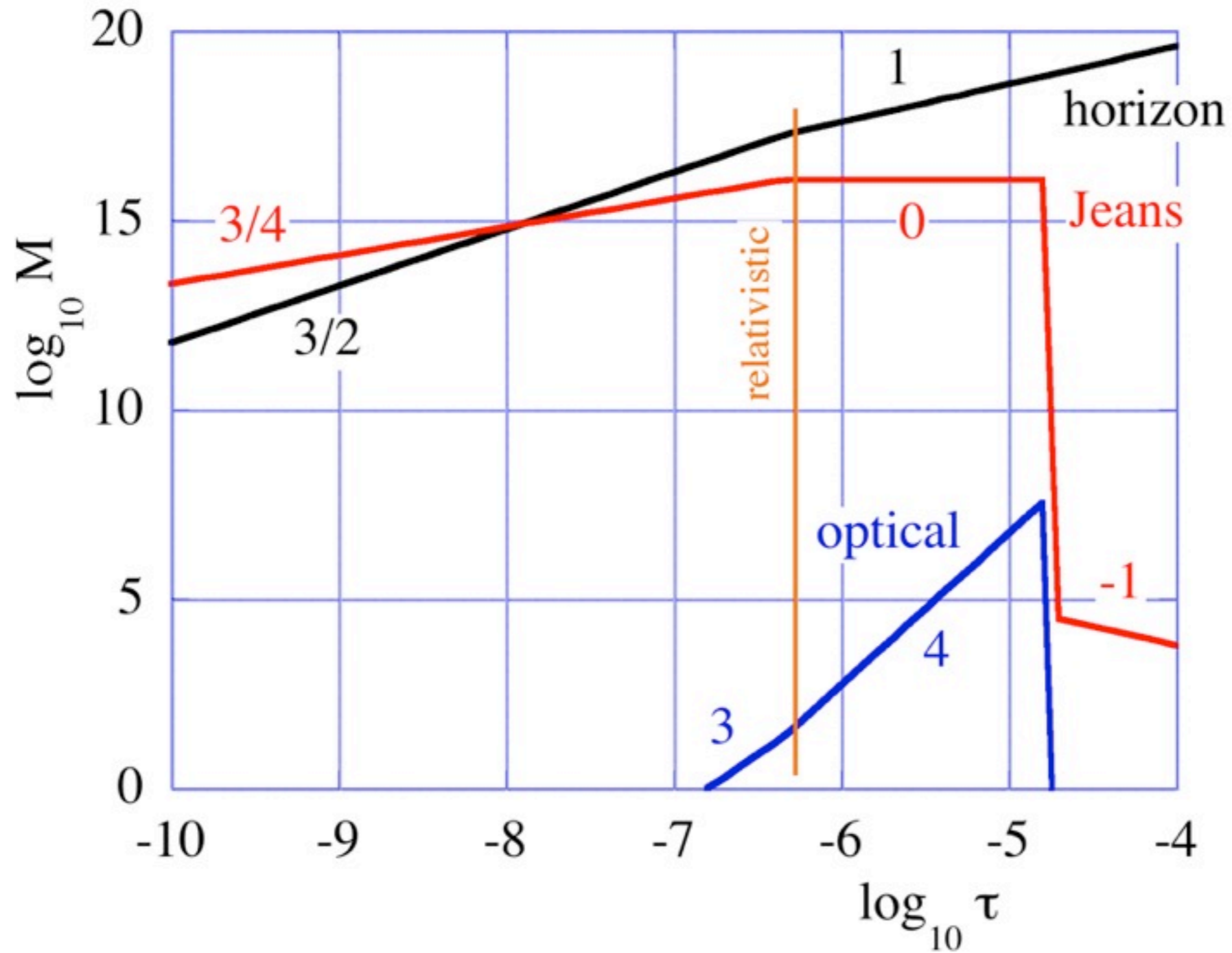
Cosmic Mass Scales



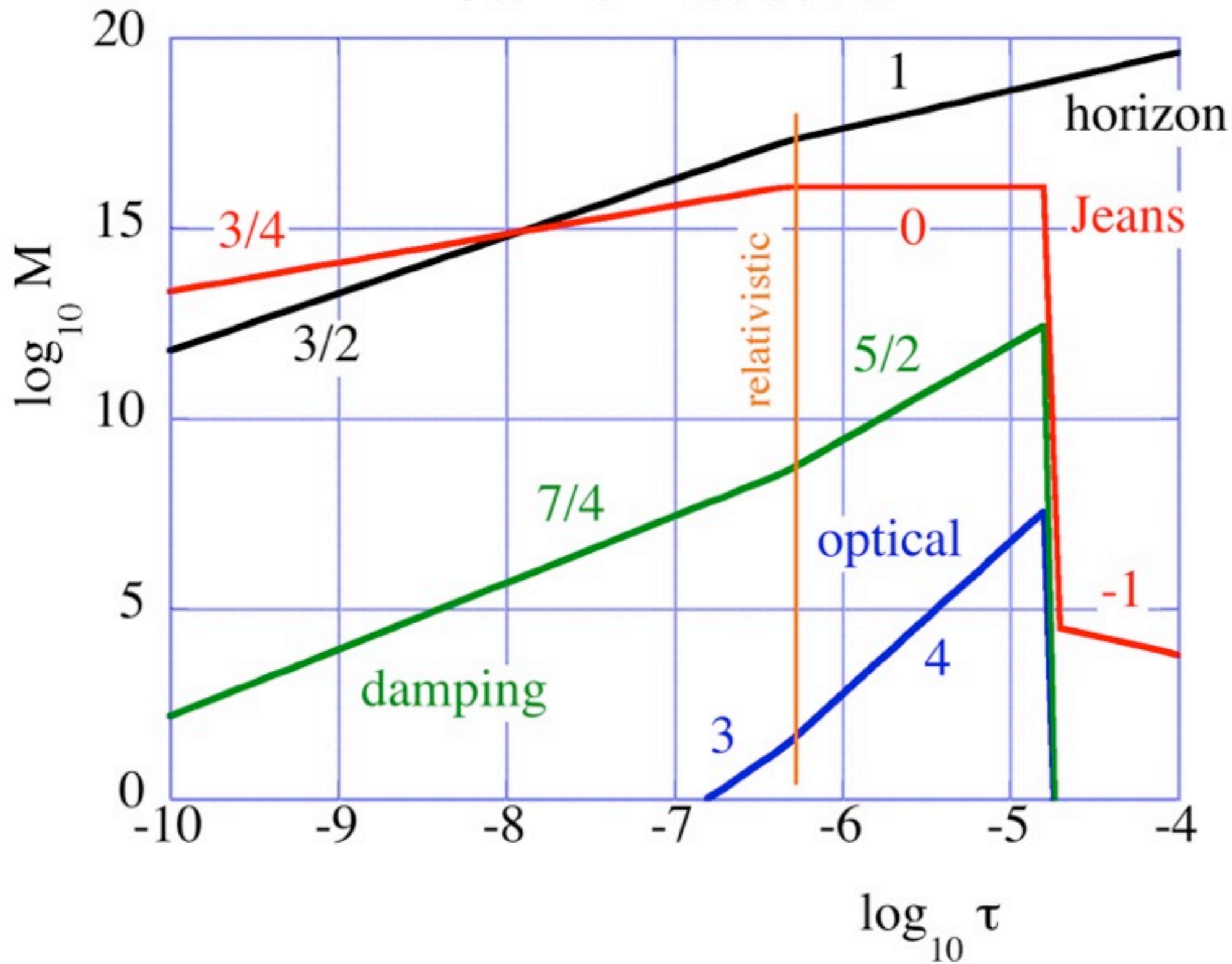
Cosmic Mass Scales



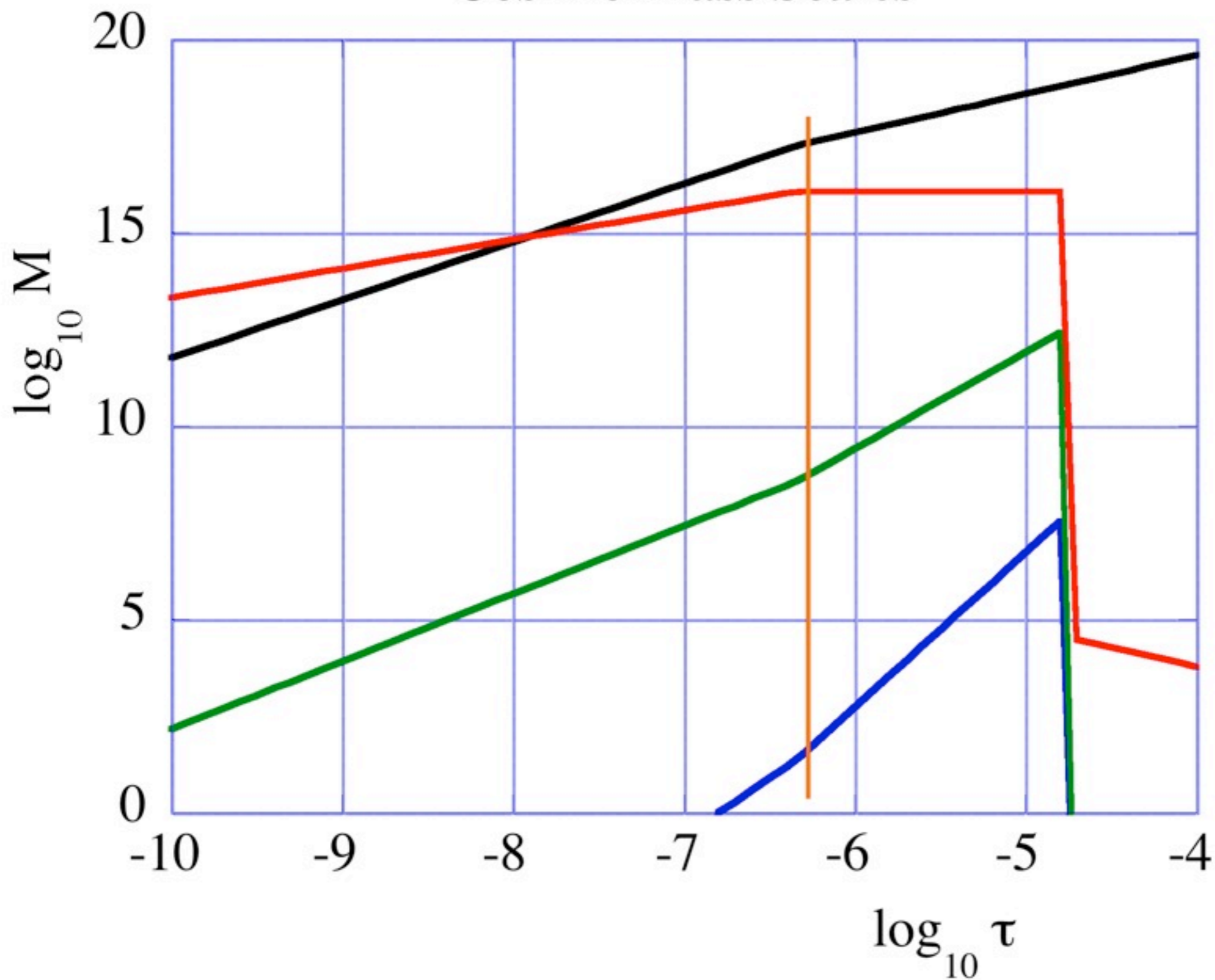
Cosmic Mass Scales



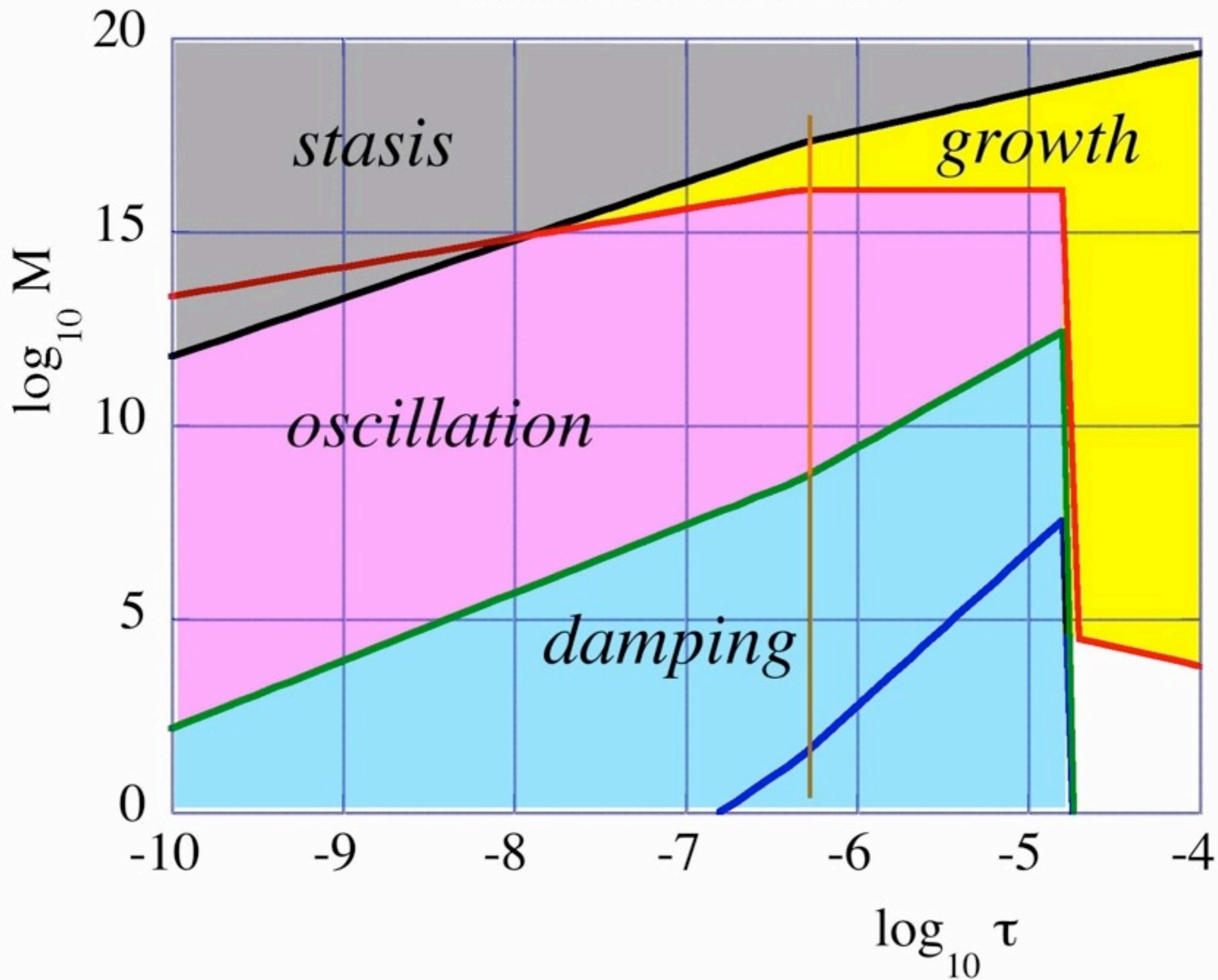
Cosmic Mass Scales



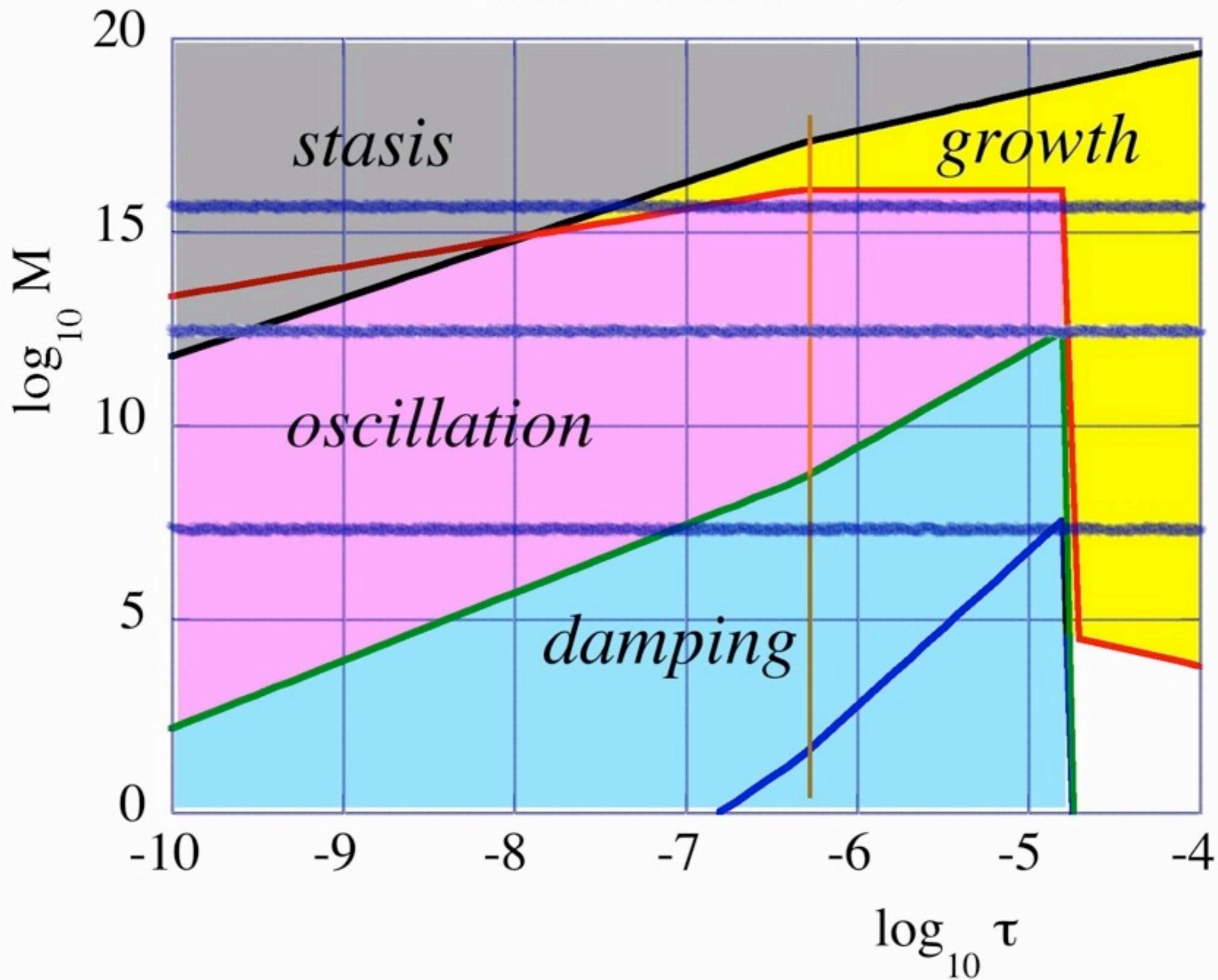
Cosmic Mass Scales



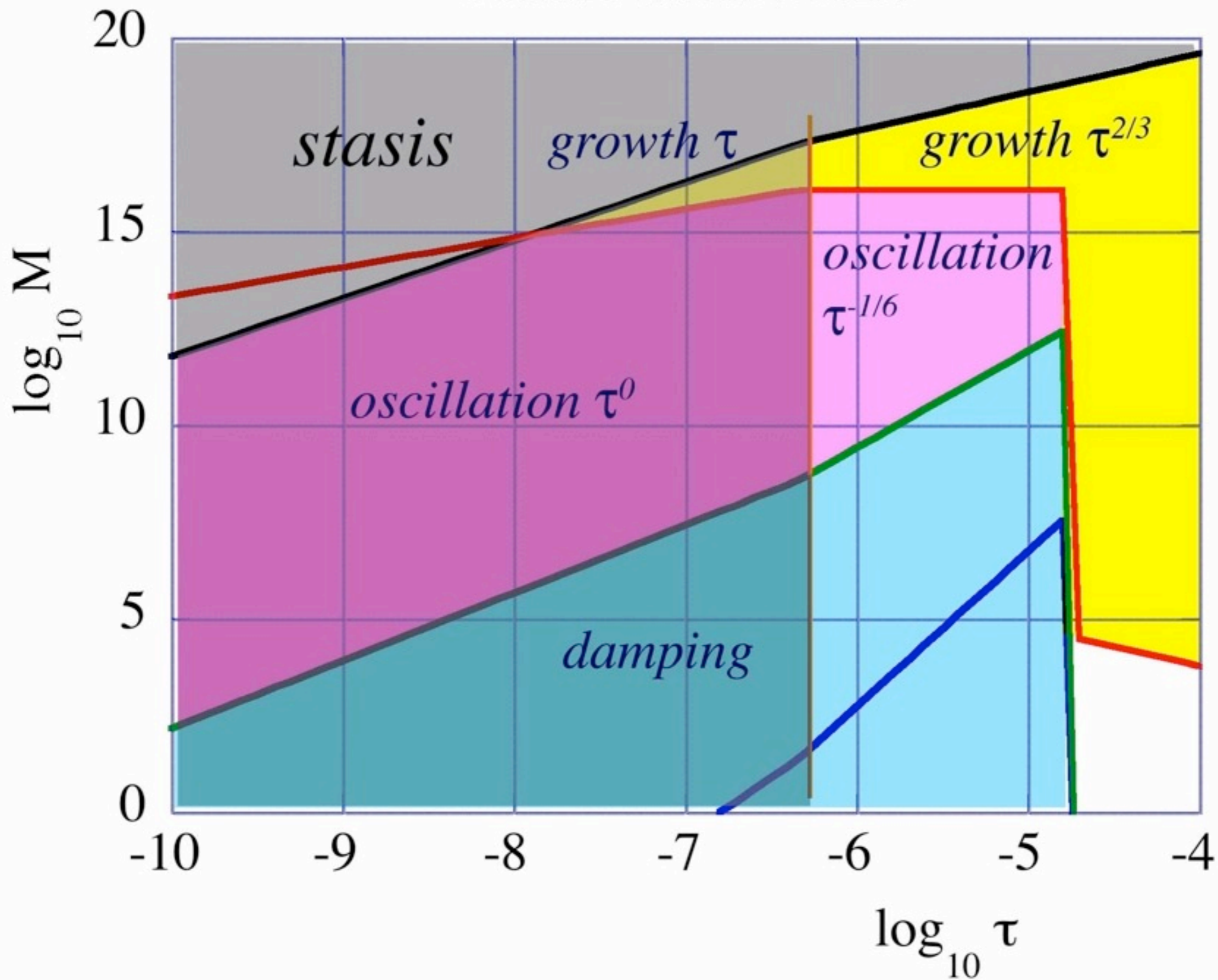
Cosmic Mass Scales



Cosmic Mass Scales

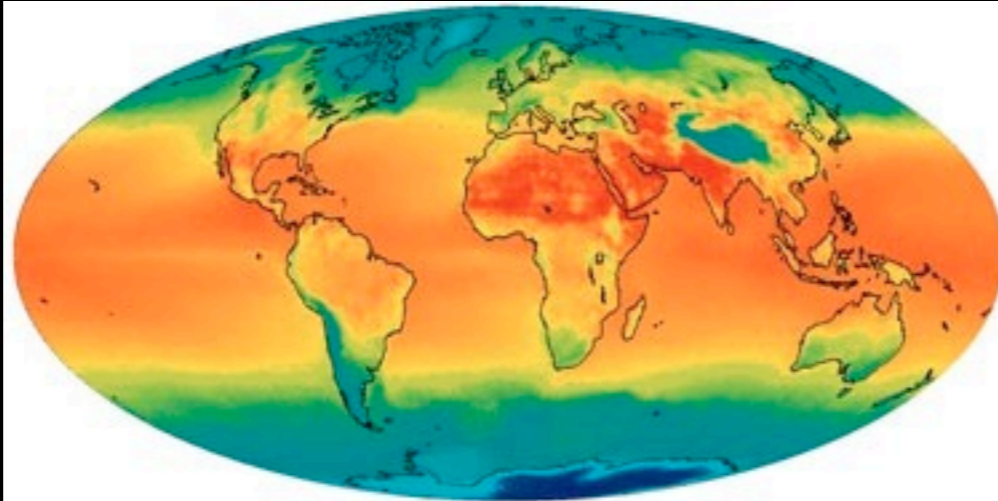


Cosmic Mass Scales

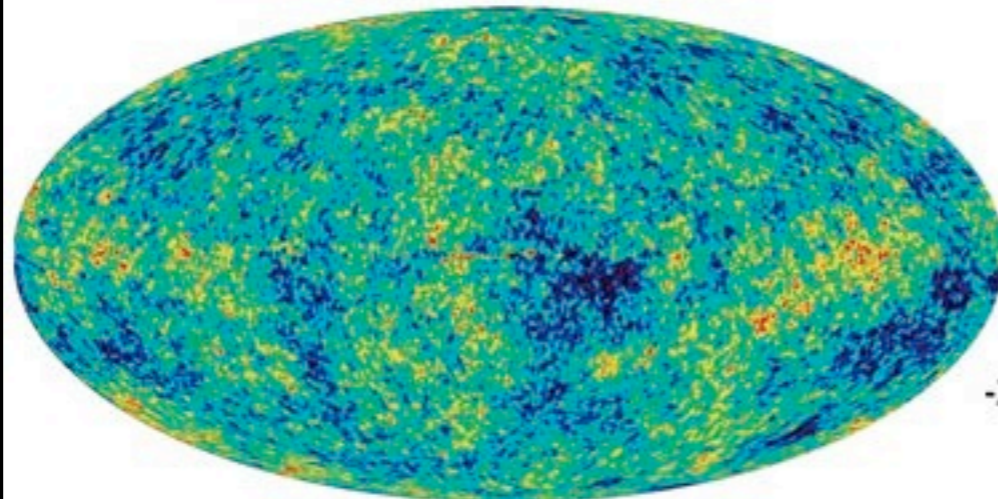


WMAP Introduction

- **Wilkinson Microwave Anisotropy Probe**
- **Long-term all-sky mapping**
- **Remove foreground signal**
- **Multipole analysis of signal on sphere**

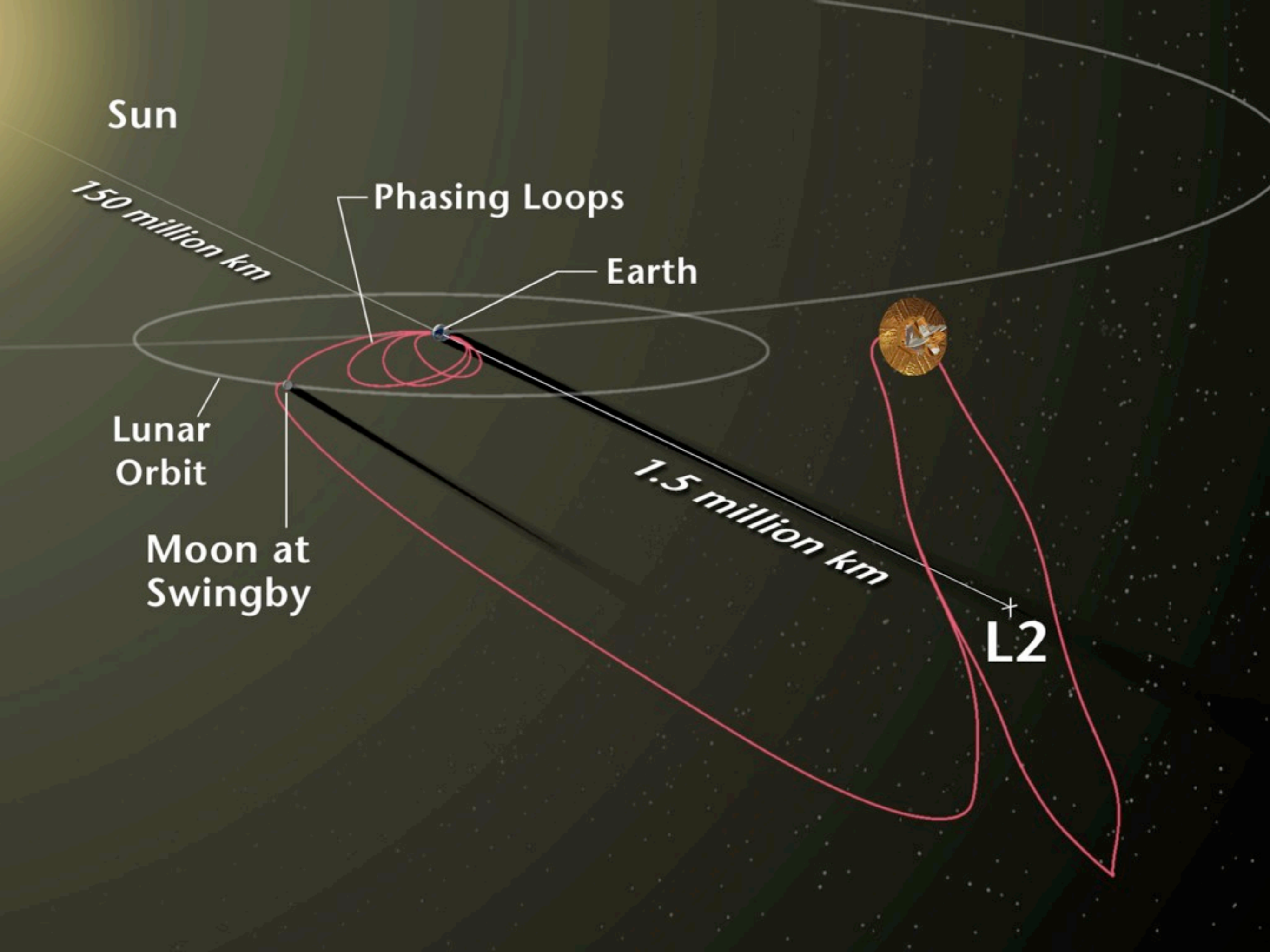


Earth
Temperatures

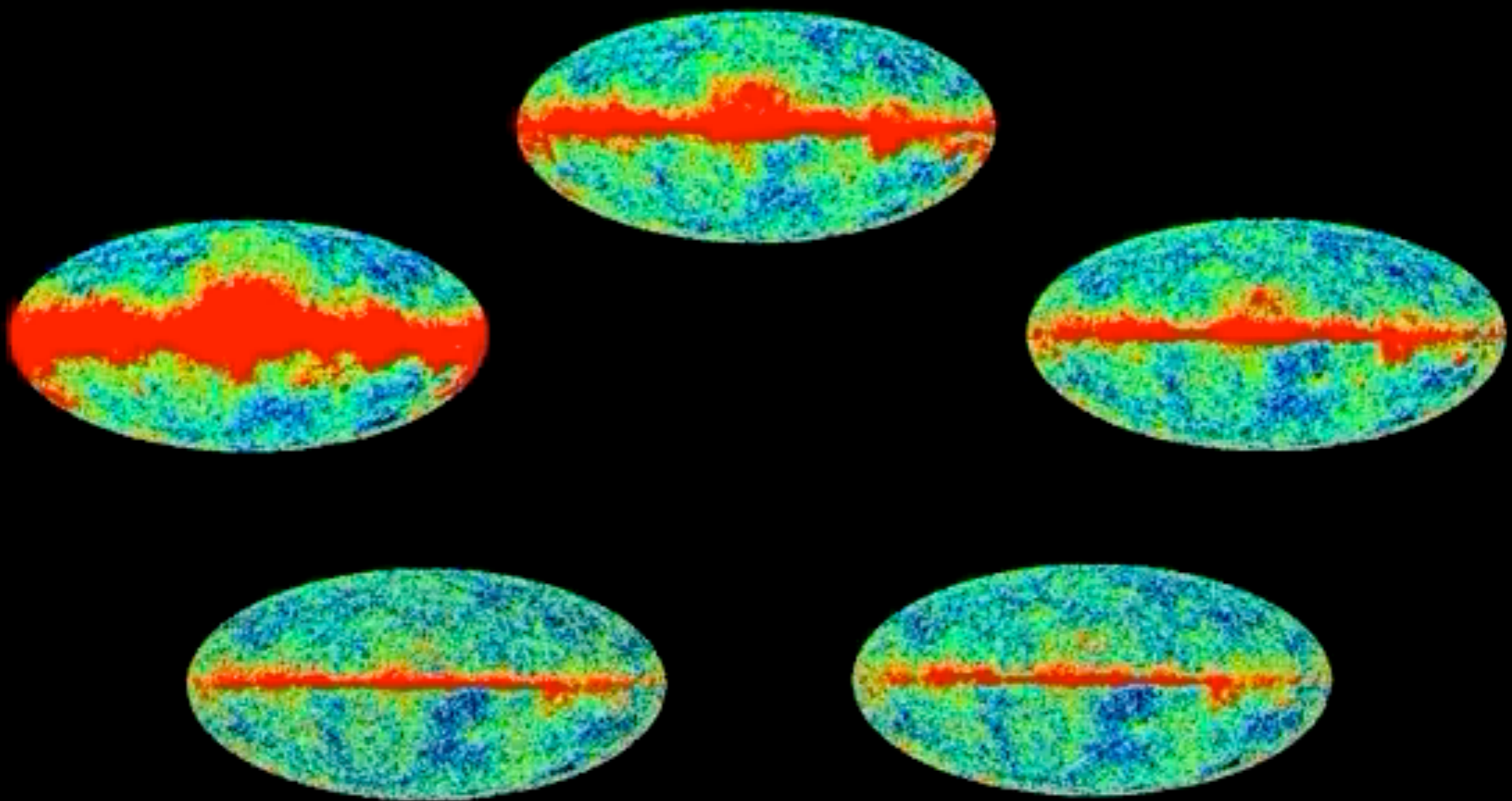


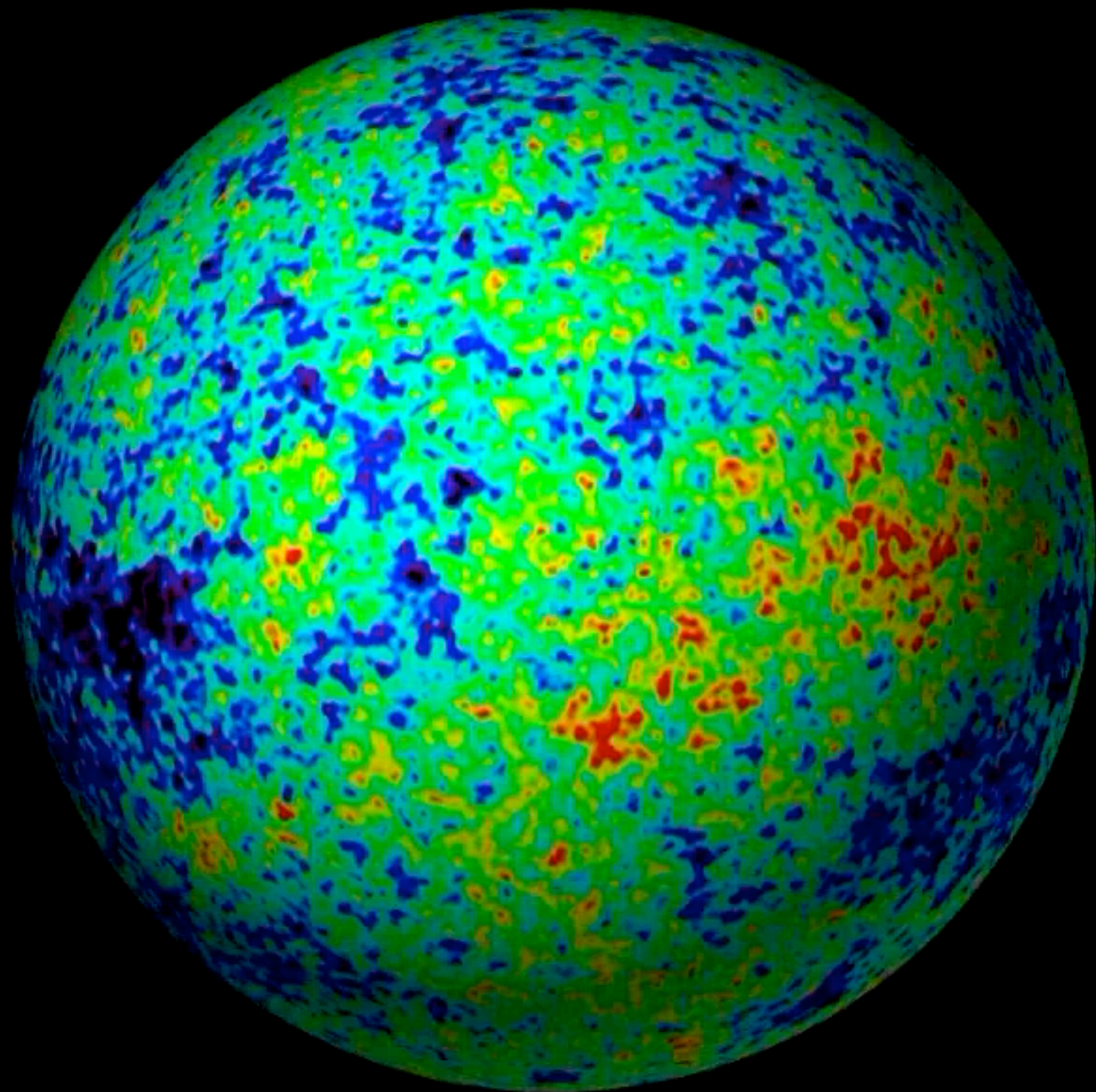
Microwave Sky
Temperatures

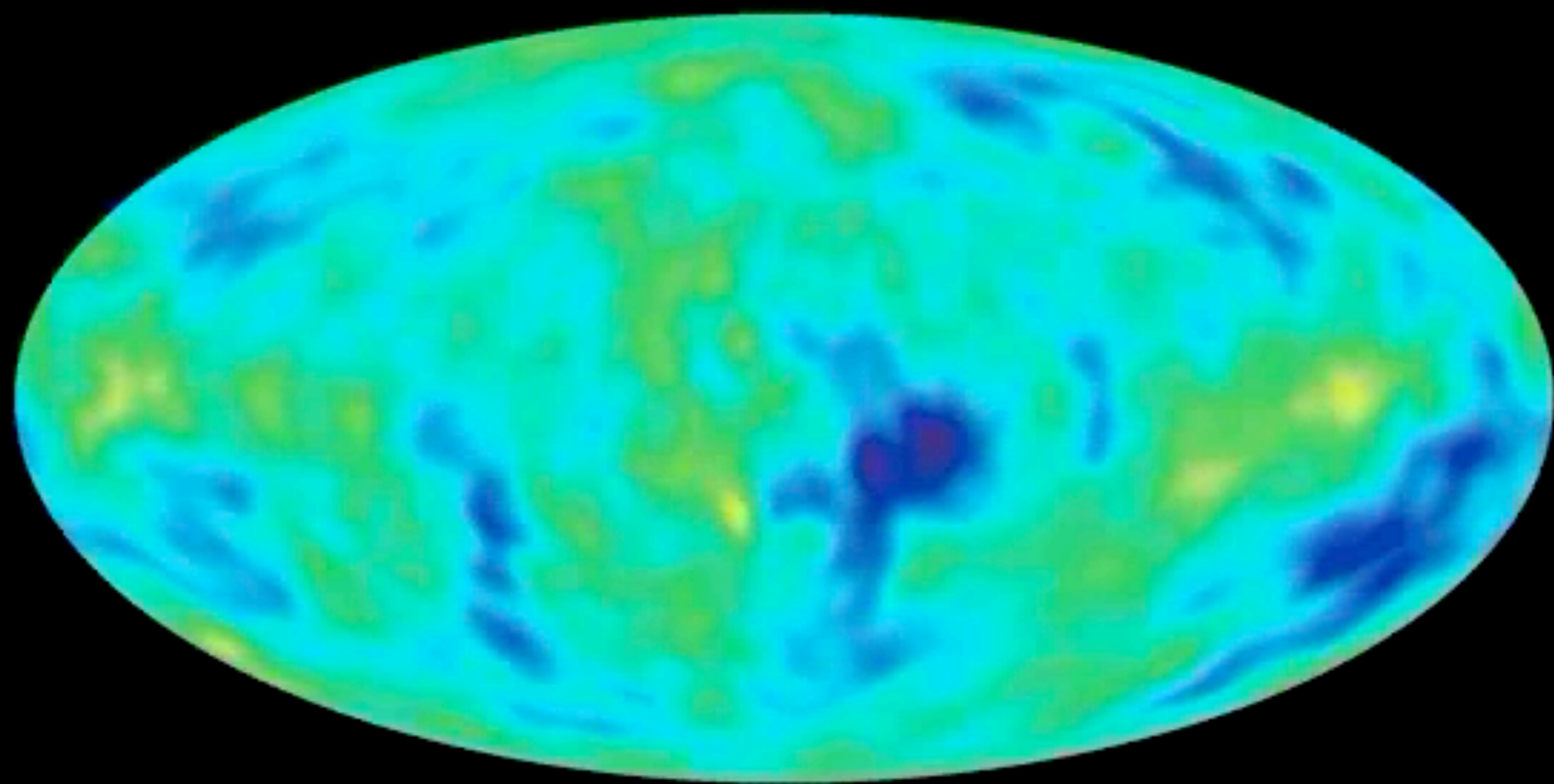


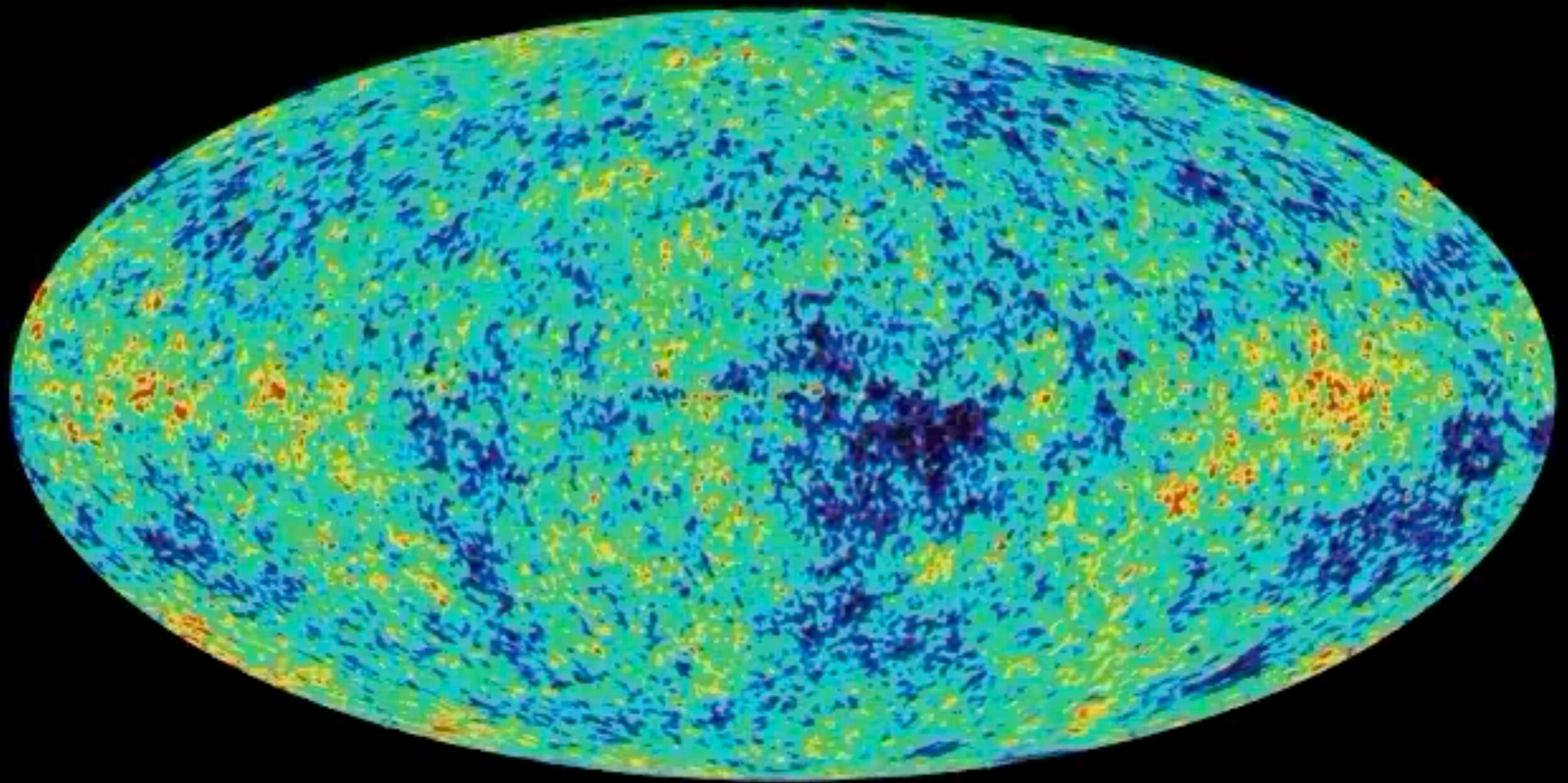


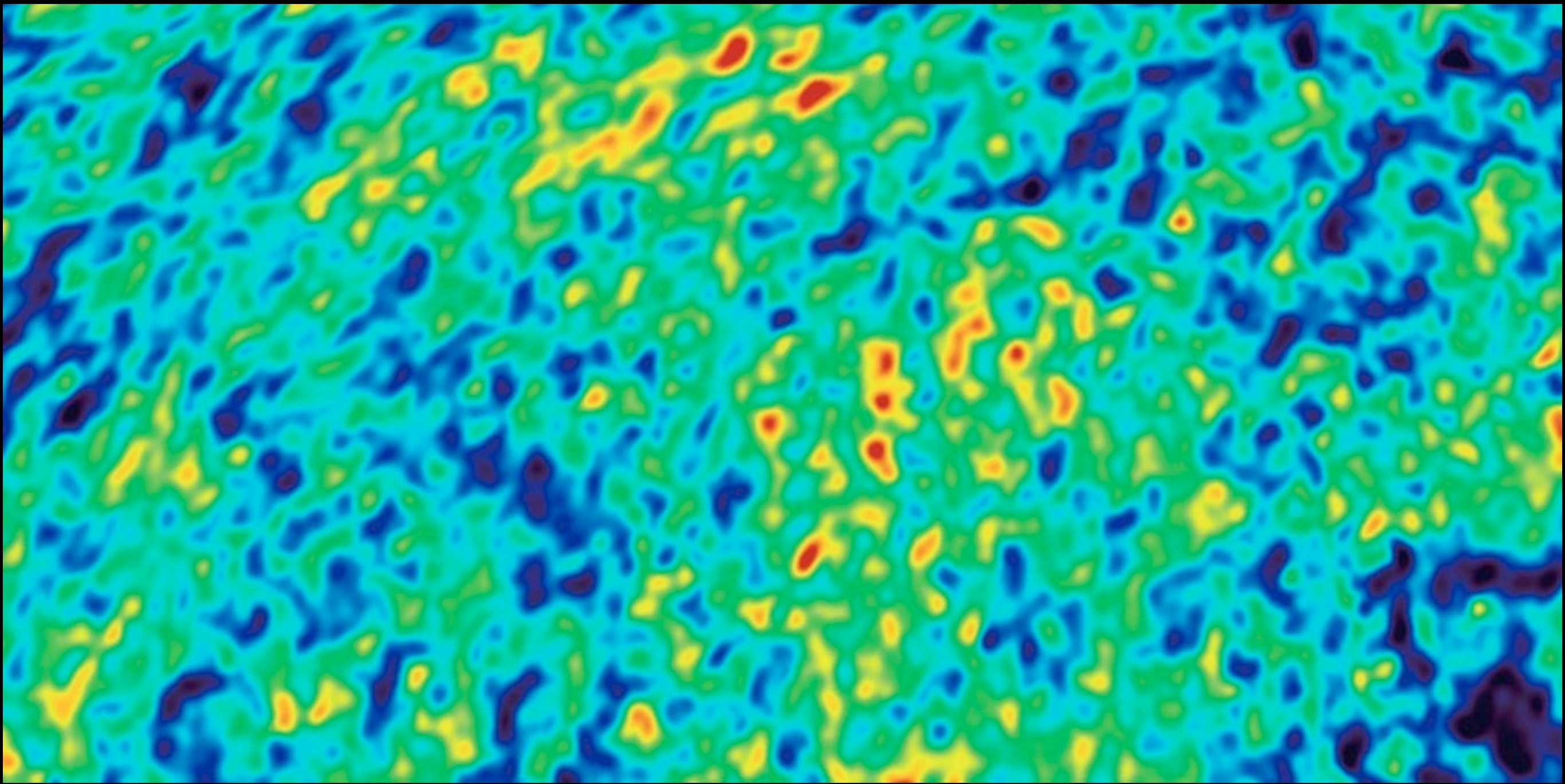


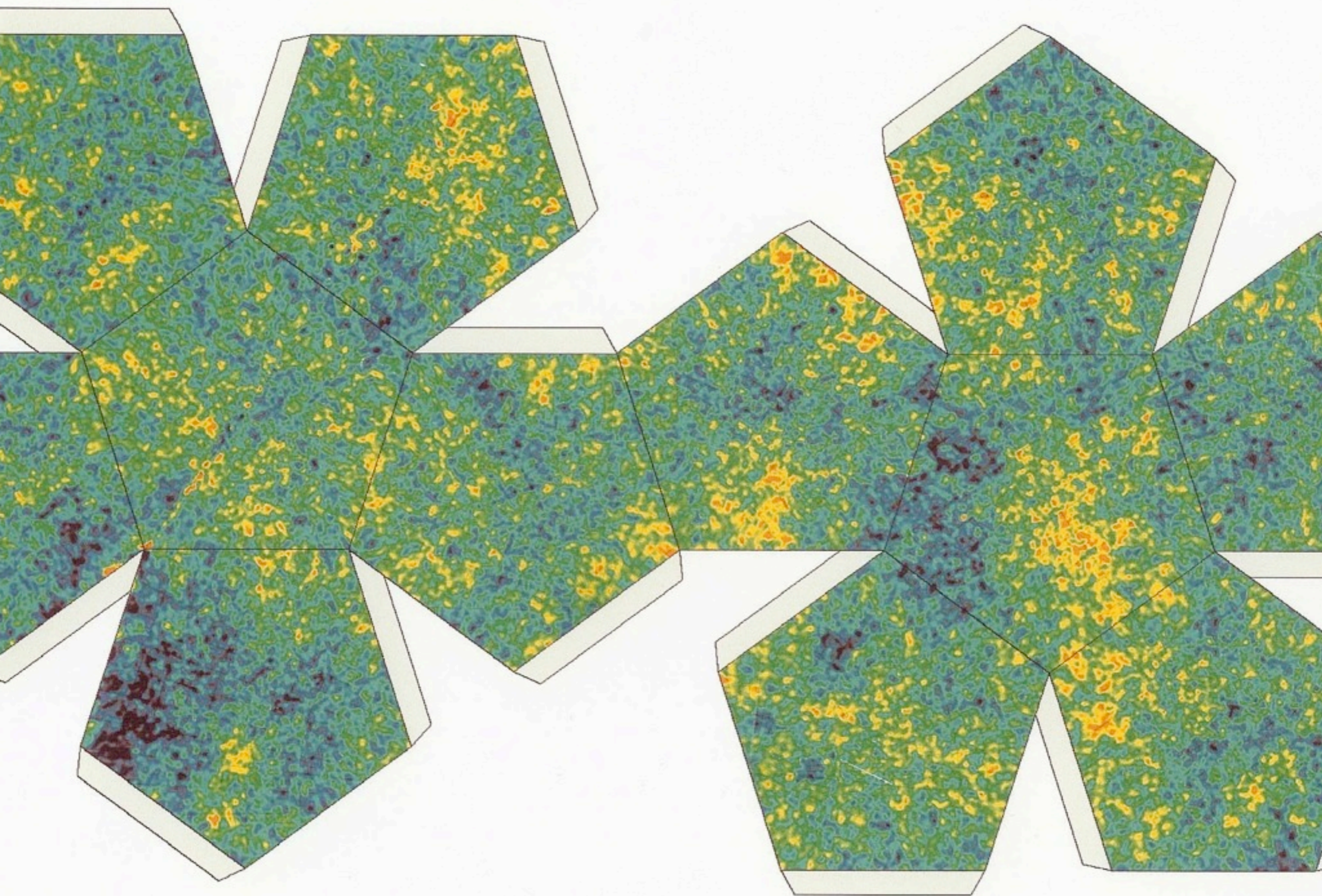








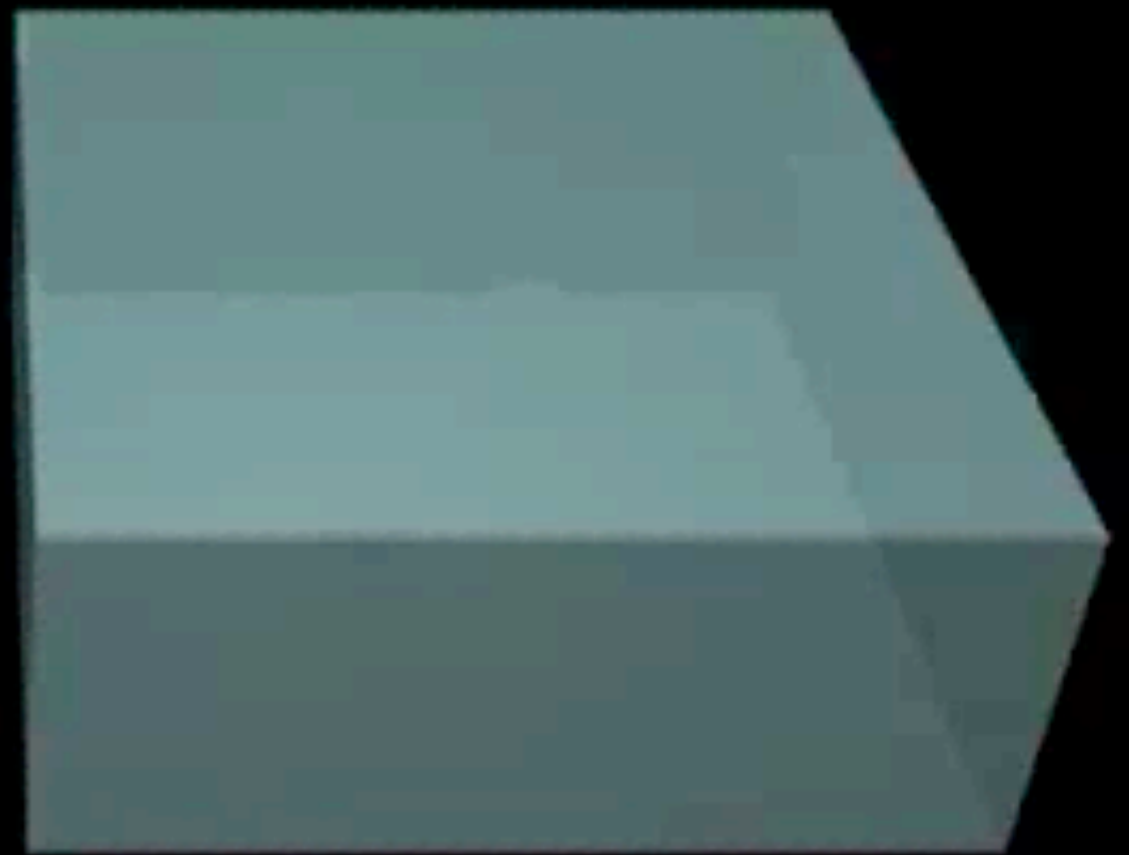




of the Cosmic Microwave Background as observed by the WMAP Satellite

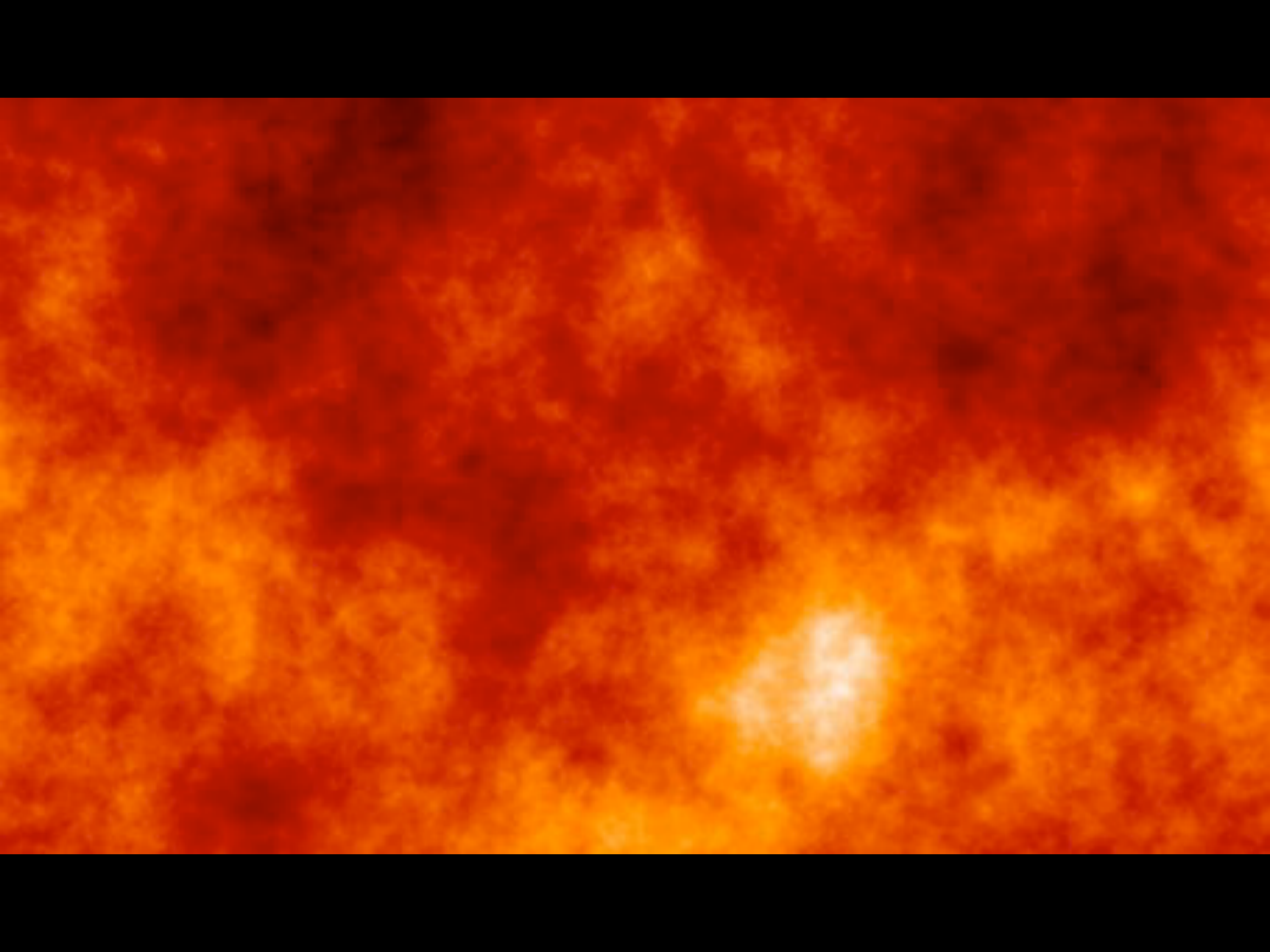
WMAP Oscillations

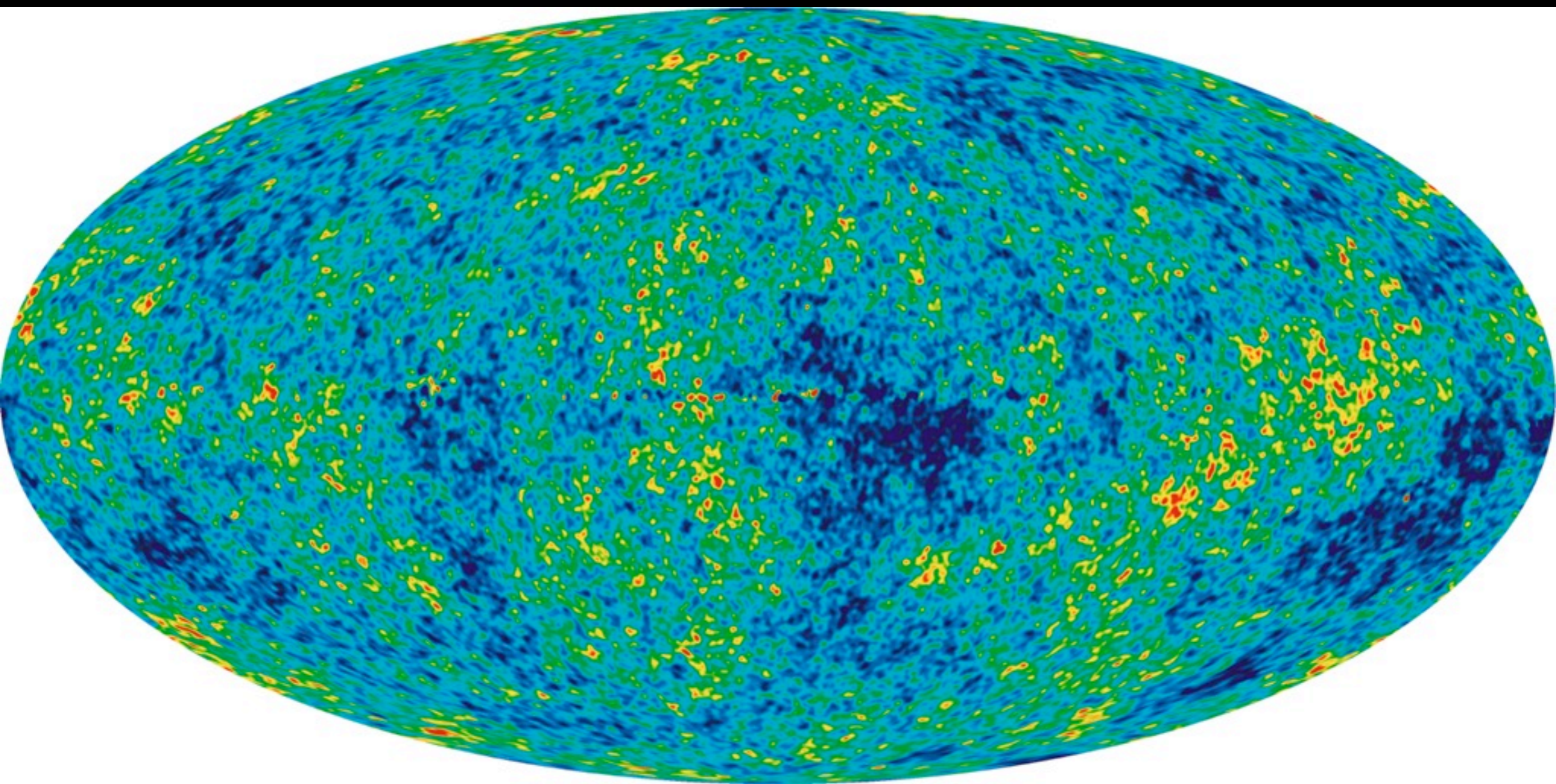
- Anisotropic cosmology
- Requires full Einstein Equations...
- ...but linearized due to smallness of anisotropy
- Probes the Equation of State
- Probes the isotropic background metric



Plasma Era:
Horizon Crossing
Silk Oscillations
Sachs-Wolfe Ripples

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Acoustic Peaks

- Inhomogeneities start to oscillate as soon as the horizon becomes larger than their size
- Amplitude at decoupling is determined by the acoustic frequency and the time until recombination
- “First acoustic peak” is a measure of the acoustic horizon at recombination
- Thus, it is a primary observable for fixing cosmic parameters

Angular Scale

90°

2°

0.5°

0.2°

6000

5000

4000

3000

2000

1000

0

Anisotropy Power (μK^2)

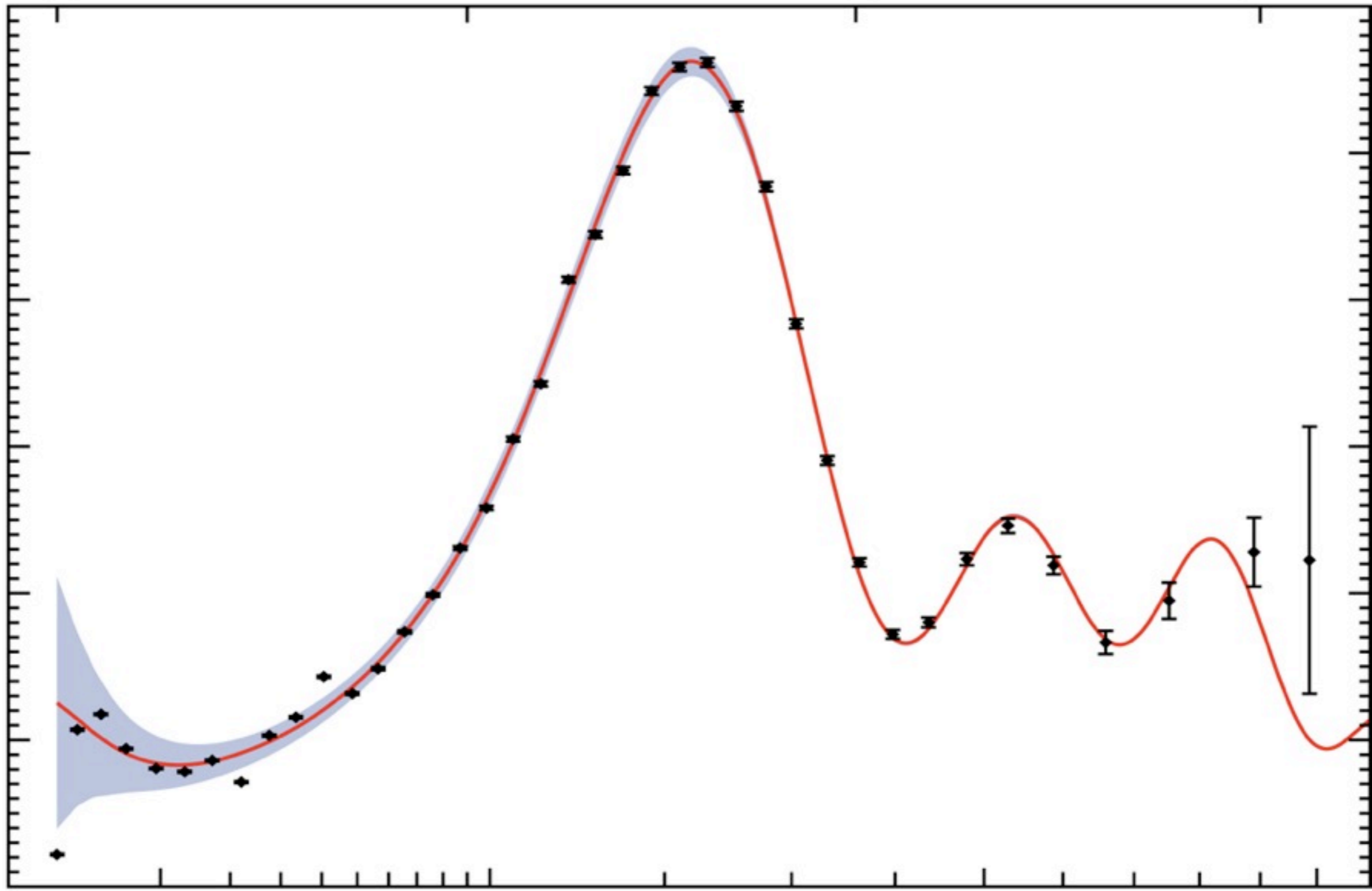
10

100

500

1000

Multipole moment (l)

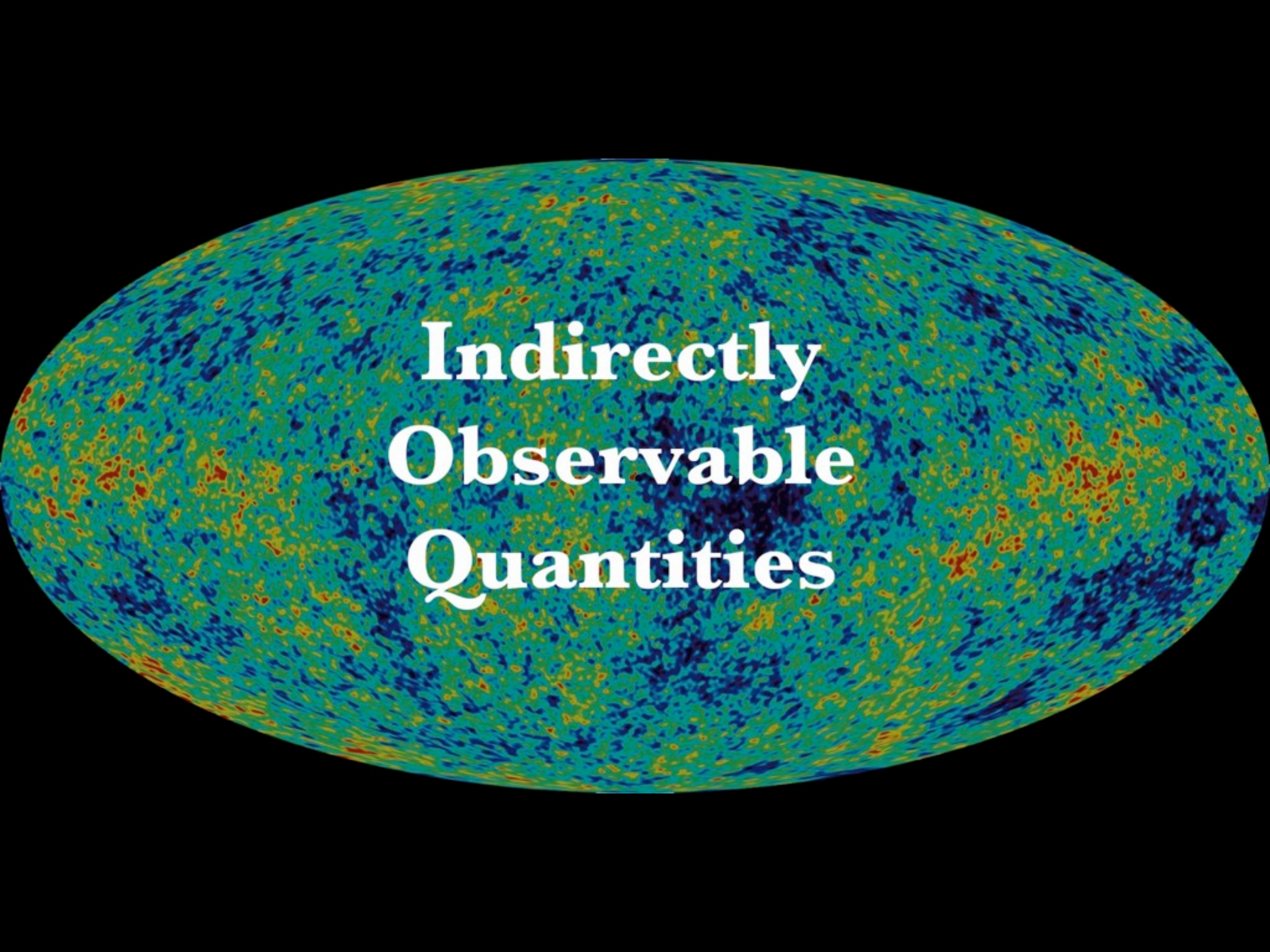


Initial conditions

- Origin of homogeneity and isotropy
- Disjunct parts of Universe have same properties
- Inflation...?
- But why are physical laws, constants etc. the same?
- Answers in interactions between particles and space-time?

WMAP Results

- Superb thermal blackbody spectrum
- Anisotropy amplitude 10^{-5}
- Harrison-Zel'dovich isocurvature noise
- Silk peaks
- Ideal baryonic gas
- Dark stuff 95%

The image features a large, horizontally-oriented oval shape with a complex, multi-colored texture. The colors range from dark blue and purple to bright yellow and orange, creating a mottled, grainy appearance. This texture is set against a solid black background. The overall effect is reminiscent of a map of the universe, such as a Cosmic Microwave Background (CMB) fluctuation map, or perhaps a detailed view of a celestial body's surface. The text is centered within this oval.

Indirectly Observable Quantities

A map of the Cosmic Microwave Background (CMB) showing temperature fluctuations across the sky. The map is an oval shape with a complex, noisy pattern of colors ranging from dark blue (cooler) to yellow and red (warmer). The text is overlaid on the map in white.

Age: 13.7 Gyr

4% baryonic matter

96% dark matter/energy

Harrison-Zel'dovich noise

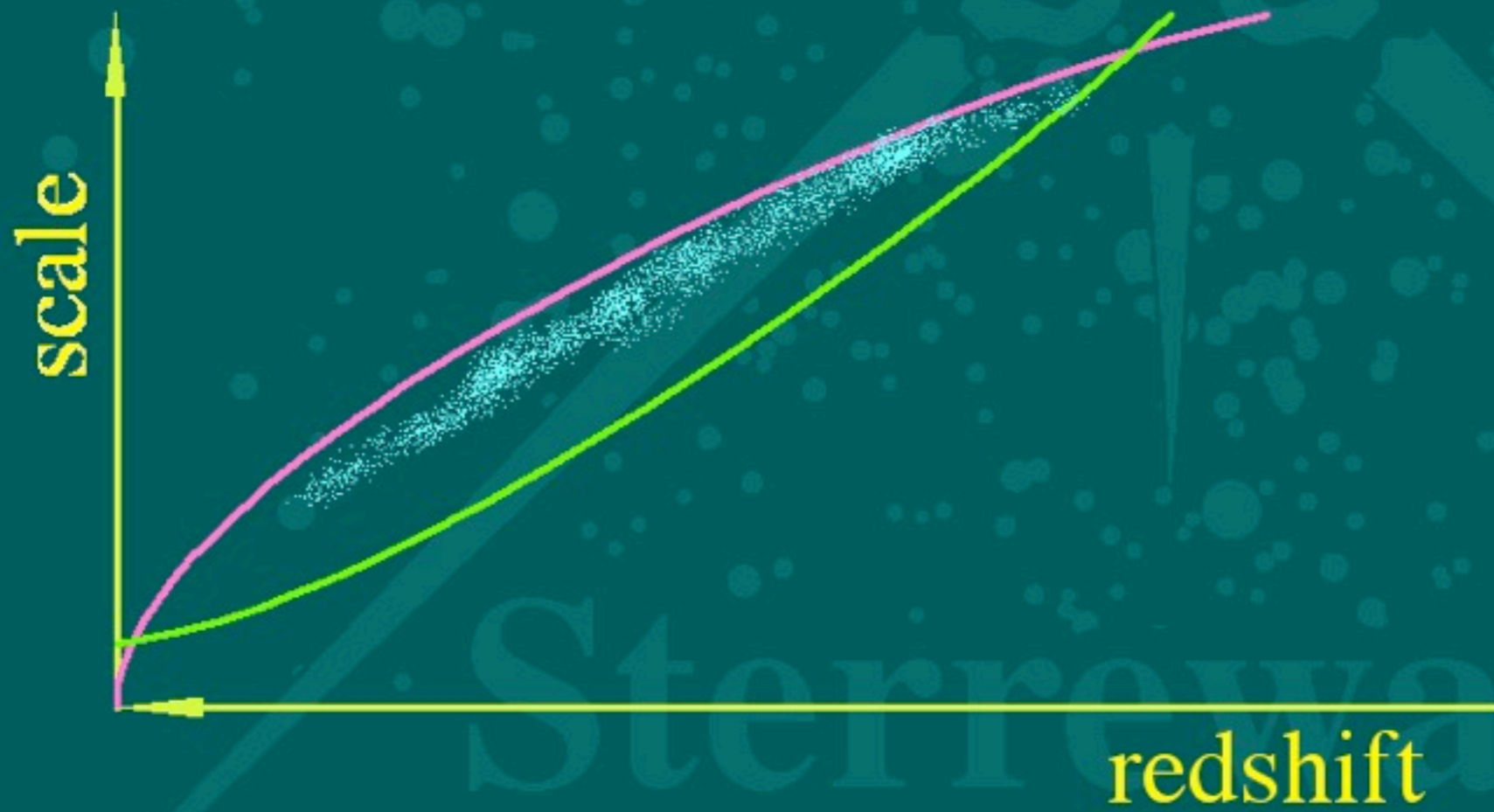
Solutions

$$\left(\frac{da}{dt}\right)^2 = \frac{8\pi}{3} G \rho a^2 - kc^2 + \frac{\Lambda}{3} a^2$$



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Observations



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